

# The LUX-ZEPLIN (LZ) Experiment



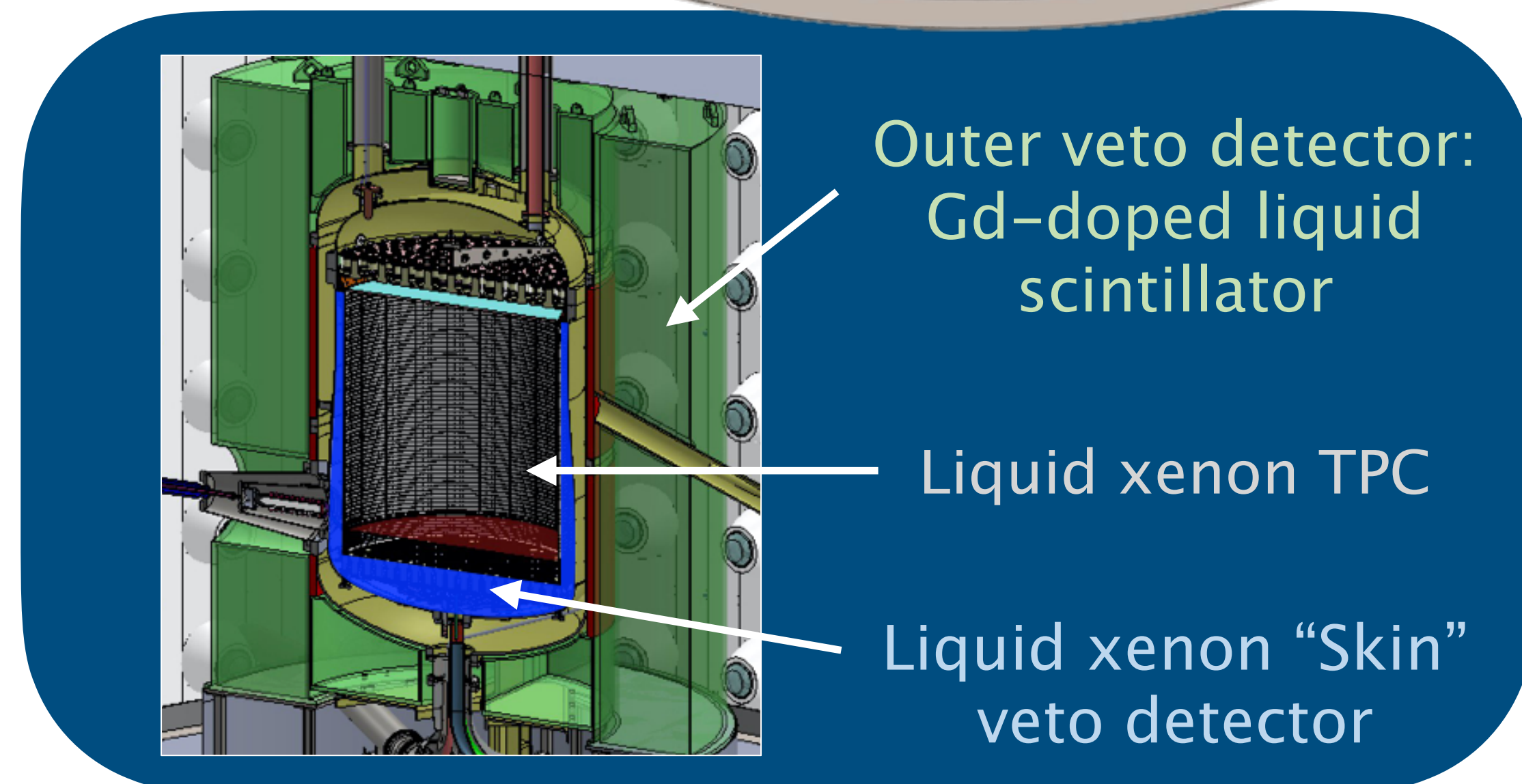
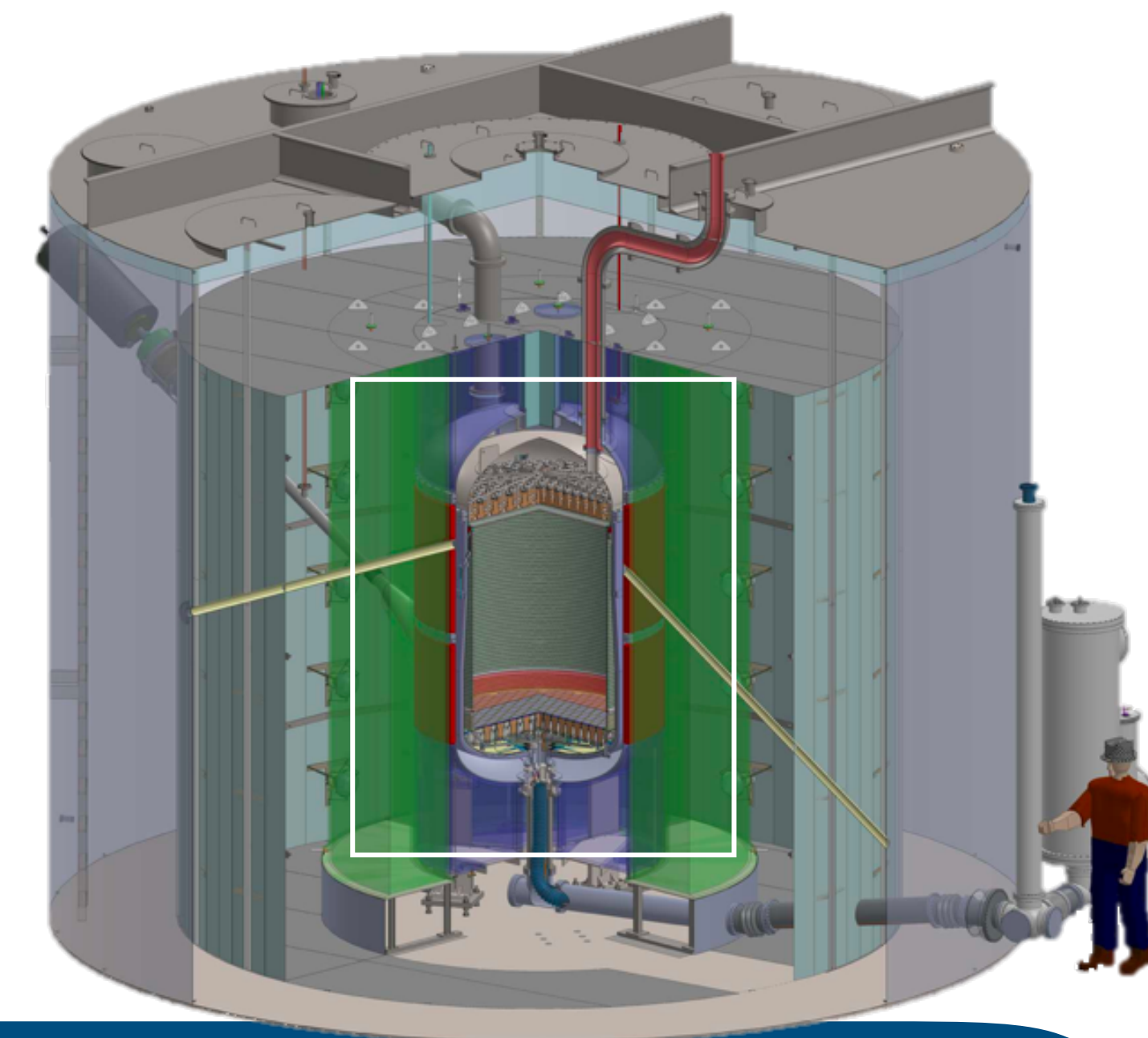
*Amy Cottle, University of Oxford*

