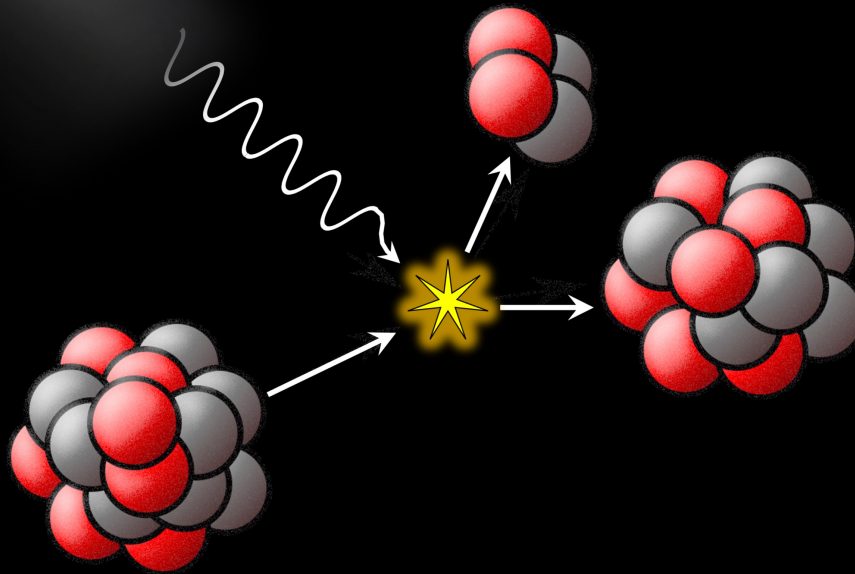


Shining a light on nuclear astrophysics with γ -beams at HI γ S



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[@UnclearPhysics](#)

UK Nuclear Physics Community Meeting– 11th January 2023

**Sheffield
Hallam
University**

Overview

- Nuclear physics at Sheffield Hallam University
- Overview of HI γ S facility
- Importance of $^{12}\text{C}(\alpha,\gamma)$ reaction in astrophysics
- Efforts to measure $^{12}\text{C}(\alpha,\gamma)$ using Optical TPC at HI γ S
 - Published results and on-going analysis
- New measurements of $^{12}\text{C}(\alpha,\gamma)$ with a new electronic TPC
 - on-going analysis

Nuclear physics at Sheffield Hallam



Robin Smith



Raed Dallal

New detector technologies



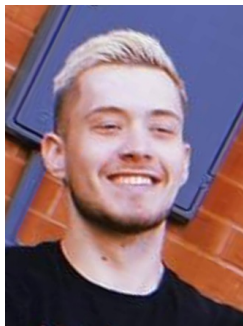
Ocean Wong

Neutron spectrum unfolding (UKAEA)



Olivia Tindle

Alpha particle clustering in light nuclei



Kris Haverson

Nuclear structure & astrophysics with TPC detectors



Kimberley Lennon

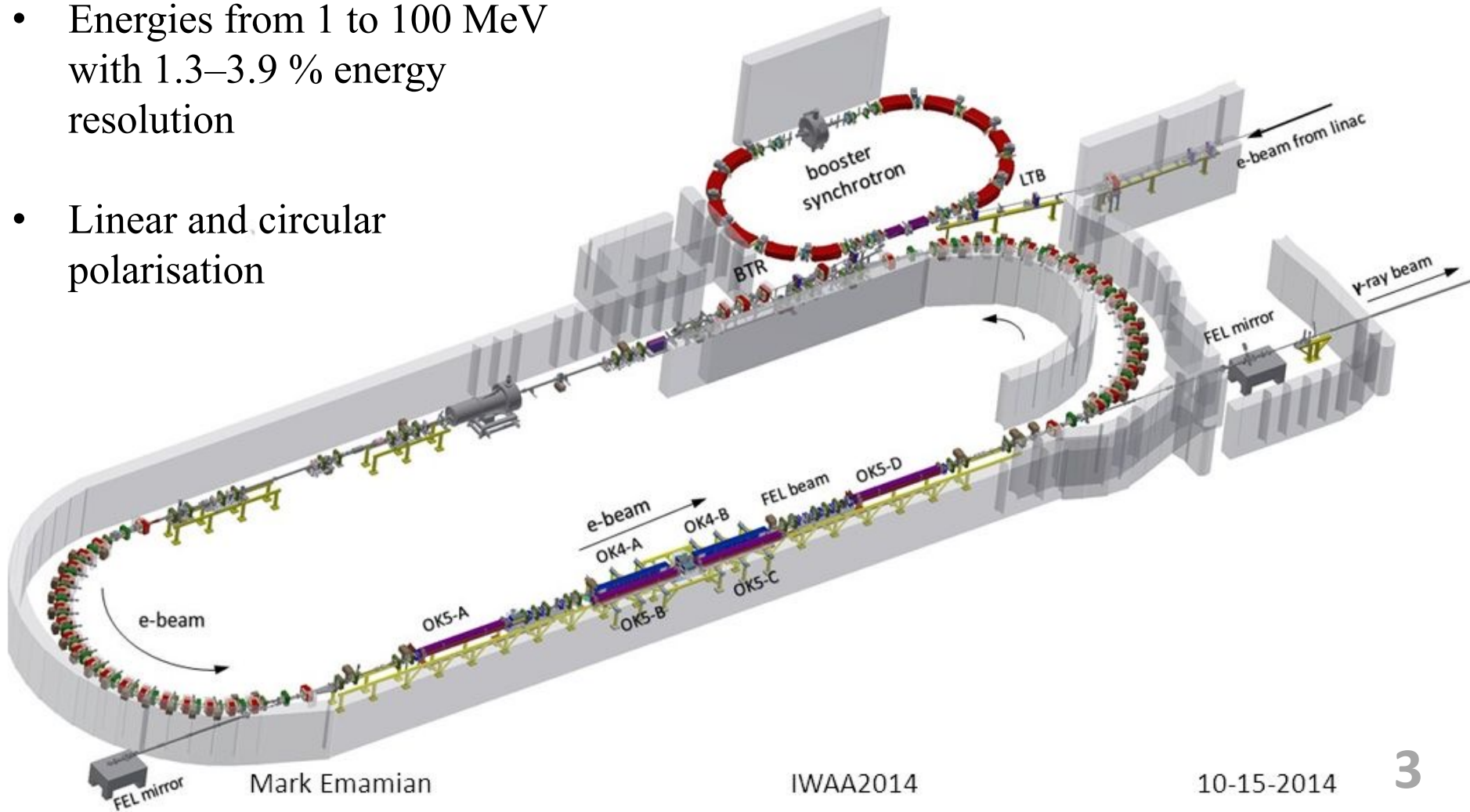
Machine learning for improvements in gamma spectroscopy (UKAEA)

HI γ S facility



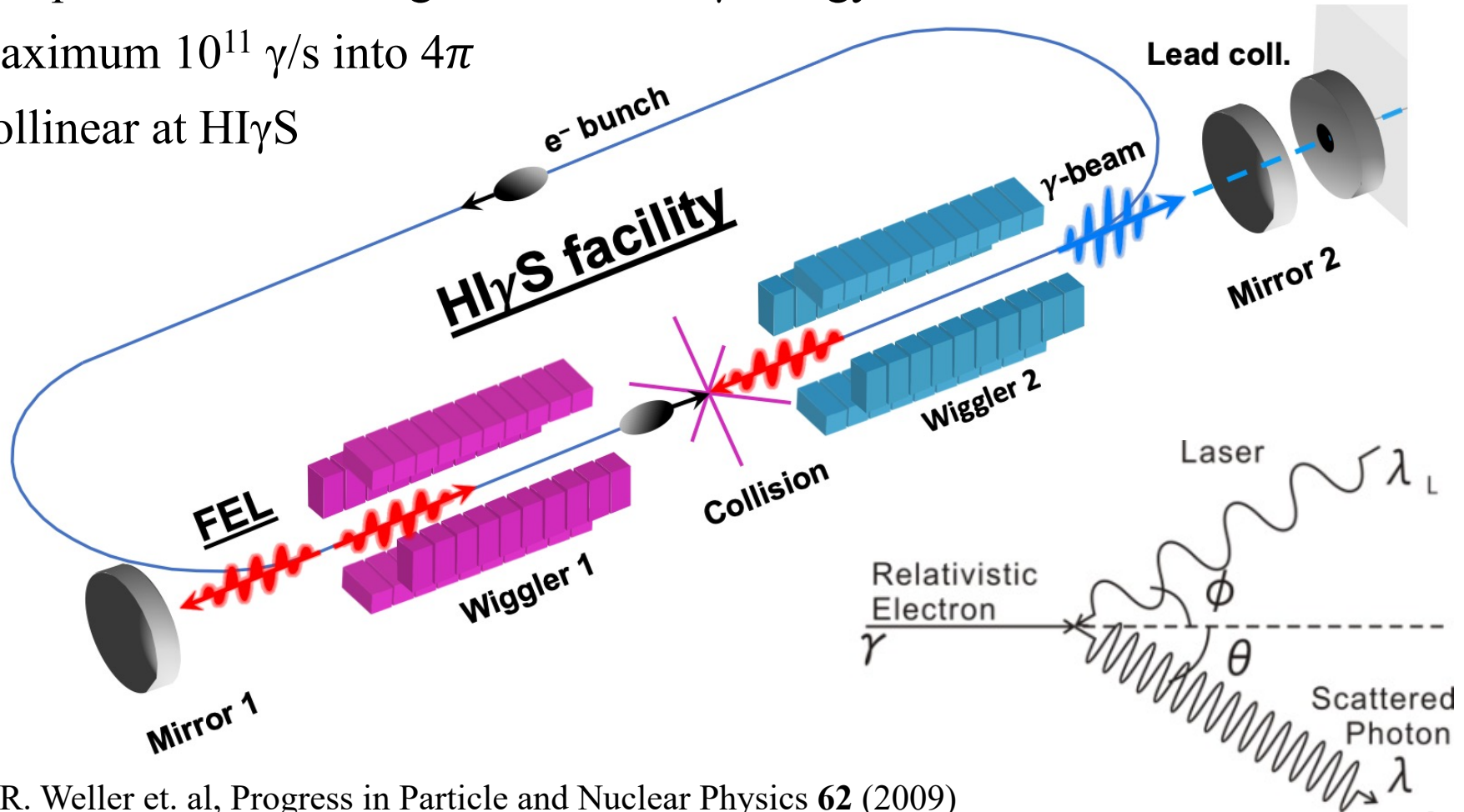
HIγS facility

- Quasi-monoenergetic γ -beams
- Energies from 1 to 100 MeV with 1.3–3.9 % energy resolution
- Linear and circular polarisation



