

The Aria and Urania Projects

Industrial-Scale Production of Radiopure Argon for the DarkSide-20k Dark Matter Experiment

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Low Radioactivity Techniques (LRT)

September 21, 2026



DarkSide-20k Overview

Next-generation liquid argon TPC hosted at LNGS (Gran Sasso, Italy).

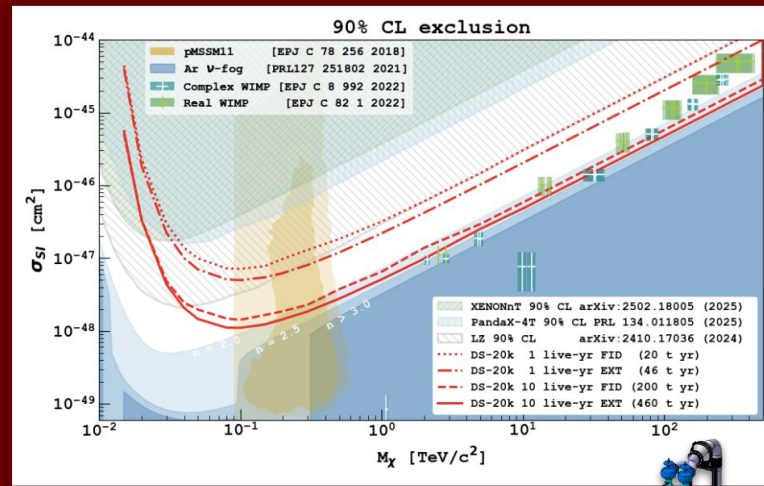
- 50-tonnes active mass (20-tonnes fiducial volume) of underground argon (UAr), depleted in ^{39}Ar .

Key features:

- Background-free design enabling exposures up to 200 t-yr.

Physics reach:

- Sensitivity to WIMP–nucleon cross section down to $\sim 10^{-48} \text{ cm}^2$.
- World-leading discovery potential for multi-TeV WIMPs.



Inner veto (IV):

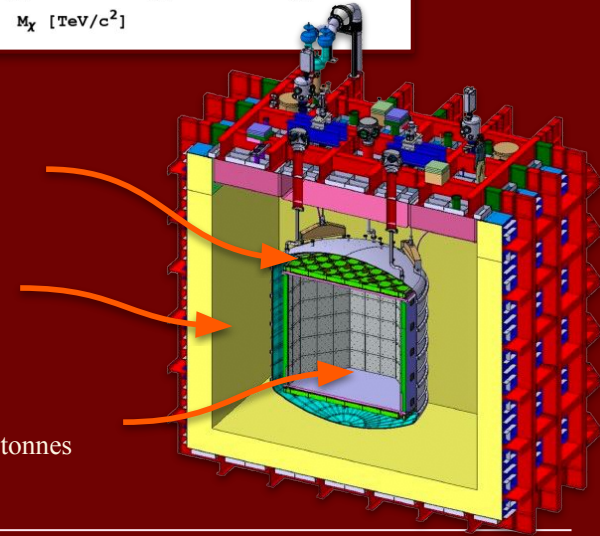
- Active UAr mass: 36 tonnes

Outer Veto (OV):

- 650 tonnes of AAr

TPC:

