

# Special Relativity

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# Overview

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- The principle of special relativity
- Lorentz transformation and consequences
- Space-time
- 4-vectors: position, velocity, momentum, invariants, covariance.
- Derivation of  $E=mc^2$
- Examples of the use of 4-vectors
- Inter-relation between  $\beta$  and  $\gamma$ , momentum and energy
- An accelerator problem in relativity
- Photons and wave 4-vector
- Relativistic particle dynamics
- Lagrangian and Hamiltonian Formulation
- Radiation from an Accelerating Charge
- Motion faster than speed of light



# Reading

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- W. Rindler: Introduction to Special Relativity (OUP 1991)
- D.F. Lawden: An Introduction to Tensor Calculus and Relativity (Dover, 2003)
- N.M.J. Woodhouse: Special Relativity (Springer 2002)
- A.P. French: Special Relativity, MIT Introductory Physics Series (Nelson Thomes)
- C.Misner, K.Thorne and J.Wheeler: Relativity (Freeman, 1973)
- C.R. Prior: Special Relativity, CERN Accelerator School (Zeegse)

# Historical Background

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- Groundwork of Special Relativity laid by **Lorentz** in studies of electrodynamics, with crucial concepts contributed by **Einstein** to place the theory on a consistent footing.
- **Maxwell's** equations (1863) attempted to explain electromagnetism and optics through wave theory
  - light propagates with speed  $c = 3 \times 10^8$  m/s in “ether” but with different speeds in other frames
  - the ether exists solely for the transport of e/m waves
  - Maxwell's equations not invariant under Galilean transformations
- To avoid setting e/m apart from classical mechanics, assume
  - light has speed  $c$  only in frames where source is at rest
  - the ether has a small interaction with matter and is carried along with astronomical objects such as the Earth



# Contradicted by Experiment

- Aberration of star light (small shift in apparent positions of distant stars)
- Fizeau's 1859 experiments on velocity of light in liquids
- Michelson-Morley 1907 experiment to detect motion of the earth through ether
- Suggestion: perhaps material objects contract in the direction of their motion

$$L(v) = L_0 \left( 1 - \frac{v^2}{c^2} \right)^{1/2}$$

This was the last gasp of ether advocates and the germ of Special Relativity led by Lorentz, Minkowski and Einstein.



# The Principle of Special Relativity

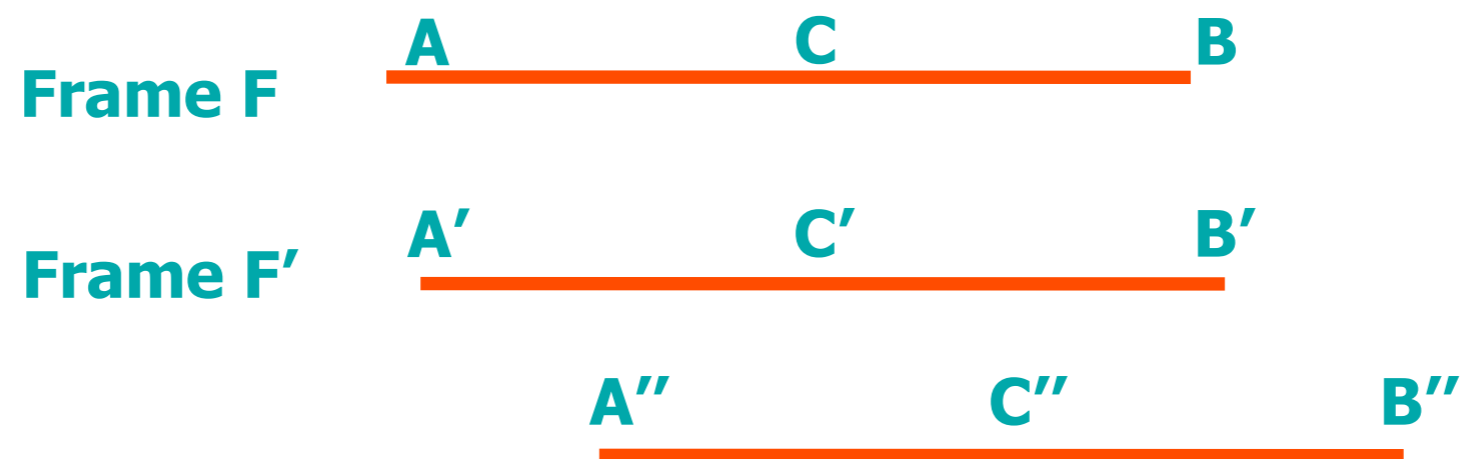
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- A frame in which particles under no forces move with constant velocity is *inertial*.
- Consider relations between inertial frames where measuring apparatus (rulers, clocks) can be transferred from one to another: *related frames*.
- Assume:
  - Behaviour of apparatus transferred from F to F' is independent of mode of transfer
  - Apparatus transferred from F to F', then from F' to F'', agrees with apparatus transferred directly from F to F''.
- *The Principle of Special Relativity states that all physical laws take equivalent forms in related inertial frames, so that we cannot distinguish between the frames.*



# Simultaneity

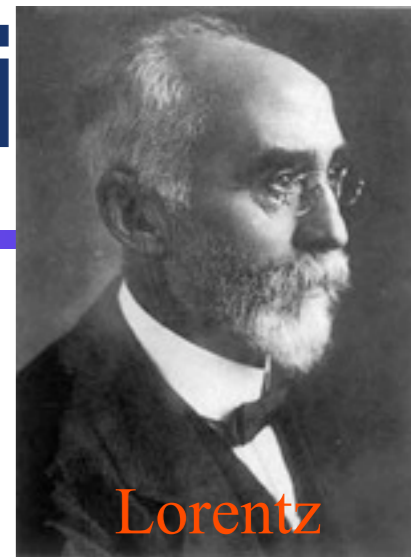
- Two clocks A and B are synchronised if light rays emitted at the same time from A and B meet at the mid-point of AB



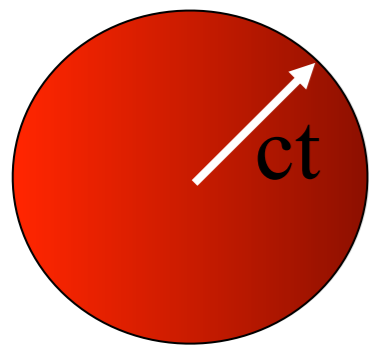
- Frame F' moving with respect to F. Events simultaneous in F cannot be simultaneous in F'.
- Simultaneity is **not absolute** but frame dependent.



# The Lorentz Transformation



- Must be linear to agree with standard Galilean transformation in low velocity limit
- Preserves wave fronts of pulses of light,



*i.e.*  $P = x^2 + y^2 + z^2 - c^2t^2 = 0$   
whenever  $Q = x'^2 + y'^2 + z'^2 - c^2t'^2 = 0$



- Solution is the **Lorentz transformation** from frame  $F(t,x,y,z)$  to frame  $F'(t',x',y',z')$  moving with velocity  $\mathbf{v}$  along the  $x$ -axis:



$$\begin{aligned}t' &= \gamma \left( t - \frac{vx}{c^2} \right) \\x' &= \gamma(x - vt) \\y' &= y \\z' &= z\end{aligned}$$

where  $\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}}$





# Outline of Derivation

Set  $t' = \alpha t + \beta x$

$$x' = \gamma x + \delta t$$

$$y' = \epsilon y$$

$$z' = \zeta z$$

where  $\alpha, \beta, \gamma, \delta, \epsilon, \zeta$   
are constants

$$kP = Q$$

$$\iff k(c^2 t^2 - x^2 - y^2 - z^2) = c^2 t'^2 - x'^2 - y'^2 - z'^2$$

$$= c^2 (\alpha t + \beta x)^2 - (\gamma x + \delta t)^2 - \epsilon^2 y^2 - \zeta^2 z^2$$

Equate coefficients of  $x, y, z, t$

Impose isotropy of space  $\implies k = k(\vec{v}) = k(|\vec{v}|) = \pm 1$

Apply some common sense (e.g.  $\epsilon, \zeta, k = +1$  and not  $-1$ )



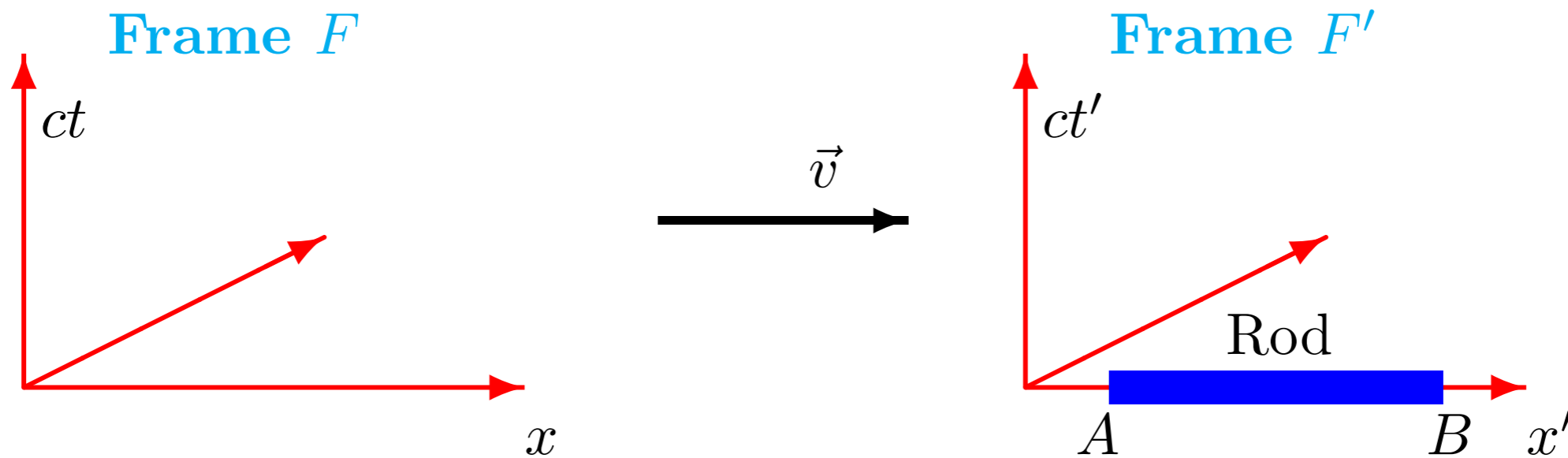
# General 3D form of Lorentz Transformation:

$$\vec{x}' = \vec{x} - \vec{v} \left( \gamma t - (\gamma - 1) \frac{\vec{v} \cdot \vec{x}}{v^2} \right)$$

$$t' = \gamma \left( t - \frac{\vec{v} \cdot \vec{x}}{c^2} \right)$$



# Consequences: length contraction



A rod  $AB$  of length  $L'$ , fixed in frame  $F'$  at  $x'_A, x'_B$ . What is its length measured in  $F$ ?

Must measure position of ends in  $F$  at same time, so events in  $F$  are  $(ct, x_A)$  and  $(ct, x_B)$ .