H γ S facility

- Free electron laser – $\lambda = 190 - 1064$ nm
- Compton backscattering increases the γ energy
- Maximum 10^{11} γ /s into 4π
- Collinear at H γ S



[1] H.R. Weller et. al, Progress in Particle and Nuclear Physics **62** (2009)

[2] A. Endo, Laser Pulses-Theory, Technology, and Applications. InTech, (2012)

Stellar helium burning – $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

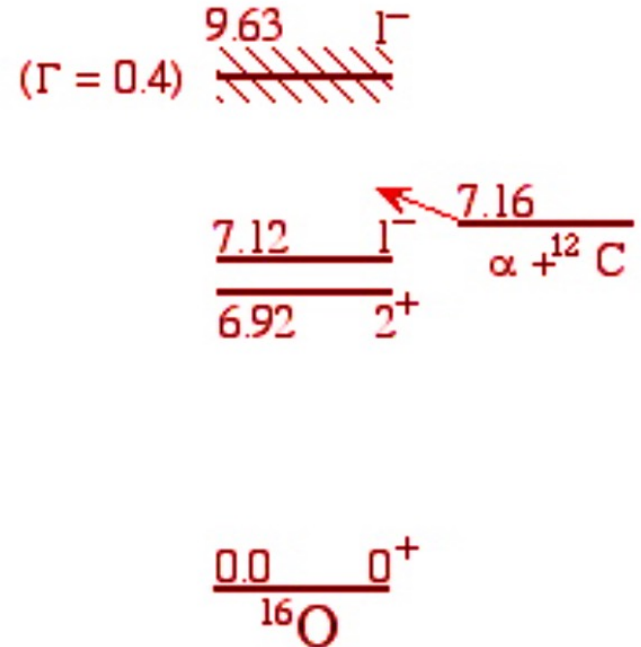
- The ratio of carbon-to-oxygen after helium burning
 - Composition of White Dwarfs
 - Yield of intermediate-mass isotopes (C, Ne, O burning etc.)
 - Explosive burning – light curves of Type Ia supernovae
 - Final states of massive stars – Type II supernovae

Stellar helium burning – $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

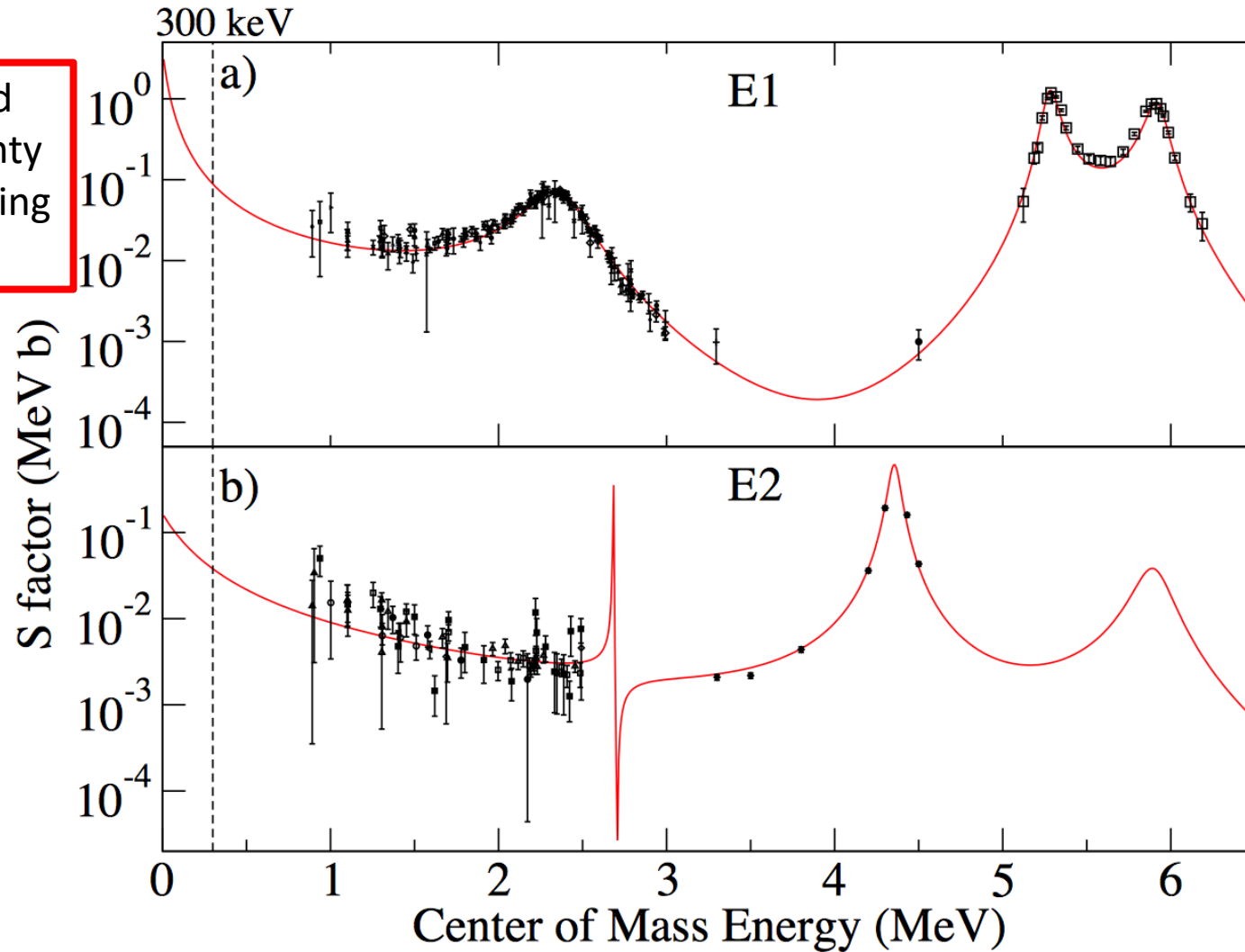
- The ratio of carbon-to-oxygen after helium burning

- Cross sections for $3\alpha \rightarrow ^{12}\text{C}$ and $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$

“The single most important nuclear physics uncertainty in astrophysics”



Data evaluation efforts



[4] R. J. deBoer et. al, The $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ reaction and its implications for stellar helium burning, Rev. Mod. Phys. **89** (2017)

New method: $^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$

- Measure reverse process – photodissociation of ^{16}O inside active target detector
- The γ -beam excites ^{16}O target to a known energy
- Measure the angular distributions of the emitted ^{12}C and α from the break-up – Time Projection Chamber

Method: $^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$

- Advantages
 - Higher cross section – detailed balance
 - Ability to measure precise angular distributions with TPC

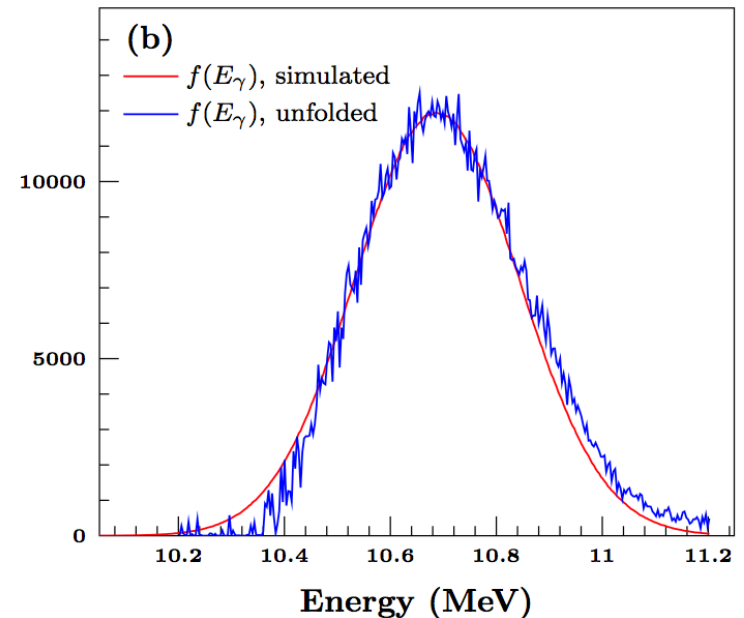
Method: $^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$

- Advantages

- Higher cross section – detailed balance
- Ability to measure precise angular distributions with TPC

- Disadvantages

- $\sigma = 130$ keV γ energy resolution
- Uncertainty in beam flux
- Rate in TPC



