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Quantum Technologies for Applied and Fundamental Physics

Tristan Valenzuela

Quantum Sensors and Technology Programme Lead

Outline

- ❖ Intro to Quantum Technologies 2.0
- ❖ Applications of QT
- ❖ Intro to Cold Atoms and Atom Interferometry
- ❖ Applications I : Gravity mapping and Atmospheric Science
- ❖ Applications II: Gravitational Waves and Ultralight Dark Matter

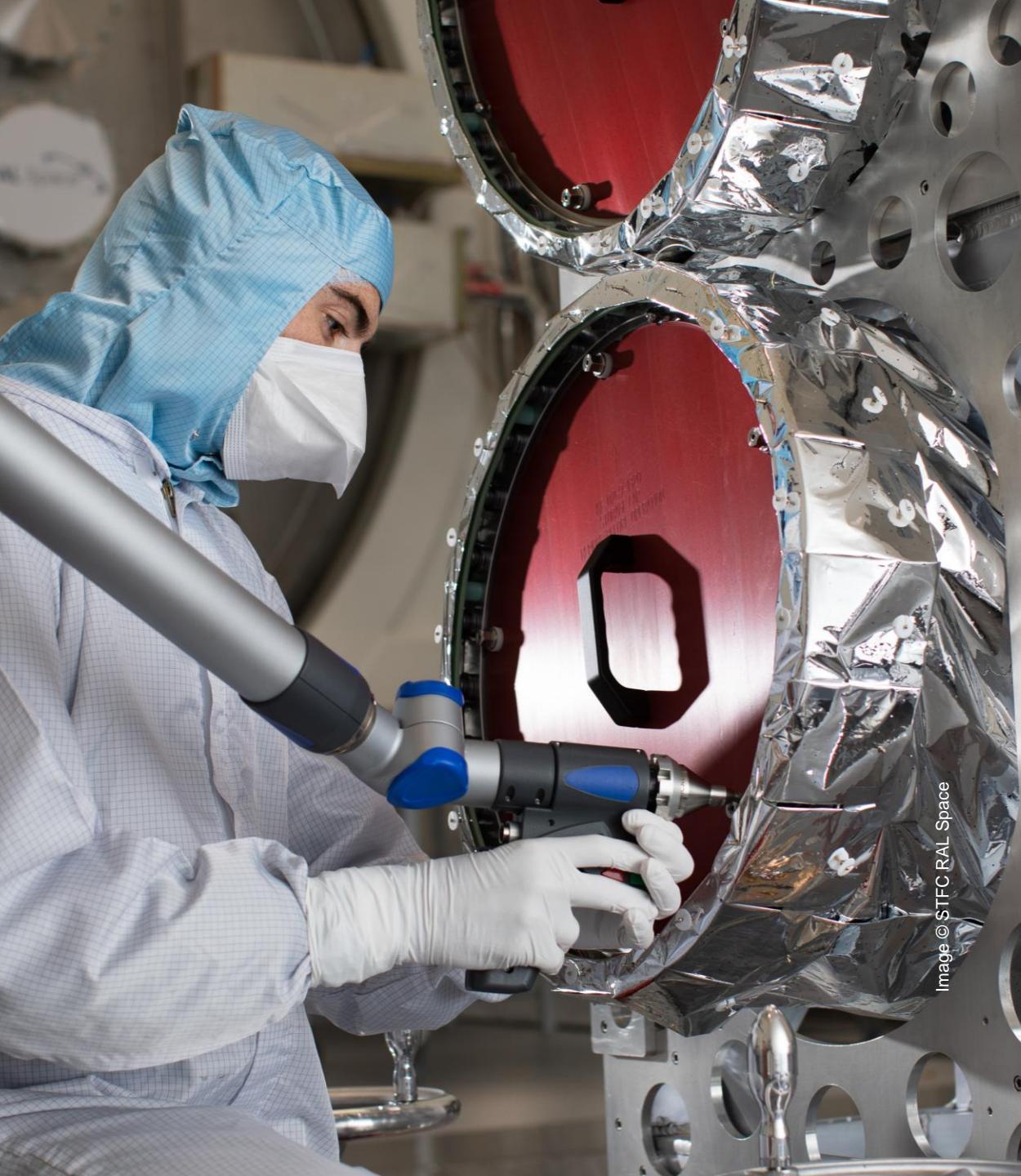


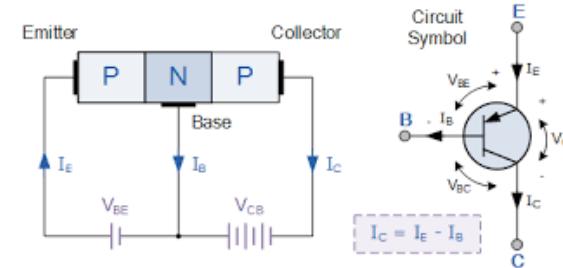
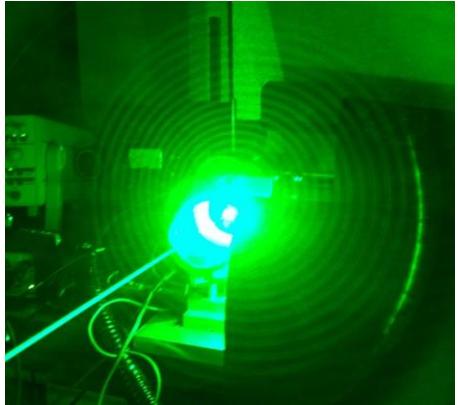
Image © STFC RAL Space



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Quantum Technologies