By Lorentz:

$$\left. \begin{aligned} x'_A &= \gamma(x_A - vt) \\ x'_B &= \gamma(x_B - vt) \end{aligned} \right\} \implies \begin{aligned} L' &= x'_B - x'_A \\ &= \gamma(x_B - x_A) \\ &= \gamma L > L \end{aligned}$$

**Moving objects appear contracted in the direction of the motion**

# Consequences: time dilation

- Clock in frame  $F$  at point with coordinates  $(x,y,z)$  at different times  $t_A$  and  $t_B$



- In frame  $F'$  moving with speed  $v$ , Lorentz transformation gives

$$t'_A = \gamma \left( t_A - \frac{vx}{c^2} \right) \quad t'_B = \gamma \left( t_B - \frac{vx}{c^2} \right)$$

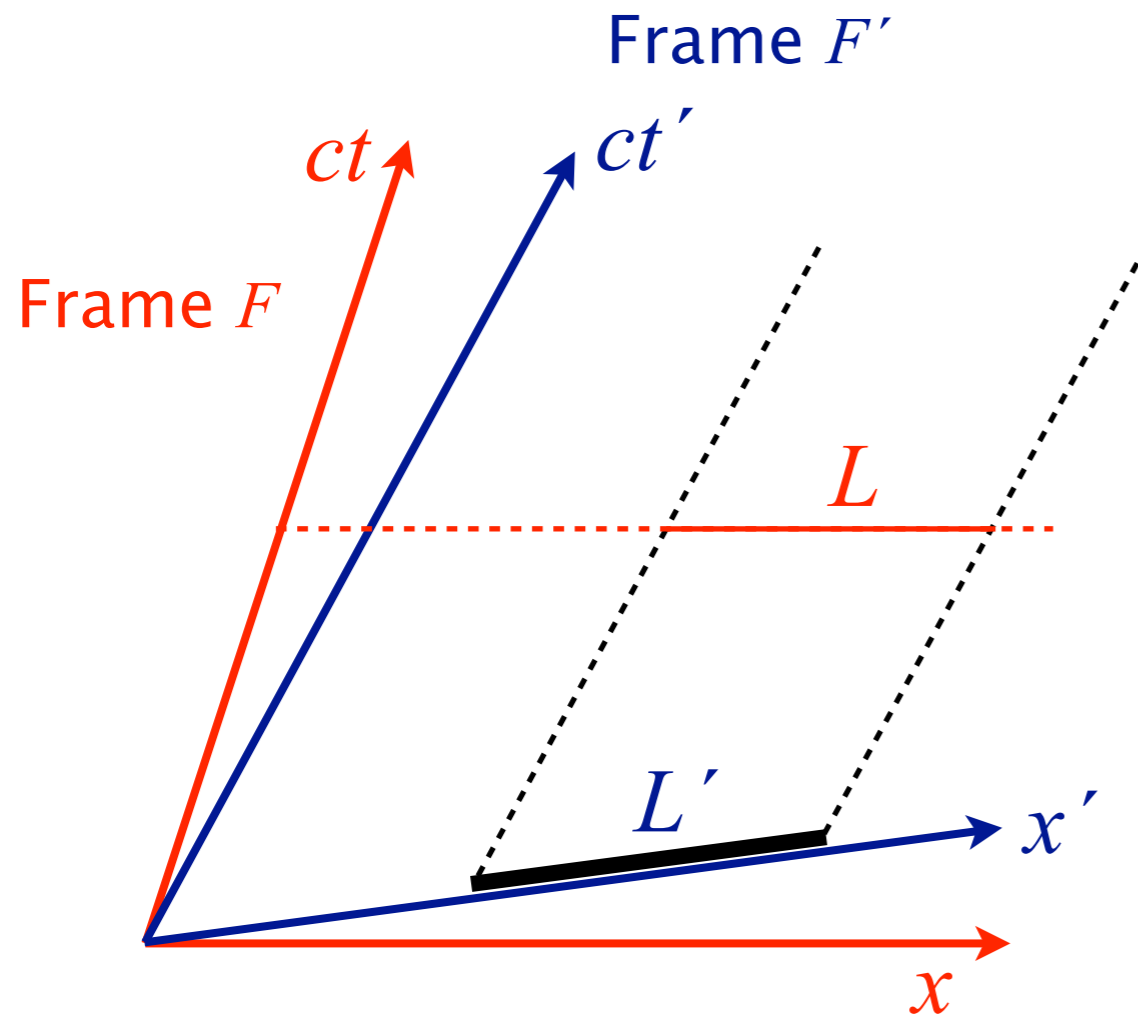
- So

$$\Delta t' = t'_B - t'_A = \gamma (t_B - t_A) = \gamma \Delta t > \Delta t$$

**Moving clocks appear to run slow**

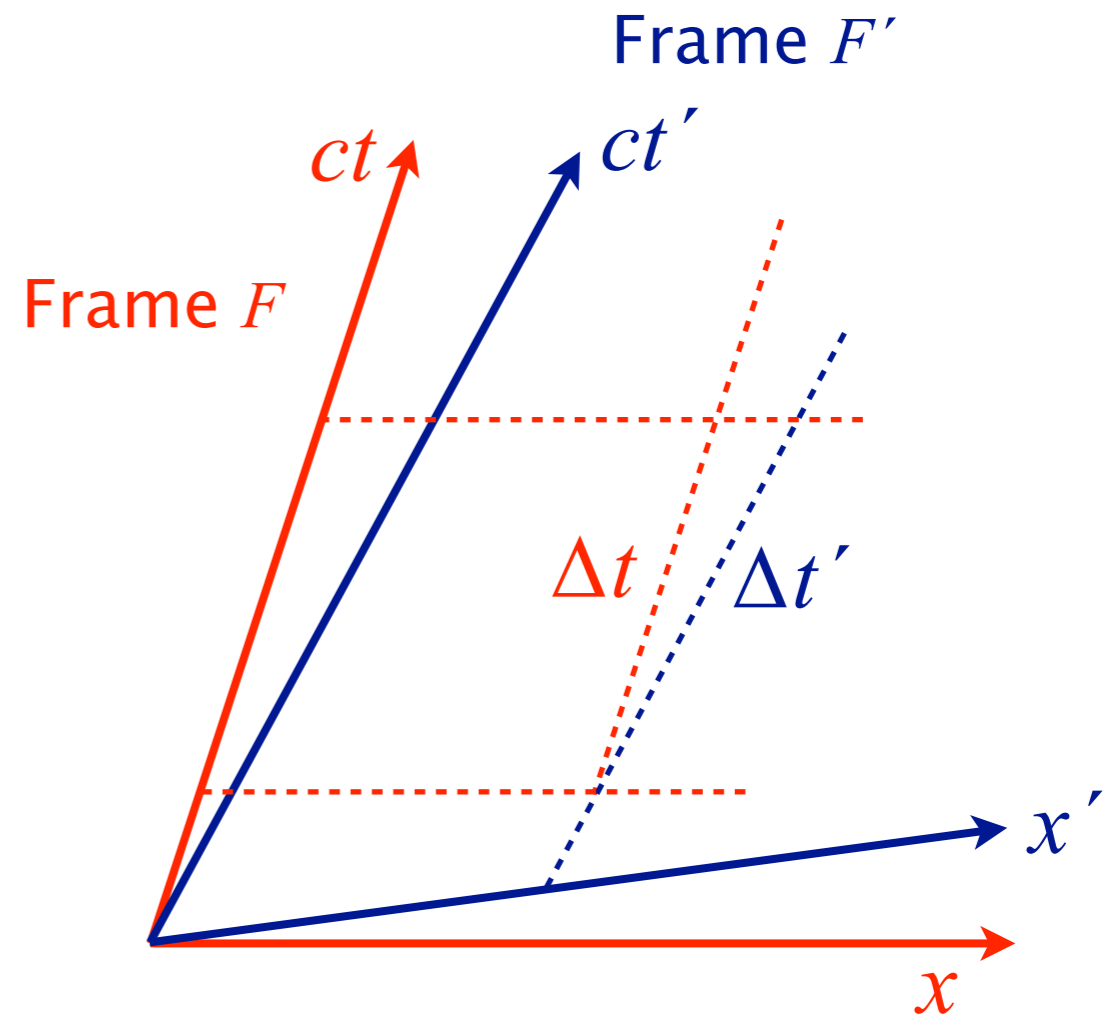


# Schematic Representation



Length contraction  $L < L'$

Rod at rest in  $F'$ . Measurements in  $F$  at a fixed time  $t$ , along a line parallel to  $x$ -axis

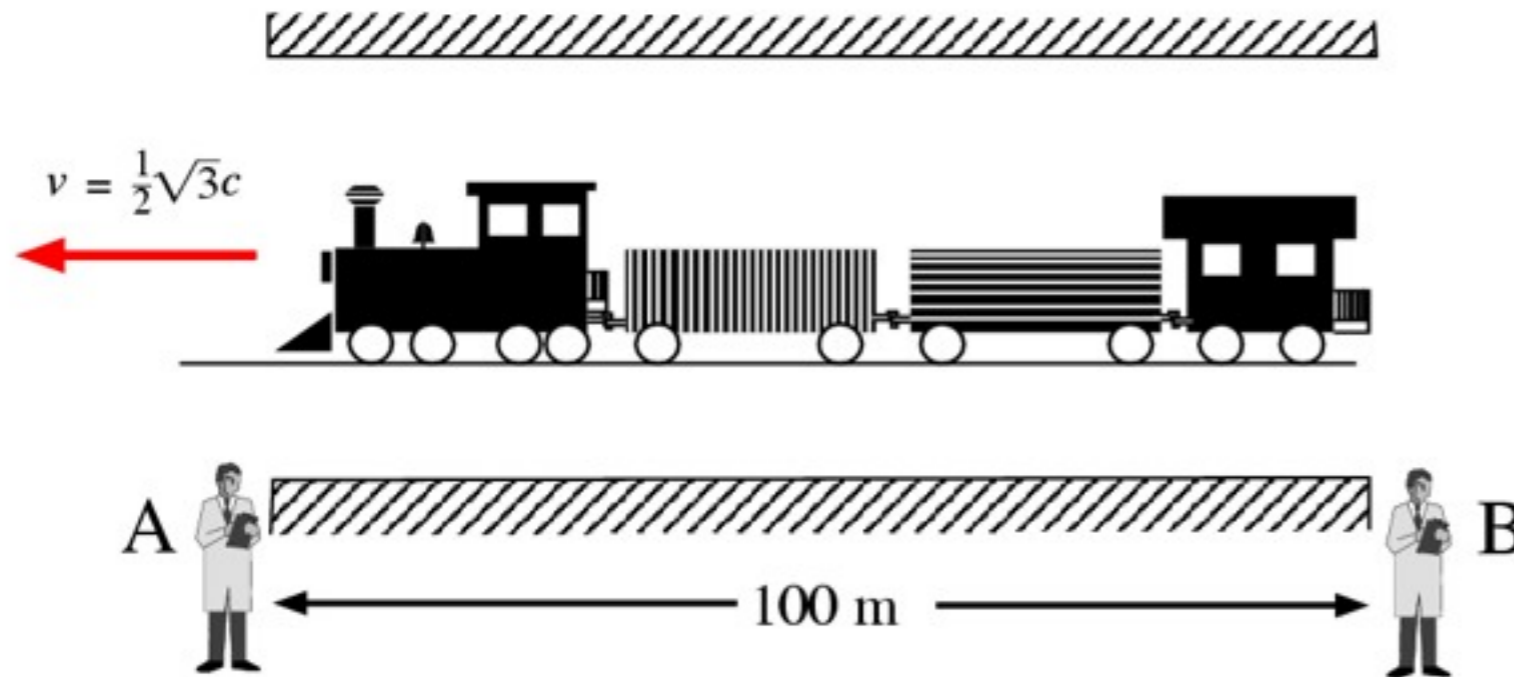


Time dilation  $\Delta t < \Delta t'$

Clock at rest in  $F$ . Time difference in  $F'$  from line parallel to  $t'$ -axis



# Example: High Speed Train



All clocks synchronised.

A's clock and driver's clock read 0 as front of train emerges from tunnel.