$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$ with an Optical TPC

- Proof-of-principle experiment
- Optical TPC operating with $\text{CO}_2 + \text{N}_2$ gas mixture
- Focused on higher energy region $E_{\text{cm}} = 2 - 3.5 \text{ MeV}$
 - cross section is higher and shapes of angular distributions vary significantly

ARTICLE



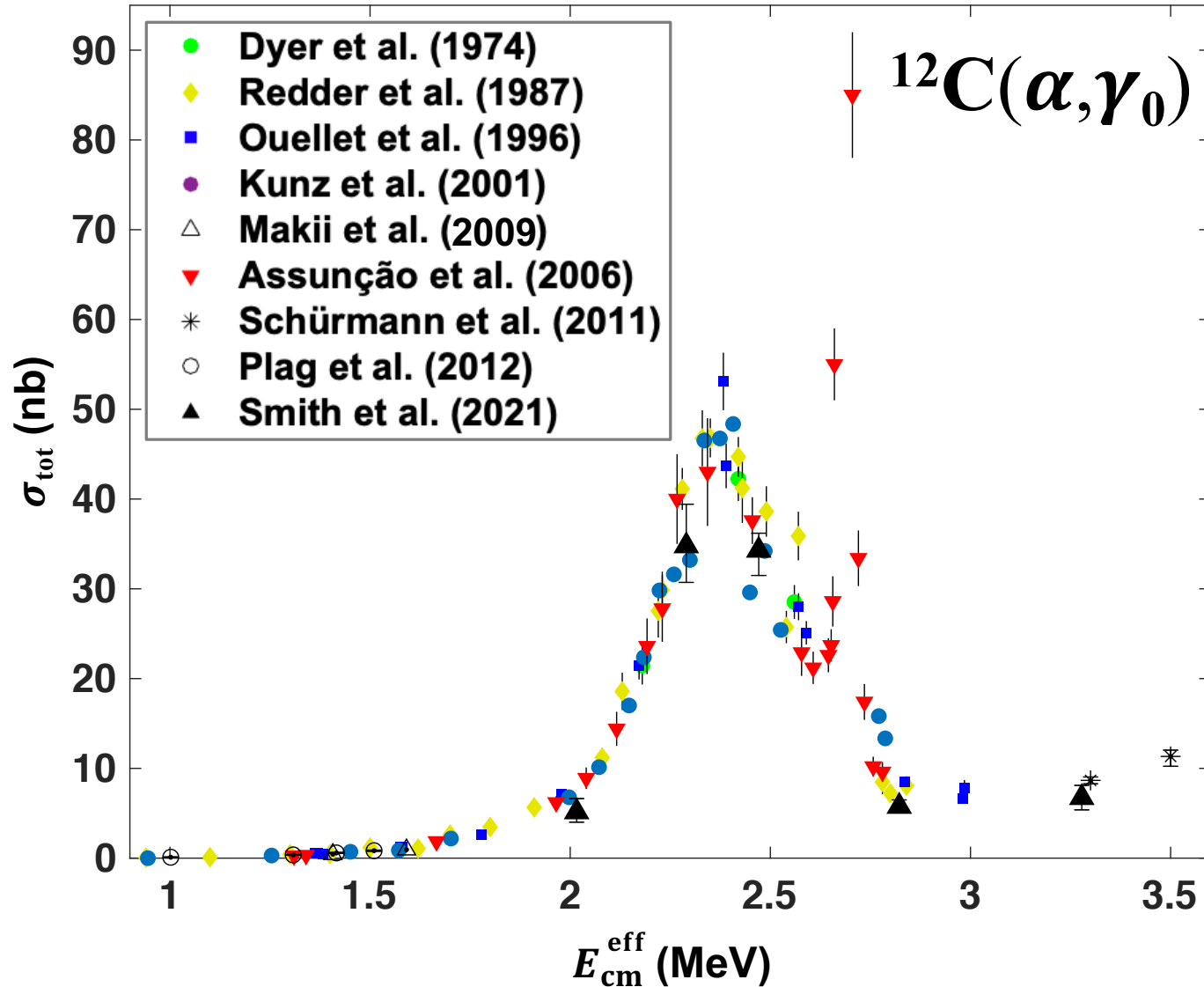
<https://doi.org/10.1038/s41467-021-26179-x>

OPEN

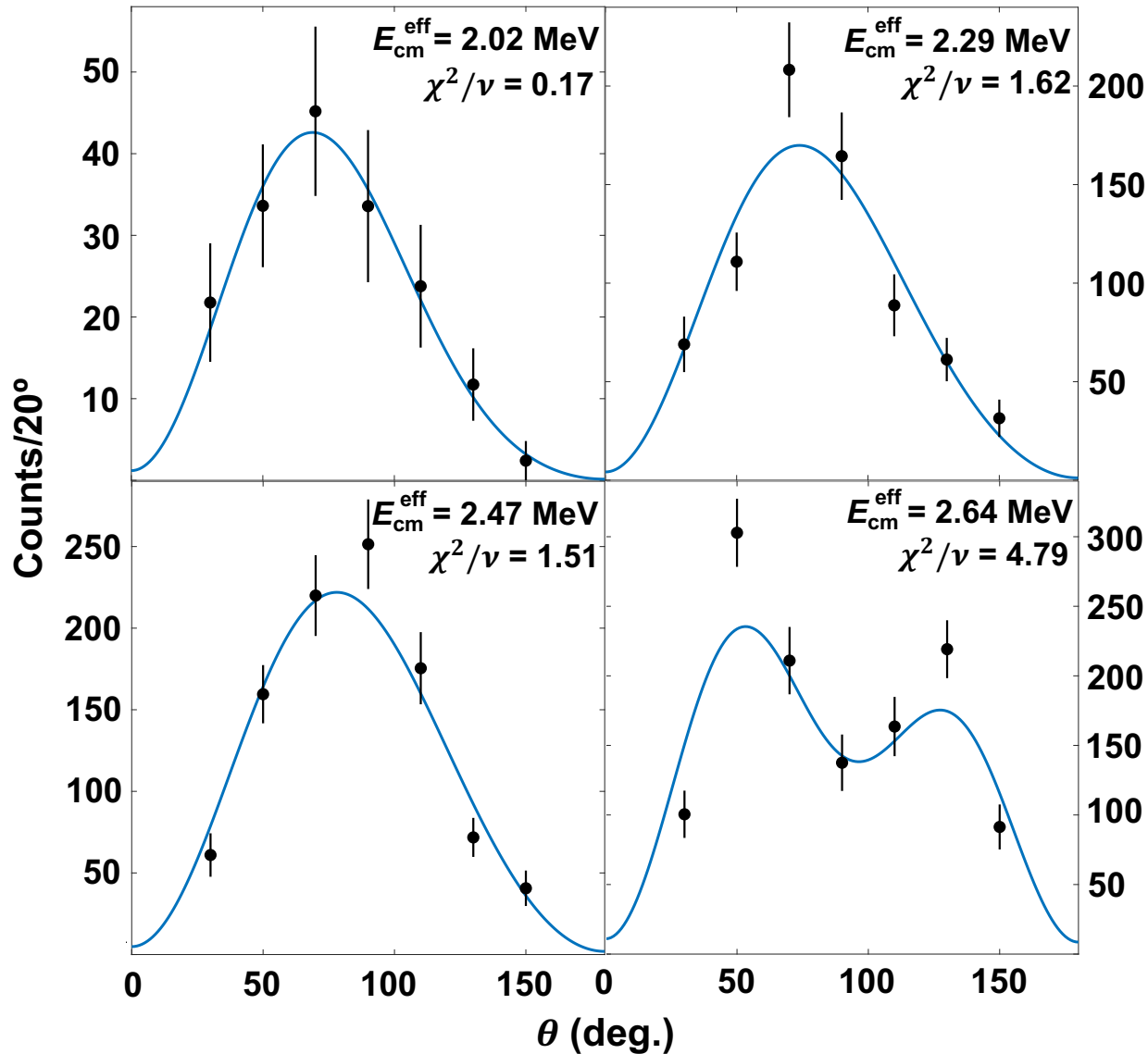
Precision measurements on oxygen formation in stellar helium burning with gamma-ray beams and a Time Projection Chamber