- Active UAr mass: 50 tonnes



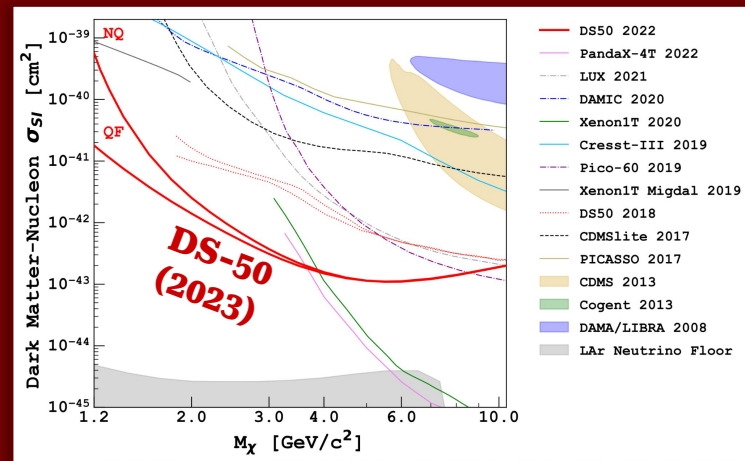
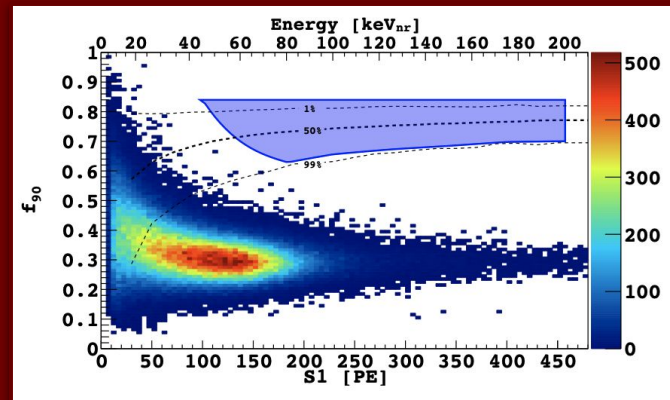
DarkSide-50 Highlights

High-mass WIMPs

- Blind analysis, 524 live-days with UAr $\rightarrow$ first 70 days used to tune cuts, rest kept blind to avoid bias.
- Total expected background in ROI: 0.09 ± 0.04 events $\rightarrow$ ER leakage dominates over neutron backgrounds, demonstrated bg-free operation with UAr fill.

Low-mass WIMPs

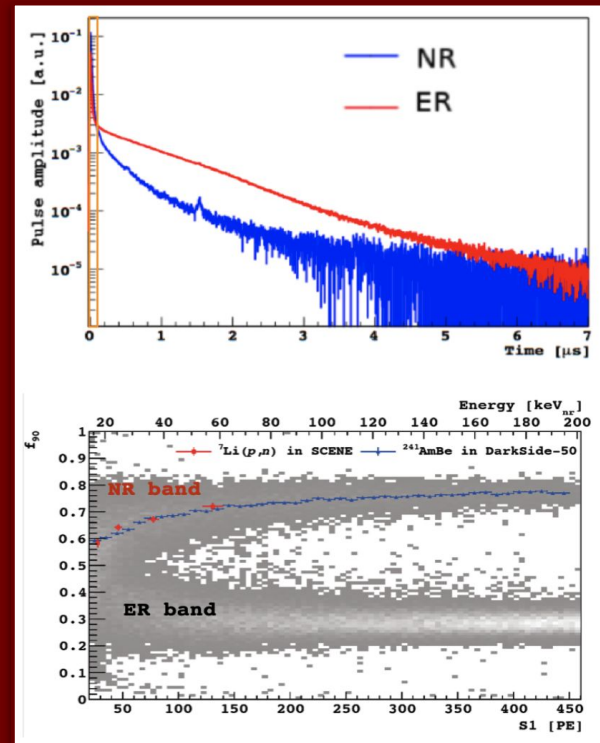
- S2-only search, bg limited (no S1 $\Rightarrow$ no PSD).
- Analysis threshold at 4 electrons (~ 0.4 keV_{NR}).
- World leading constraints at ~ 1 GeV.



Pulse Shape Discrimination in LAr

Fast (singlet, ~ 6 ns) vs. slow (triplet, ~ 1.3 – 1.5 μ s) scintillation components in LAr.