- Observers A and B at exit and entrance of tunnel say the train is moving, has contracted and has length

$$\frac{100}{\gamma} = 100 \times \left(1 - \frac{v^2}{c^2}\right)^{\frac{1}{2}} = 100 \times \left(1 - \frac{3}{4}\right)^{\frac{1}{2}} = 50\text{m}$$

- But the tunnel is moving relative to the driver and guard on the train and they say the train is 100 m in length but the tunnel has contracted to 50 m

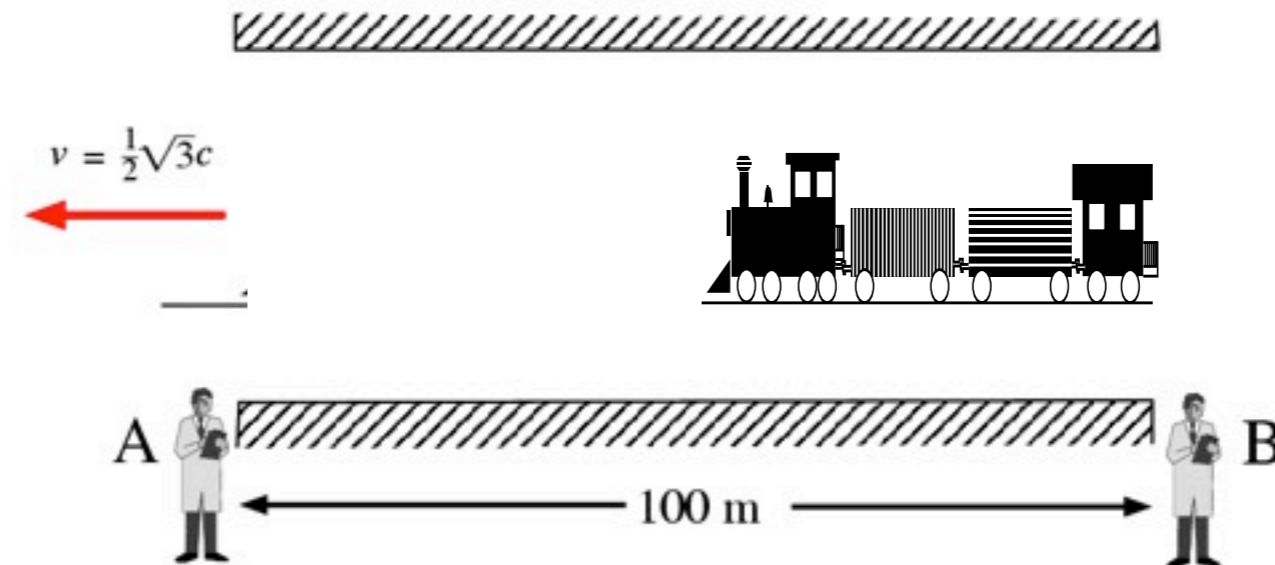


# Question 1



A's clock and the driver's clock read zero as the driver exits tunnel.

What does B's clock read when the guard goes in?



*Moving train length 50m, so driver has still 50m to travel before he exits and his clock reads 0. A's clock and B's clock are synchronised. Hence the reading on B's clock is*

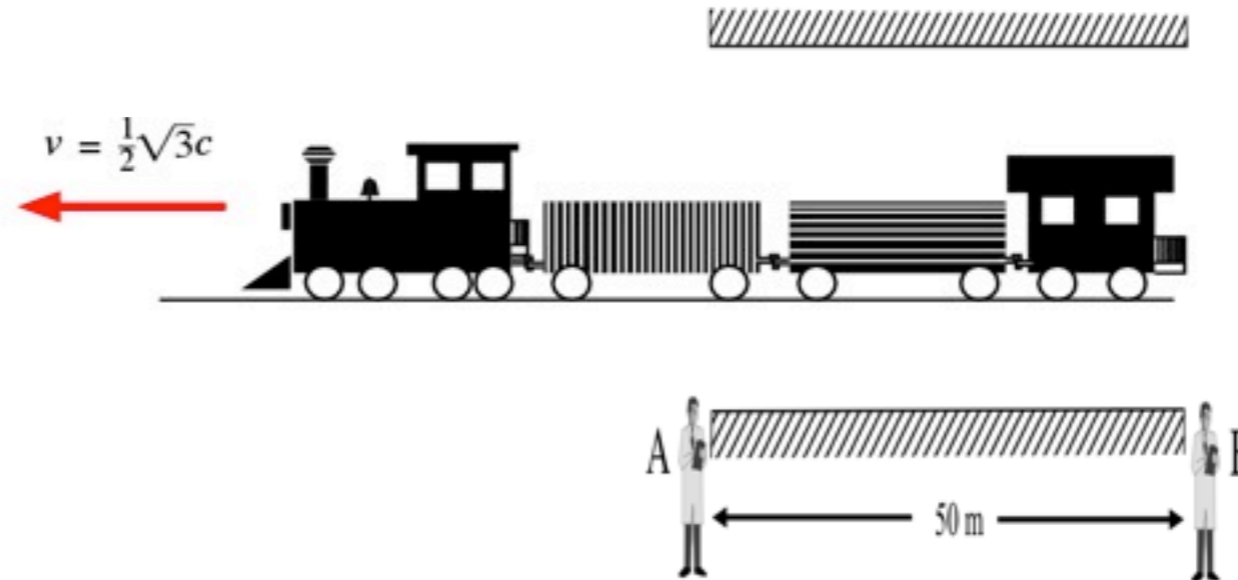
$$-\frac{50}{v} = -\frac{100}{\sqrt{3}c} \approx -200 \text{ ns}$$



# Question 2



What does the guard's clock read as he goes in?



*To the guard, tunnel is only 50m long, so driver is 50m past the exit as guard goes in. Hence clock reading is*

$$+\frac{50}{v} = +\frac{100}{\sqrt{3}c} \approx +200 \text{ ns}$$

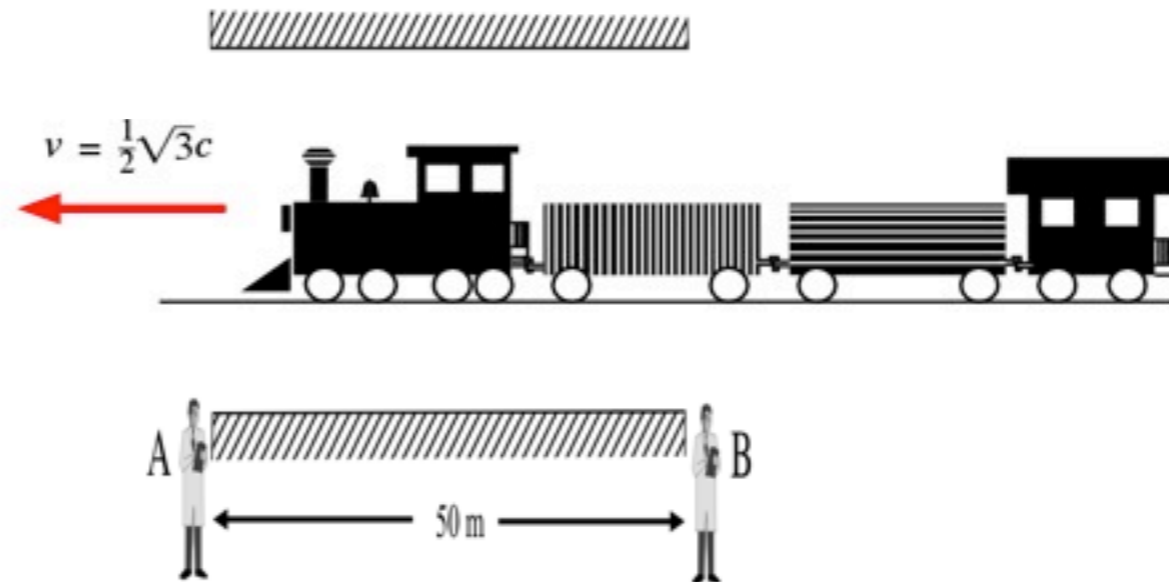




# Question 3



Where is the guard when his clock reads 0?



*Guard's clock reads 0 when driver's clock reads 0, which is as driver exits the tunnel. To guard and driver, tunnel is 50m, so guard is 50m from the entrance in the train's frame, or 100m in tunnel frame.*

*So the guard is 100m from the entrance to the tunnel when his clock reads 0.*



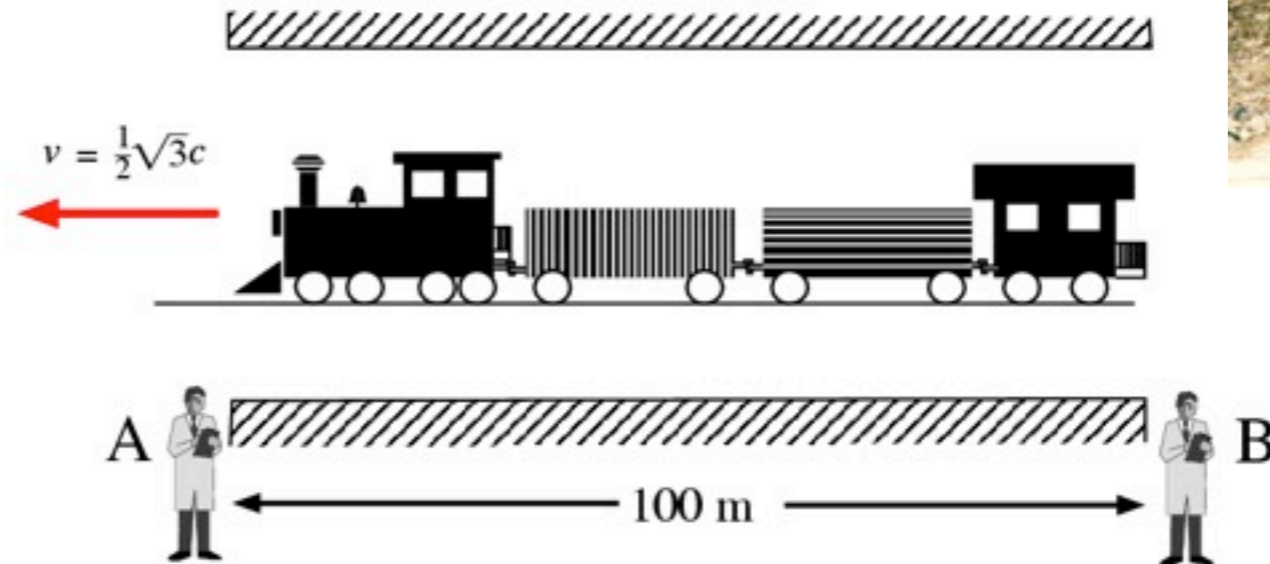
Repeat within framework of Lorentz transformation

# Question 1



A's clock and the driver's clock read zero as the driver exits tunnel.

What does B's clock read when the guard goes in?



$F(t,x)$  is frame of A and B,  $F'(t',x')$  is frame of driver and guard.

We know  $x_A = 0$ ,  $x_B = 100$ ,  $x'_D = 0$ ,  $x'_G = 100$ . We want  $t_B$  when G and B coincide.

$$x = \gamma(x' - vt') \quad t = \gamma \left( t' - \frac{vx'}{c^2} \right)$$

$$x' = \gamma(x + vt) \quad t' = \gamma \left( t + \frac{vx}{c^2} \right)$$

$$100 = \gamma(100 + vt_B)$$

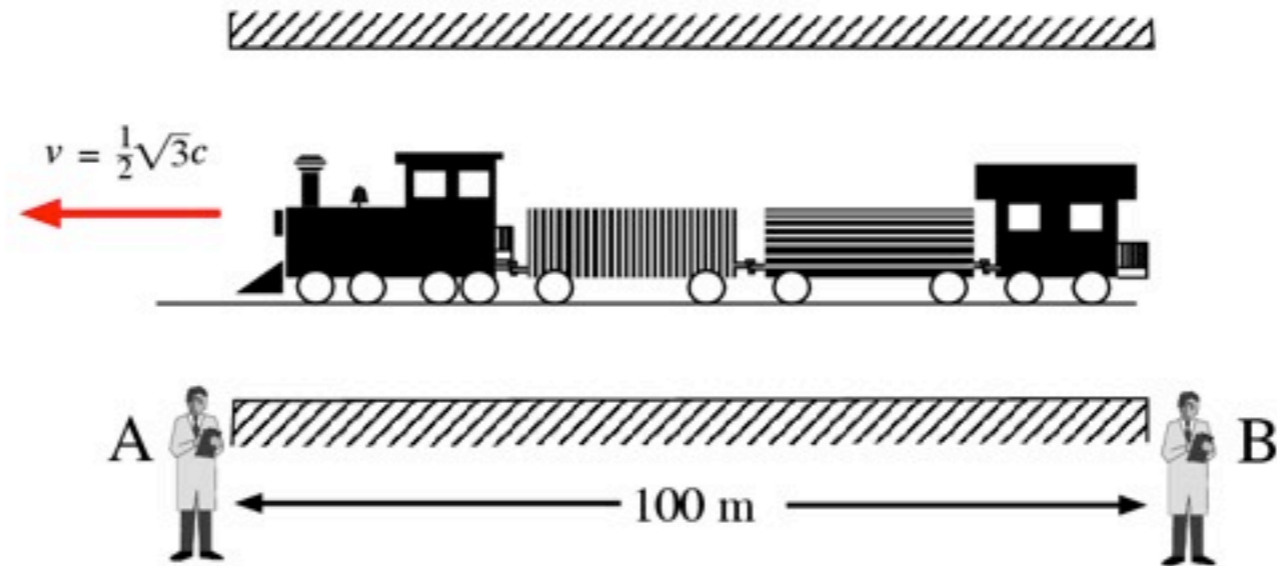
$$\implies t_B = 100 \frac{1 - \gamma}{\gamma v} = -\frac{50}{v}$$



# Question 2



What does the guard's clock read as he goes in?



