

(today & future)



Seminar @ RAL (UK) November 2020

Reactor Neutrino Potential in the era of the novel LiquidO Detection Technology

■ Wednesday 18 Nov 2020, 11:30 → 12:30 Europe/Londor **9** 700n

Anatael Cabrera CNRS/IN2P3

IJCLab @ Orsay LNCA @ Chooz



CNIS UNIVERSITE PARIS-SACLAY DES SCIENCES

FACULTÉ

D'ORSAY

Université de Paris

"A long time ago in a galaxy far, far away..."

Reines & Cowan (et al) around 1950

discover the neutrino (upon 1930's Pauli's hypothesis) [Nobel prize 1995]

pave much of today's technological ground [even ~70 years later, dominant today]

the v discovery (1950's)...

ic

inverse- β decay (IBD) interaction...

IBD: anti- $v_e + p \rightarrow e^+ + n$



no e+ PID implies $\mathbf{y} \approx \mathbf{e}^{-} \approx \mathbf{e}^{+} \approx \mathbf{\alpha} \approx \mathbf{p}$ -recoil (fast-n)



Double Chooz reactor-VETO: signal analysis...

rate(I reactor) ≈ IBD per 3 min

6



cool to the sun off for a while?

(very cool \rightarrow a dream for a few)

(fast) v oscillations reminder...

ingredients for neutrino oscillations...



Anatael Cabrera (CNRS-IN2P3 & APC)

9



where are we now (~2020)?

status on neutrino oscillation knowledge...

Standard Model(3 families)

 $\mathbb{PMNS}_{3\times3}(\theta_{12},\theta_{23},\theta_{13})$

 $\pm \Delta m^2 \& \pm \delta m^2$

no conclusive sign of any extension so far!!

(inconsistencies vs uncertainties)

must measure all parameters→characterise & test (i.e. over-constrain) Standard Model

	today			≥2030			
	best kno	owledge	NuFIT4.0	foreseen	dominant	technique	
θ12	3.0 %	sno	2.3 %	≲1.0%	JUNO	reactor	
θ23	5.0 %	NOvA	2.0 %	≲1.0%	DUNE⊕HK	beam (octant)	
θιз	1.8 %	DYB	1.5 %	I.5 %	DC⊕ <u>DYB</u> ⊕RENO	reactor	
+δm²	2.5 %	KamLAND	2.3 %	≲1.0%	JUNO	reactor	
∆m ²	3.0 %	T2K & DYB	1.3 %	≲1.0%	JUNO⊕DUNE⊕HK	reactor⊕beam	
sign(∆m²)	unknown	(SK et al)	NO @ ~3σ	@5σ	JUNO&DUNE&HK	reactor⊕beam	
CPV	unknown	(T2K et al)	3/2π @ ~2σ	@5σ?	DUNE⊕HK⊕ALL	beam driven	
			(Nov 2018)			(reactor-beam)	

essentially JUNO \oplus DUNE \oplus HK will lead most of the field (goal CPV) \rightarrow except θ_{13} !

NOTE: ORCA \oplus PINGU \oplus IceCube complementary (Mass Ordering & Δ m² measurements)

12

status on neutrino oscillation knowledge...

Standard Model(3 families)

[leptons & quarks]
&
PMNS_{3×3}(
$$\theta_{12}, \theta_{23}, \theta_{13}$$
)
&
 $\pm \Delta m^2 \& \pm \delta m^2$



(inconsistencies vs uncertainties)

must measure all parameters→characterise & test (i.e. over-constrain) Standard Model

	today				
	best kno	NuFIT4.0			
θ12	3.0 %	sk⊕sno	2.3 %		
θ23	5.0 %	NOvA+T2K	2.0 %		
θιз	1.8 %	DYB+DC+RENO	1.5 %		
+δm²	2.5 %	KamLAND	2.3 %		
Δm²	3.0 %	T2K+NOvA & DYB	1.3 %		
sign(∆m²)	unknown	SK et al	NO @ ~3σ		
CPV	unknown	T2K	3/2π @ ≲ 2σ		
			(now)		

13

(reactor-beam)

JUNO \oplus DUNE \oplus HK will lead precision in the field (\rightarrow CPV) except θ_{13} !

