

Opportunities in nonlinear classical and quantum physics with intense XFELs



Tom Blackburn Department of Physics, University of Gothenburg www.gu.se/en/research/plasma-physics 19th January 2023 UK XFEL Townhall – Fundamental Physics, Quantum Computing and AI



- Introduction: key parameters, the nonlinear regimes
- XFELs as probes: vacuum birefringence
- XFELs as backgrounds: gamma-ray sources for nuclear photonics
- XFELs as drivers: wakefields in solid-density plasmas...
- In and Schwinger pair creation in vacuum

Introduction Classical nonlinearity parameter, a₀



Blackburn, RMPP 2020

 Compare the characteristic frequency of the emitted radiation to the orbital (cyclotron) frequency:

$$rac{\omega'}{\omega_c} \simeq a_0^3$$

 Harmonic order of the emitted radiation, or number of participating photons. Introduction Classical nonlinearity parameter, *a*₀



Alternatively...

- Electron does not interact with one photon of the radiation field, but many.
- Determines the importance of higher order terms in a perturbative expansion, which only works if a₀ is much smaller than one.

Introduction Classical nonlinearity parameter, a₀



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- Electron does not interact with one photon of the radiation field, but many.
- Determines the importance of higher order terms in a perturbative expansion, which only works if a₀ is much smaller than one.
- If we swap out the photon energy...

Introduction GOTHENBURG Quantum nonlinearity parameter $\chi_{e,\gamma}$



- ... we see if the field does enough work over a Compton length to create an electron,
- or the field strength in the particle rest frame, in units of the critical (Schwinger) field.
- Then quantum effects must be important – even if the field amplitude in the lab frame is far below critical.

Introduction Quantum nonlinearity parameter $\chi_{e,\gamma}$



 For a photon, vacuum polarization and electron-positron pair creation are controlled by

$$\chi_{\gamma} = \frac{e\sqrt{-(F_{\mu\nu}k^{\nu})^2}}{m^3} = \frac{\omega|\vec{E}_{\perp} + \vec{n} \times \vec{B}|}{mE_{\rm crit}}$$

 For an electron, photon emission (and recoil effects) are controlled by

$$\chi_e = \frac{e\sqrt{-(F_{\mu\nu}p^{\nu})^2}}{m^3} \simeq \frac{\gamma |\vec{E}_{\perp} + \vec{v} \times \vec{B}|}{E_{\rm crit}}$$



Introduction Parameter space

high intensity, quasistatic, tunnelling, field can be treated as being locally constant





Introduction Wavelength matters



- Which of nonlinear classical or quantum effects kicks in first depends on the wavelength of the strong field (among other things).
- Most research has focused on highpower optical lasers.
- In this talk I will present some ideas, all of which rely upon the co-location of intense XFELs with other particle and radiation sources.



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- Photons can interact with each other via fluctuating electron-positron pairs.
- Astrophysical applications; laboratory tests of strong-field QED.
- Many of these cross sections scale positively with increased frequency.
- Top to bottom: Delbruck scattering, photon splitting, four-wave mixing.



XFELs as probes Vacuum birefringence



- Cross section for four-photon scattering $\sigma [\text{cm}^{-2}] = 10^{-31} (\omega/m)^6 (\omega = \text{photon} \text{energy in ZMF, natural units hereon}).$
- 11 orders of magnitude larger for XFEllaser than laser-laser (33 for XFEL-XFEL vs laser-laser).
- Use optical laser as "vacuum polarizer" (coherence!) and XFEL as probe [HIBEF, SEL etc].



XFELs as probes Vacuum birefringence



- Scattering visible in helicity change of the probe X rays, as if the vacuum refractive index were polarization dependent.
- Initially linearly polarized probe acquires a small ellipticity $\delta = \omega_{\rm X} L \Delta n/2$, where $\Delta n = 2 \alpha I_{\rm laser}/I_{\rm cr}$.
- Why X rays? High polarization purity and precision of measurement.