*$F(t,x)$  is frame of A and B,  $F'(t',x')$  is frame of driver and guard.*

We know  $x_A = 0$ ,  $x_B = 100$ ,  $x'_D = 0$ ,  $x'_G = 100$ . We want  $t'_G$  when B and G coincide.

$$x = \gamma(x' - vt')$$

$$t = \gamma\left(t' - \frac{vx'}{c^2}\right)$$

$$x' = \gamma(x + vt)$$

$$t' = \gamma\left(t + \frac{vx}{c^2}\right)$$

$$100 = \gamma(100 - vt'_G)$$

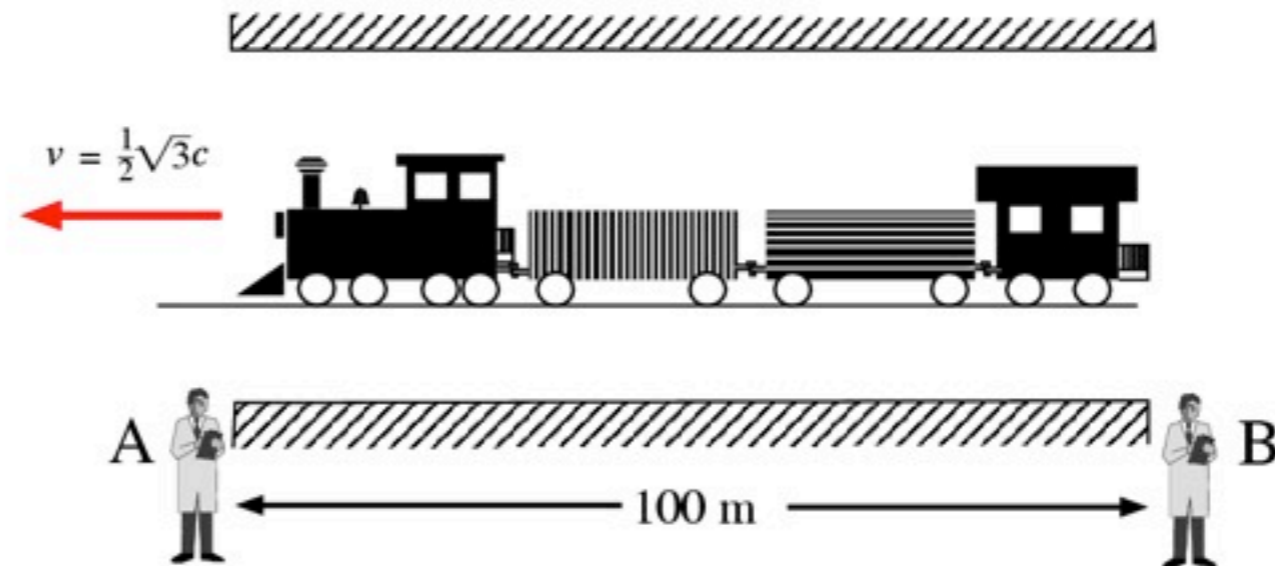
$$\Rightarrow t'_G = 100 \frac{\gamma - 1}{\gamma v} = + \frac{50}{v}$$



# Question 3



Where is the guard when his clock reads 0?



*$F(t,x)$  is frame of A and B,  $F'(t',x')$  is frame of driver and guard.*

We know  $x_A = 0$ ,  $x_B = 100$ ,  $x'_D = 0$ ,  $x'_G = 100$ .

We want  $x$  corresponding to  $x'_G = 100$  given  $t'_G = 0$ .

$$x = \gamma(x' - vt') \quad t = \gamma\left(t' - \frac{vx'}{c^2}\right)$$

$$x = \gamma(100 - v \times 0) = 200 \text{ m}$$

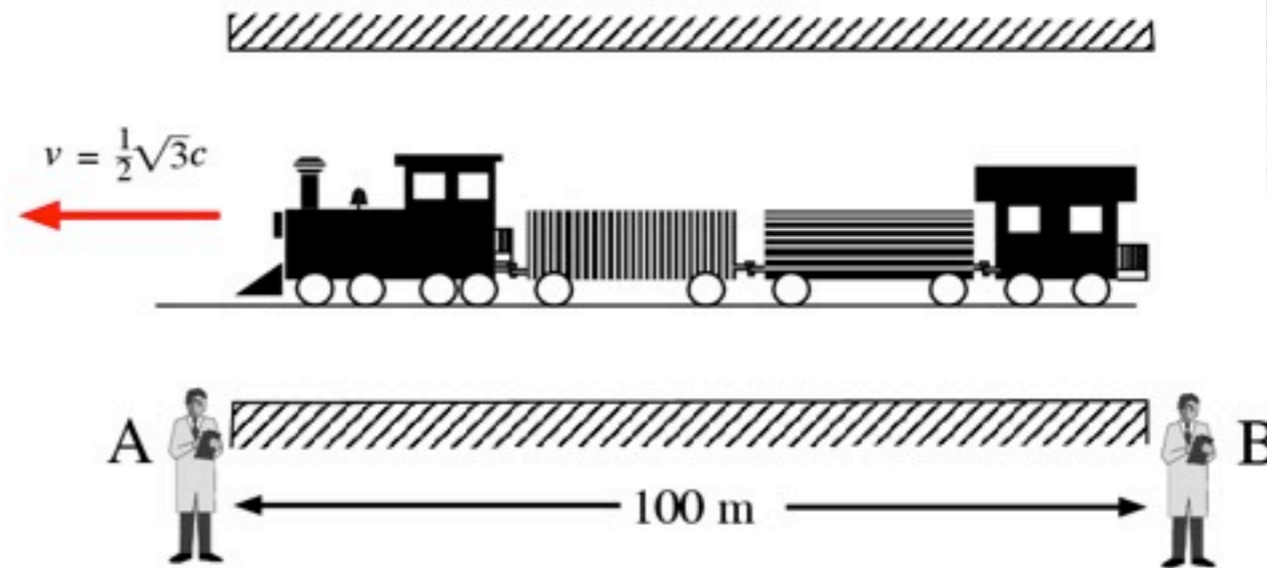
$$x' = \gamma(x + vt) \quad t' = \gamma\left(t + \frac{vx}{c^2}\right)$$

Or 100m from the entrance to the tunnel

# Question 4



Where was the driver when his clock reads the same as the guard's when he enters the tunnel?



$F(t,x)$  is frame of A and B,  $F'(t',x')$  is frame of driver and guard.

We know  $x_A = 0$ ,  $x_B = 100$ ,  $x'_D = 0$ ,  $x'_G = 100$ . We want  $x_D$  when  $t'_D = \frac{50}{v}$ .

$$x = \gamma(x' - vt')$$

$$t = \gamma\left(t' - \frac{vx'}{c^2}\right)$$

$$x' = \gamma(x + vt) \quad t' = \gamma\left(t + \frac{vx}{c^2}\right)$$

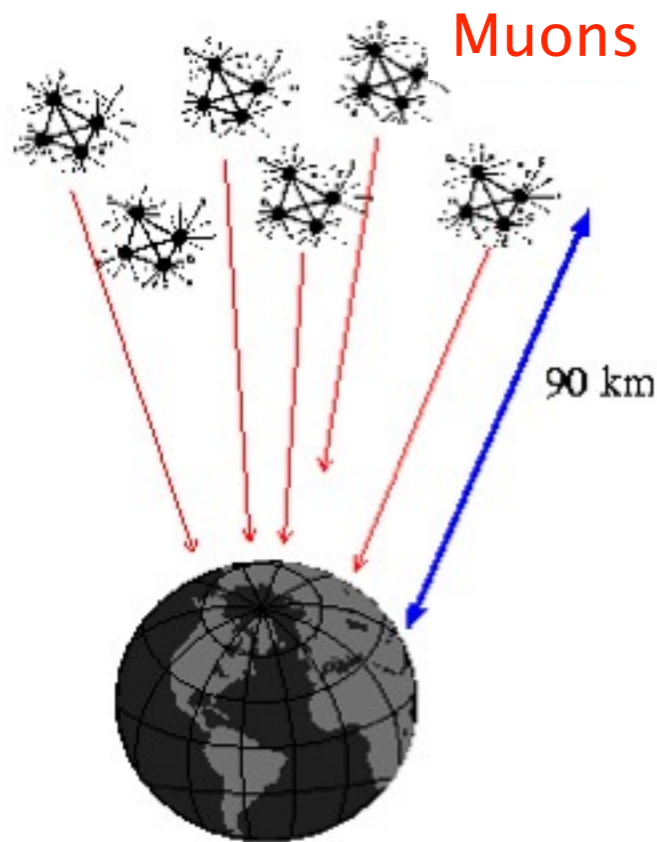
$$x = \gamma(x'_D - vt'_D)$$

$$= -50\gamma = -100 \text{ m}$$

The driver is already out of the tunnel by 100m

# Example: Cosmic Rays

- Muons are created in the upper atmosphere, 90km from earth. Their half life is  $\tau=2 \mu\text{s}$ , so they can travel at most  $2 \times 10^{-6}c=600 \text{ m}$  before decaying. So **how do more than 50% reach the earth's surface?**



Muons see distance contracted by  $\gamma$ , so

$$v\tau \approx \frac{90}{\gamma} \text{ km}$$

Earthlings say muons' clocks run slow so their half-life is  $\gamma\tau$  and

$$v(\gamma\tau) \approx 90 \text{ km}$$

- Both give

$$\frac{\gamma v}{c} = \frac{90 \text{ km}}{c\tau} = 150, \quad v \approx c, \quad \gamma \approx 150$$



# Space-time

- An invariant is a quantity that has the same value in all inertial frames.
- Lorentz transformation is based on invariance of

$$c^2 t^2 - (x^2 + y^2 + z^2) = (ct)^2 - \vec{x}^2$$

- 4D-space with coordinates  $(t,x,y,z)$  is called **space-time** and the point  $(t,x,y,z)=(t,\mathbf{x})$  is called an **event**.
- Fundamental invariant (preservation of speed of light):

$$\begin{aligned} \Delta s^2 = c^2 \Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2 &= c^2 \Delta t^2 \left( 1 - \frac{\Delta x^2 + \Delta y^2 + \Delta z^2}{c^2 \Delta t^2} \right) \\ &= c^2 \Delta t^2 \left( 1 - \frac{v^2}{c^2} \right) = c^2 \left( \frac{\Delta t}{\gamma} \right)^2 \end{aligned}$$

$\tau = \int \frac{dt}{\gamma}$  is called proper time, the time in the instantaneous rest-frame and an invariant.  $\Delta s$  is called the **separation** between two events.