NOTE: ORCA \oplus PINGU \oplus IceCube complementary (Mass Ordering & Δ m² measurements)

status on neutrino oscillation knowledge...

Standard Model (3 families)

14

[leptons & quarks] & PMNS_{3×3}($\theta_{12}, \theta_{23}, \theta_{13}$) & &± Δm^2 & + δm^2

	today			≥2030			
	best kno	owledge	NuFIT4.0	foreseen	dominant	technique	
θ12	3.0 %	SK⊕SNO	2.3 %	<1.0%	JUNO & SC	reactor	
θ ₂₃	5.0 %	NOvA+T2K	2.0 %	≲ .0%	DUNE⊕HK [⊕SC??]	beam (octant)	
θιз	1.8 %	DYB+DC+RENO	1.5 %	<1.0%	Super Chooz (SC)	reactor	
+δm²	2.5 %	KamLAND	2.3 %	≲ .0%	JUNO	reactor	
∆m²	3.0 %	T2K+NOvA & DYB	1.3 %	≲ .0%	JUNO⊕DUNE⊕HK⊕ SC	reactor⊕beam	
Mass Ordering	unknown	SK et al	NMO @ <u>≤</u> 3σ	@5o	<u>JUNO</u> ⊕DUNE⊕HK (NOvA⊕T2K)	reactor⊕beam	
CPV	unknown	T2K+NOvA	3/2π @ <mark>≤2σ</mark>	@5σ?	DUNE⊕HK⊕SC	beam driven	
CPTV	assumed			< %?	SC?? [studying]	reactor+solar	
Unitarity	assumed			< %?	SC?? [studying]	reactor+solar	

(reactor+solar+beam)

all done?

by 2030, mixing @ ~1% level.. (**no unknowns**)



despite major success so far... **challenges** leads **discoveries** (and fun)!!





reactor no direct CPV, but...

T2K⊕reactor best knowledge CP-Violation...



19

θΙ3 implications CPV phase vs θΙ3 [constrained by reactor]

CPV phase vs θ23

[octant ambiguity]

CPV phase vs (Atmospheric) Mass Ordering [T2K blinded]



consider matrix structure (not just composition)

why shape?

•large mixing but a small one!

largest CP-Violation (SM)

•any symmetry behind? [Nature's caprice?]



[next slides]

20

CKM vs PMNS...



d s b



elegance (symmetry)



A. De Gouvea, H. Murayama, hep-ph/0301050; PLB, 2015.L. Hall, H. Murayama, N. Weiner, hep-ph/9911341.

Unitarity: the last discovery?



since no CPV (yet) ⇒ test PMNS Unitarity via "each row"

$$U_{l1}|^{2} + |U_{l2}|^{2} + |U_{l3}|^{2} = 1$$

 $|U_{e1}|^2 + |U_{e2}|^2 + |U_{e3}|^2 = 1 \Rightarrow$ explore "electron top-row"

only " θ_{12} " and " θ_{13} "

unitarity violation implications...



non-standard v states and/or non-standard v interaction well definition theory/experimental problem

perfect prediction ("symmetry")
experimentally precision accessible?

•neutrino: direct & clean probe [no "slippery" corrections?]

major!! discovery potential [building blocks of SM]

(if discovery) possible experimental redundancy

Unitarity violation test...

today's status.



2020 world status in Θ 13...