*IOP HEPP 2022*

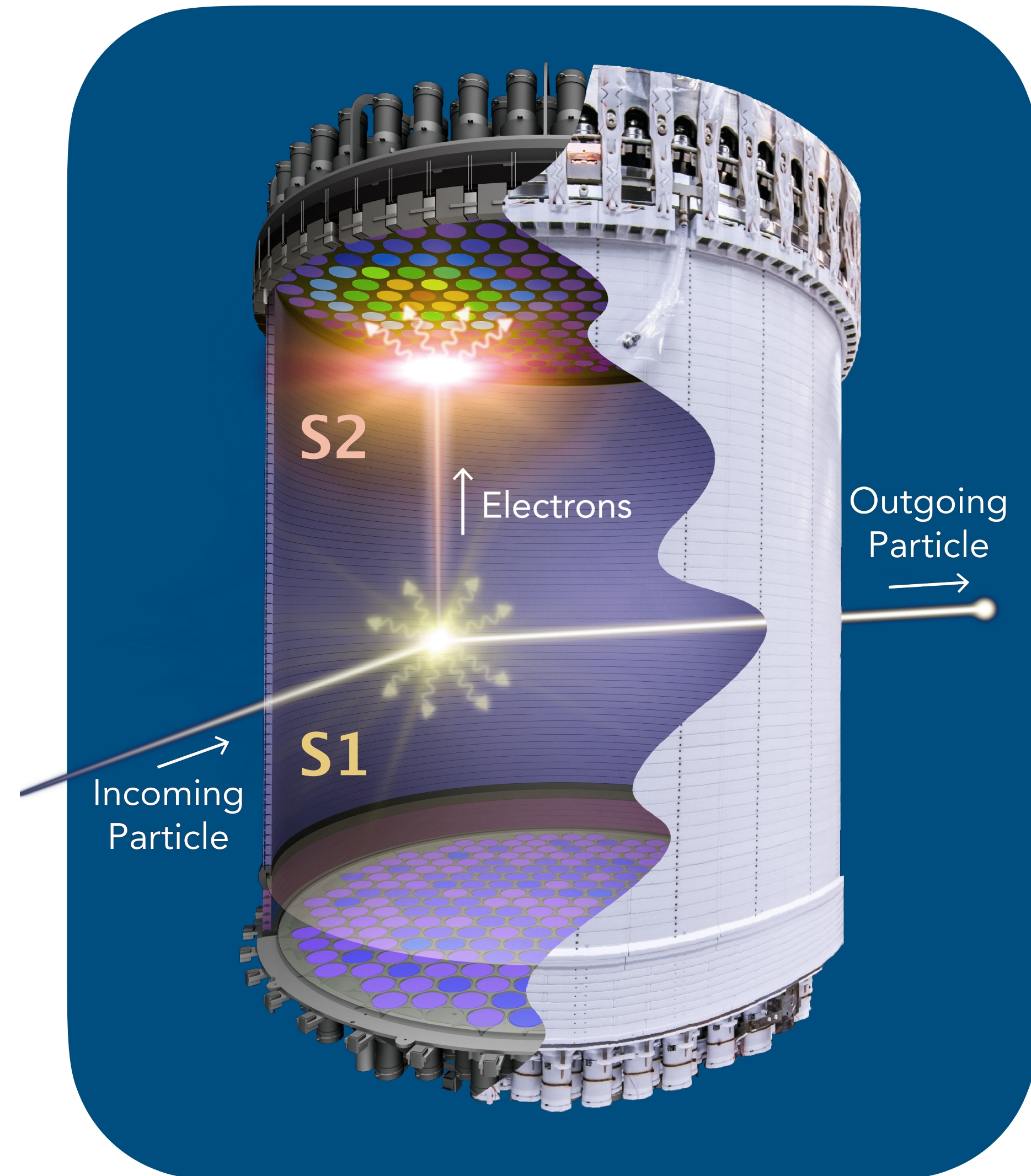


# Introduction to LZ

- LZ is based 4850 ft underground at the Sanford Underground Research Facility (SURF) in Lead, SD
- LZ is a dark matter direct detection experiment
  - Primarily designed for WIMPs, but has considerable sensitivity to other new physics
- Central detector: dual-phase xenon time projection chamber (TPC) with 7 t active xenon
- “Skin” & outer detector (OD) active vetoes
- The detectors have been commissioned and science data-taking is underway



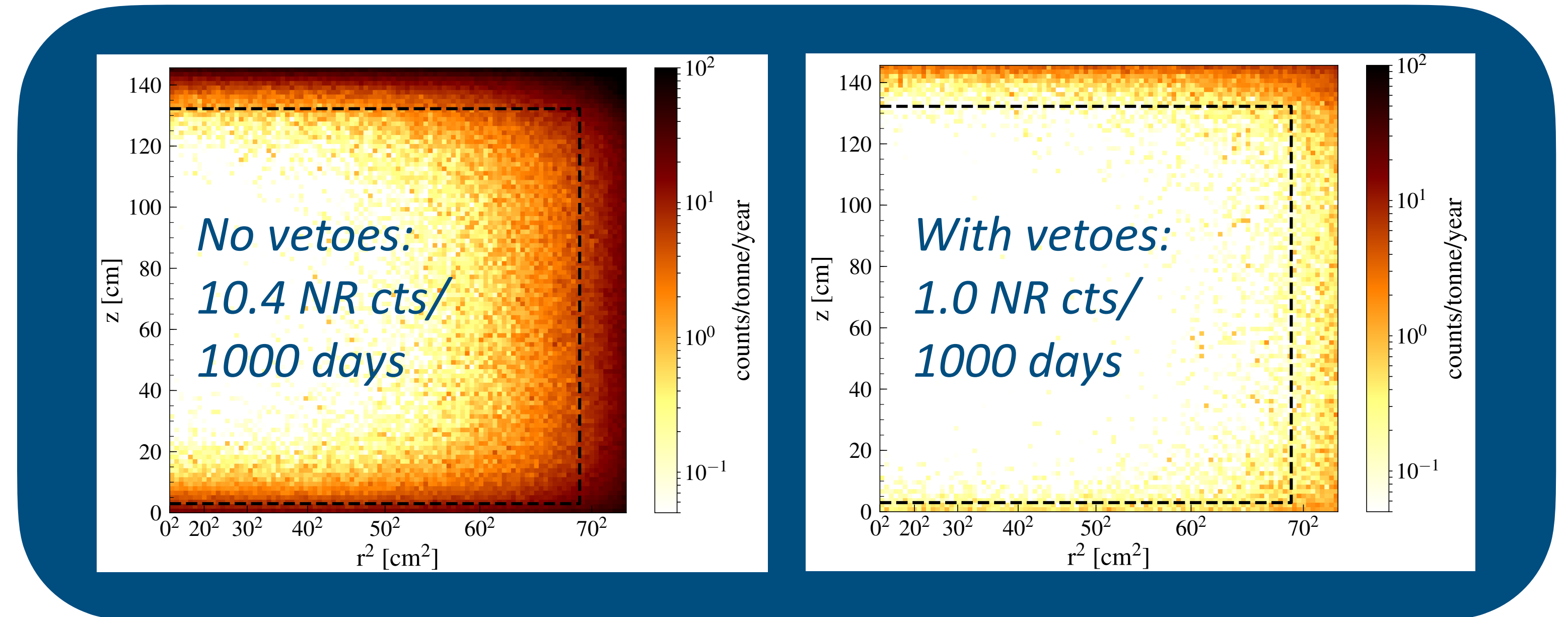
# TPC Detection Principle



- Interactions in the xenon create
  - Light – prompt scintillation – **S1**
  - Charge – electrons drifted and extracted into gas -> proportional scintillation – **S2**
- Excellent 3D position reconstruction (~mm)
  - Z from time difference between S1 and S2
  - XY from S2 hit pattern on top PMT array
- S2:S1 ratio – discriminate electronic recoils (ERs) from potential WIMP nuclear recoils (NRs)