# 4-Vectors

The Lorentz transformation can be written in matrix form as

$$\begin{aligned} t' &= \gamma \left( t - \frac{vx}{c^2} \right) \\ x' &= \gamma (x - vt) \\ y' &= y \\ z' &= z \end{aligned} \quad \Rightarrow \quad \begin{pmatrix} ct' \\ x' \\ y' \\ z' \end{pmatrix} = \begin{pmatrix} \gamma & -\frac{\gamma v}{c} & 0 & 0 \\ -\frac{\gamma v}{c} & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix}$$

$$X' = LX$$

Lorentz matrix  $L$

Position 4-vector  $X$

*An object made up of 4 elements which transforms like  $X$  is called a 4-vector*

*(analogous to the 3-vector of classical mechanics)*





# 4-Vector Invariants

Basic invariant:

$$c^2t^2 - x^2 - y^2 - z^2 = (ct, x, y, z) \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \begin{pmatrix} ct \\ x \\ y \\ z \end{pmatrix} = X^t g X = X \cdot X$$

Inner product of two four vectors  $A = (a_0, \vec{a})$ ,  $B = (b_0, \vec{b})$ :

$$A \cdot B = A^T g B = a_0 b_0 - a_1 b_1 - a_2 b_2 - a_3 b_3 = a_0 b_0 - \vec{a} \cdot \vec{b}$$

Invariance:

$$A' \cdot B' = (LA)^T g (LB) = A^T (L^T g L) B = A^T g B = A \cdot B$$

In particular  $A \cdot A = a_0^2 - \vec{a}^2$ .



# 4-Vectors in S.R. Mechanics

- Velocity:  $V = \frac{dX}{d\tau} = \gamma \frac{dX}{dt} = \gamma \frac{d}{dt}(ct, \vec{x}) = \gamma(c, \vec{v})$
- Note invariant:  $V \cdot V = \gamma^2(c^2 - \vec{v}^2) = \frac{c^2 - \vec{v}^2}{1 - \vec{v}^2/c^2} = c^2$
- Momentum:  $P = m_0 V = m_0 \gamma(c, \vec{v}) = (mc, \vec{p})$

$m = m_0 \gamma$  is the relativistic mass

$p = m_0 \gamma \vec{v} = m \vec{v}$  is the relativistic 3-momentum



# 4-Force

From Newton's 2<sup>nd</sup> Law expect 4-Force given by

$$\begin{aligned} F &= \frac{dP}{d\tau} = \gamma \frac{dP}{dt} \\ &= \gamma \frac{d}{dt} (mc, \vec{p}) = \gamma \left( c \frac{dm}{dt}, \frac{d\vec{p}}{dt} \right) \\ &= \gamma \left( c \frac{dm}{dt}, \vec{f} \right) \end{aligned}$$

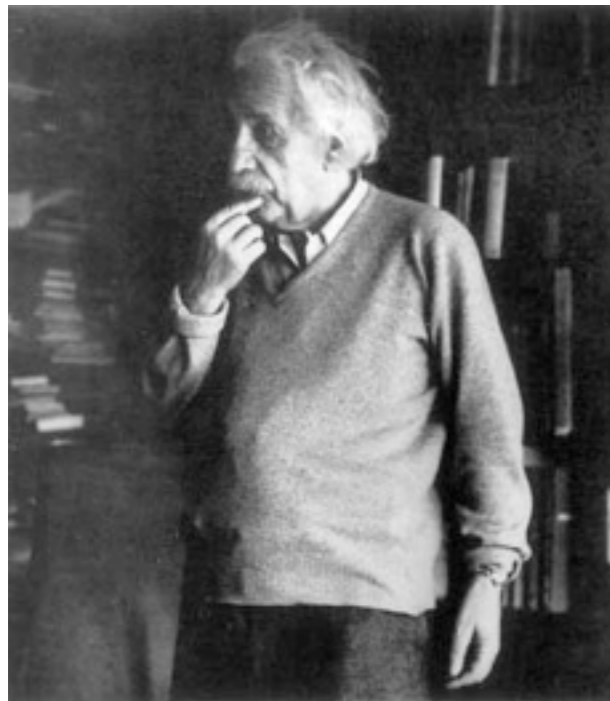
Note: 3-force equation:  $\vec{f} = \frac{d\vec{p}}{dt} = m_0 \frac{d}{dt} (\gamma \vec{v})$



# Einstein's Relation: Energy and Mass

• Momentum invariant  $P \cdot P = m_0^2 V \cdot V = m_0^2 c^2$

• Differentiate  $P \cdot \frac{dP}{d\tau} = 0 \implies V \cdot \frac{dP}{d\tau} = 0 \implies V \cdot F = 0$



$$\implies \gamma(c, \vec{v}) \cdot \gamma \left( c \frac{dm}{dt}, \vec{f} \right) = 0$$

$$\implies \frac{d}{dt}(mc^2) - \vec{v} \cdot \vec{f} = 0$$

$$\begin{aligned} \vec{v} \cdot \vec{f} &= \text{rate at which force does work} \\ &= \text{rate of change of kinetic energy} \end{aligned}$$

Therefore kinetic energy is

$$T = mc^2 + \text{constant} = m_0 c^2 (\gamma - 1)$$

***$E=mc^2$  is total energy***



# Summary of 4-Vectors

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Position  $X = (ct, \vec{x})$

Velocity  $V = \gamma(c, \vec{v})$

Momentum  $P = m_0 V = m(c, \vec{v}) = \left( \frac{E}{c}, \vec{p} \right)$

Force  $F = \gamma \left( c \frac{dm}{dt}, \vec{f} \right) = \gamma \left( \frac{1}{c} \frac{dE}{dt}, \vec{f} \right)$



# Example: Addition of Velocities

An object has velocity  $\vec{u} = (u_x, u_y)$  in frame  $F'$ , which moves with velocity  $\vec{v} = (v, 0)$  with respect to frame  $F$ .

The 4-velocity  $U = \gamma_u(c, u_x, u_y)$  has to be Lorentz transformed to  $F$ , resulting in a 4-velocity  $W = \gamma_w(c, w_x, w_y)$ :

$$\begin{pmatrix} c\gamma_w \\ \gamma_w w_x \\ \gamma_w w_y \end{pmatrix} = \begin{pmatrix} \gamma & \gamma v/c & 0 \\ \gamma v/c & \gamma & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} c\gamma_u \\ \gamma_u u_x \\ \gamma_u u_y \end{pmatrix}$$

$$\gamma_w = \gamma \gamma_u \left( 1 + \frac{v u_x}{c^2} \right)$$

$$\gamma_w w_x = \gamma \gamma_u (v + u_x)$$

$$\gamma_w w_y = \gamma_u u_y$$

$$w_x = \frac{v + u_x}{\left( 1 + \frac{v u_x}{c^2} \right)}$$

$$w_y = \frac{u_y}{\gamma \left( 1 + \frac{v u_x}{c^2} \right)}$$

# Using Invariants

A neater way of finding the speed of the particle as measured in frame  $F$ :

An observer in  $F$  has 4-velocity  $V$  and the object has 4-velocity  $U$

$$U \cdot V \text{ is invariant}$$

$$\text{Evaluated in frame } F' : \quad U = \gamma_u(c, u_x, u_y), \quad V = \gamma(c, -v, 0)$$

$$\implies U \cdot V = \gamma\gamma_u(c^2 + vu_x)$$

$$\text{Evaluated in frame } F : \quad U = (c, 0, 0), \quad V = \gamma_w(c, w_x, w_y)$$

$$\implies U \cdot V = c^2\gamma_w$$

Hence

$$\gamma_w = \gamma\gamma_u \left( 1 + \frac{vu_x}{c^2} \right)$$



# Basic Quantities used in Accelerator Calculations

**Relative velocity**  $\beta = \frac{v}{c}$

**Velocity**  $v = \beta c$

**Momentum**  $p = mv = m_0\gamma v = m_0\gamma\beta c$

**Kinetic energy**  $T = mc^2 - m_0c^2 = (\gamma - 1)m_0c^2 = (\gamma - 1)E_0$

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}} = (1 - \beta^2)^{-\frac{1}{2}}$$

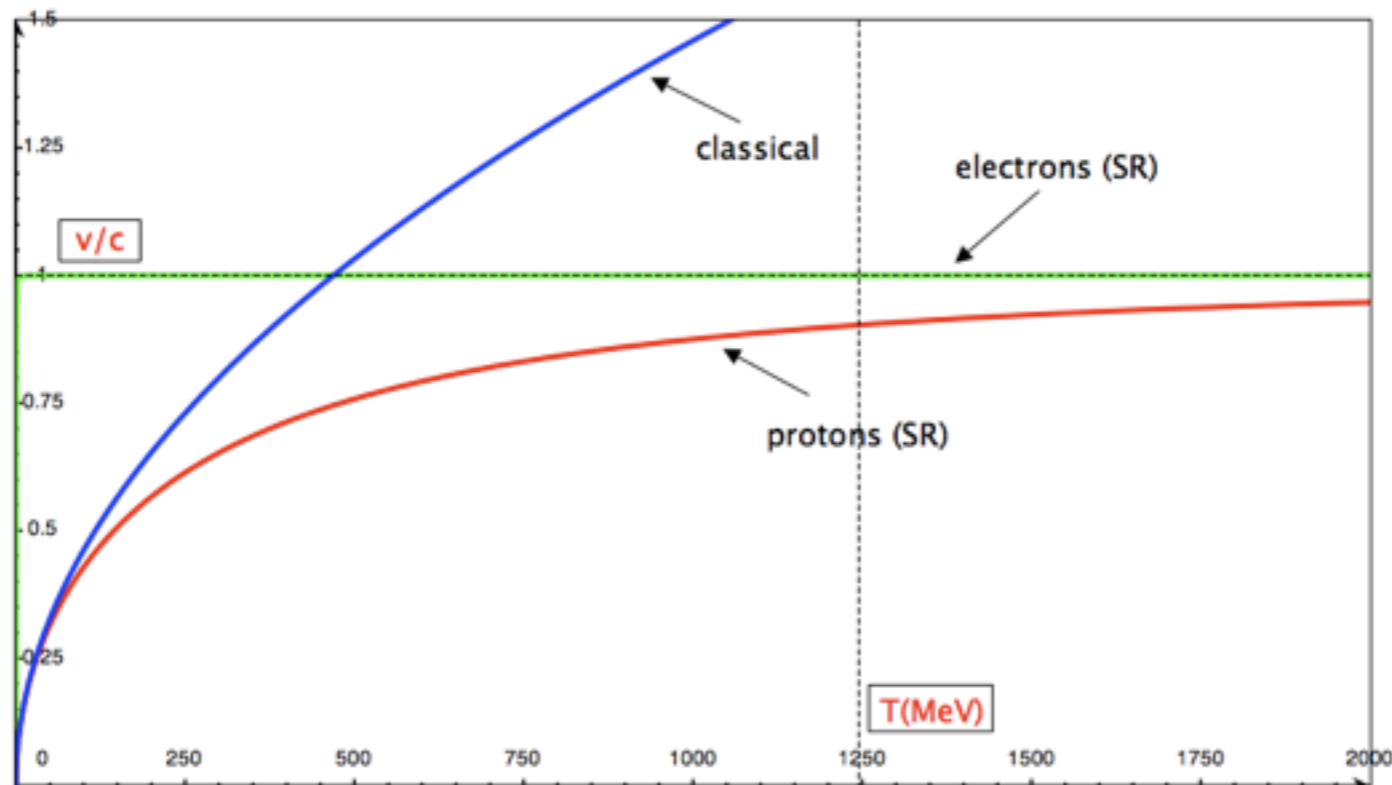
$$\implies (\beta\gamma)^2 = \frac{\gamma^2 v^2}{c^2} = \gamma^2 - 1 \implies \beta^2 = \frac{v^2}{c^2} = 1 - \frac{1}{\gamma^2}$$