ARTICLE

First Double Chooz θ_{13} Measurement via Total Neutron Capture Detection

Hervé de Kerret et al (arXiv:1901.094451)

²⁹ summary on today's θ I 3 knowledge/experiments...

reactor-θI3 experiments [DC⊕DYB⊕RENO]

statistics: ~10⁵ (far) [<10⁶]
systematics: ~0.1% (each)
energy control: <1% precision

	<2010	tod	ay [2010-20	cancellation		
	total	total	rate-only	shape-only	methodology	
statistics	few %	~0.1%			~100/day @ 1.5km	
flux	~2.2%	~0.1%	~0.1%	<0.1%	near-to-far monitor (ideal: iso-flux)	
BG	few %	~0.1%	~0.1%	<0.1%	overburden→few/day	
detection	2.0 %	~0.1%	~0.1%		identical detectors	
energy	few %	~0.5%		~0.5%	identical detectors	

"naively extrapolating" from reactor-θ|3 experiments...
 •statistics: ~|0^x? (far) [>|0⁶]
 •systematics: ~0.0|%??!! (each)





JUNO location...



32

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JUNO a photocathode colosso→ yield energy resolution!

Anatael Cabrera (CNRS-IN2P3 & APC)

³⁴ "solar" oscillation measured by both PMT systems...

SPMT sees the "solar" oscillation

(fast oscillation washes out)



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LPMT vs SPMT comparison...



readout explore $\theta | 2 \oplus \delta m^2$ to per-mille precision ($\leq | \%$)


(reactor flux)

37

reactor flux uncertainty...



MC normalised to DYB-2017 (MCSpF per isotope)



ARTICLE

First Double Chooz θ_{13} Measurement via Total Neutron Capture Detection

38



Unitarity?

Φ(reactor) ⊕ θΙ3(now) ⊕ θΙ2(JUNO)

today's e-row unitarity knowledge...

–I. Nunokawa e*t al* (arXiv:1609.08623v2)

41



even with JUNO, sub-perfect explorations IMPOSSIBLE!

Mass Ordering? reactors' role?

PREPARED FOR SUBMISSION TO JHEP

The fate of hints: updated global analysis of three-flavor neutrino oscillations

Ivan Esteban,^{*a*} M. C. Gonzalez-Garcia,^{*a,b,c*} Michele Maltoni,^{*d*} Thomas Schwetz,^{*e*} Albert Zhou^{*e*}

- ^aDepartament de Fisíca Quàntica i Astrofísica and Institut de Ciencies del Cosmos, Universitat de Barcelona, Diagonal 647, E-08028 Barcelona, Spain
- ^bInstitució Catalana de Recerca i Estudis Avançats (ICREA), Pg. Lluis Companys 23, 08010 Barcelona, Spain.
- ^cC.N. Yang Institute for Theoretical Physics, State University of New York at Stony Brook, Stony Brook, NY 11794-3840, USA
- ^dInstituto de Física Teórica UAM/CSIC, Calle de Nicolás Cabrera 13–15, Universidad Autónoma de Madrid, Cantoblanco, E-28049 Madrid, Spain
- e Institut für Kernphysik, Karlsruher Institut für Technologie (KIT), D-76021 Karlsruhe, Germany

E-mail: ivan.esteban@fqa.ub.edu, maria.gonzalez-garcia@stonybrook.edu,michele.maltoni@csic.es, schwetz@kit.edu, albert.zhou@kit.edu

ABSTRACT: Our herein described combined analysis of the latest neutrino oscillation data presented at the Neutrino2020 conference shows that previous hints for the neutrino mass ordering have significantly decreased, and normal ordering (NO) is favored only at the 1.6σ level. Combined with the χ^2 map provided by Super-Kamiokande for their atmospheric neutrino data analysis the hint for NO is at 2.7σ . The CP conserving value $\delta_{\rm CP} = 180^{\circ}$ is within 0.6σ of the global best fit point. Only if we restrict to inverted mass ordering, CP violation is favored at the $\sim 3\sigma$ level. We discuss the origin of these results – which are driven by the new data from the T2K and NOvA long-baseline experiments–, and the relevance of the LBL-reactor oscillation frequency complementarity. The previous 2.2σ tension in Δm_{21}^2 preferred by KamLAND and solar experiments is also reduced to the 1.1σ level after the inclusion of the latest Super-Kamiokande solar neutrino results. Finally we present updated allowed ranges for the oscillation parameters and for the leptonic Jarlskog determinant from the global analysis.

