

# HNL Prompt Progress: Improving Kinematic Reconstruction

RAL ATLAS Physics Discussion, Zoom Land

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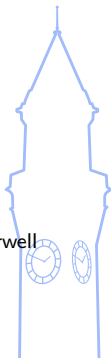
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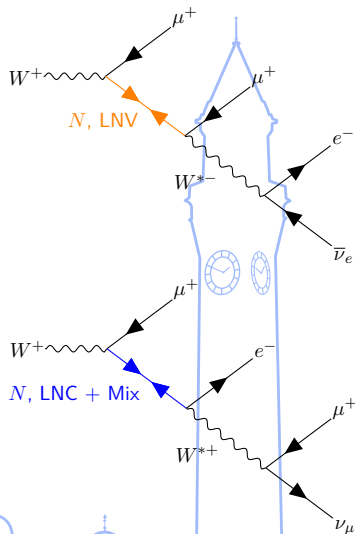
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**ATLAS**  
EXPERIMENT



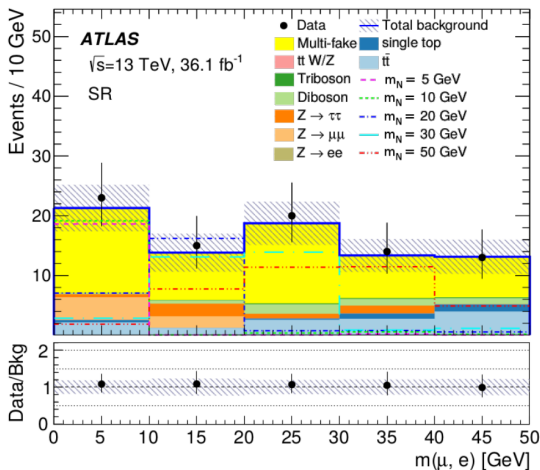
- HNL or type-I seesaw
- Current Run 2 analysis looking for a resonance below the mass of the W via displaced tracks at low mass or high mass LFV (prompt)
- Prompt relies on Same Sign Same Flavour pair
- Possible LNV vs 'LNC + mix' signal extraction under study (spin correlation effects)
- With Giacomo from Pavia I have been studying using MET to improve the reconstruction of the event
- Work in progress, useful plot ideas appreciated



# Previous analysis



- Previous analysis used a reasonably simple set of cuts to optimise signal significance with limited combined variables.
- The signal region was binned in mass of the lowest pt leptons to separate mass hypothesis. Correlated but not particularly sharp.

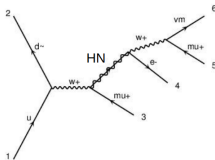


- Not particularly complex: to leading order you only have one missing neutrino and your signal events should come from a  $W$ . Solve for  $p_{z,W}$
- Once have full kinematics, have to pick an assignment for the HNL decay products
- Quadratic ambiguity ( $\cos \theta_{\text{HNL}}^* = \pm a$ )/no root has small effect mostly at small  $m_N$

## Kinematics for heavy neutrino production

$pp \rightarrow W, W \rightarrow \text{HN l1}, \text{HN} \rightarrow \text{l2 l3 n1}$  **Final state:**

- 3 leptons (3,4,5)
- 1 neutrino (6)



### Measured quantities

- 3-momenta for 3 leptons
- $p_x^{\text{miss}}$  = neutrino  $p_x$
- $p_y^{\text{miss}}$  = neutrino  $p_y$

$W$  is on-shell,  $\rightarrow$  mass constraint:

$$(p_{l1} + p_{l2} + p_{l3} + p_{n1})^2 = m_W^2$$

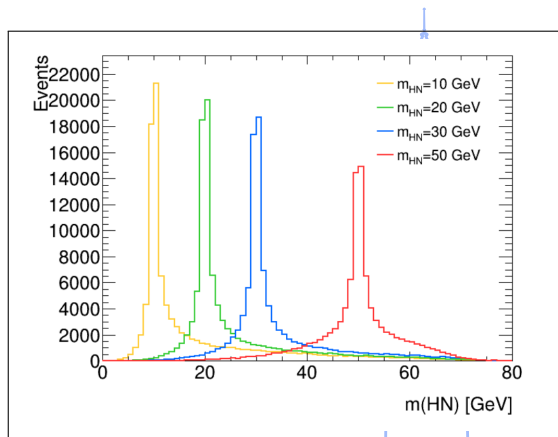
Only neutrino  $p_z$  unknown, use the  $W$  mass constraint to calculate it