# Background Mitigation

*EPJC, Vol 80: 1044 (2020)*



- Material selection – ~2000 assays; radon emanation & neutron activation analysis
- Cleanliness protocols – Rn-reduced cleanroom assembly → limit surface contamination
- Xenon purification – charcoal chromatography; online radon reduction & getter usage
- Analysis cuts – fiducialisation, single scatter, energy ROI, veto anti-coincidence

# WIMP Sensitivity

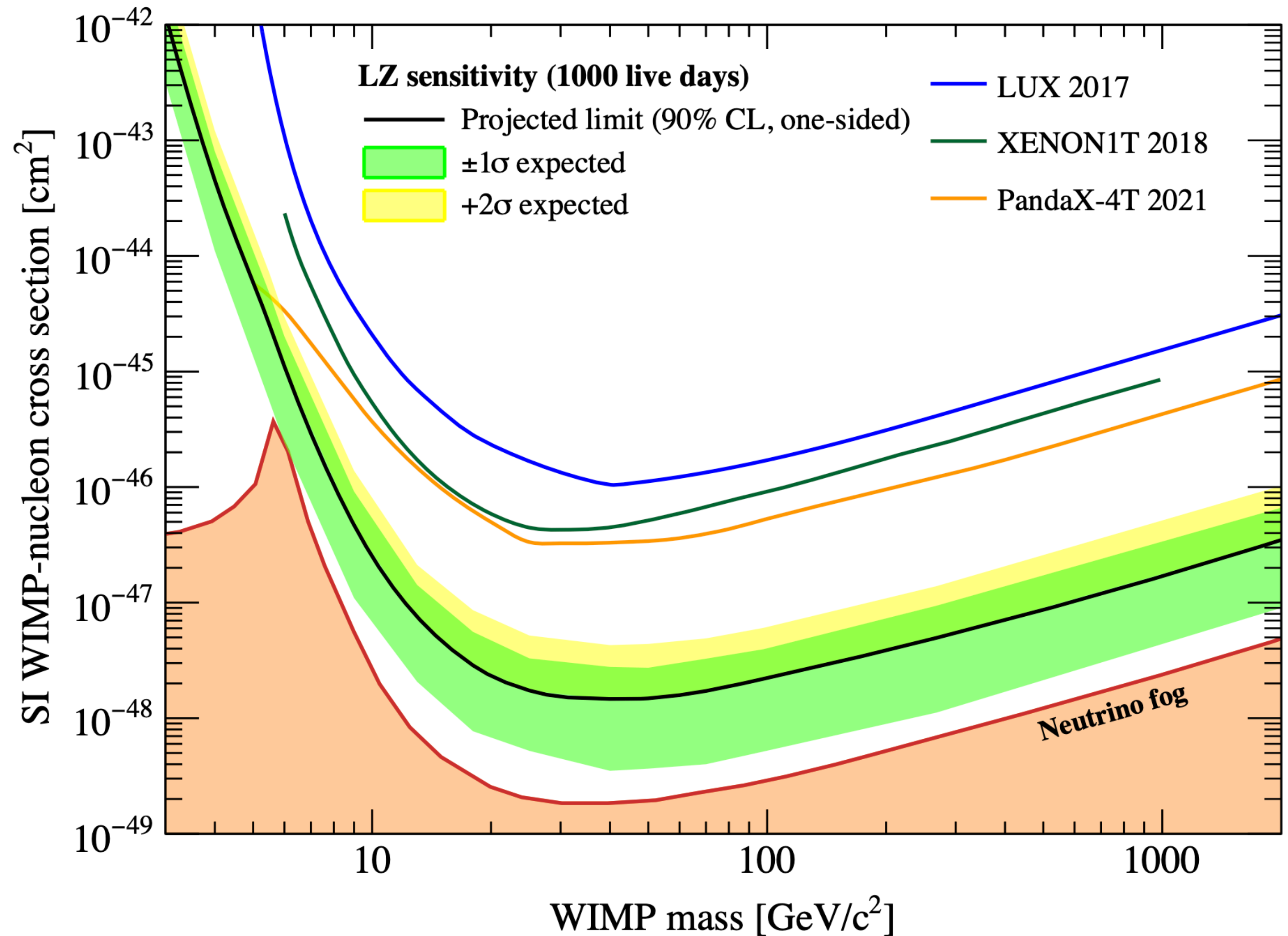
*PRD 101, 052002 (2020)*  
*j.astropartphys.2020.102480*

*For full exposure (1000 days \* 5.6 t):*

Background counts after analysis cuts in the 40 GeV/c<sup>2</sup> WIMP ROI

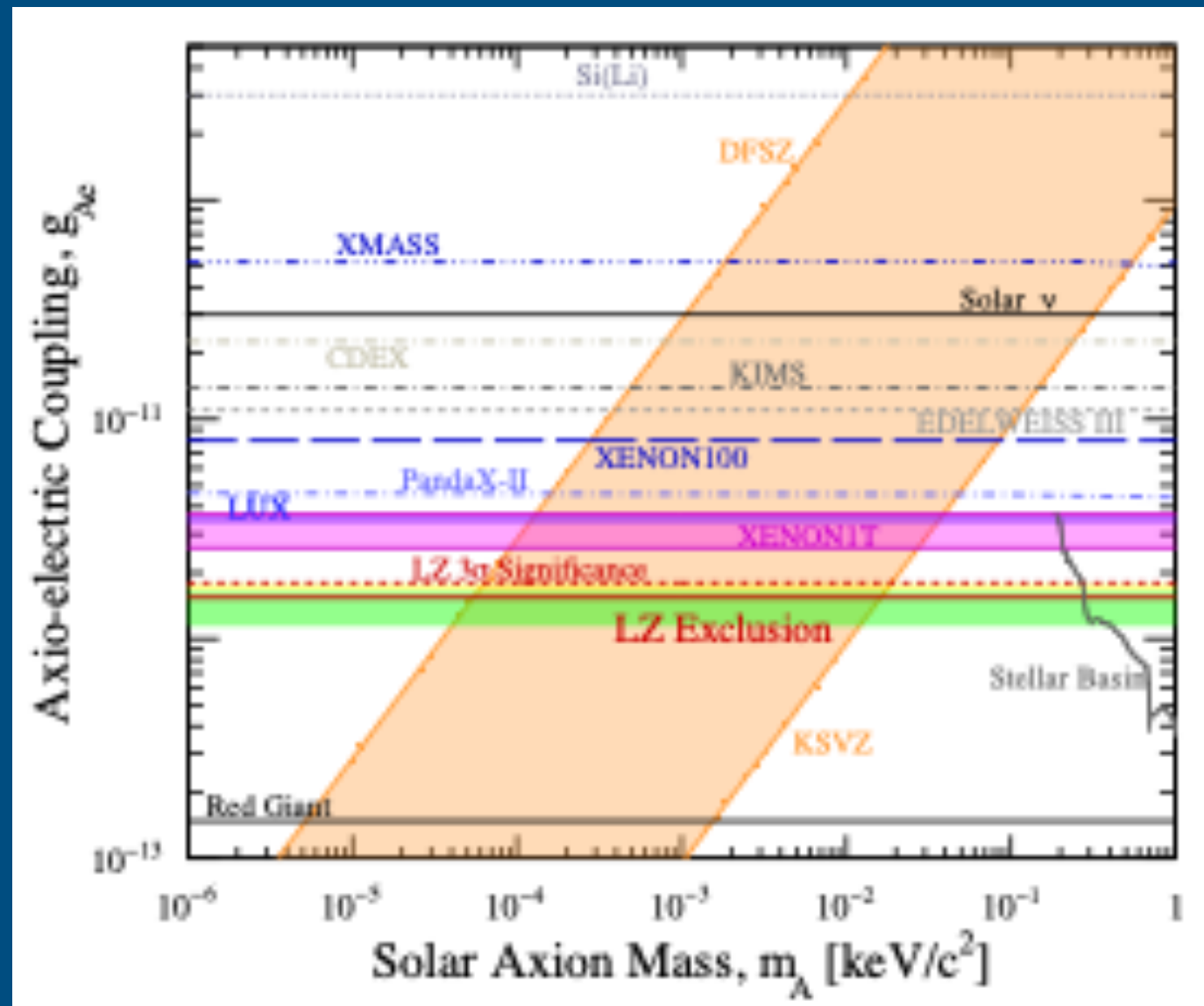
Source	ER [cts]	NR [cts]
Total	1131	1.03
+ 99.5% ER discrimination, 50% NR efficiency	5.66	0.52