KEYWORDS: neutrino oscillations, solar and atmospheric neutrinos

today's world data leads to...

NMO favoured to $\sim 2.7\sigma$ (2020)

main experiments so far...

- •SK
- •NOvA⊕T2K
- •DC⊕DYB⊕RENO

JUNO will provide a ~3σ result in ≥2028

why important?
too late & too little?
ever enough?

NOTE: almost impossible to $\ge 3\sigma$ alone!

today's MO status & JUNO..

Earliest Resolution to the Neutrino Mass Ordering?

Anatael Cabrera^{*1,2,4}, Yang Han^{†1,2}, Michel Obolensky¹, Fabien Cavalier², João Coelho², Diana Navas-Nicolás², Hiroshi Nunokawa^{‡2,7}, Laurent Simard², Jianming Bian³, Nitish Nayak³, Juan Pedro Ochoa-Ricoux³, Bedřich Roskovec³, Pietro Chimenti⁵, Stefano Dusini^{6a}, Marco Grassi^{6b}, Mathieu Bongrand^{8,2}, Rebin Karaparambil⁸, Victor Lebrin⁸, Benoit Viaud⁸, Frederic Yermia⁸, Lily Asquith⁹, Thiago J. C. Bezerra⁹, Jeff Hartnell⁹, Pierre Lasorak⁹, Jiajie Ling¹⁰, Jiajun Liao¹⁰, and Hongzhao Yu¹⁰

¹APC, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité University, 75205 Paris Cedex 13, France
 ²IJCLab,, Université Paris-Saclay, CNRS/IN2P3, 91405 Orsay, France
 ³Department of Physics and Astronomy, University of California at Irvine, Irvine, California 92697, USA
 ⁴LNCA Underground Laboratory, CNRS/IN2P3 - CEA, Chooz, France
 ⁵Departamento de Física, Universidade Estadual de Londrina, 86051-990, Londrina - PR, Brazil
 ^{6a}INFN, Sezione di Padova, via Marzolo 8, I-35131 Padova, Italy
 ^{6b}Dipartimento di Fisica e Astronomia, Università di Padova, via Marzolo 8, I-35131 Padova, Italy
 ⁷Department of Physics, Pontifícia Universidade Católica do Rio de Janeiro, Rio de Janeiro, RJ, 22451-900, Brazil
 ⁸SUBATECH, CNRS/IN2P3, Université de Nantes, IMT-Atlantique, 44307 Nantes, France
 ⁹Department of Physics and Astronomy, University of Sussex, Falmer, Brighton BN1 9QH, United Kingdom
 ¹⁰Sun Yat-sen University, NO. 135 Xingang Xi Road, Guangzhou, China, 510275

August 27, 2020 - v3.5

when can we resolve (≥5σ) the neutrino Mass Order? [earliest time scale]

which experiments (many planned, but <u>minimal set</u>) can yield the full resolution?

what physics exploited to yield the full resolution?

implications beyond the Standard Model?

NuFit5.0 (July 2020): Normal-MO favoured to ~3σ

our paper's goal.

44



45

combination (JUNO NOvA T2K)

• first MO measurement @ $\geq 5\sigma$ possible ($\geq 90\%$ CL)

combined both vacuum⊕matter MO information
 [less clean but powerful]

•JUNO schedule: ≥2028 [T2K / NOvA stopped ≤2026]

•preparation 3rd beam generation (DUNE & HyperK)
 → DUNE the most powerful standalone MO experiment

combination (JUNO⊕HyperK⊕DUNE)

DUNE (CPV depends, but any — >2years data) \rightarrow matter only $\geq 5\sigma$ MO

JUNO \oplus **HyperK** \oplus **DUNE**(CPV rather insensitive) \rightarrow **vacuum only** \geq **5** σ **MO** (only Δ m²₃₂ disappearance)

new physics? yes, if discrepancies found! [→discovery!]