# Velocity and Energy

Relativistic Velocity as a function of Kinetic Energy



$$T = m_0 c^2 (\gamma - 1) = E_0 (\gamma - 1)$$

$$\gamma = 1 + \frac{T}{m_0 c^2} = 1 + \frac{T}{E_0}$$

$$\beta = \sqrt{1 - \frac{1}{\gamma^2}}$$

$$p = m_0 \beta \gamma c = \frac{E_0}{c} \beta \gamma$$

$$\text{For } v \ll c, \quad \gamma = \left(1 - \frac{v^2}{c^2}\right)^{-\frac{1}{2}} \approx 1 + \frac{1}{2} \frac{v^2}{c^2} + \frac{3}{8} \frac{v^4}{c^4} + \dots$$

$$\text{so } T = m_0 c^2 (\gamma - 1) \approx \frac{1}{2} m_0 v^2$$

# Energy-Momentum

- Important invariant:

$$P = (E/c, \mathbf{p}) \implies P \cdot P = E^2/c^2 - \mathbf{p}^2$$

$$\text{and } P = m_0 V \implies P \cdot P = m_0^2 V \cdot V = m_0^2 c^2 = E_0^2/c^2$$

$$\frac{E^2}{c^2} = \mathbf{p}^2 + m_0^2 c^2 \quad \text{or} \quad E^2 = \mathbf{p}^2 c^2 + E_0^2$$

$$\implies p^2 c^2 = E^2 - E_0^2 = (E - E_0)(E + E_0) = T(T + 2E_0)$$

**Example:** ISIS at RAL accelerates protons ( $E_0 = 938 \text{ MeV}$ ) to  $800 \text{ MeV}$

$$\begin{aligned} \implies pc &= \sqrt{800 \times (800 + 2 \times 938)} \text{ MeV} \\ &= 1.463 \text{ GeV} \end{aligned}$$

$$\beta\gamma = \frac{m_0 \beta \gamma c^2}{m_0 c^2} = \frac{pc}{E_0} = 1.56$$

$$\gamma^2 = (\beta\gamma)^2 + 1 \implies \gamma = 1.85$$

$$\beta = \frac{\beta\gamma}{\gamma} = 0.84$$

# Relationships between small variations in parameters $\Delta E$ , $\Delta T$ , $\Delta p$ , $\Delta\beta$ , $\Delta\gamma$

$$\begin{aligned} & (\beta\gamma)^2 = \gamma^2 - 1 \\ \implies & \beta\gamma\Delta(\beta\gamma) = \gamma\Delta\gamma \\ \implies & \beta\Delta(\beta\gamma) = \Delta\gamma \end{aligned} \quad (1)$$

$$\begin{aligned} & \frac{1}{\gamma^2} = 1 - \beta^2 \\ \implies & \frac{1}{\gamma^3}\Delta\gamma = \beta\Delta\beta \end{aligned} \quad (2)$$

$$\begin{aligned} \frac{\Delta p}{p} &= \frac{\Delta(m_0\beta\gamma c)}{m_0\beta\gamma c} = \frac{\Delta(\beta\gamma)}{\beta\gamma} \\ &= \frac{1}{\beta^2} \frac{\Delta\gamma}{\gamma} = \frac{1}{\beta^2} \frac{\Delta E}{E} \\ &= \gamma^2 \frac{\Delta\beta}{\beta} \\ &= \frac{\gamma}{\gamma + 1} \frac{\Delta T}{T} \quad (\text{exercise}) \end{aligned}$$

Note: valid to first order only



	$\frac{\Delta\beta}{\beta}$	$\frac{\Delta p}{p}$	$\frac{\Delta T}{T}$	$\frac{\Delta E}{E} = \frac{\Delta\gamma}{\gamma}$
$\frac{\Delta\beta}{\beta} =$	$\frac{\Delta\beta}{\beta}$	$\frac{1}{\gamma^2} \frac{\Delta p}{p}$	$\frac{1}{\gamma(\gamma+1)} \frac{\Delta T}{T}$	$\frac{1}{\beta^2 \gamma^2} \frac{\Delta\gamma}{\gamma}$
		$\frac{\Delta p}{p} - \frac{\Delta\gamma}{\gamma}$		$\frac{1}{\gamma^2 - 1} \frac{\Delta\gamma}{\gamma}$
$\frac{\Delta p}{p} =$	$\gamma^2 \frac{\Delta\beta}{\beta}$	$\frac{\Delta p}{p}$	$\frac{\gamma}{\gamma+1} \frac{\Delta T}{T}$	$\frac{1}{\beta^2} \frac{\Delta\gamma}{\gamma}$
$\frac{\Delta T}{T} =$	$\gamma(\gamma+1) \frac{\Delta\beta}{\beta}$	$\left(1 + \frac{1}{\gamma}\right) \frac{\Delta p}{p}$	$\frac{\Delta T}{T}$	$\frac{\gamma}{\gamma-1} \frac{\Delta\gamma}{\gamma}$
$\frac{\Delta E}{E} =$	$(\beta\gamma)^2 \frac{\Delta\beta}{\beta}$	$\beta^2 \frac{\Delta p}{p}$	$\left(1 - \frac{1}{\gamma}\right) \frac{\Delta T}{T}$	$\frac{\Delta\gamma}{\gamma}$
		$(\gamma^2 - 1) \frac{\Delta\beta}{\beta}$		

Table 1: Incremental relationships between energy, velocity and momentum.



# 4-Momentum Conservation

- Equivalent expression for 4-momentum

$$P = m_0 \gamma(c, \vec{v}) = (mc, \vec{p}) = \left( \frac{E}{c}, \vec{p} \right)$$

- Invariant  $m_0^2 c^2 = P \cdot P = \frac{E^2}{c^2} - \vec{p}^2 \implies \boxed{\frac{E^2}{c^2} = \vec{p}^2 + m_0^2 c^2}$

- Classical conservation laws:

- conservation of total mass
- conservation of total 3-momentum

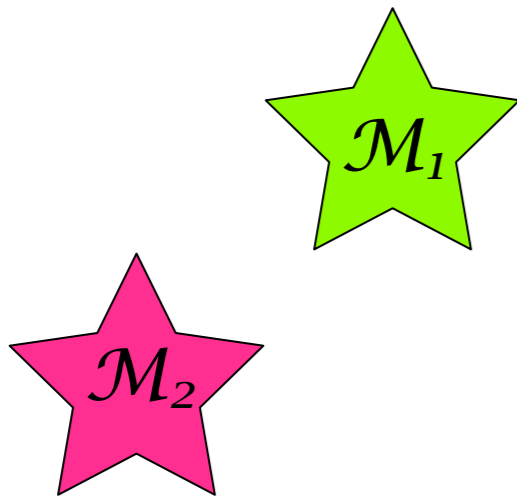
- In relativity, mass equivalent to energy (incl. rest energy)

$$\implies \sum_{\text{particles } i} E_i \quad \text{and} \quad \sum_{\text{particles } i} \mathbf{p}_i \text{ constant}$$

$\implies$  total 4-momentum

$$\sum_{\text{particles } i} P_i \text{ constant}$$

# Problem



A body of mass  $M$  disintegrates while at rest into two parts of rest masses  $M_1$  and  $M_2$ .

Show that the energies of the parts are given by

$$E_1 = c^2 \frac{M^2 + M_1^2 - M_2^2}{2M}, \quad E_2 = c^2 \frac{M^2 - M_1^2 + M_2^2}{2M}$$



# Solution

Before: 

$$P = (Mc, \vec{0})$$

After:

$$P_2 = \left( \frac{E_2}{c}, -\vec{p} \right)$$



$$P_1 = \left( \frac{E_1}{c}, \vec{p} \right)$$

Conservation of 4-momentum:

$$P = P_1 + P_2 \Rightarrow P - P_1 = P_2$$

$$\Rightarrow (P - P_1) \cdot (P - P_1) = P_2 \cdot P_2$$

$$\Rightarrow P \cdot P - 2P \cdot P_1 + P_1 \cdot P_1 = P_2 \cdot P_2$$

$$\Rightarrow M^2 c^2 - 2ME_1 + M_1^2 c^2 = M_2^2 c^2$$

$$\Rightarrow E_1 = \frac{M^2 + M_1^2 - M_2^2}{2M} c^2$$



# Example of use of invariants

---

- Two particles have equal rest mass  $m_0$ .
  - Frame 1: one particle at rest, total energy is  $E_1$ .
  - Frame 2: centre of mass frame where velocities are equal and opposite, total energy is  $E_2$ .

Problem: Relate  $E_1$  to  $E_2$





$$P_1 = \left( \frac{E_1 - m_0 c^2}{c}, \vec{p} \right) \quad P_2 = (m_0 c, \vec{0})$$

Total energy  $E_1$   
(Fixed target experiment)

$$P_1 = \left( \frac{E_2}{2c}, \vec{p}' \right) \quad P_2 = \left( \frac{E_2}{2c}, -\vec{p}' \right)$$

Total energy  $E_2$   
(Colliding beams experiment)

Invariant:  $P_2 \cdot (P_1 + P_2)$

$$m_0 c \times \frac{E_1}{c} - 0 \times p = \frac{E_2}{2c} \times \frac{E_2}{c} + p' \times 0$$

$$\Rightarrow 2m_0 c^2 E_1 = E_2^2$$



# Collider Problem

---

In an accelerator, a proton  $p_1$  with rest mass  $m_0$  collides with an anti-proton  $p_2$  (with the same rest mass), producing two particles  $W_1$  and  $W_2$  with equal rest mass  $M_0=100m_0$

- **Experiment 1:**  $p_1$  and  $p_2$  have equal and opposite velocities in the lab frame. Find the minimum energy of  $p_2$  in order for  $W_1$  and  $W_2$  to be produced.
- **Experiment 2:** in the rest frame of  $p_1$ , find the minimum energy  $E'$  of  $p_2$  in order for  $W_1$  and  $W_2$  to be produced.

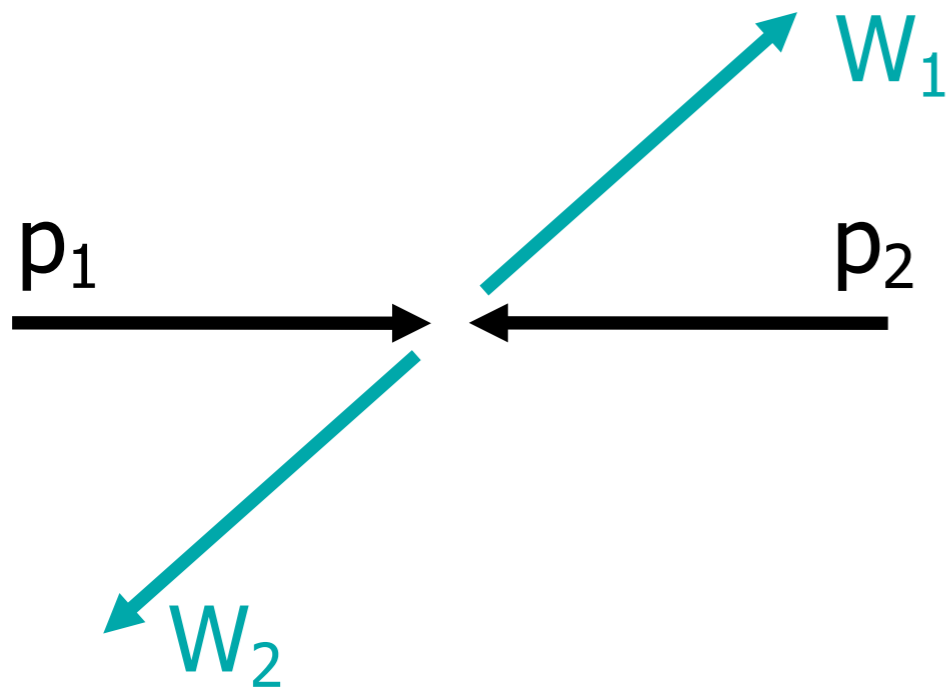


# Experiment 1

$$\frac{E^2}{c^2} = \vec{p}^2 + m_0^2 c^2$$

$\implies$

Particles with same rest-mass and same momentum have same energies.