- **Discrimination parameter: prompt fraction (f_{prompt})** — ratio of light collected in an early window (~ 60 – 200 ns) to total S1; NR and ER populate well-separated bands in f_{prompt} vs. energy.
- **DEAP-3600** achieves rejection $> 10^9$ in the energy ROI for WIMPs [arXiv:2302.14484v1].



Underground Argon

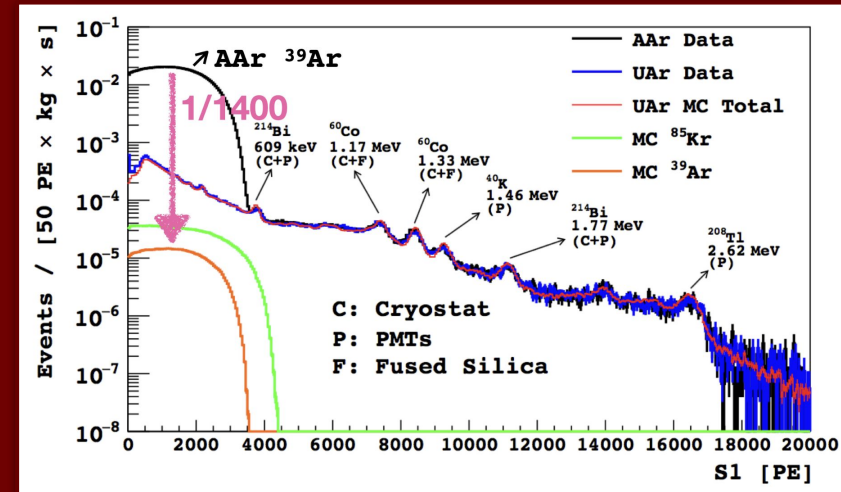
Intrinsic ^{39}Ar ($Q=565$ keV) is the dominant background in argon-based detectors:

~1 Bq/kg in atmospheric argon (AAr)

Sets the low-energy detection threshold, where PSD is least effective, and causes event pile-up in multi-tonne detectors.

^{39}Ar is produced by cosmic rays in the atmosphere → argon from underground wells, shielded from cosmic rays, is strongly depleted.

Reaction	Estimated ^{39}Ar Production rate [atoms/(kg _{Ar} day)]	Fraction of total AAr [%]
$^{40}\text{Ar}(n, 2n)^{39}\text{Ar}$ + $^{40}\text{Ar}(n, d)^{39}\text{Cl}$	759 ± 128	72.3
$^{40}\text{Ar}(\mu, n)^{39}\text{Cl}$	172 ± 26	16.4
$^{40}\text{Ar}(\gamma, n)^{39}\text{Ar}$	89 ± 19	8.5
$^{40}\text{Ar}(\gamma, p)^{39}\text{Cl}$	23.8 ± 8.7	2.3
$^{40}\text{Ar}(p, 2p)^{39}\text{Cl}$	<0.1	< 0.01
$^{40}\text{Ar}(p, pn)^{39}\text{Ar}$	3.6 ± 2.2	0.3
$^{38}\text{Ar}(n, \gamma)^{39}\text{Ar}$	$\ll 0.1$ (UAr) 1.1 ± 0.3 (AAr)	- 0.1
Total	1048 ± 133	100



The Argon Purity Requirements for DarkSide

1) Isotopic purity — remove ^{39}Ar :

- Atmospheric argon contains **1 Bq/kg** of ^{39}Ar β -decays (endpoint 565 keV, $\tau_{1/2} = 268$ yr) — a major source of background and pile-up.
- Reduced **x1400 (potentially more)** by sourcing **UAr (Urania)**, and could potentially be further suppressed via **isotopic distillation (Aria)** by a factor of 10.

2) Chemical purity — remove impurities (N_2 , O_2 , CH_4 , etc.):

- Achieved in-situ with a **getter system**, continuously removing contaminants down to the **ppt–ppb** level.
- Getters only work effectively if the input argon is already very pure — so **chemical distillation (Aria)** is performed upstream.
- Contaminants would **quench scintillation light** and degrade pulse-shape discrimination.