R. Smith ^{1,2}✉, M. Gai ², S. R. Stern ², D. K. Schweitzer ² & M. W. Ahmed^{3,4}

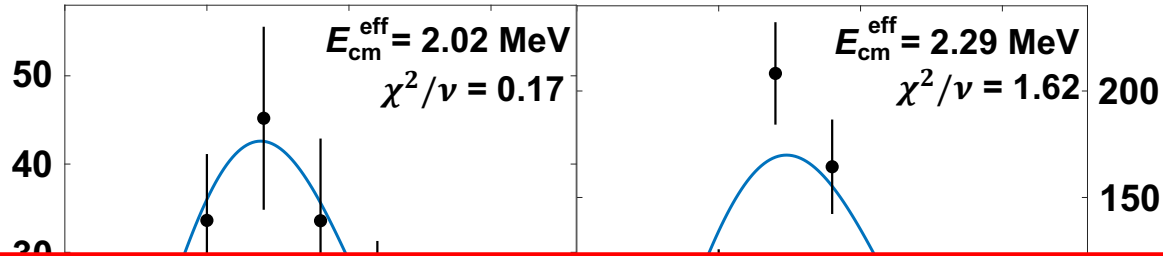
Total cross section



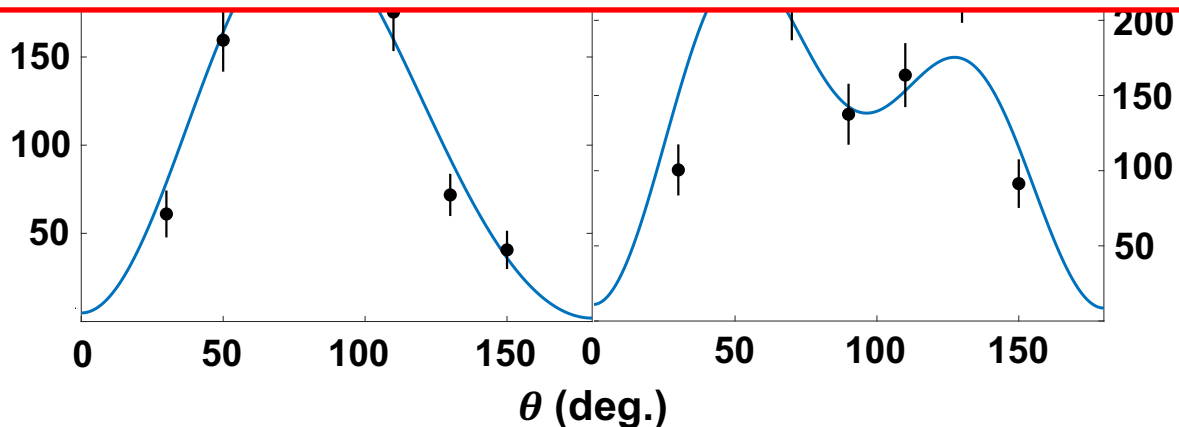
Angular distributions



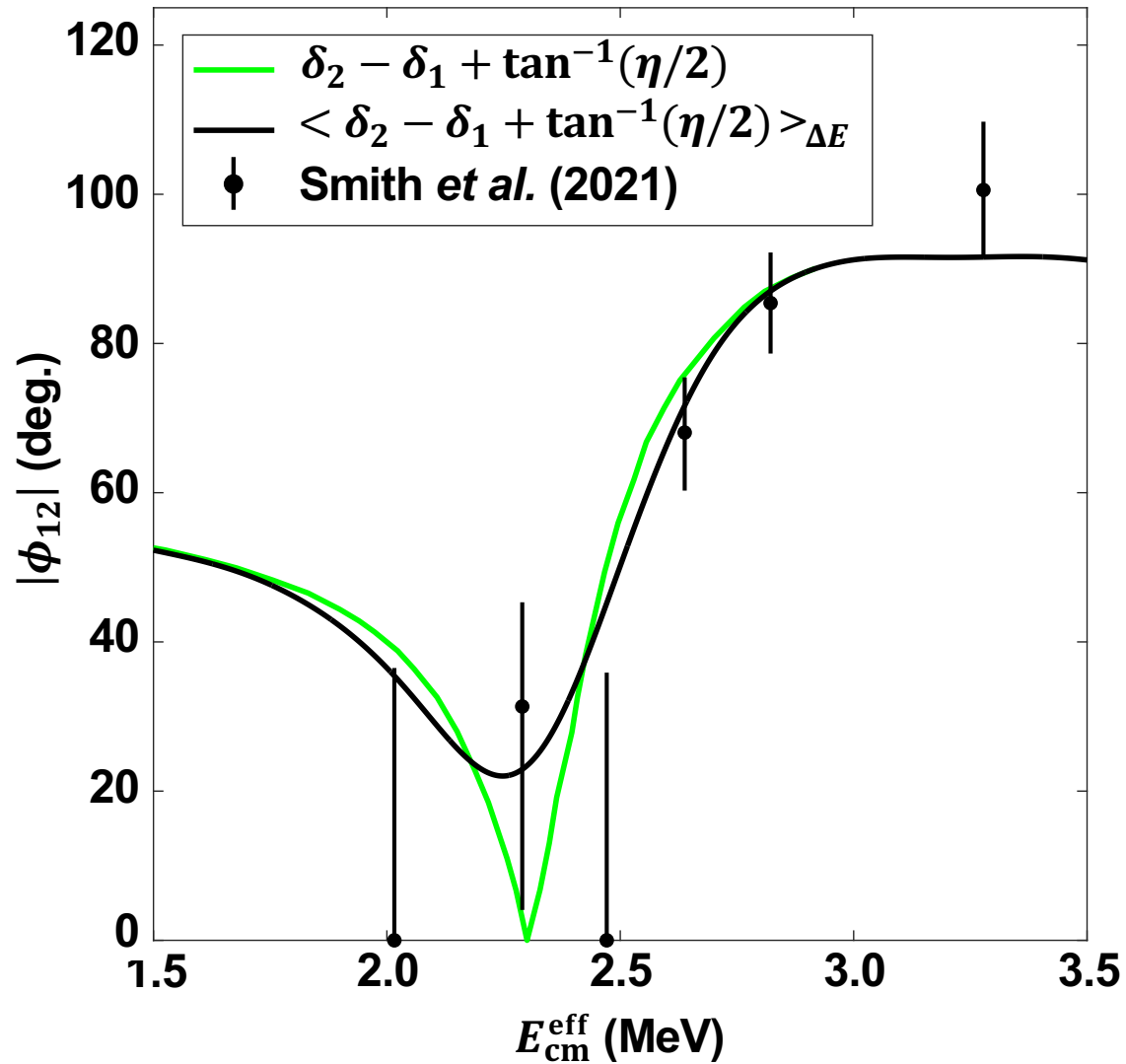
Angular distributions



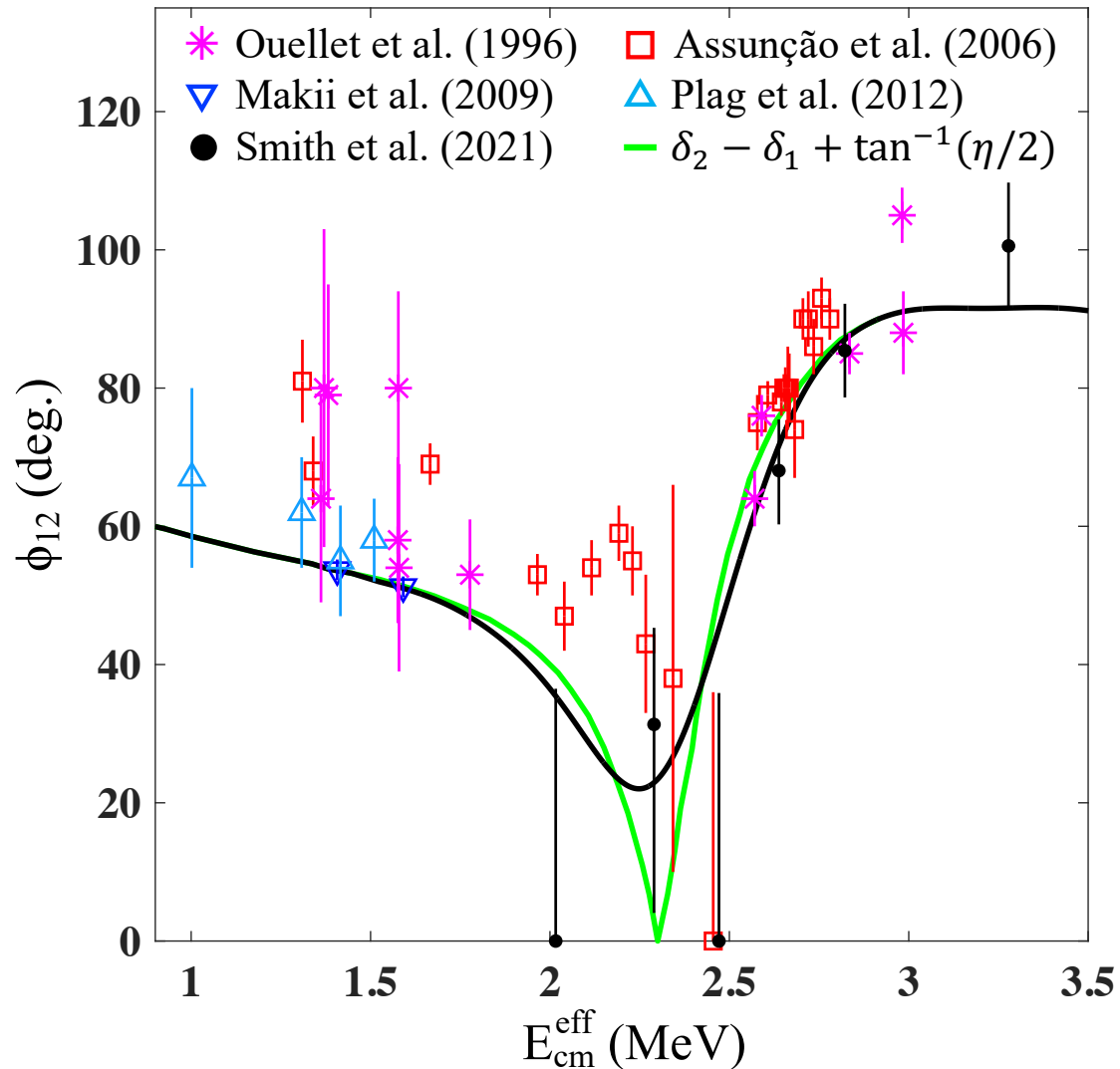
$$\begin{aligned}
 W(\theta) = & (3|A_{E1}|^2 + 5|A_{E2}|^2)P_0(\cos \theta) \\
 & + (25/7|A_{E2}|^2 - 3|A_{E1}|^2)P_2(\cos \theta) \\
 & - 60/7|A_{E2}|^2P_4(\cos \theta) \\
 & + 6\sqrt{3}|A_{E1}||A_{E2}|\cos \phi_{12} [P_1(\cos \theta) - P_3(\cos \theta)]
 \end{aligned}$$