Total 3-momentum is zero before collision, so is zero afterwards

4-momenta before collision:

$$P_1 = \left( \frac{E}{c}, \vec{p} \right) \quad P_2 = \left( \frac{E}{c}, -\vec{p} \right)$$

4-momenta after collision:

$$P_1 = \left( \frac{E'}{c}, \vec{q} \right) \quad P_2 = \left( \frac{E'}{c}, -\vec{q} \right)$$

**Total energy is conserved  $\implies 2E = 2E'$**

**$\implies E = E' > \text{rest energy} = M_0 c^2 = 100 m_0 c^2$**

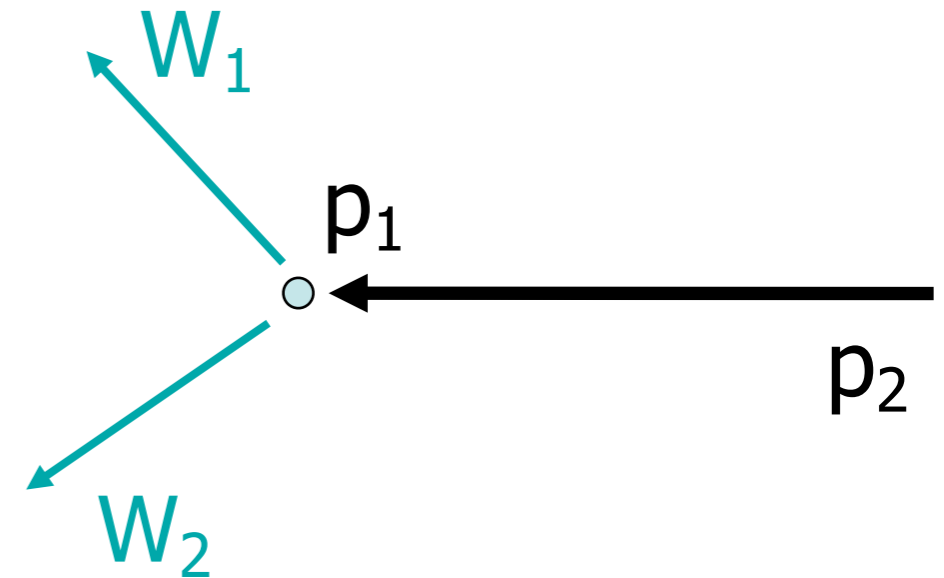


# Experiment 2

Before collision

$$P_1 = (m_0c, \vec{0}), \quad P_2 = \left( \frac{E'}{c}, \vec{p} \right)$$

Total energy is  $E_1 = E' + m_0c^2$



Use previous result  $2m_0c^2 E_1 = E_2^2$  to relate  $E_1$  to total energy  $E_2$  in the centre of mass frame

$$2m_0c^2 E_1 = E_2^2$$

$$\implies 2m_0c^2 (E' + m_0c^2) = (2E)^2 > (200m_0c^2)^2$$

$$\implies E' > (2 \times 10^4 - 1)m_0c^2 \approx 20,000 m_0c^2$$

# Photons and Wave 4-Vectors

- Monochromatic plane wave:  $\sin(\omega t - \vec{k} \cdot \vec{x})$
- $\vec{k}$  is the *wave vector*,  $|\vec{k}| = \frac{2\pi}{\lambda}$ ;  $\omega$  is the *angular frequency*,  $\omega = 2\pi\nu$
- The phase  $\frac{1}{2\pi}(\omega t - \vec{k} \cdot \vec{x})$  is the number of wave crests passing an observer

- **Invariant:** 
$$\omega t - \vec{k} \cdot \vec{x} = (ct, \vec{x}) \cdot \left(\frac{\omega}{c}, \vec{k}\right)$$

Position 4-vector                      Wave 4-vector

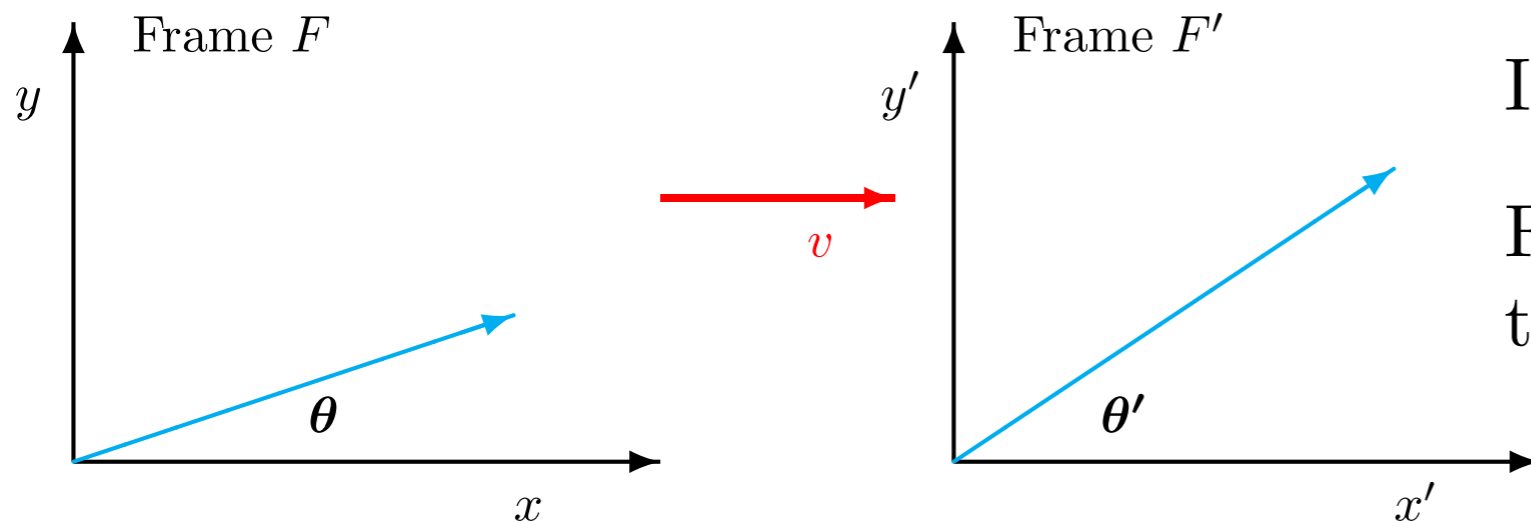
- 4-momentum  $P = \left(\frac{E}{c}, p\right) = \hbar \left(\frac{\omega}{c}, k\right) = \hbar K$



# Relativistic Doppler Shift

For light rays  $\omega = c|\vec{k}|$  so  $K = \left(\frac{\omega}{c}, \vec{k}\right)$  is *a null vector*

and can be written  $K = \frac{\omega}{c}(1, \vec{n})$  where  $|\vec{n}| = 1$ .



In  $F$ ,  $K = \frac{\omega}{c}(1, \cos \theta, \sin \theta)$

For  $F'$ , use Lorentz transformation:  $K' = LK$

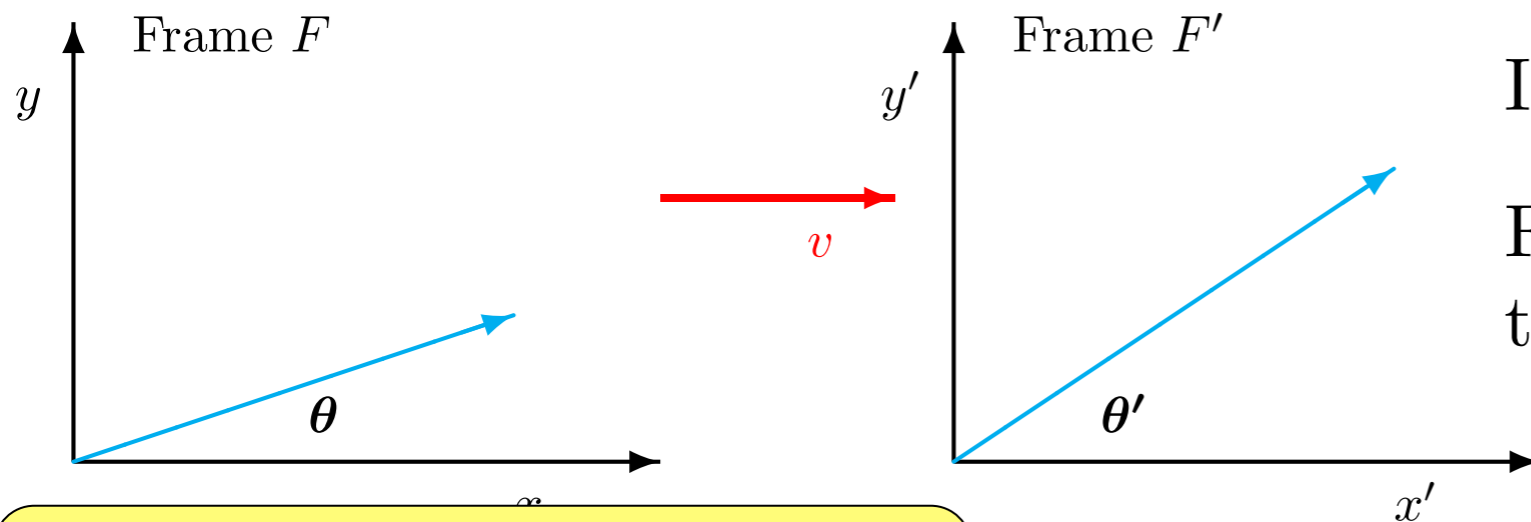
$$\begin{pmatrix} \omega'/c \\ (\omega'/c) \cos \theta' \\ (\omega'/c) \sin \theta' \end{pmatrix} = \begin{bmatrix} \gamma & -\gamma v/c & 0 \\ -\gamma v/c & \gamma & 0 \\ 0 & 0 & 1 \end{bmatrix} \begin{pmatrix} \omega/c \\ (\omega/c) \cos \theta \\ (\omega/c) \sin \theta \end{pmatrix}$$



# Relativistic Doppler Shift

For light rays  $\omega = c|\vec{k}|$  so  $K = \left(\frac{\omega}{c}, \vec{k}\right)$  is *a null vector*

and can be written  $K = \frac{\omega}{c}(1, \vec{n})$  where  $|\vec{n}| = 1$ .



In  $F$ ,  $K = \frac{\omega}{c}(1, \cos \theta, \sin \theta)$

For  $F'$ , use Lorentz transformation:  $K' = LK$

$$\omega' = \gamma \left( \omega - \frac{v\omega \cos \theta}{c} \right)$$

$$\omega' \cos \theta' = \gamma \left( \omega \cos \theta - v \frac{\omega}{c} \right)$$

$$\omega' \sin \theta' = \omega \sin \theta$$

$$\omega' = \gamma \omega \left( 1 - \frac{v}{c} \cos \theta \right)$$

$$\tan \theta' = \frac{\sin \theta}{\gamma \left( \cos \theta - \frac{v}{c} \right)}$$

Note there is a transverse Doppler effect even when  $\theta = \frac{1}{2}\pi$



# 4-Acceleration

- 4-Acceleration=rate of change of 4-Velocity

$$A = \frac{dV}{d\tau} = \gamma \frac{d}{dt} (\gamma c, \gamma \vec{v})$$

- Use  $\frac{1}{\gamma^2} = 1 - \frac{\vec{v} \cdot \vec{v}}{c^2} \implies \frac{1}{\gamma^3} \frac{d\gamma}{dt} = \frac{\vec{v} \cdot \dot{\vec{v}}}{c^2} = \frac{\vec{v} \cdot \vec{a}}{c^2}$

$$A = \gamma \left( \gamma^3 \frac{\vec{v} \cdot \vec{a}}{c}, \gamma \vec{a} + \gamma^3 \left( \frac{\vec{v} \cdot \vec{a}}{c^2} \right) \vec{v} \right)$$

- In instantaneous rest-frame  $A = (0, \vec{a}), \quad A \cdot A = -|\vec{a}|^2$





# Radiation from an accelerating charge

- Rate of radiation,  $R$ , known to be invariant and proportional to  $|\vec{a}|^2$  in instantaneous rest frame.