90% CL minimum:  
 $1.4 \times 10^{-48} \text{ cm}^2$  at 40 GeV/c<sup>2</sup>

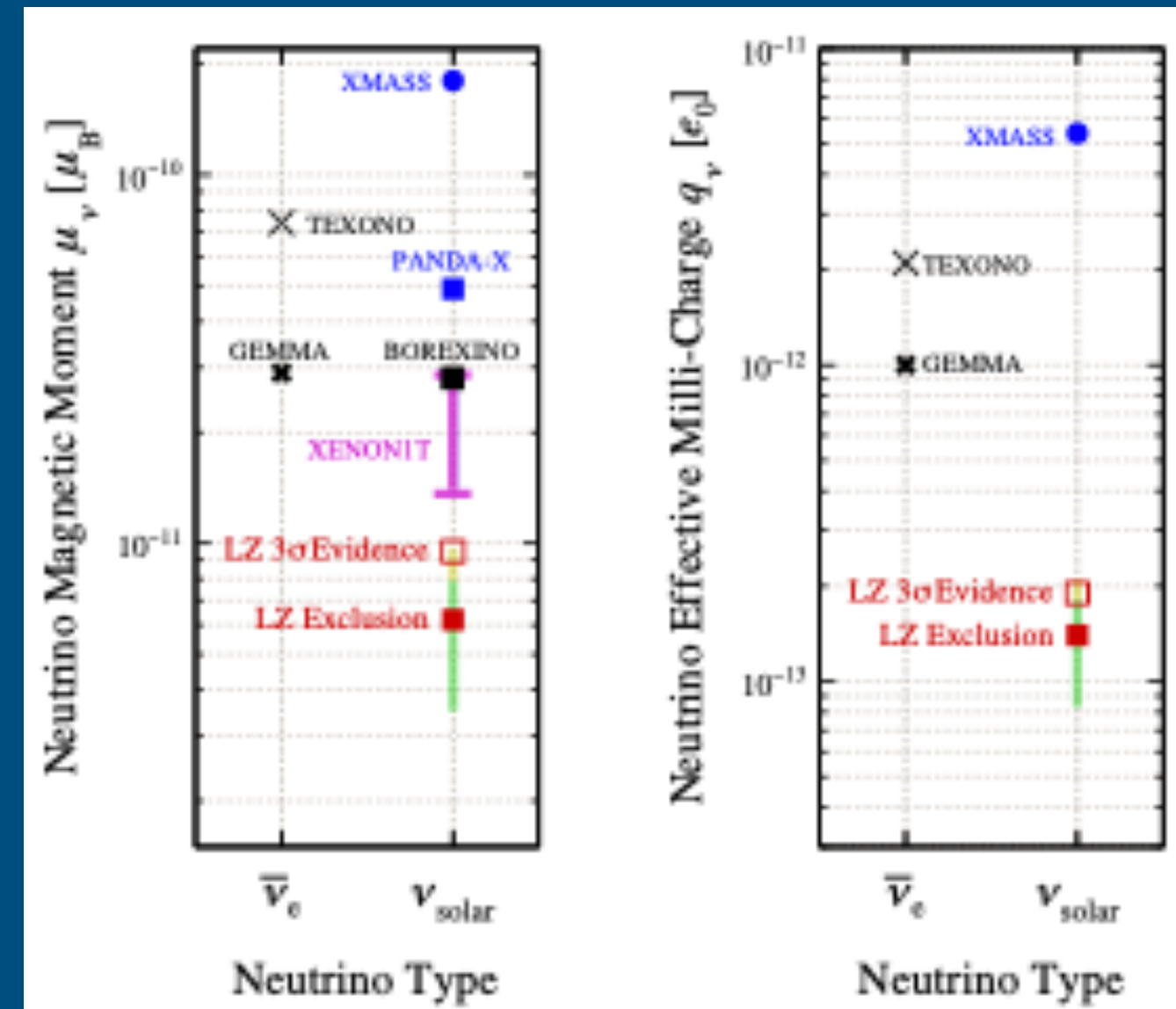


# Physics Reach Beyond WIMPs

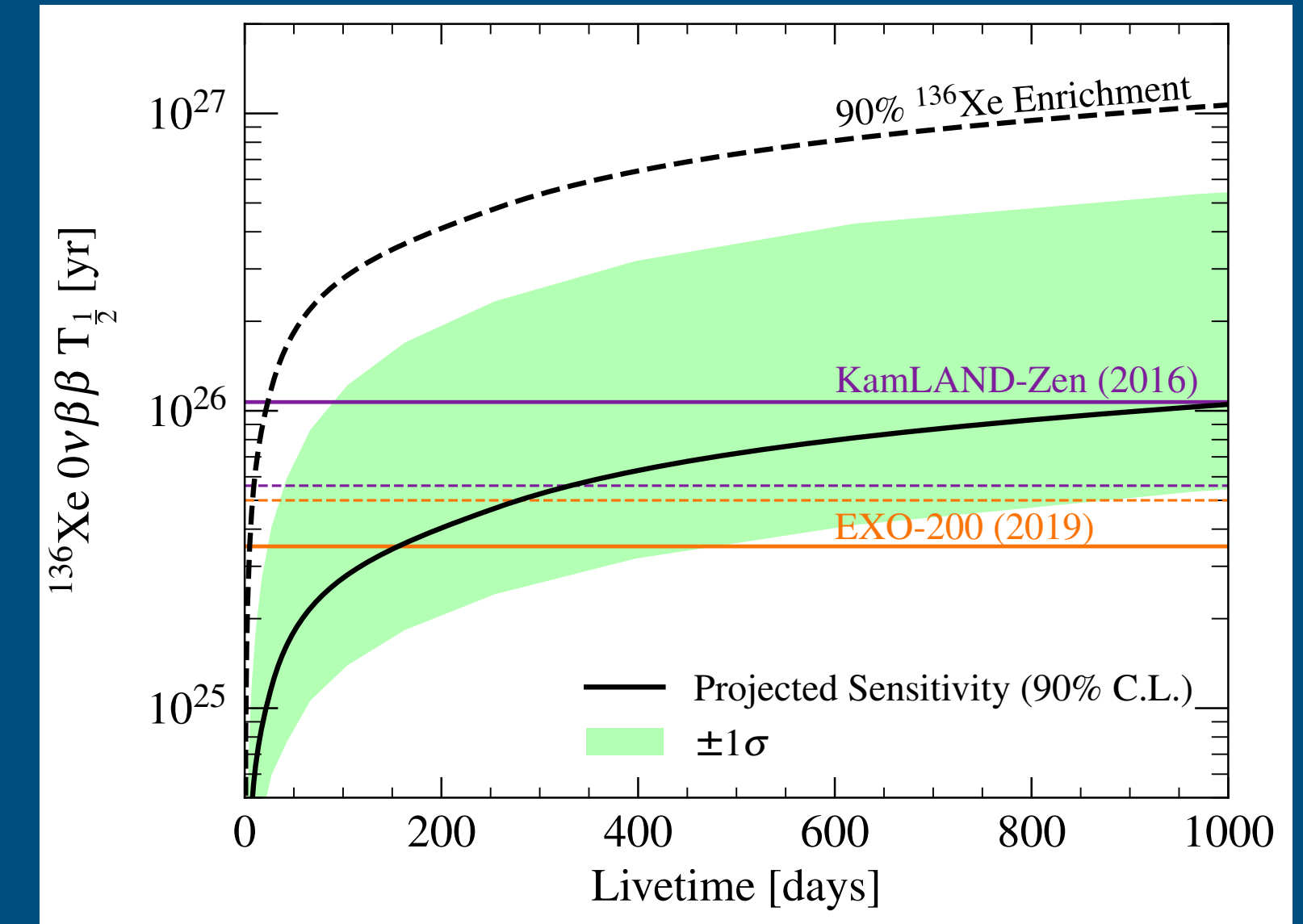
Solar Axions



Neutrino Magnetic Moment & Effective Millicharge



$^{136}\text{Xe}$  Neutrinoless Double Beta Decay



- Low energy ER – competitive for solar axions; neutrino  $\mu_\nu$ ,  $q_\nu$ ; axion-like particles + more
- High energy ER –  $^{134}\text{Xe}$   $2\nu\beta\beta$  &  $0\nu\beta\beta$ ,  $^{136}\text{Xe}$   $0\nu\beta\beta$  + exotic decay modes (see Z. Tong's talk)