ϕ_{12} values



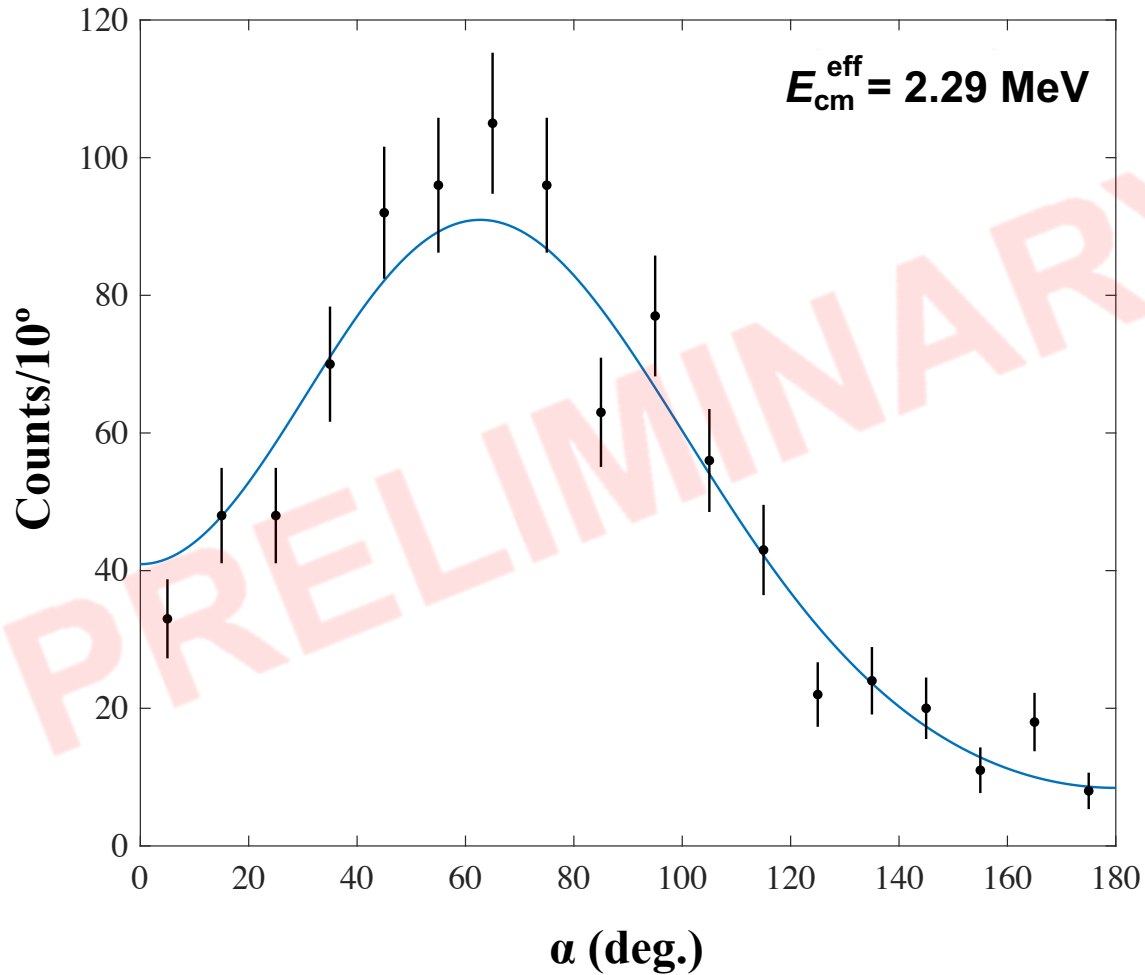
ϕ_{12} values



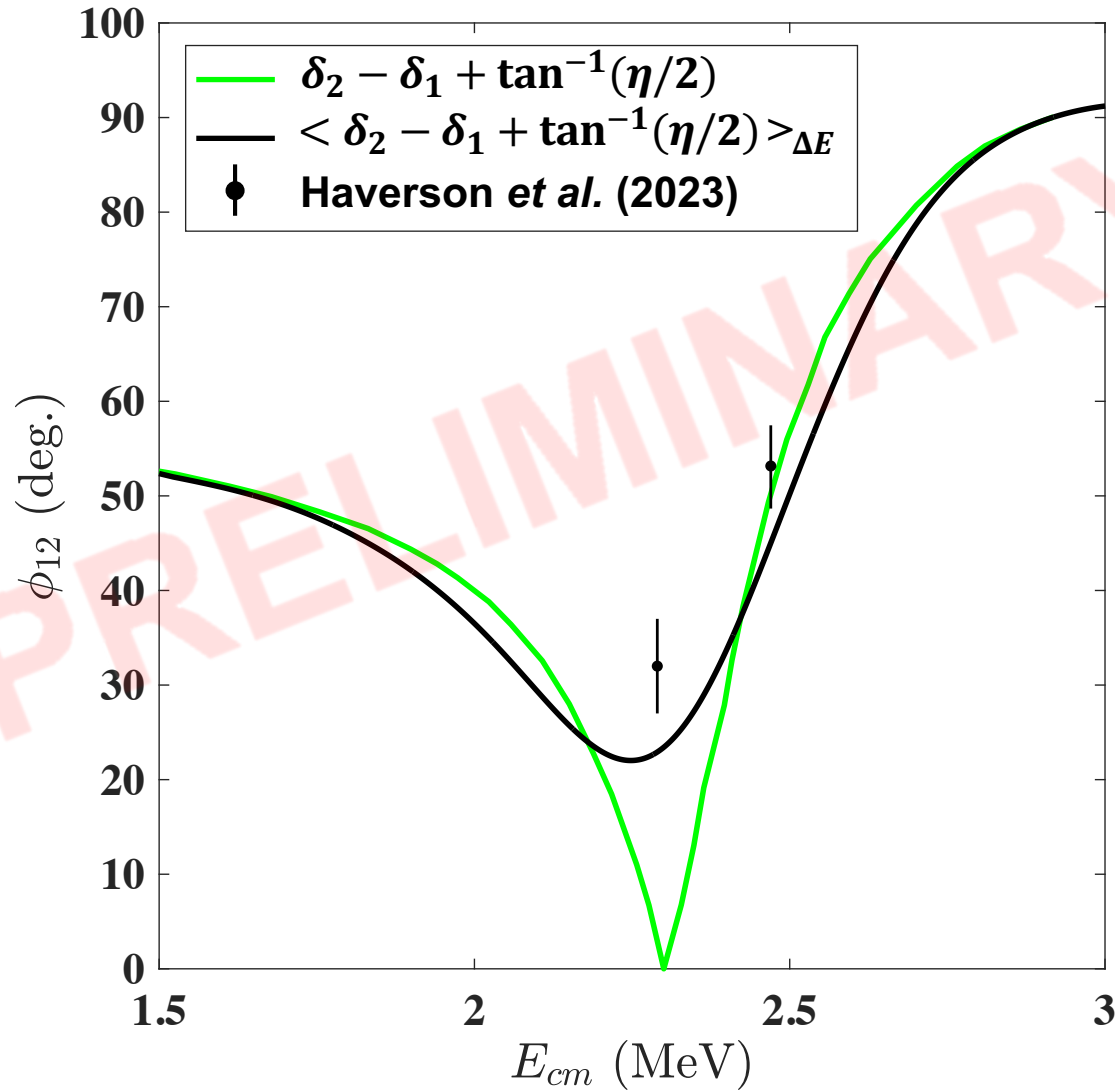
$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$ with an Optical TPC

- Later experiment performed – analysis by Kris Haverson (Ph.D. at Sheffield Hallam)
- Optical TPC operating with $\text{N}_2\text{O} + \text{N}_2$ gas mixture
- Focused on high energy region $E_{\text{cm}} = 2 - 3 \text{ MeV}$
- Higher statistics
- No carbon in the target – should permit simpler analysis