## Giacomo's Talk

# Why use MET



- *At particle/parton level,* get very strong peaks
- I am investigating how well this setup works on real data and how the data is smeared
- It is possible that if you put in all the kinematics into a decent sig vs bkg classifier it would unpack this but useful to understand and feed in sensible values

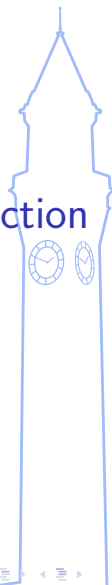


Giacomo's Talk



## Part II

# First attempts on ATLAS reconstruction



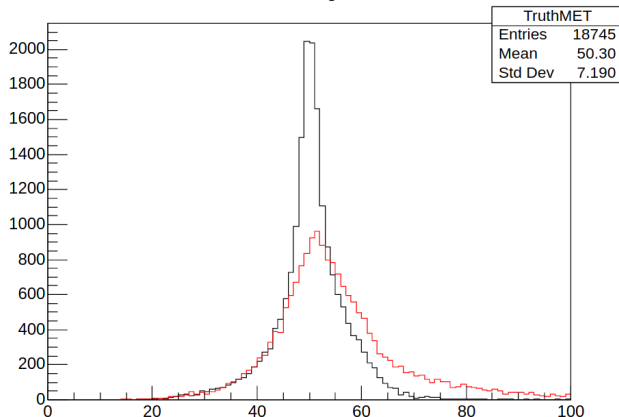


- Some caveats:
  - ▶ The DAODs I use do use a generator without spin correlations which may have a small effect, however this is nearing readiness
  - ▶ No samples greater than 50GeV(yet)
  - ▶ All  $\mu$  mixing on next plots, but  $e$  on later slides with similar picture
- Ran Giacomo's reconstruction on our ATLAS simulation data.
- The major loss of reconstruction power by inspecting the data is truth vs **reco** MET

# 50 GeV $m_{\text{HN}}$ distribution



Reconstructed Mass of HNL using Truth vs Reco Event MET

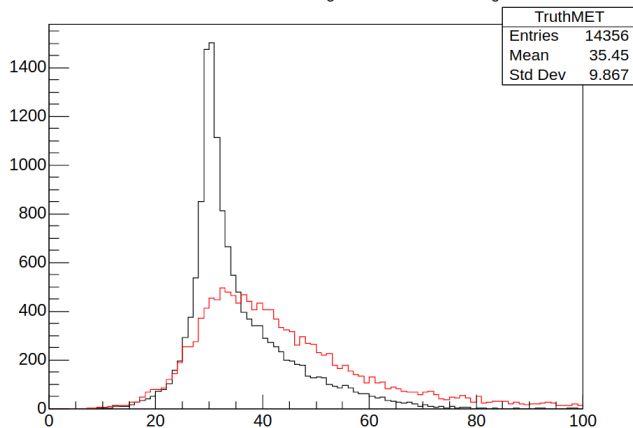


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# 30 GeV $m_{\text{HNL}}$ distribution



Reconstructed Mass of HNL using Truth vs Event Missing MET

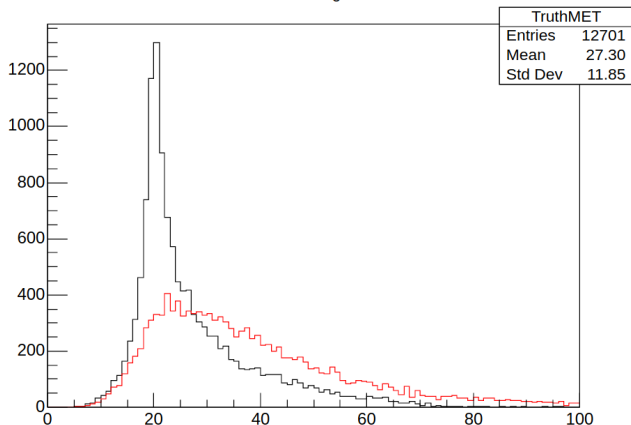


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# 20 GeV $m_{\text{HN}}$ distribution