[PRD 104, 092009 \(2021\)](#), [PRC 102, 014602 \(2020\)](#), [PRC 104, 065501 \(2021\)](#)

# Construction & Commissioning Timeline

TPC & Skin integration in the Surface Assembly Laboratory (SAL)



OD acrylic tanks underground



Inner cryostat vessel (ICV), housing TPC, underground



2018

2019

# Construction & Commissioning Timeline

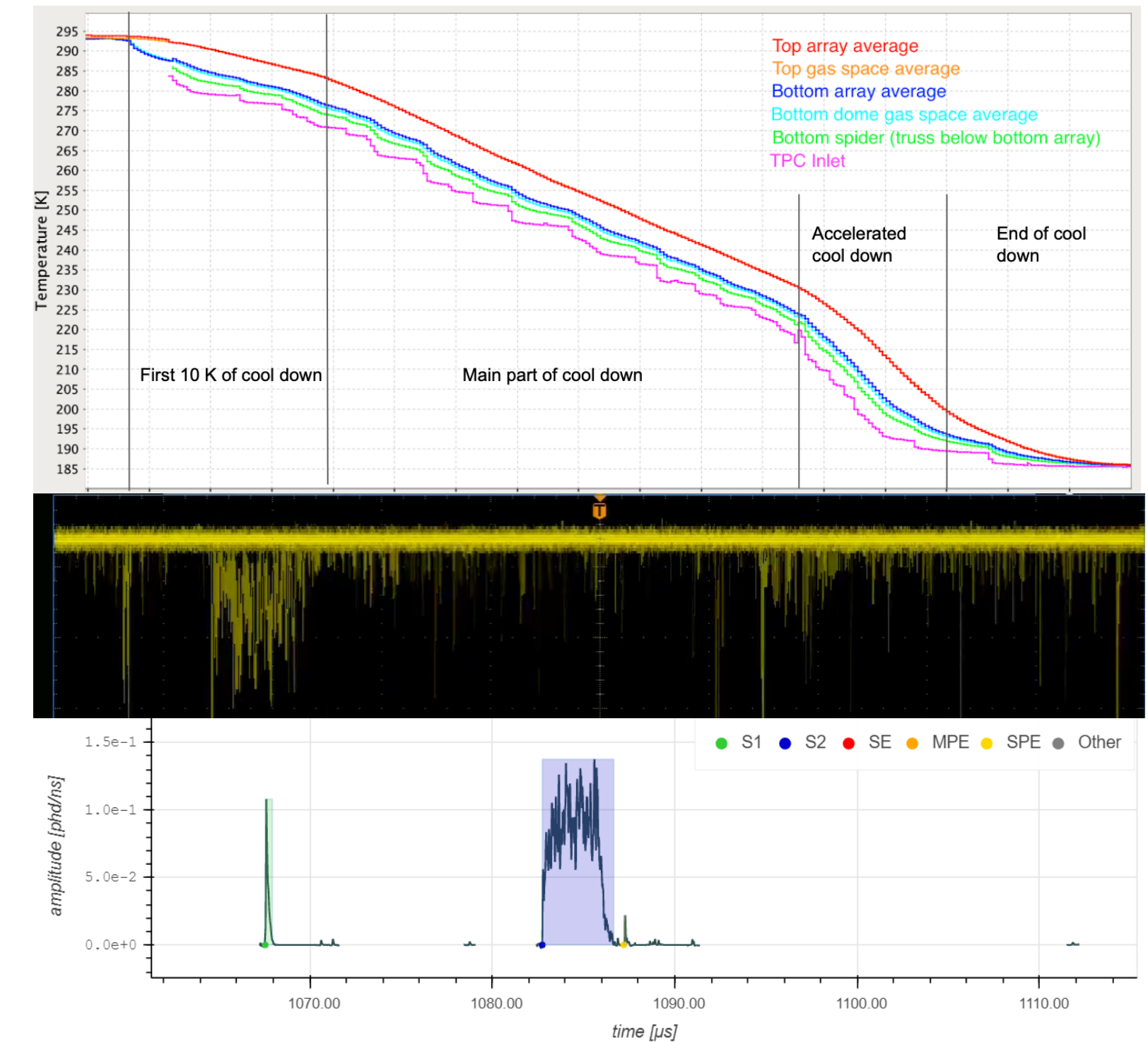
High voltage installed;  
ICV sealed & under vacuum