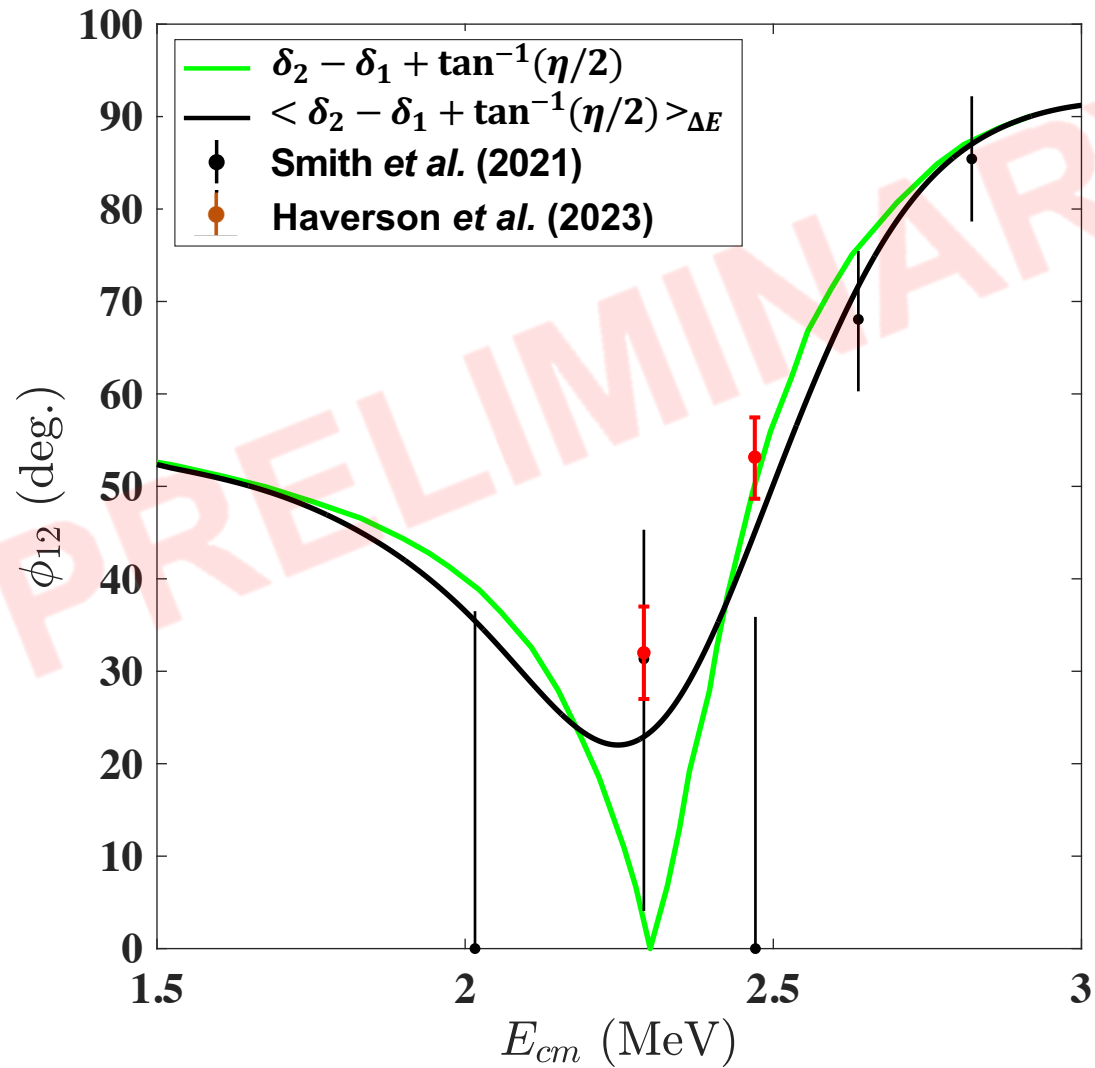
Angular distribution



ϕ_{12} values



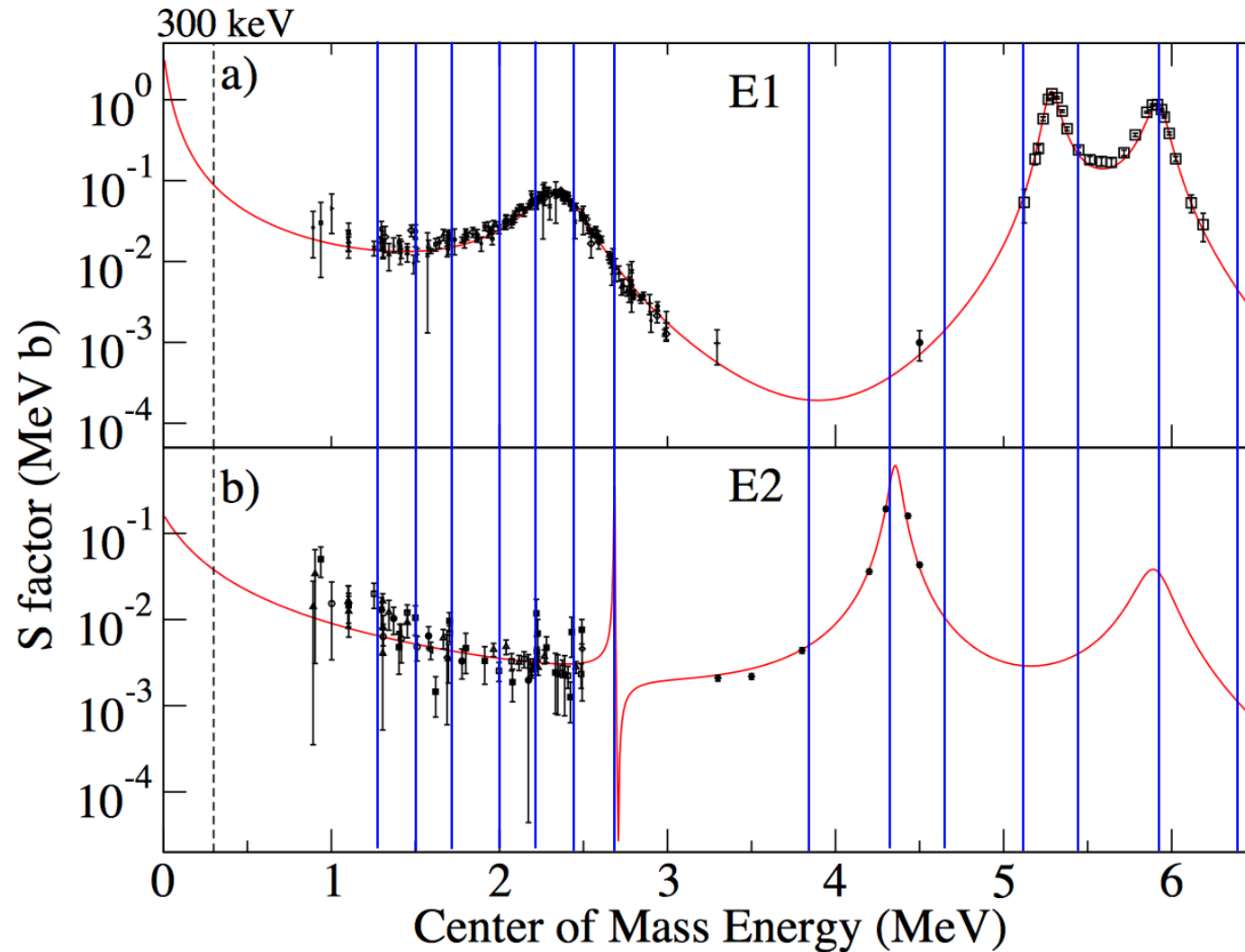
ϕ_{12} values



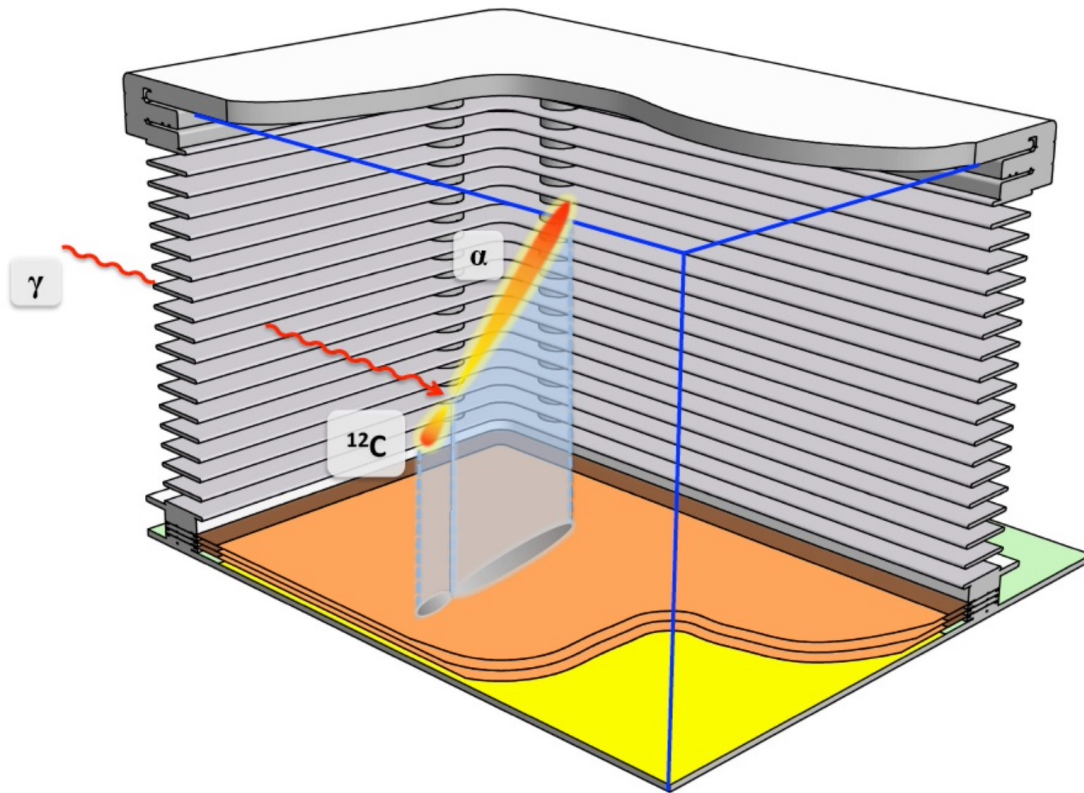
$^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$ with an electronic TPC

- New experiment performed at HIγS in 2022
- Electronic TPC operating with pure CO_2 gas
- TPC built by University of Warsaw
- $^{16}\text{O}(\gamma,\alpha)^{12}\text{C}$ measured at nominal $E_{\text{cm}} = 1.35 - 6.7$ MeV
- 275 hours of beamtime during 2022