The DarkSide Argon Procurement

Step 1 — Urania (Cortez, CO): extraction and purification of 120-t of UAr from CO₂ well gas, achieving ~99.9% purity (nitrogen $\lesssim 10^{-4}$, other contaminants $\lesssim 10^{-6}$).

Step 2 — Aria (Sardinia): further chemical purification + possible **isotopic depletion** of ³⁹Ar via cryogenic distillation, exploiting small differences in relative volatility between argon isotopes.

Step 3 — DArT (LSC, Spain): detector that measures the ³⁹Ar **depletion factor** of UAr batches.



Urania

- Urania plant productivity: ~250-300 kg/day of UAr from CO₂.
- Kinder Morgan's Doe Canyon facility, pulling CO₂ from the Earth's mantle as part of its natural gas mining.
- Contained within this CO₂ stream coming from those deep underground wells is a small amount of low-radioactivity argon.

Our low-radioactivity argon is a byproduct in their production of natural gas.



Aria: Cryogenic Distillation Plant

Cryogenic distillation column (total height ~350 m):

28 Central modules: 12 m each

Outer Diameter of the column: 32.3 cm

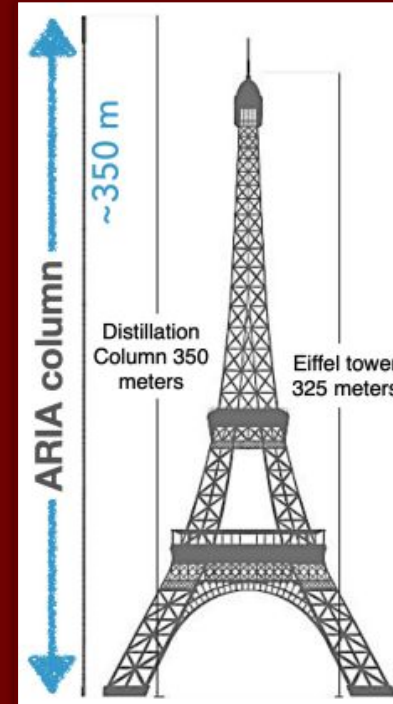
Inner Diameter of the column: 31.8 cm

Outer Diameter of cold box: 71.2 cm

Number of theoretical stages: ~2870

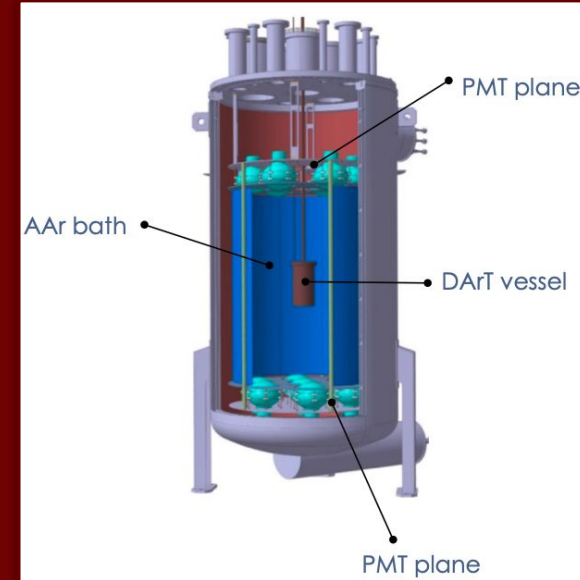
Total column active height: ~ 287 m

Structured packing



DArT—Measuring the ^{39}Ar Depletion Factor

- **DArT:** Single phase 1.35 kg of UAr detector that measures the ^{39}Ar **depletion factor** of UAr batches from Urania and Aria.
- **Located inside ArDM** at the **Canfranc Underground Laboratory (LSC, Spain)** — ArDM's argon volume serves as an active veto.
- **Method:** ^{39}Ar β -decays in DArT are the signal; events also seen in ArDM are tagged and removed as background.
- DArT will be sensitive to very high depletion factors of ^{39}Ar , which is of the order of 1000, reaching 10% statistical uncertainty in 1 week of data-taking.