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XFELs as backgrounds Gamma-ray source





- Compton scattering of electrons is a source of polarized, monochromatic photons.
- The 10 to 100 MeV range useful for nuclear photonics: excitation of resonances, transitions etc
- e.g. 720 MeV electrons + 1 eV optical laser (ELI-NP)
- or 1 GeV + 1 to 5 eV FEL (HIGS)



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- To achieve the same photon energies by scattering XFEL light requires lower electron energies.
- Example: 100 MeV or 1 GeV electrons + 200 eV light @ 10¹⁹ W/cm² (300 cycles in duration)





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- Example: 100 MeV or 1 GeV electrons + 200 eV light @ 10¹⁹ W/cm² (300 cycles in duration)
- Gammas inherit polarization properties of the XFEL light: linear, circular etc
- Monochromatise by angle selection





- How far can one go?
- Multi-GeV polarized gammas needed for hadron studies in strange sector / photoproduction of QCD exotics
- LEPS2 @ Spring-8 (ICS with 8-GeV electrons + VUV), GlueX (coherent bremstrahlung)
- At 10 GeV, Compton edge at 9.7 GeV.

XFELs as backgrounds Radiation reaction in the deep quantum regime



- Recoil effects on the electron beam are significant!
- Investigation of radiation reaction with optical lasers usually in the regime where single-photon recoil is much less important than the accumulated recoil from many emissions.
- Importance of "spin light" increased.



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XFELs as drivers XWFA



- Solid-density plasmas become underdense at XUV/X-ray wavelengths: $\lambda_p \ [\mu m] = n^{-1/2} \ [10^{21} \ cm^{-3}]$
- As an example, a relativistically intense XUV pulse could drive wakefield acceleration of electrons [Tajima EPJST 2014].

• $a_0 > 1 \Rightarrow I > 10^{24} (\lambda_X \text{ [nm]})^{-2} \text{ W/cm}^2$





Wettervik, PoP 2018

- Scaling wakefield acceleration to XUV wavelengths relies on similarity theory $(S = n_e/a_0 n_{cr})$.
- But not all physics scales with this parameter, e.g. radiation generation (betatron or nonlinear Compton)
 [Zhang PRAB 2016].
- PIC simulations show it holds to a wavelength of 1 nm.



XFELs as drivers XWFA

Quantity	Scaling
Timescale	λ
Spatial dimensions of cavity	λ
Density	a_0/λ^2
Current density	a_0/λ^2
Current	a_0
Trapped charge	λa_0
Electron energy	a_0
Electromagnetic fields	a_0/λ
Pulse energy	λa_0^2
Ion motion	a_0
Quantum parameter, χ	a_0^2/λ
Radiated energy fraction	$\lambda^{1-lpha} a_0^{2lpha-1}$
Photon energy $(\chi \ll 1)$	a_0^3/λ
Photon energy $(\chi \sim 1)$	a_0
Time between emissions $(\chi \ll 1)$	T/a_0
Time between emissions $(\chi \sim 1)$	$T/(\lambda a_0)^{1/3}$

Wettervik, PoP 2018

- Motivation might be accelerating gradient (TeV/cm)...
- or perhaps the current density or electron bunch duration.
- For densities of 10²³ to 10²⁵ cm⁻³, cavity size implies an equivalent duration of 10⁻¹⁷ to 10⁻¹⁵ seconds.



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XFELs as drivers How do fields interact?



- Maxwell's equations are linear, at but ultrahigh field strengths this is not an adequate description anymore.
- Vacuum polarisation (photon-photon scattering, birefringence, dichroism)
- Schwinger pair creation/vacuum pair creation



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XFELs as drivers Schwinger effect



- The action for a static electric field acquires an imaginary part that is nonperturbative in the electric field strength.
- The characteristic field strength required is $E_{\rm crit} = 1.3 \times 10^{18} \, {\rm V/m...}$
- ... which corresponds to an intensity of 2×10²⁹ W/cm²





- $2 \times 10^{29} \text{ W/cm}^2$?
 - Aside from the technological challenges(!), it's likely impossible to reach an intensity anywhere near this with optical lasers because of pair avalanches [Bulanov et al, PRL 2010; Fedotov et al, PRL 2010].
- Routes toward "extremer" extreme field science rely on some means of conversion to higher frequencies.





- Takes advantage of smaller focal spot sizes at shorter wavelength
- ... and suppression of pair avalanche growth due to volume factors.
- In principle, Schwinger pair creation observable with a terawatt XFEL... focused to nm focal spot sizes [Ringwald, PLB 2001]





- Rate depends on invariants E² B² and E.B, not the electric field strength alone.
- Theory question: does a quasistatic model of pair creation work at high frequency/field gradient?
- "Assisted" Schwinger pair creation: reduce field strength required by combining the driver with other highfrequency EM waves.



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