OD PMT & Tyvek  
Installation



## Detector Commissioning

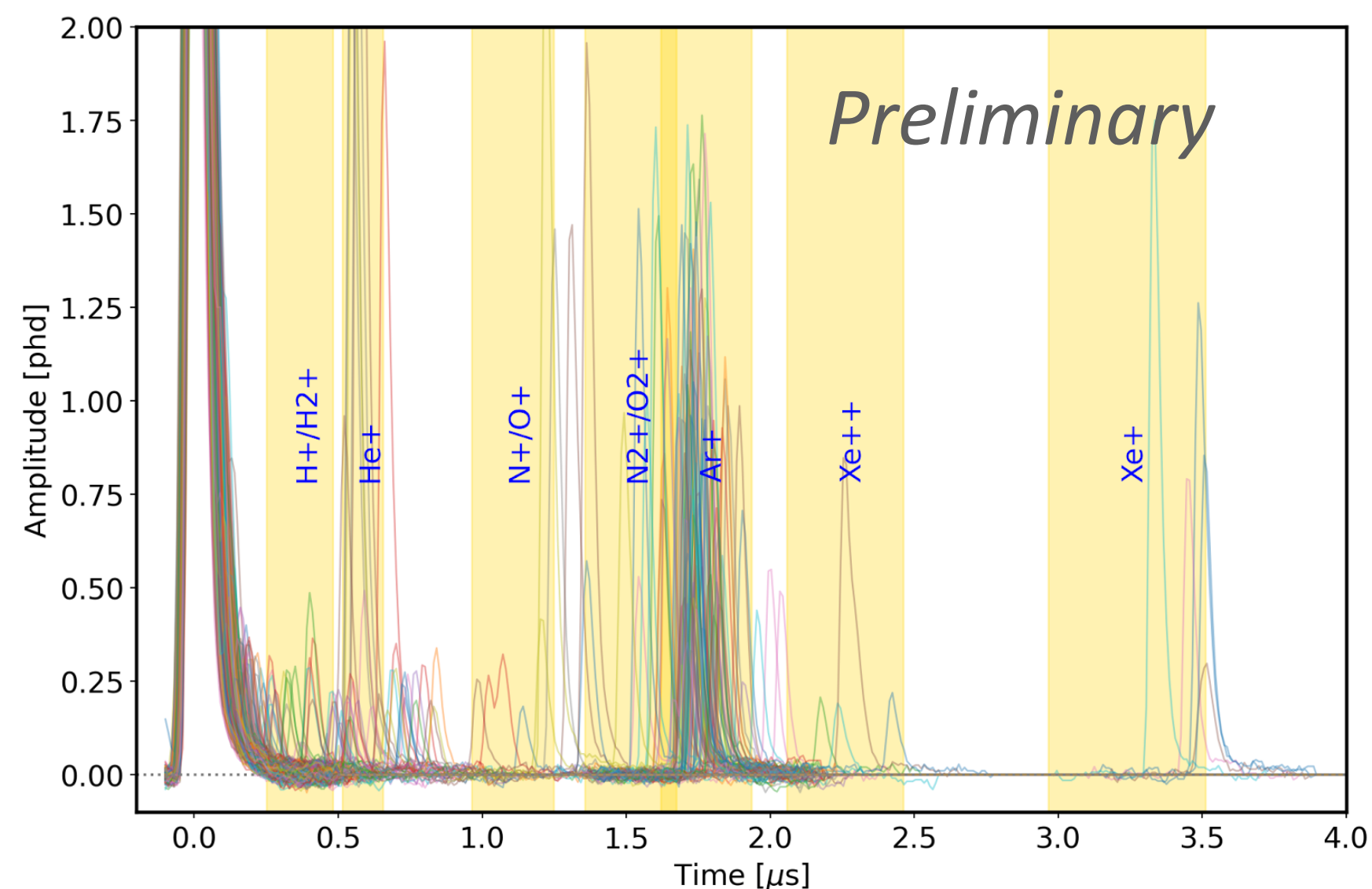
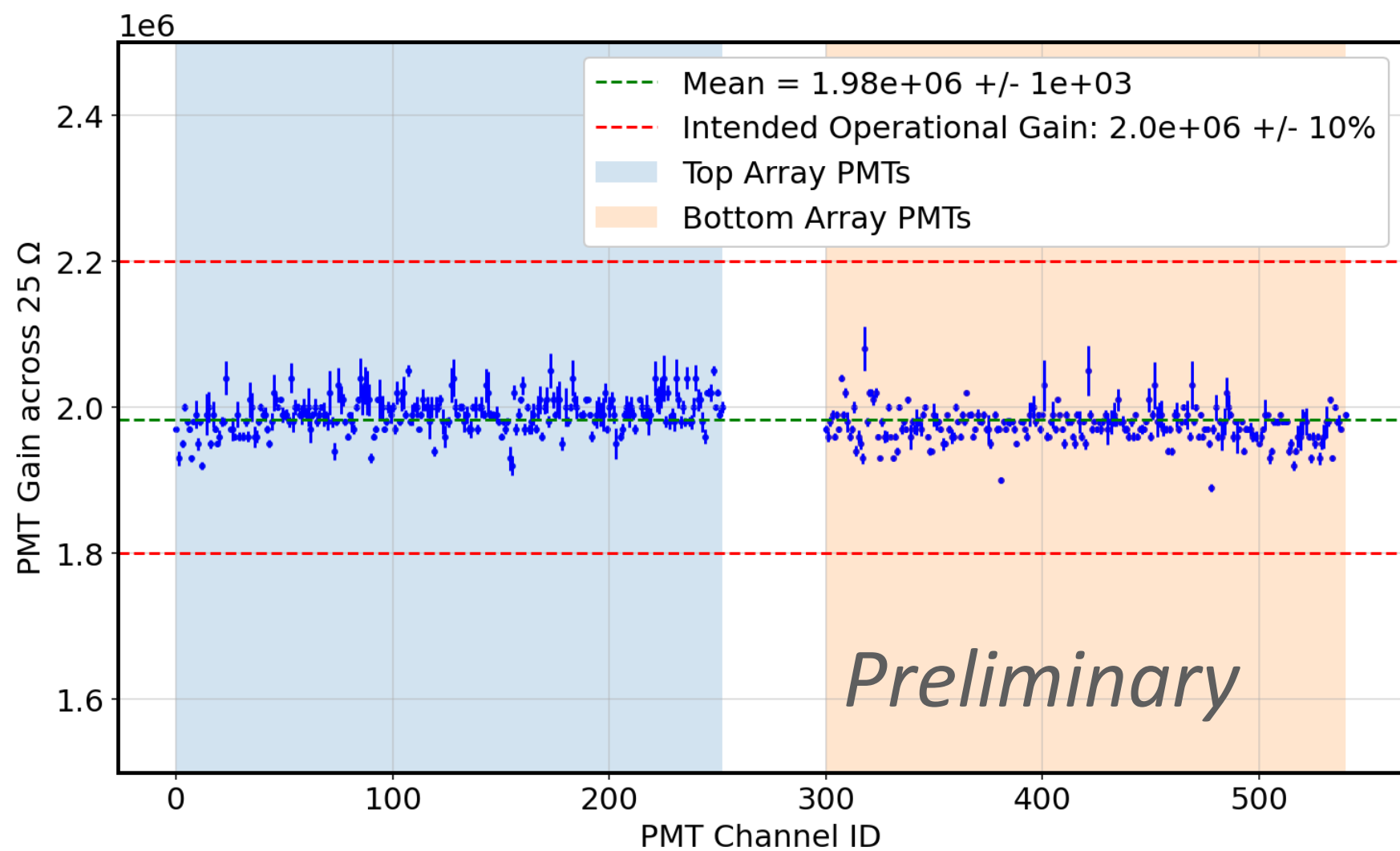