Transport of argon from Urania to Aria

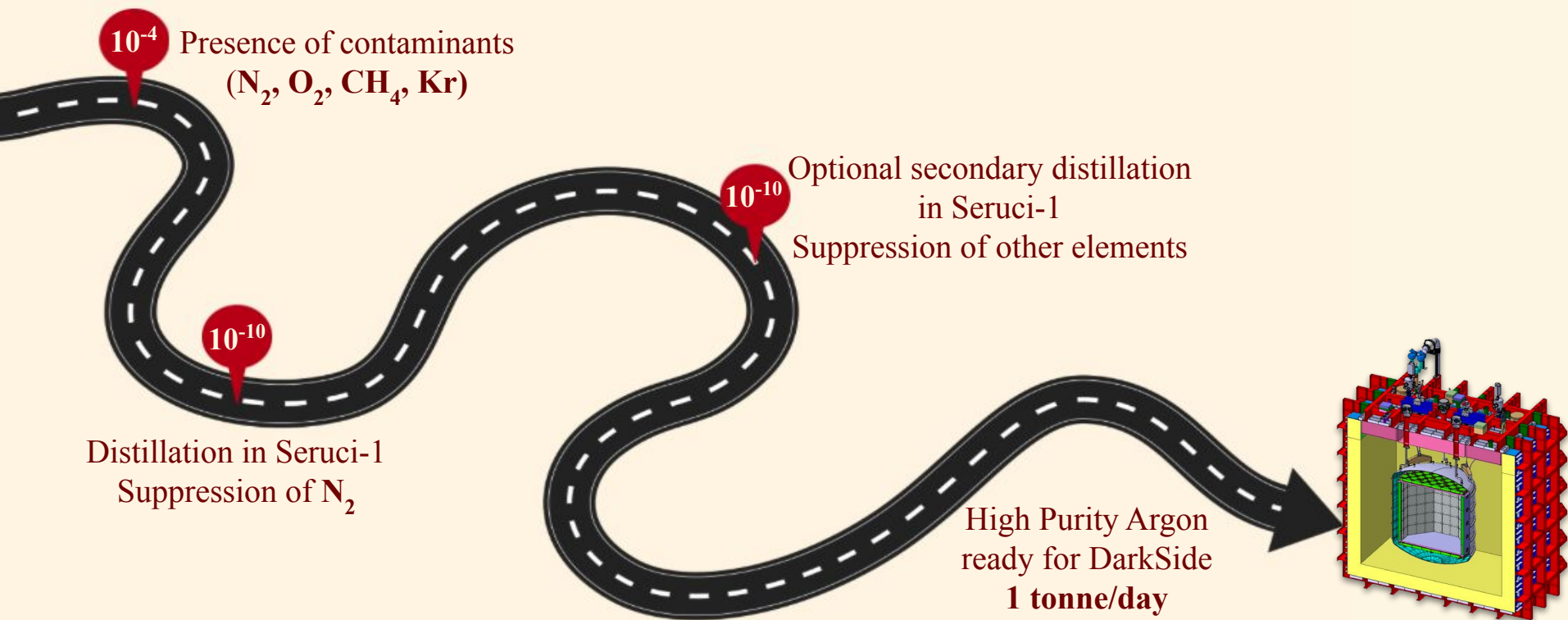
~15-20 tonne shipping and storage vessels will be used to transport liquid argon from Urania to Aria.

- repeating the whole shipment procedure for arrays of 8 skids.