Reconstructed Mass of HNL using Truth vs Reco Event MET

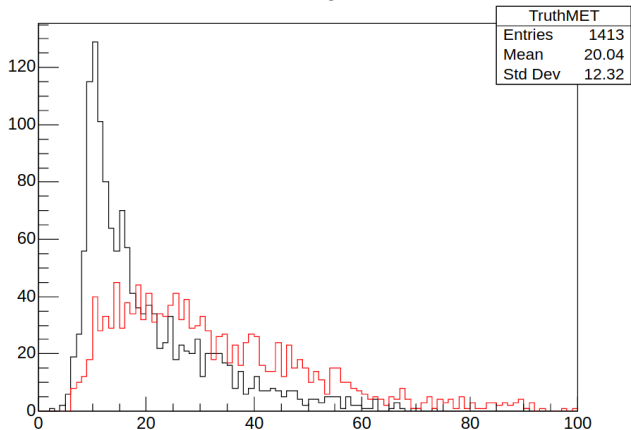


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# 10 GeV $m_{\text{HN}}$ distribution



Reconstructed Mass of HNL using Truth vs Reco Event MET

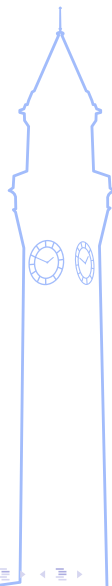


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## Part III

### A bit more on MET



# So what does the smearing and shifting?

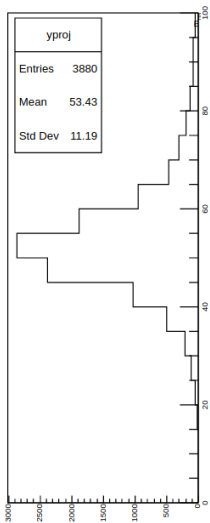
- Two driving factors:

- ① Low MET resolution is clearly going to cause issue for low mass HNLs
  - ★ I have some plots that show MET resolution error of  $\sim 10\text{GeV}$  for truth vs reco, ran out of time putting this together.
- ② Higher hadronic activity ( $H_T$ )
  - ★ The following plots show the effect of  $H_T$  on the mass

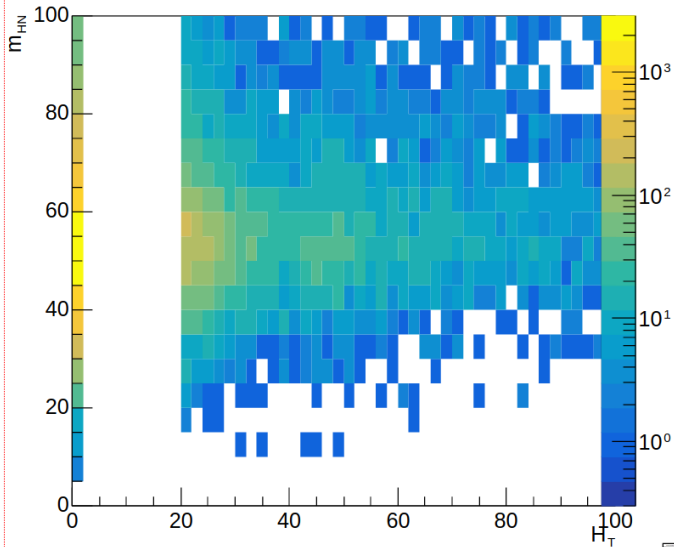


# Breaking down this TH2...

## Y projection



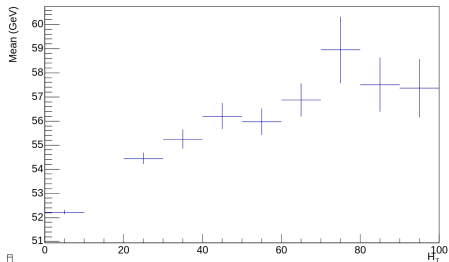
Reconstructed Mass of HNL using reco MET vs  $H_T$