2020

2021



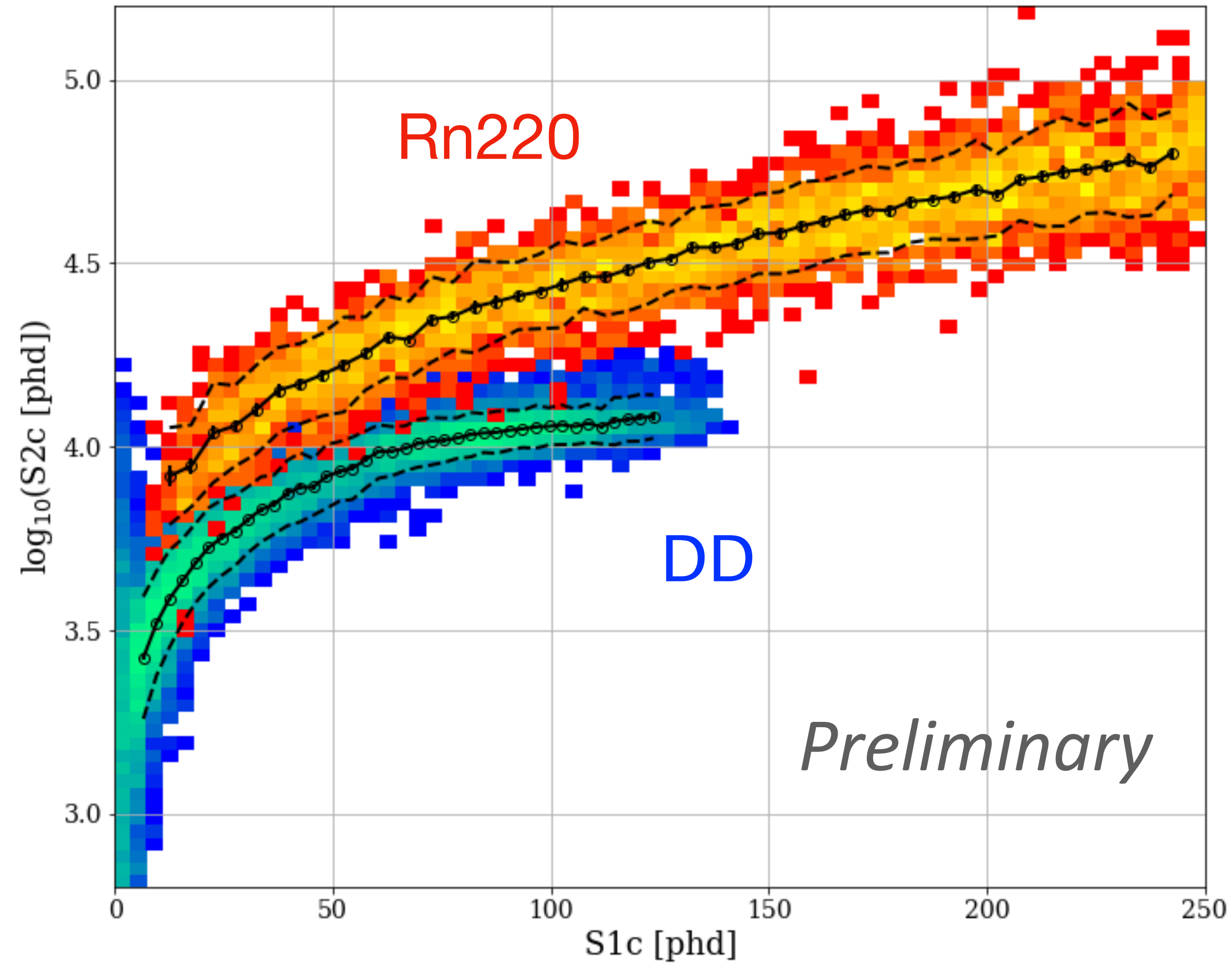
# Commissioning Activities



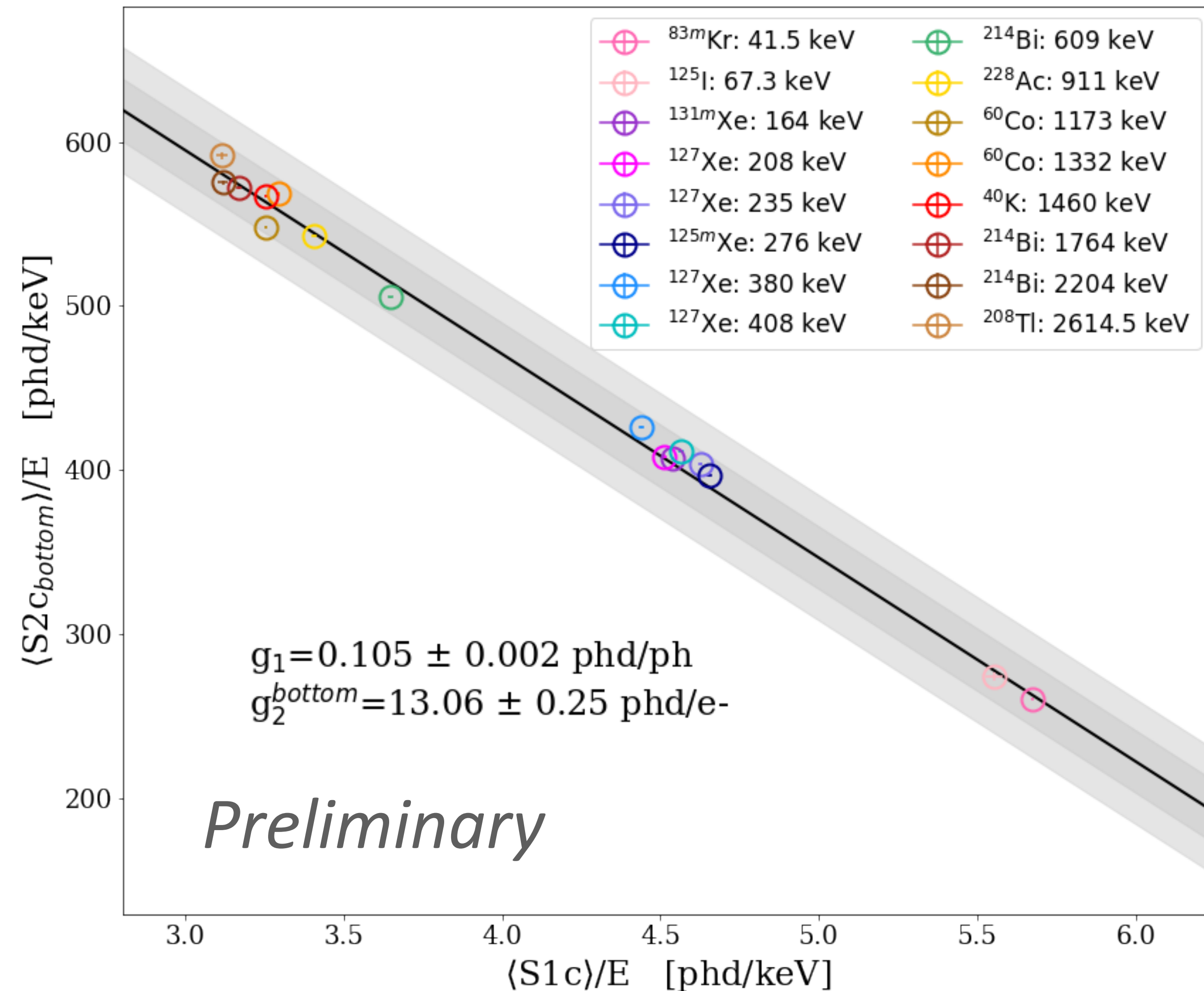
- TPC detector levelled & grids biased in liquid
  - $\sim 190\ \text{V/cm}$  drift &  $\sim 7.5\ \text{kV/cm}$  gas fields established
- Data processing chain exercised with first S1+S2s
- Data acquisition & trigger settings tuned
- PMT operations & characterisation
  - LED measurements for e.g. afterpulsing and single photoelectron (SPE) studies
  - PMTs gain-matched and gain drifts monitored
  - Dark count & DPE analyses (see A. Baker's talk)

# Calibrations

- Different deployment systems available
  - Circulation panel for injection into Xe
  - Vertical source tubes between TPC & Skin for commercial rod sources
  - DD neutron generator + TPC conduits (see J. Orpwood's poster)
- Calibrations can be used to inform
  - Energy scale in all three detectors
  - Inter-detector timings
  - NR & ER bands in the TPC