The activity of ^{39}Ar induced during extraction, purification and transport on surface, in baseline conditions, is evaluated to be <5% of the UAr activity measured in DarkSide-50, and thus considered acceptable.



Chemical Purification of Argon

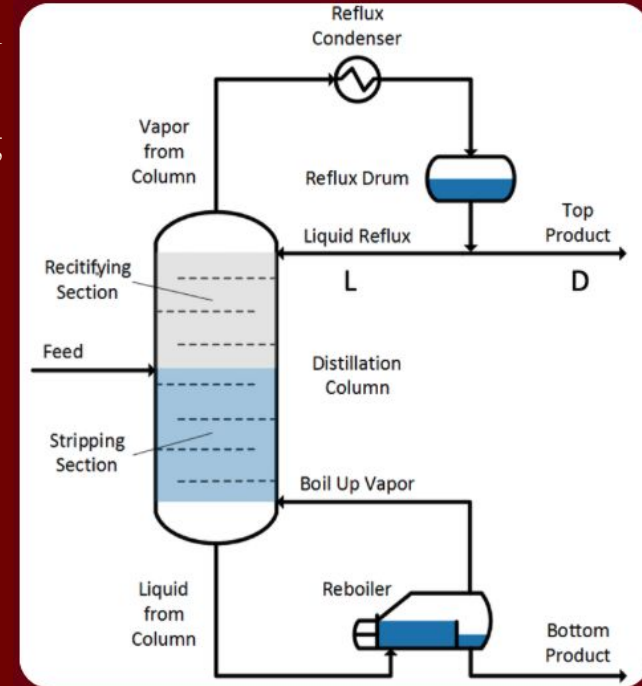


Isotopic Purification of Argon *(from Simulations)*



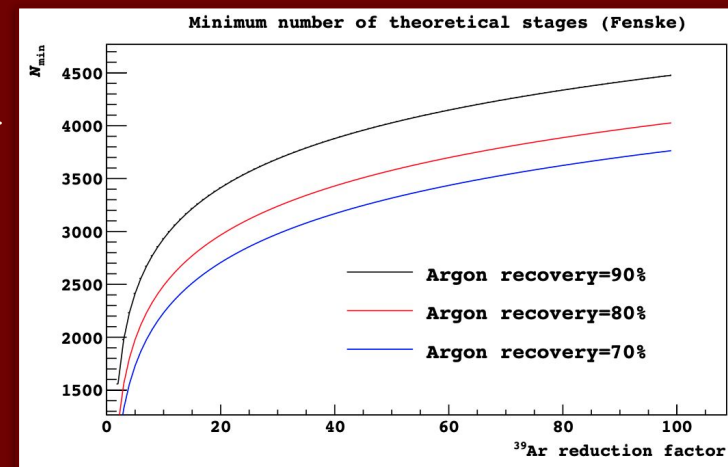
Cryogenic Distillation

- Isotopes separated by tiny boiling point differences in countercurrent exchange.
- Mass and heat transfer between descending liquid and ascending vapour drives enrichment.
- Structured packing maximises equilibrium stages.



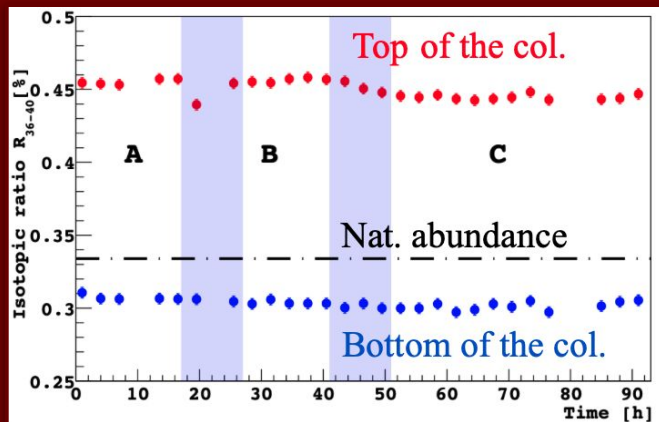
Relative Volatility $^{39}\text{Ar}/^{40}\text{Ar}$