- But in instantaneous rest-frame  $A \cdot A = -|\vec{a}|^2$

- Deduce  $R \propto A \cdot A = -\gamma^6 \left( \left( \frac{\vec{v} \cdot \vec{a}}{c} \right)^2 + \frac{1}{\gamma^2} \vec{a}^2 \right)$

- Rearranged:

$$R = \frac{e^2}{6\pi\epsilon_0 c^3} \gamma^6 \left[ |\vec{a}|^2 - \frac{(\vec{a} \times \vec{v})^2}{c^2} \right]$$

**Relativistic  
Larmor  
Formula**

If  $\vec{a} \parallel \vec{v}$ ,  $R \propto \gamma^6$ , but if  $\vec{a} \perp \vec{v}$ ,  $R \propto \gamma^4$



# Motion under constant acceleration; world lines

- Introduce *rapidity*  $\rho$  defined by

$$\beta = \frac{v}{c} = \tanh \rho \quad \Longrightarrow \quad \gamma = \frac{1}{\sqrt{1 - \beta^2}} = \cosh \rho$$

- Then  $V = \gamma(c, v) = c(\cosh \rho, \sinh \rho)$

- And  $A = \frac{dV}{d\tau} = c(\sinh \rho, \cosh \rho) \frac{d\rho}{d\tau}$

- So constant acceleration satisfies

$$a^2 = |\vec{a}|^2 = -A \cdot A = c^2 \left( \frac{d\rho}{d\tau} \right)^2 \quad \Longrightarrow \quad \frac{d\rho}{d\tau} = \frac{a}{c}, \quad \text{so } \rho = \frac{a\tau}{c}$$



# Particle Paths

$$\frac{dx}{d\tau} = \gamma \frac{dx}{dt} = \gamma v = c \sinh \rho = c \sinh \frac{a\tau}{c}$$

$$\Rightarrow x = x_0 + \frac{c^2}{a} \left( \cosh \frac{a\tau}{c} - 1 \right)$$

$$\frac{dt}{d\tau} = \gamma = \cosh \rho = \cosh \frac{a\tau}{c}$$

$$\Rightarrow t = \frac{c}{a} \sinh \frac{a\tau}{c}$$

$$\cosh^2 \rho - \sinh^2 \rho = 1$$

$$\Rightarrow \left( x - x_0 + \frac{c^2}{a} \right)^2 - c^2 t^2 = \frac{c^4}{a^2}$$

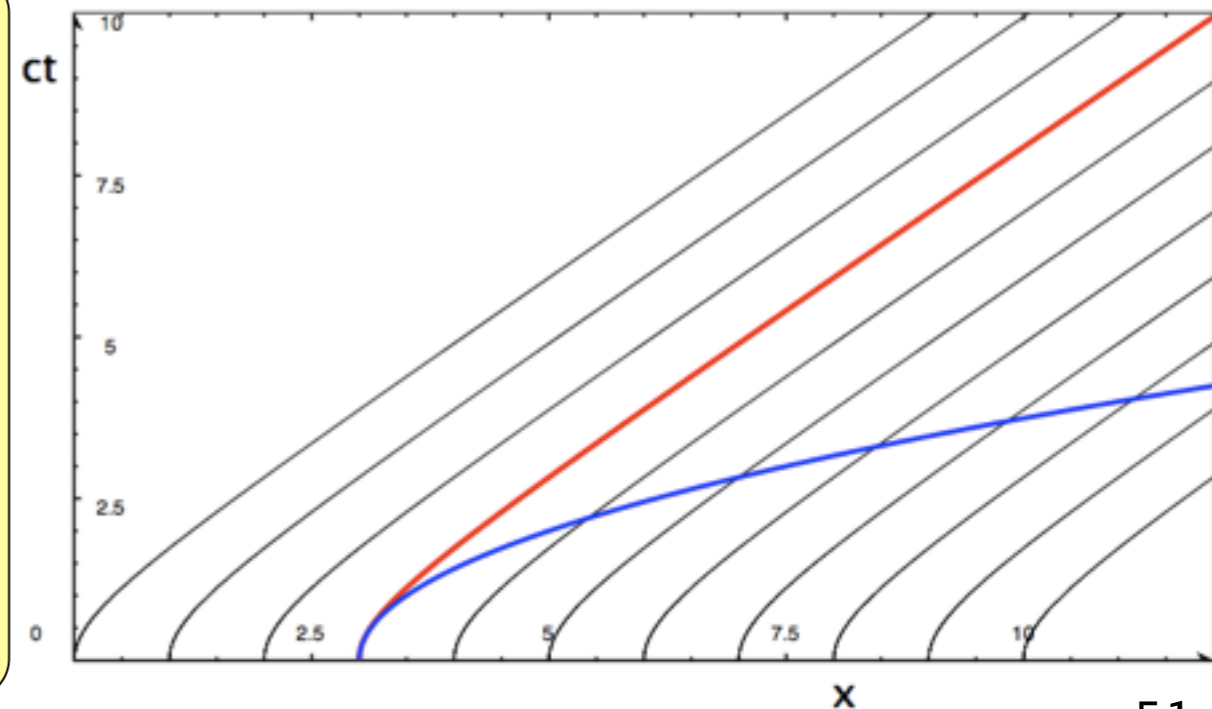
**Relativistic paths  
are hyperbolic**

$$x = x_0 + \frac{c^2}{a} \left( 1 + \frac{1}{2} \frac{a^2 \tau^2}{c^2} + \dots - 1 \right)$$

$$\approx x_0 + \frac{1}{2} a \tau^2$$

**Non-relativistic  
paths are  
parabolic**

$$t \approx \frac{c}{a} \times \frac{a\tau}{c} = \tau$$



# Relativistic Lagrangian and Hamiltonian Formulation

4-force equation of motion under a potential  $V$ :

$$\vec{f} = -\nabla V = \frac{d\vec{p}}{dt} = m_0 \frac{d}{dt}(\gamma\vec{v})$$

$$\implies m_0 \frac{d}{dt} \left( \frac{\dot{x}}{\sqrt{1 - v^2/c^2}} \right) = -\frac{\partial V}{\partial x} \quad \text{etc., where } v^2 = \dot{x}^2 + \dot{y}^2 + \dot{z}^2$$

Compare with standard Lagrangian formulation:  $\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{x}} \right) = \frac{\partial L}{\partial x}$ .

Deduce **Relativistic Lagrangian**:

$$\mathcal{L} = -m_0 c^2 \left( 1 - \frac{v^2}{c^2} \right)^{\frac{1}{2}} - V = -\frac{m_0 c^2}{\gamma} - V.$$

**Note:**  $\mathcal{L} \neq T - V$



Hamiltonian  $\mathcal{H}$ , by definition:

$$\begin{aligned}\mathcal{H} &= \sum_{x,y,z} \dot{x} \frac{\partial \mathcal{L}}{\partial \dot{x}} - \mathcal{L} = \frac{m_0}{\sqrt{1 - v^2/c^2}} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) - \mathcal{L} \\ &= m_0 \gamma v^2 + \frac{m_0 c^2}{\gamma} + V = m_0 \gamma c^2 + V \\ &= T + V, \quad \text{Total Energy}\end{aligned}$$

Since  $E^2 = \vec{p}^2 c^2 + m_0^2 c^4$ ,

$$\mathcal{H} = c (\vec{p}^2 + m_0^2 c^2)^{\frac{1}{2}} + V$$

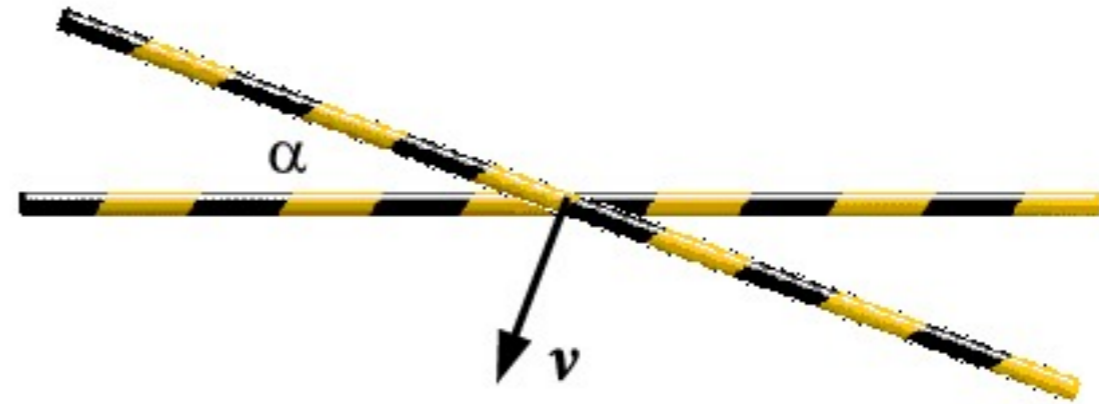
Then Hamilton's equations of motion are:

$$\dot{\vec{p}} = -\frac{\partial \mathcal{H}}{\partial \vec{x}}, \quad \dot{\vec{x}} = \frac{\partial \mathcal{H}}{\partial \vec{p}}$$

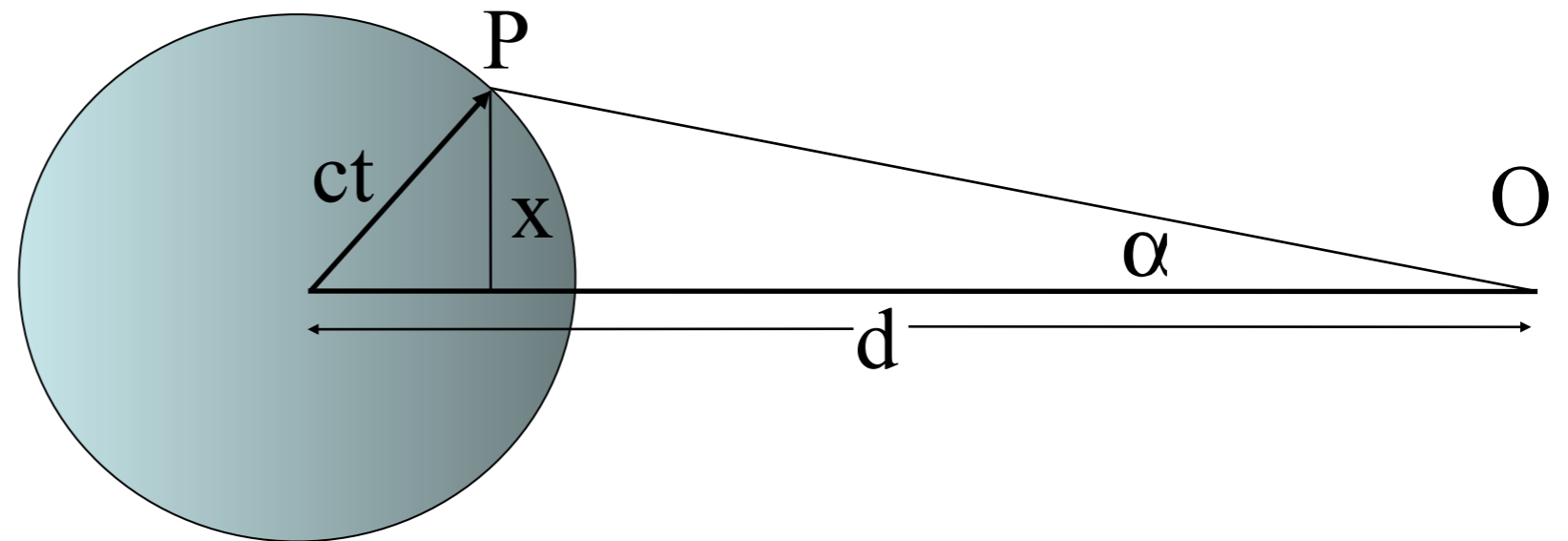


# Motion faster than light

- Two rods sliding over each other. Speed of intersection point is  $v/\sin\alpha$ , which can be made greater than  $c$ .



- Explosion of planetary nebula. Observer sees bright spot spreading out. Light from P arrives  $t=d\alpha^2/2c$  later.



$$t = \frac{d\alpha^2}{2c} \approx \frac{x}{c} \frac{\alpha}{2} \ll \frac{x}{c}$$