$JUNO \oplus NOVA \oplus T2K: MO \ge 5\sigma by 2028$

powerful synergy JUNO vs NOvA \oplus T2K: high precision disappearance $\Delta m^{2}_{32...}$



JUNO MO sensitivity boosted $3\sigma \rightarrow \geq 5\sigma$ [leading order effect] **physics:** extra discriminator due to Δm_{32}^2 solutions slightly different (i.e. synergy) between reactor-accelerator but only one true MO solution forces equality \rightarrow powerful boosting with precision of Δm_{32}^2 .

Mass Ordering JUNO boosting...

Anatael Cabrera CNRS-IN2P3 / IJCLab (Orsay) - LNCA (Chooz) Laboratories

PREPARED FOR SUBMISSION TO JHEP

The fate of hints: updated global analysis of three-flavor neutrino oscillations

Ivan Esteban,^{*a*} M. C. Gonzalez-Garcia,^{*a,b,c*} Michele Maltoni,^{*d*} Thomas Schwetz,^{*e*} Albert Zhou^{*e*}

- ^aDepartament de Fisíca Quàntica i Astrofísica and Institut de Ciencies del Cosmos, Universitat de Barcelona, Diagonal 647, E-08028 Barcelona, Spain
- ^bInstitució Catalana de Recerca i Estudis Avançats (ICREA), Pg. Lluis Companys 23, 08010 Barcelona, Spain.
- ^cC.N. Yang Institute for Theoretical Physics, State University of New York at Stony Brook, Stony Brook, NY 11794-3840, USA
- ^dInstituto de Física Teórica UAM/CSIC, Calle de Nicolás Cabrera 13–15, Universidad Autónoma de Madrid, Cantoblanco, E-28049 Madrid, Spain
- e Institut für Kernphysik, Karlsruher Institut für Technologie (KIT), D-76021 Karlsruhe, Germany

E-mail: ivan.esteban@fqa.ub.edu, maria.gonzalez-garcia@stonybrook.edu, michele.maltoni@csic.es, schwetz@kit.edu, albert.zhou@kit.edu

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KEYWORDS: neutrino oscillations, solar and atmospheric neutrinos

today's world data leads to...

NMO favoured to $\sim 2.7\sigma$ (2020)

main experiments so far...

- •SuperK
- •NOvA⊕T2K
- •DC⊕DYB⊕RENO

JUNO will provide a ~3σ result in ≥2028

JUNO critical upon combination •key player for discovery (≥5σ)! •unique vacuum boost to ≥5σ! •unique explorations BSM? (thinking)

NOTE: almost impossible to $\ge 3\sigma$ alone!

today's MO status & JUNO..

and beyond... the future...?

v's left Europe next decade...?



how to reduce BG with no more overburden?

esson: avoid civil construction (reactor!!)...

LIQUID

a novel neutrino detection

antimatter (e+) tagging

BG active rejection

[only "mundane" matter]

LiquidO active **Particle-ID**...

Imaging→powerful Particle-IDentification (PID)



LiquidO ≈ PID ⊕ (high) Doping

physics beyond detector "native composition" (H,C)

diffusion —> shaper images!

LiquidO event-wise imaging...



opaque scintillator→stochastic light confinement (self-segmentation)

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53

(30m overburden)



54

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most basic principle.



LiquidO: unique stochastic light confinement (mainly lossless)

scattering...a "milky business"!