- **Fieschi–Terzi model**: predicts relative vapor pressure between argon isotopes $\rightarrow$ separation factor α .
- **Validated on $^{36}\text{Ar}/^{40}\text{Ar}$** (Boato–Scoles data), then extrapolated to $^{39}\text{Ar}/^{40}\text{Ar}$.
- $\alpha \approx 1.0015$, consistent across 84.4–91.8 K.
- **Fenske equation** $\rightarrow$ thousands of theoretical stages required.



Proof of Concept–Seruci-0 Prototype

- 26 m prototype: one module of the full 350 m column.
- First demonstration of argon isotopic separation — all three stable isotopes (^{36}Ar , ^{38}Ar , ^{40}Ar) separated simultaneously.
- Column performance matches design and simulation, validating the full-scale plant.



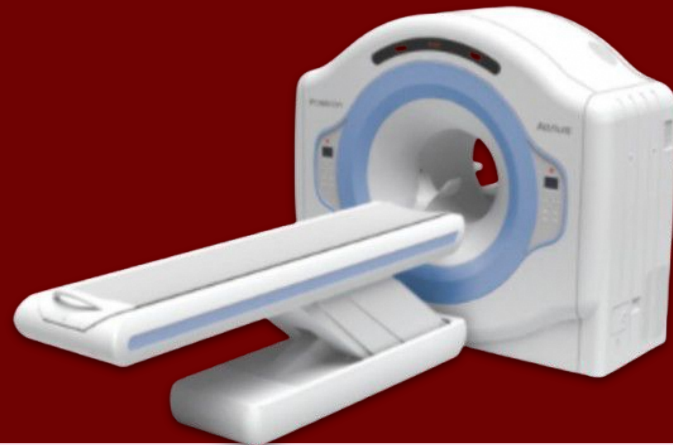
Beyond DarkSide-20k with Aria

Isotopic separation for medical applications

Production of high-purity stable oxygen isotopes as precursors for Positron Emission Tomography (PET).

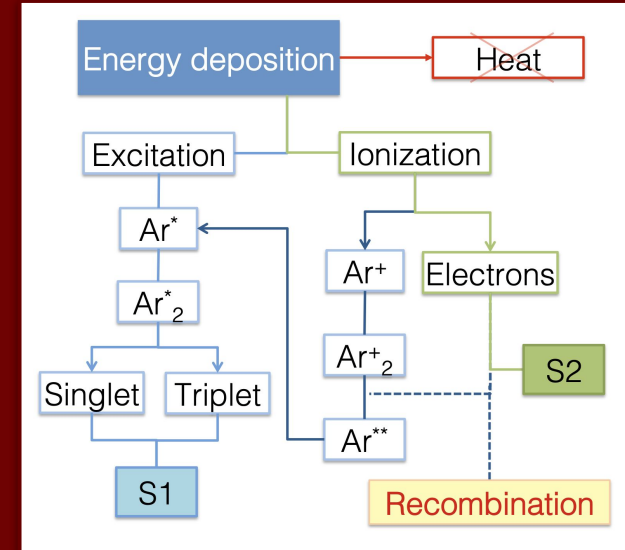
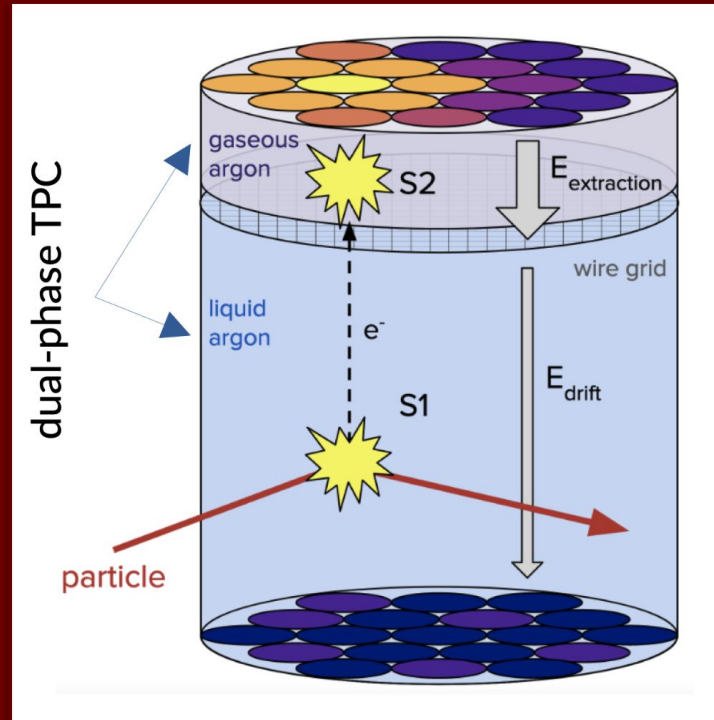
Fractionation of the $^{16}\text{O}_2$ / $^{16}\text{O}^{18}\text{O}$ binary mixture, overcoming the low natural abundance of ^{18}O (0.205%) and a relative volatility of only $\alpha \approx 1.006$.

