

First Demonstration of Event-by-Event Directional Reconstruction in the SNO+ Scintillator Phase

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Motivation

- Directional information can be a very powerful tool for background rejection in neutrino experiments
 - Solar studies can benefit greatly from event selection based on direction
 - This information is often obtained through Cherenkov cone identification
- Scintillator experiments benefit from high light yield
 - This leads to high precision energy and position reconstruction
 - However, scintillator light is isotropic and gives no directional information
- If Cherenkov light can be isolated from the scintillation signal, directional reconstruction in scintillator could be possible
 - Slow scintillators have shown promise on a bench-top scale (Nucl. Instrum. Meth. A 972, 164106)

The SNO+ Detector

12 m diameter Acrylic Vessel (AV), filled with 780 T of liquid scintillator

> 9,300 inward facing PMTs for ~50% effective coverage



PMT Support Structure (PSUP)

The SNO+ Experiment 2021 JINST 16 P08059 Water shielding 1700 T inside PSUP 5300 T outside PSUP

Timing Separation

- SNO+ uses 2,5-diphenyloxazole (PPO) as a primary fluor in linear alkylbenzene (LAB)
- Lower concentration leads
 to a slower scintillator
- Easier separation of instantaneous Cherenkov light



SNO+ Partial Fill Phase

- A "Partial Fill" stage of SNO+ contained 0.6 g/L PPO
 - 5th April 24th October 2020



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Cherenkov Peak Isolation

- Two ways to separate Cherenkov light
- Time Residual $t_{res} = t_{hit} t_{event} t_{flight}$
 - Cherenkov will be earlier than scintillation
- Photon Angle:
 - Cherenkov will have a peak in θ_{γ}
- Cherenkov peak can be isolated by plotting the detected light using these parameters



2D Plot - Simulation



6 MeV electrons simulated

- Perfect-state full-fill detector
- Full volume, isotropic directions
- 0.6 g/L PPO scintillator
- Clear Cherenkov peak seen
- "Backwards" peak also seen
 - Caused by bias in positional reconstruction
- Can be used as a PDF for a directional likelihood fitter

MC Predictions

- ⁸B solar ν_{e} MC
 - Run conditions of the Partial Fill detector
 - Cuts: z > 1 m above equator, r < 5.5 m from centre of AV
- Grid search likelihood fitter
- $\cos(\theta_{Sun})$ angle between reconstructed and solar direction







- Interdependent effects
 - By reconstructing direction, it may be possible to correct for drive
 - By correcting for drive, we could more accurately reconstruct direction
- Work is ongoing for improvements to this reconstruction

Summary

- First demonstration of event-by-event* direction reconstruction
 < 10 MeV** in a liquid scintillator detector!
 - Also the first demonstration in a high yield scintillator**
 - Direction has been reconstructed for Solar events > 5 MeV in the SNO+ Partial Fill Phase
- Even more improvements to come!

- *Statistical separation of solar directionality has been shown by Borexino using ~20,000 events [1]
- ** LSND[2] MiniBooNE[3] have previously used directional reconstruction at higher energy scales using lower yield scintillators.

Thank you for listening

SNG

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References

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- 2. C Athanassopoulos et al. The liquid scintillator neutrino detector and LAMPF neutrino source. Nucl. Instrum. Meth. A, 388(1):149–172, 1997.
- 3. R.B. Patterson and E.M. Laird and Y. Liu and P.D. Meyers and I. Stancu and H.A. Tanaka. The extended-track event reconstruction for MiniBooNE. Nucl. Instrum. Meth. A, 608(1):206–224, 2009.