$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$ with an electronic TPC



$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$ with an electronic TPC



Active volume

33 x 20 cm² (readout) x 20 cm (drift)

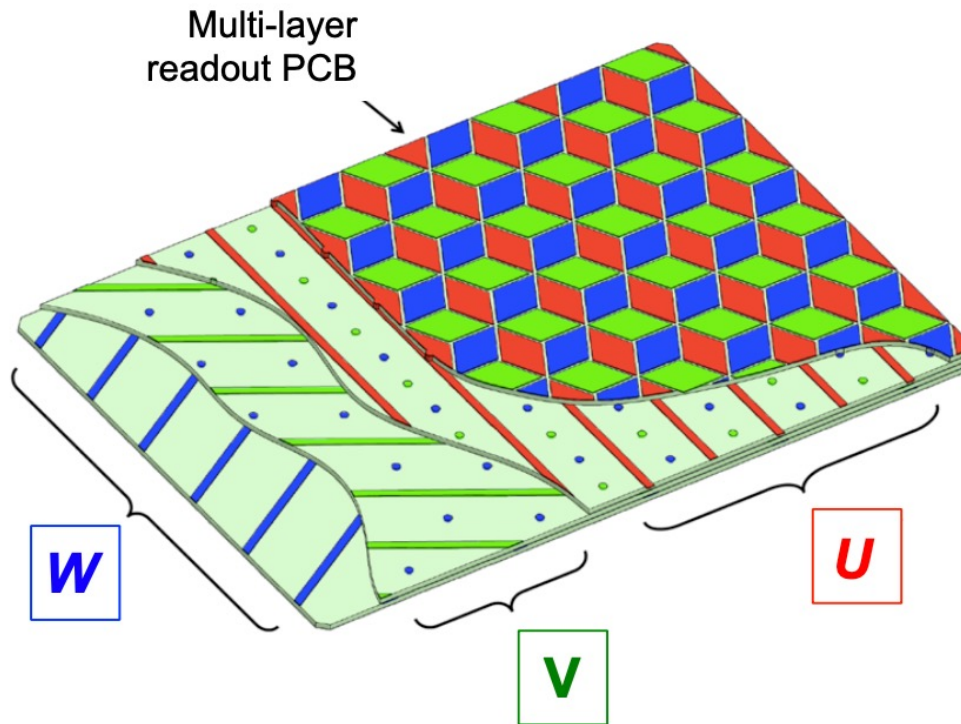
Charge amplification

Gas Electron Multiplier (GEM) structures

Readout

Planar, 3-coordinate, redundant strip arrays, ~1000 channels
GET electronics

$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$ with an electronic TPC



Active volume

33 x 20 cm² (readout) x 20 cm (drift)

Charge amplification

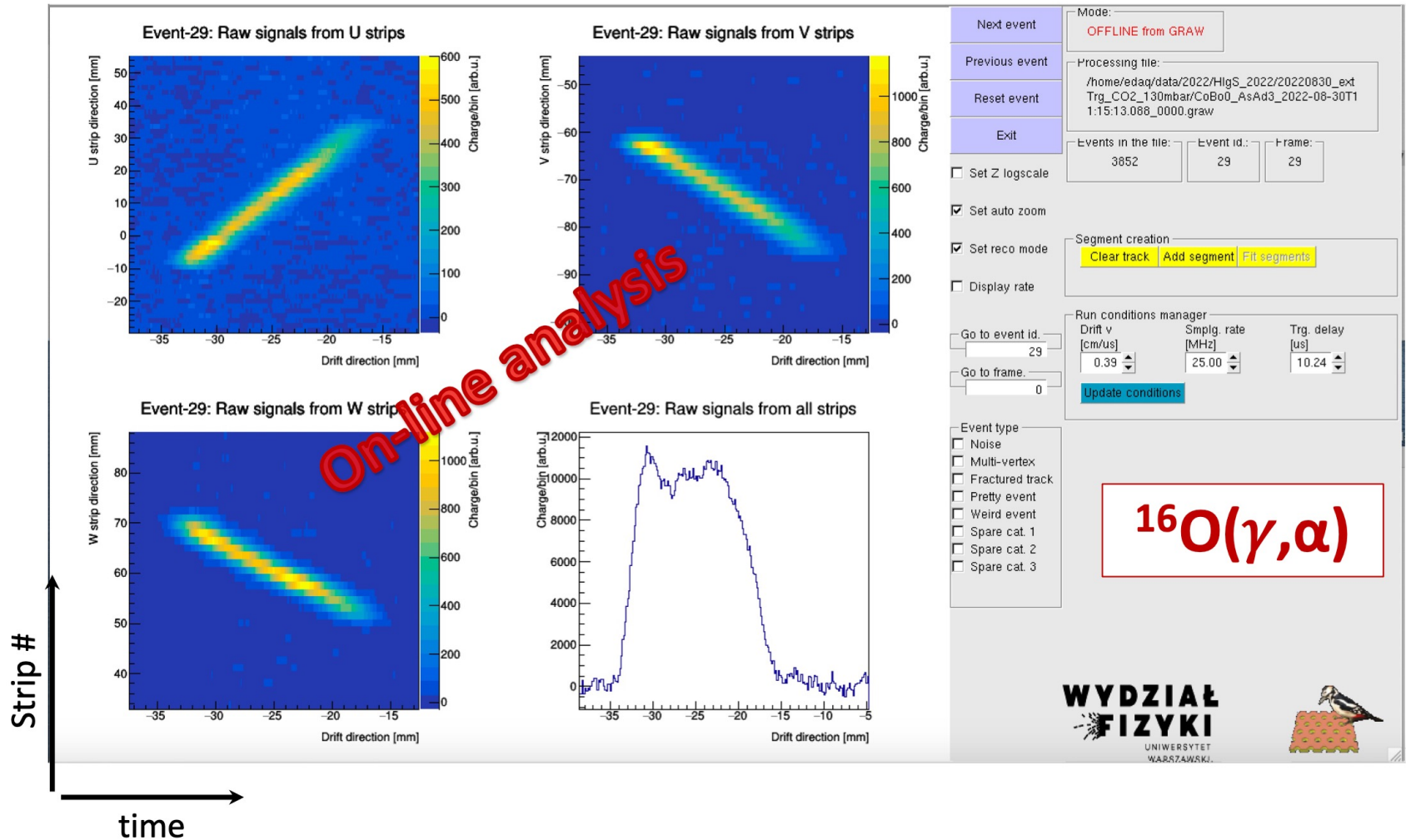
Gas Electron Multiplier (GEM) structures

Readout

Planar, 3-coordinate, redundant strip arrays, ~1000 channels
GET electronics

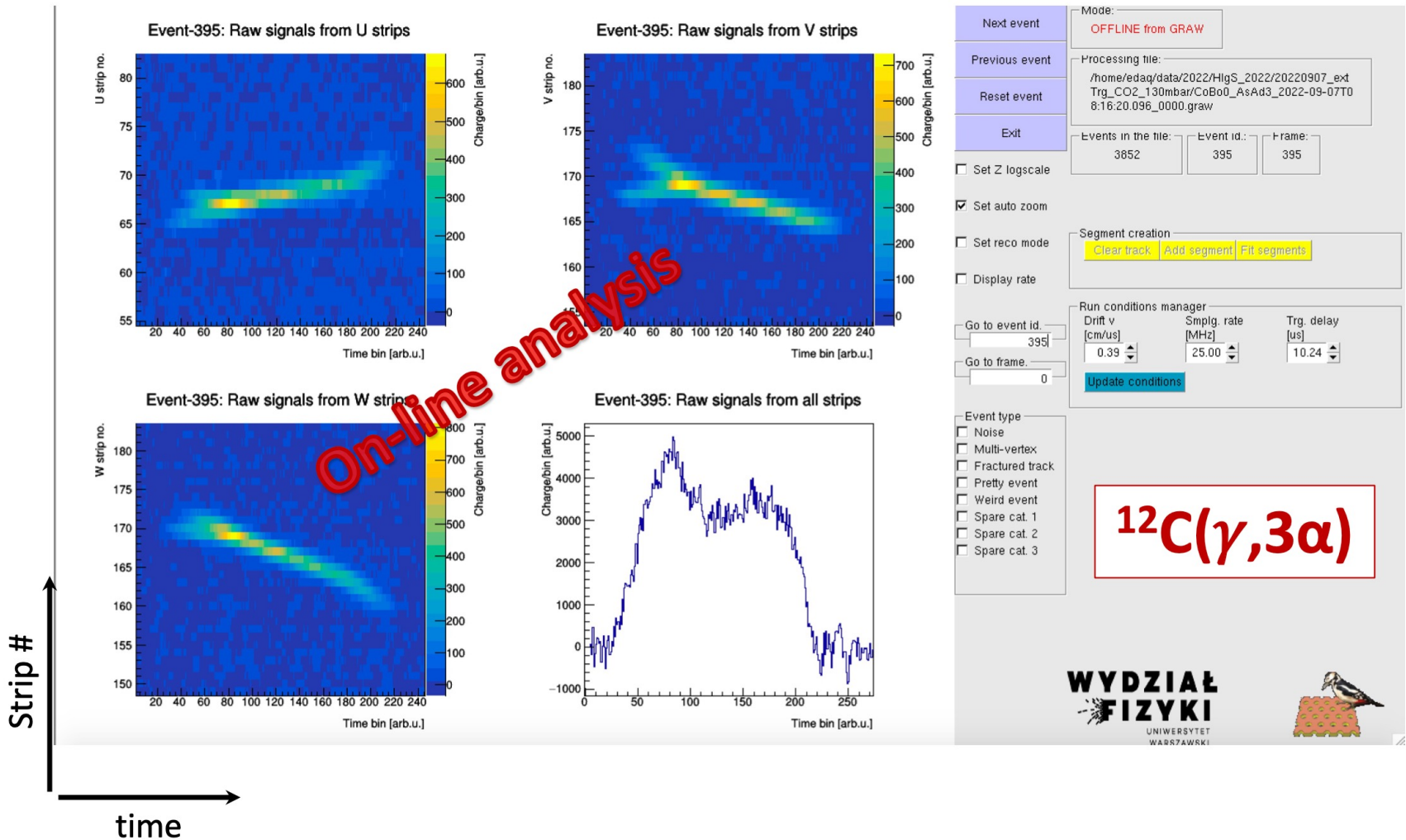
$^{16}\text{O}(\gamma, \alpha)^{12}\text{C}$ preliminary results