Direct transfer of capability from fundamental physics to medicine.



BackUp

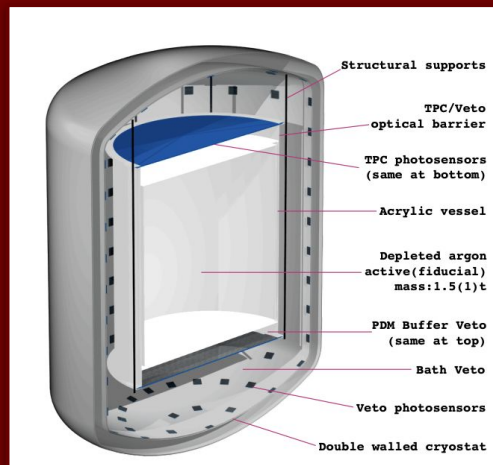
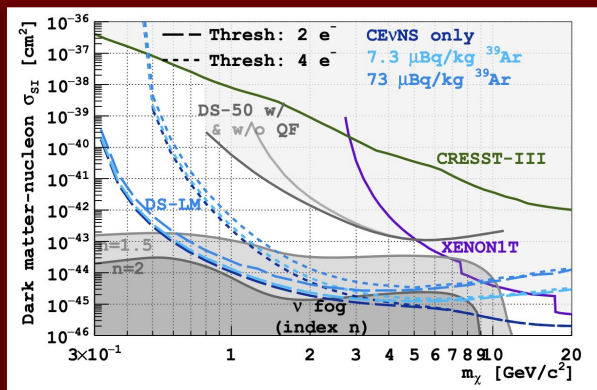
Dual Phase Time Projection Chamber



DarkSide-LowMass → Possible Future Project

Cryogenic distillation enables isotopic separation at unprecedented scale, paving the way for new applications beyond DarkSide-20k.

- Further depletion of ^{39}Ar by up to an additional factor of 10 per pass (≈ 7 kg/day throughput) could support **DarkSide-LowMass**, a 1.5-ton detector optimized for GeV-scale dark matter searches.



The Global Argon Dark Matter Collaboration

GADMC unites predecessors → DarkSide, DEAP, MiniCLEAN, ArDM into one staged argon program.

Argon: dense, scalable, high ionization, strong scintillation, excellent ER/NR discrimination.

DarkSide-20k, background-free via ultra-low backgrounds + active rejection (neutron veto + PSD).

Future: ARGO (SNOLAB, ~300 t fiducial).