Computing the Scattering Properties of Participating Media Using Lorenz-Mie Theory

Jeppe Revall Frisvad¹

Niels Jørgen Christensen¹

Henrik Wann Jensen²

¹Informatics and Mathematical Modelling, Technical University of Denmark ²University of California, San Diego



Figure 1: Rendered images of the components in milk as well as mixed concentrations. The optical properties of the components and the milk have been computed using the generalization of the Lorenz-Mie theory presented in this paper. From left to right the glasses contain: Water, water and vitamin B2, water and protein, water and fat, skimmed milk, regular milk, and whole milk.

Mie scattering (well known) used to study samples

LiquidO theorem...

''milky'' / ''cloudy'' / ''waxy'' scintillator/cherenkov (gas↔liquid↔solid)

Anatael Cabrera (CNRS-IN2P3 @ LAL - LNCA)



LiquidO's multi-axes...

LiquidO

up to 3 axes (unlike drift-TPC)→**needed**?



Transparent Scintillator •Fibres

Anatael Cabrera (CNRS-IN2P3 @ LAL - LNCA)

light readout via ''collectors''...

60



(beyond chemical stability)



our digitisation electronics.

scintillation+Cherenkov

few-PE's pulses (100's of ns sampling)

~150ps per sample

(expected) time resolution: $\leq 100 \text{ps/PE}$ (i.e. $\leq 3 \text{cm/PE}$ @ speed of light)

IEEE NSS (2014): The SAMPIC Waveform and Time to Digital Converter NIM paper (2016): Measurements of timing resolution of ultra-fast silicon detectors with the SAMPIC Waveform Digitizer

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(instrumentation-wise)

LiquidO ≈ "light" TPC ⊕ 4π-ToF

LiquidO "bread & butter" physics...



today's technology



LiquidO technology

"ephemeral foam" of optical photons

light ball size: scattering⊕fibres (stochastic light confinement)

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LiquidO full release 2019...

Neutrino Physics with an Opaque Detector

A. Cabrera^{*1,9,10}, A. Abusleme¹⁵, J. dos Anjos^{†3}, T. J. C. Bezerra¹⁸, M. Bongrand⁹, C. Bourgeois⁹, D. Breton⁹, C. Buck¹², J. Busto⁶, E. Calvo⁵, E. Chauveau⁴, M. Chen¹⁶, P. Chimenti¹¹, F. Dal Corso¹³, G. De Conto¹¹, S. Dusini¹³, G. Fiorentini^{7a,7b}, C. Frigerio Martins¹¹, A. Givaudan¹, P. Govoni^{2a,2b}, B. Gramlich¹², M. Grassi^{1,9}, Y. Han^{1,9} J. Hartnell¹⁹, C. Hugon⁶, S. Jiménez⁵, H. de Kerret^{‡1}, A. Le Nevé⁹, P. Loaiza⁹, J. Maalmi⁹, F. Mantovani^{7a,7b} L. Manzanillas⁹, C. Marquet⁴, J. Martino¹⁸, D. Navas⁵, H. Nunokawa¹⁴, M. Obolensky¹, J. P. Ochoa-Ricoux^{8,15} G. Ortona²⁰, C. Palomares⁵, F. Pessina¹⁴, A. Pin⁴, M. S. Pravikoff⁴, M. Roche⁴, B. Roskovec⁸, N. Roy⁹, C. Santos¹ A. Serafini^{7a,7b}, L. Simard⁹, M. Sisti^{2a,2b}, L. Stanco¹³, V. Strati^{7a,7b}, J.-S. Stutzmann¹⁸, F. Suekane^{*§1,17}, A. Verdugo⁵, B. Viaud¹⁸, C. Volpe¹, C. Vrignon¹, S. Wagner¹, and F. Yermia¹⁸ ¹APC, CNRS/IN2P3, CEA/IRFU, Observatoire de Paris, Sorbonne Paris Cité University, 75205 Paris Cedex 13, France