# Detector Response Characterisation



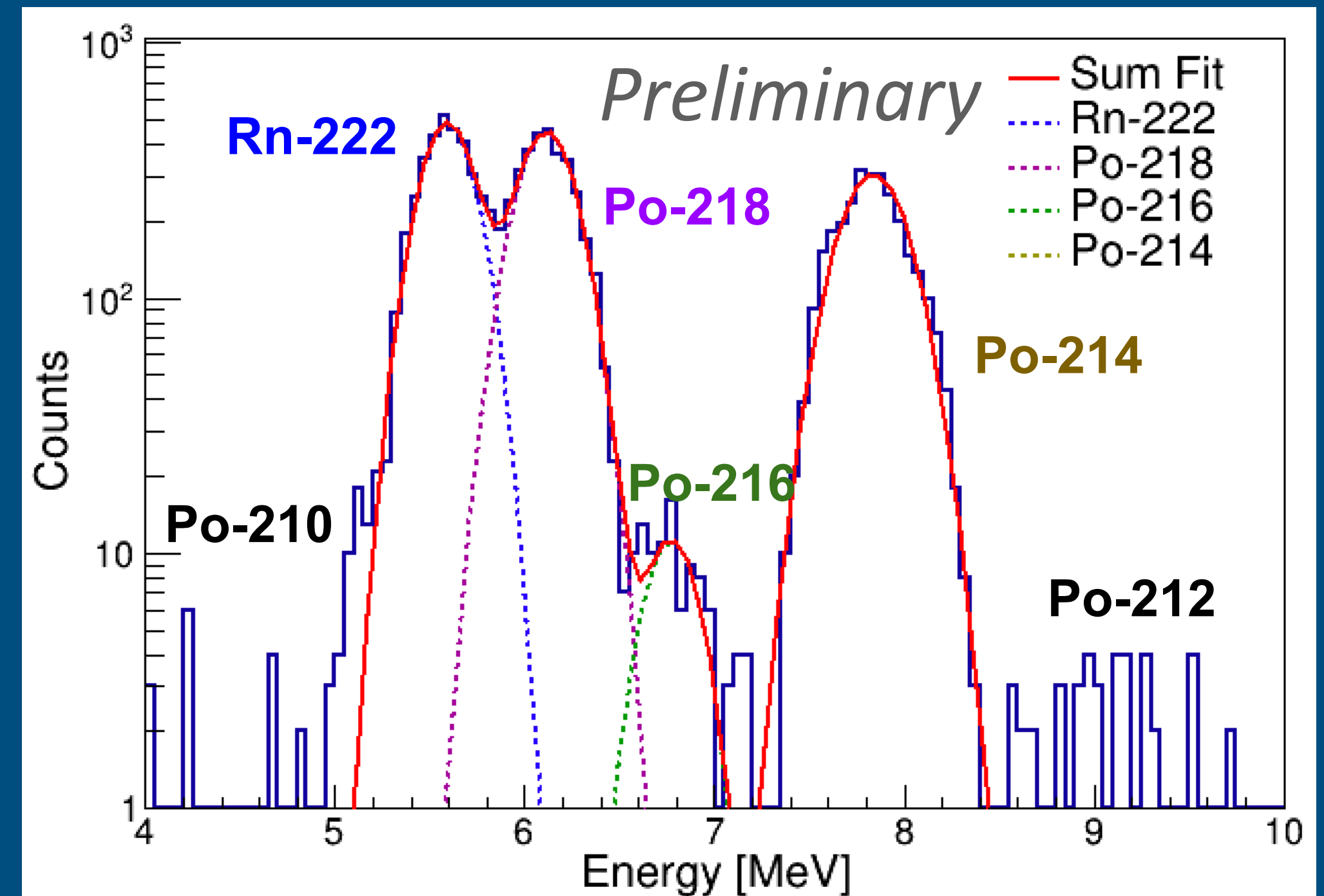
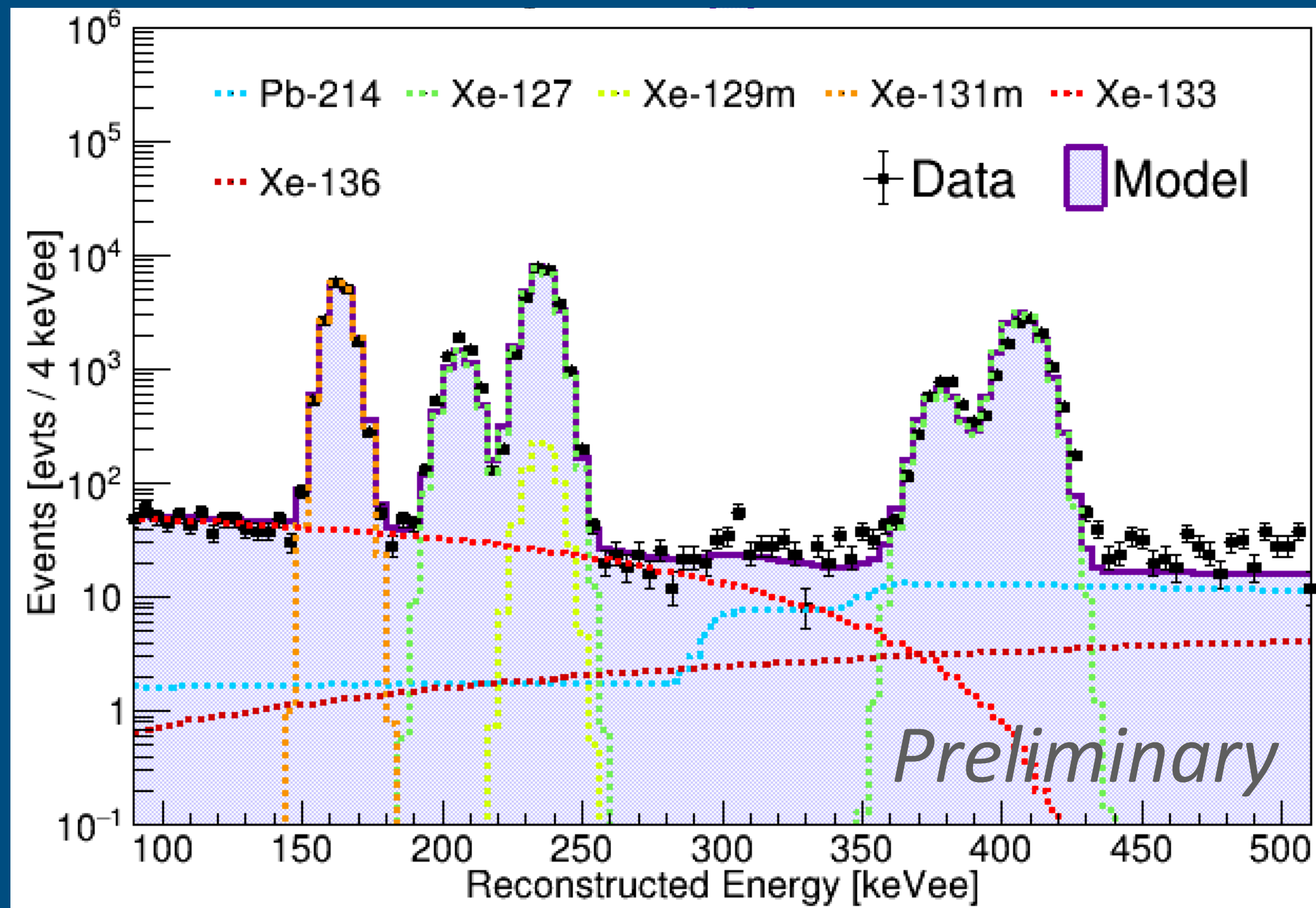
- Mono-energetic ER peaks can be used to find
  - $g_1$ , photons detected (phd) per prompt scintillation photon
  - $g_2$ , phd per ionisation electron

via the relation 
$$E = W \left( \frac{S1_c}{g_1} + \frac{S2_c}{g_2} \right),$$

where  $E$  = energy,  $W$  = W-value  
 $c$  denotes position-corrected signals

- Saturation of higher energy signals problematic (see A. Al Musalhi's talk)

# Background Analyses



- Trial fits of background simulations to data attempted, matching mono-energetic peaks
- Rn222 & Rn220 chain alpha populations identified & constrained (see N. Angelides' talk)
- Investigating non-xenon sources of charge/light e.g. Cherenkov (see I. Khurana's talk)

# Conclusions

- LZ is a multi-physics experiment, primed for the detection of WIMPs
- The experiment has started taking science data, and extensive analyses are underway

- Radon in LZ – N. Angelides
- Saturation corrections – A. Al Musalhi
- VUV detection – A. Baker
- Majoron searches – Z. Tong
- Cherenkov backgrounds – I. Khurana
- Multiple scatter studies in DD data – J. Orpwood





@lzdarkmatter

<https://lz.lbl.gov/>



# Acknowledgements



**LZ (LUX-ZEPLIN) Collaboration**  
**34 Institutions: 250 scientists, engineers, and technical staff**