$E_\gamma = 8.66 \text{ MeV}$ ($E_{\text{cm}} = 1.5 \text{ MeV}$)

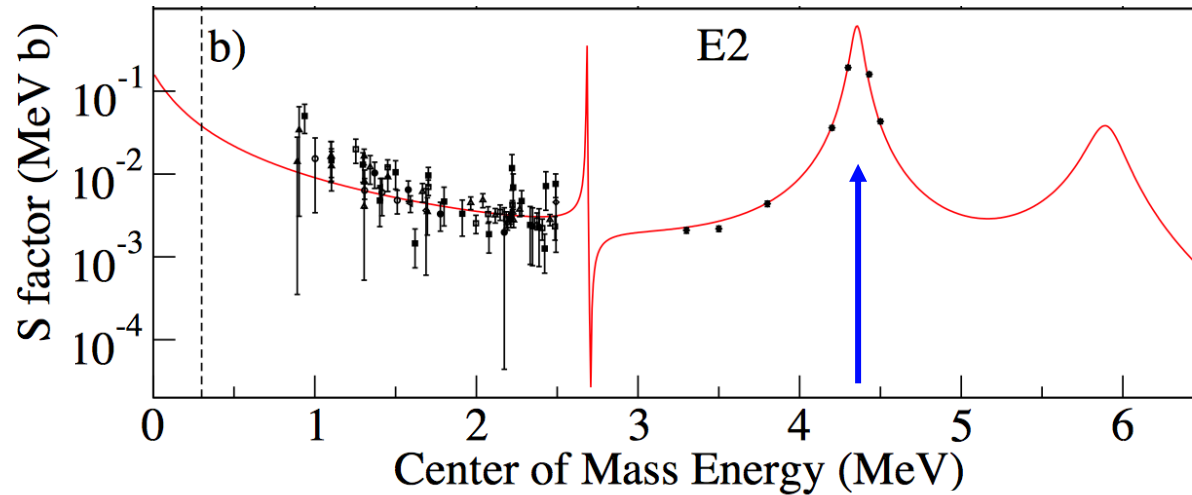


$^{12}\text{C}(\gamma, 3\alpha)$ preliminary results

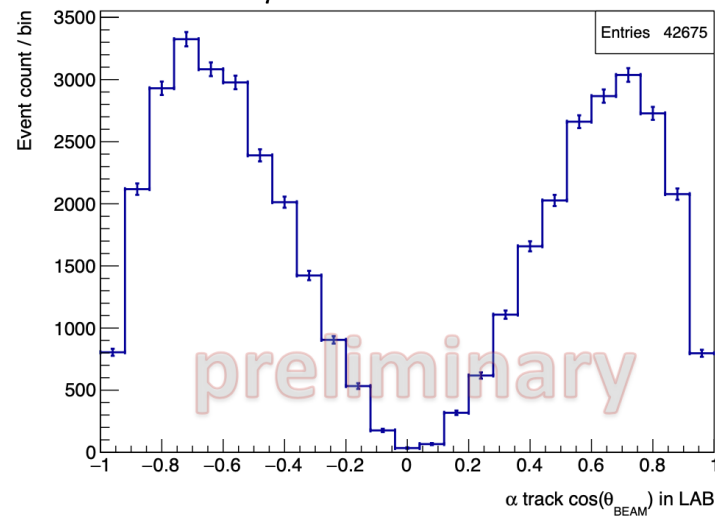
$E_\gamma = 8.51 \text{ MeV}$ ($E_{\text{cm}} = 1.35 \text{ MeV}$)



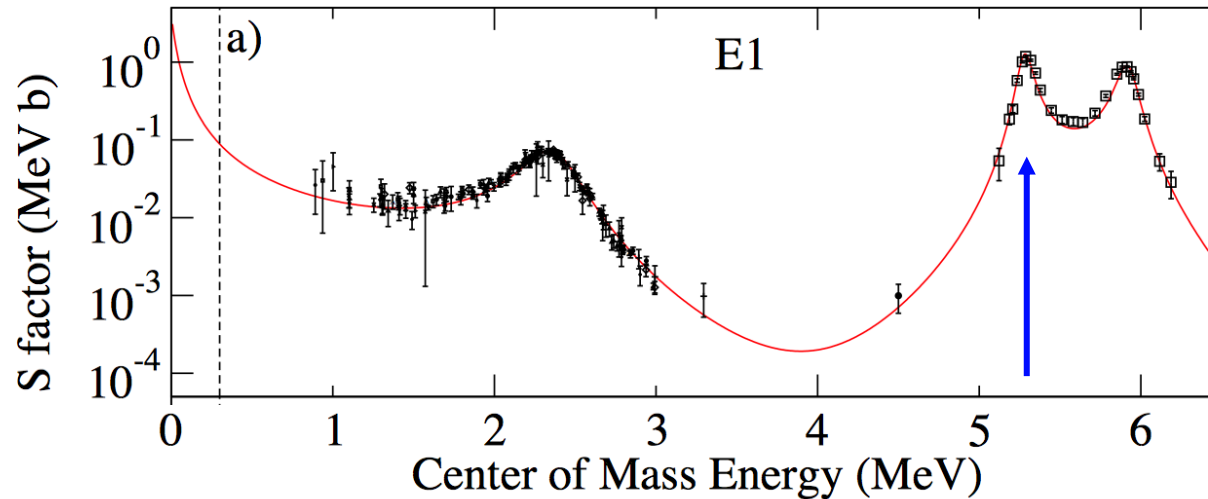
$^{16}\text{O}(\gamma,\alpha)$ prelim. ang. distributions



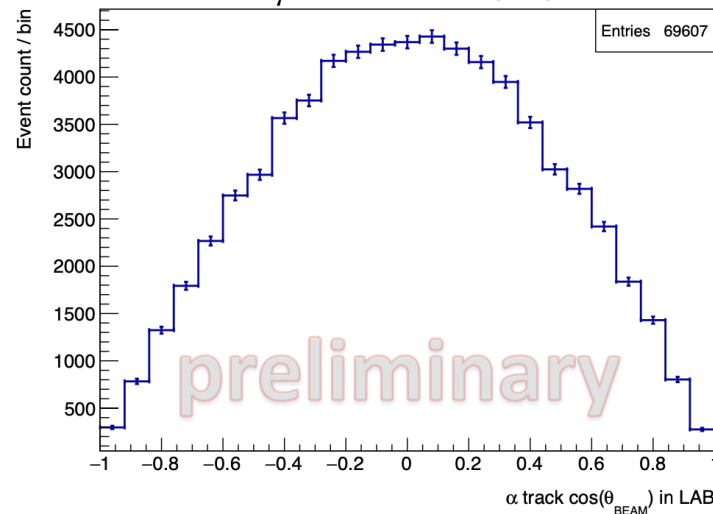
$E_\gamma = 11.5\text{MeV}$ (E2)



$^{16}\text{O}(\gamma,\alpha)$ prelim. ang. distributions



$E_\gamma = 12.3\text{MeV}$ (E1)



Summary

- Sheffield Hallam involved in several exciting studies in nuclear astrophysics using γ -beams and TPC detectors
 - Leading analysis of existing data
 - New experiments at HI γ S
- Publication: [R. Smith et al. Nature Communications 12, 5920 \(2021\)](#).
 - Viability of γ -beam experiments for measuring $^{12}\text{C}(\alpha,\gamma)$ using TPCs
 - Newer Optical TPC data with N_2O gas mixture are promising
- State-of-the-art electronic TPC data are undergoing analysis
 - Data analysis led by Warsaw with contributions from SHU and UConn
- Nuclear structure studies on carbon-12 – $^{12}\text{C}(\gamma,\alpha)$

Collaborators

Mikołaj Ćwiok¹, Wojciech Dominik¹, Aleksandra Fijałkowska¹, Mateusz Fila¹, Zenon Janas¹, Artur Kalinowski¹, Krzysztof Kierzkowski¹, Magdalena Kuich¹, Chiara Mazzocchi^{1,}, Wojciech Okliński¹, Marcin Zaremba¹, Moshe Gai², Deran K. Schweitzer², Sarah R. Stern², Sean Finch^{3,4}, Udo Friman-Gayer^{3,4,**}, Samantha R. Johnson^{5,4}, Tyler M. Kowalewski^{5,4}, Dimiter L. Balabanski⁶, Catalin Matei⁶, Adrian Rotaru⁶, Kristian C.Z. Haverson⁷, Robin Smith⁷, Ross A.M. Allen⁸, Mark R. Griffiths⁸, Stuart Pirrie⁸, and Pedro Santa Rita Alcibia⁸*

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⁸School of Physics and Astronomy, University of Birmingham, Birmingham, UK