^{2a}Università di Milano-Bicocca, I-20126 Milano, Italy ^{2b}INFN, Sezione di Milano-Bicocca, I-20126 Milano, Italy ³Centro Brasileiro de Pesquisas Físicas (CBPF), Rio de Janeiro, RJ, 22290-180, Brazil ⁴CENBG, UMR5797, Université de Bordeaux, CNRS/IN2P3, F-33170, Gradignan, France ⁵CIEMAT, Centro de Investigaciones Energéticas, Medioambientales y Tecnológicas (CIEMAT), E-28040 Madrid, Spain ⁶Aix Marseille Univ, CNRS/IN2P3, CPPM, Marseille, France ^{7a}Department of Physics and Earth Sciences, University of Ferrara, Via Saragat 1, 44122 Ferrara, Italy ^{7b}INFN, Ferrara Section, Via Saragat 1, 44122 Ferrara, Italy ⁸Department of Physics and Astronomy, University of California at Irvine, Irvine, California 92697, USA ⁹LAL, Univ. Paris-Sud, CNRS/IN2P3, Université Paris-Saclay, Orsay, France ¹⁰LNCA Underground Laboratory, CNRS/IN2P3 - CEA, Chooz, France ¹¹Departamento de Física, Universidade Estadual de Londrina, 86051-990, Londrina – PR, Brazil $^{12}\mathrm{Max}\text{-}\mathrm{Planck}\text{-}\mathrm{Institut}$ für Kernphysik, 69117 Heidelberg, Germany ¹³INFN, Sezione di Padova, via Marzolo 8, I-35131 Padova, Italy ¹⁴Department of Physics, Pontifícia Universidade Católica do Rio de Janeiro, Rio de Janeiro, RJ, 22451-900, Brazil ¹⁵Pontificia Universidad Católica de Chile, Santiago, Chile ¹⁶Department of Physics, Engineering Physics & Astronomy, Queen's University, Kingston, Ontario K7L3N6, Canada ¹⁷RCNS, Tohoku University, 6-3 AzaAoba, Aramaki, Aoba-ku, 980-8578, Sendai, Japan ¹⁸SUBATECH, CNRS/IN2P3, Université de Nantes, IMT-Atlantique, 44307 Nantes, France ¹⁹Department of Physics and Astronomy, University of Sussex, Falmer, Brighton BN1 9QH, United Kingdom ²⁰INFN, Sezione di Torino, I-10125 Torino, Italy

August 9, 2019

The discovery of the neutrino by Reines & Cowan in 1956 revolutionised our understanding of the uni-verse at its most fundamental level and provided a with the conventional paradigm of transparency by with the conventional paradigm of transparency by confining and collecting light near its creation point with an opaque scintillator and a dense array of finew probe with which to explore the cosmos. Furthermore, it laid the groundwork for one of the most successful and widely used neutrino detection technologies to date: the liquid scintillator detector. In these detectors, the light produced by particle interactions propagates across transparent scintillator volumes to surrounding photo-sensors. This article introduces a new approach, called LiquidO, that breaks

[†]Also at Observatório Nacional. Rio de Janeiro, Brasil [‡]Deceased.

[§]Blaise Paschal Chaire Fellow.

bres. The principles behind LiquidO's detection technique and the results of the first experimental validation are presented. The LiquidO technique provides high-resolution imaging that enables highly efficient identification of individual particles event-by-event. Additionally, the exploitation of an opaque medium gives LiquidO natural affinity for using dopants at unprecedented levels. With these and other capabilities, LiquidO has the potential to unlock new opportunities in neutrino physics, some of which are discussed here.

Seminar@CERN — June 2019

Web: https://indico.cern.ch/event/823865/

Igniting publication — Aug 2019

LiquidO @ arXiv:1908.02859

 new detection principle first experimental proof-of principle •vast neutrino physics prospect

Submitted for Publication

5

20

First Opaque Liquid Scintillator @ arXiv:1908.03334

^{*}Contact: anatael@in2p3.fr and suekane@awa.tohoku.ac.jp.

a novel neutrino detection

what can LiquidO do for us?