# Fitting Slices

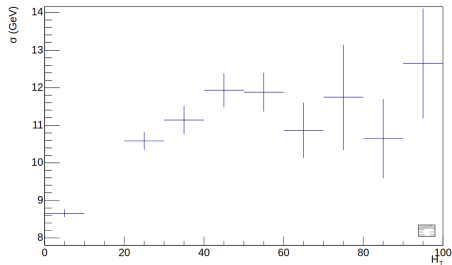


Mean from gaussian fit against  $H_T$ ,  $\mu$  mixing



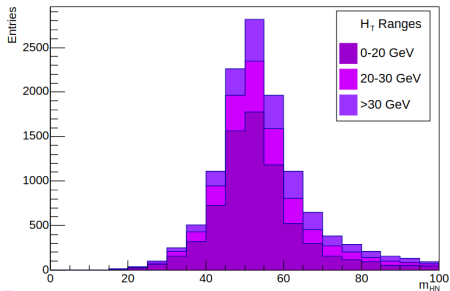
Expected trend of spread shift

Standard error from gaussian fit against  $H_T$ ,  $\mu$  mixing

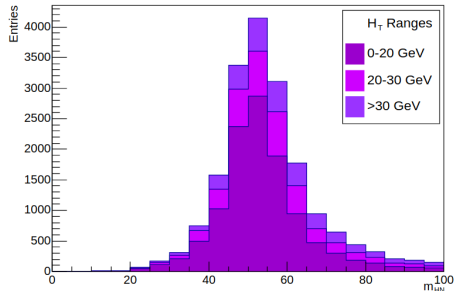


# Some more thoughtful slices: 50 GeV

$m_{\text{HN}}$  distribution stacked in  $H_T$ ,  $m_N = 50\text{GeV}$  e mixing

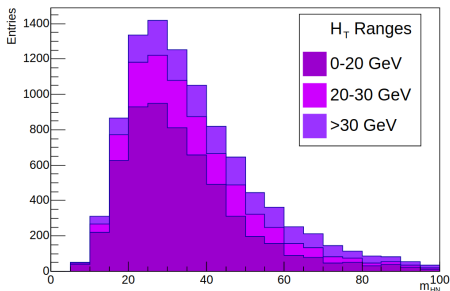


$m_{\text{HN}}$  distribution stacked in  $H_T$ ,  $m_N = 50\text{GeV}$   $\mu$  mixing

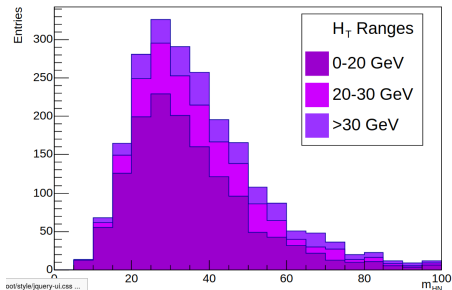


# Some more thoughtful slices: 20 GeV

$m_{HN}$  distribution stacked in  $H_T$ ,  $m_N = 20\text{GeV}$  e mixing



$m_{HN}$  distribution stacked in  $H_T$ ,  $m_N = 20\text{GeV}$   $\mu$  mixing



# Thoughts/Conclusions

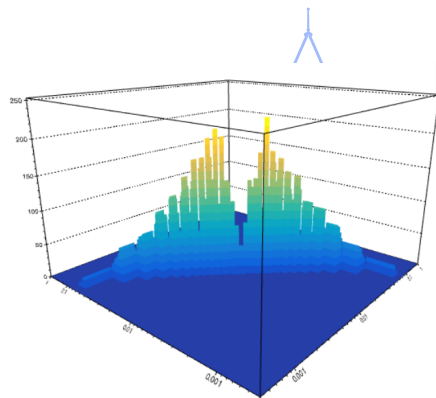
- Understand the detector resolution effects of trying to reconstruct the mass of the HNL better
- Can start to think about if/how to use  $H_T$  for better sensitivity
- Next steps:
  - ▶ Start using background samples to see rejection power
  - ▶ Examine the CMS angular distributions to try and obtain further insight



# A bit more on bjorken...



- Getting  $m_{\text{HN}}$  gets you the boost of the W and or a naive guess of  $x_1, x_2$
- This gives you a hint to direction of origin of a quark vs anti quark via PDFs
- This marginally untangles your angular part of the differential cross sections
- Could this be used a way to observe BDT selection or add in as parameter for LNC LNV? Effect likely marginal but want to have a look.

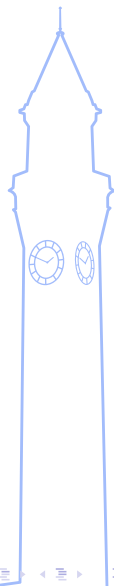


Already computed the distribution



# Part IV

## Backup





## HNLReco