EDF CNPE Chooz-B



Chooz-B 2x N4 Reactors

2x N4 Reactors: 8.4GW(thermal)→ ~10²¹v/s]

CNrs

67

Double Chooz

0

new HUGE site @ Chooz! (very poor overburden)

les montagnes des Ardennes

Europe's best reactor V site...

Anatael Cabrera CNRS-IN2P3 / IJCLab (Orsay) - LNCA (Chooz) Laboratories

a secret underground...





two huge caverns already built of the size of **Super-Kamiokande** just next to **Chooz reactors**! (unique site in France / Europe / World?)

recycling Chooz-A?

⁷⁰ Super-Kamiokande (50kton)

~50m

Super-KamiokaNDE @ Japan (Nobel prize 2015)

~14,000 PMTs (20" diameter)

Super Cool

Super Chooz since the 60's...






EDF CNPE Chooz-B

"Ultra Near"? [≤20m]

Chooz-B 2x N4 Reactors

2x N4 Reactors: 8.4GW(thermal)→ ~10²¹v/s]

CNrs

73

les montagnes des Ardennes

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site "Super Chooz"?

Double Chooz

0

leptonic sector unitarity with LiquidO?



Conference @ HEP-European Physics Society (July 2019 @ Ghent Belgium) Web: https://indico.cern.ch/event/577856/contributions/3421609/

solar neutrinos too...

Super Chooz = telescope of the sun's fusion!



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MENU

neutrino reactor 013 et Δm^2 [WB]

neutrino solar θ I 2 [WB!] — et δm²?

direct CPT violation? [WB-v & BSM] direct Unitarity violation? [WB? & BSM]

all channels [WB?]

proton decay multi-cannel (model independent) [WB? & BSM]

> **Super Chooz** (LiquidO ~10kton)

76

WB = world best ("?": under study still)

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- LiquidO would reveal GeV-neutrino interactions in extremely powerful way:





proton-decay @ Super Chooz...

status on neutrino oscillation knowledge...

Standard Model (3 families)

79

[leptons & quarks] & PMNS_{3×3}($\theta_{12}, \theta_{23}, \theta_{13}$) & &± Δm^2 & + δm^2

	today			≥2030		
	best knowledge		NuFIT4.0	foreseen	dominant	technique
θ12	3.0 %	SK⊕SNO	2.3 %	<1.0%	JUNO & SC	reactor
θ23	5.0 %	NOvA+T2K	2.0 %	≲ .0%	DUNE⊕HK [⊕SC??]	beam (octant)
θιз	1.8 %	DYB+DC+RENO	1.5 %	<1.0%	Super Chooz (SC)	reactor
+δm²	2.5 %	KamLAND	2.3 %	≲ .0%	JUNO	reactor
Δm ²	3.0 %	T2K+NOvA & DYB	1.3 %	≲ .0%	JUNO⊕DUNE⊕HK⊕ SC	reactor⊕beam
Mass Ordering	unknown	SK et al	NMO @ <u>≤3</u> σ	@5o	<u>JUNO</u> ⊕DUNE⊕HK (NOvA⊕T2K)	reactor⊕beam
CPV	unknown	T2K+NOvA	3/2π @ <mark>≤2σ</mark>	@5σ?	DUNE⊕HK⊕ SC	beam driven
СРТУ	assumed			< %?	SC?? [studying]	reactor+solar
Unitarity	assumed			< %?	SC?? [studying]	reactor+solar

(reactor+solar+beam)

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stunning opportunity...

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EDF+CNRS exploring...

v's back to Europe...?



since the v discovery, reactor v's remain <u>one of the most powerful tools</u>...

future knowledge (strongly) shaped by reactor v...

Super Chooz: a powerful opportunity in Europe?



anatael@in2p3.fr

merci... спасибі... ありがとう... danke... 고맙습니다...

obrigado... Спасибо... grazie... 谢谢... hvala...

nvaia... gracias... شکرا thanks...

₄ LiquidO (imaging⊕energy flow) ⇒ never before!