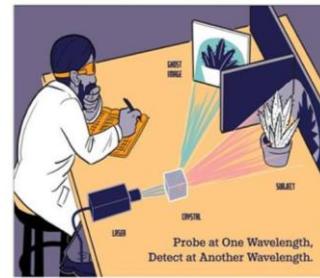
Quantum Technologies 2.0

Exploit quantum superposition and/or entanglement

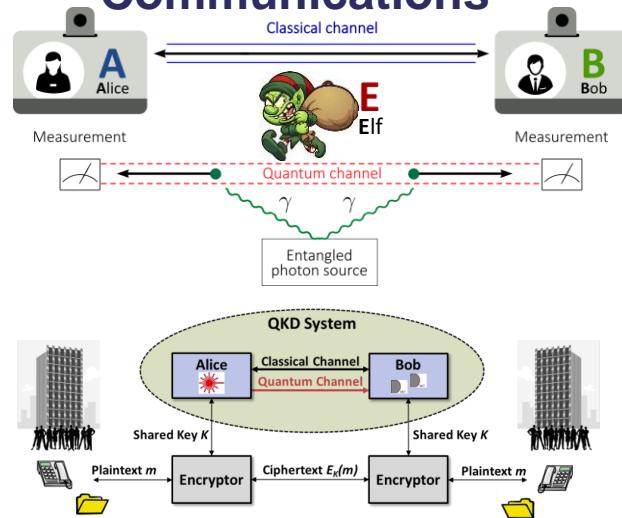
Sensing



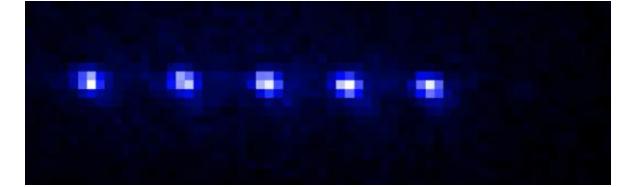
Imaging



Communications



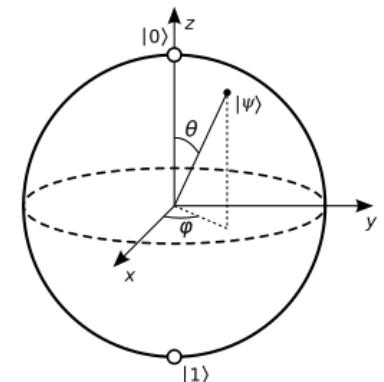
Computing



$|1\rangle$ and $|0\rangle$



$$\Psi = \alpha|0\rangle + \beta|1\rangle$$



Applications of Quantum Technologies

From now on, in this talk

$$\text{QT} = \text{QT-QC}$$

Applications of Quantum Technologies

Earth Observation/Planetary Science/Space Weather

Atom Interferometers (Accelerometers) & Magnetometers & Radiometers

- ❖ Geodesy
 - ❖ Earth & planetary interiors
- ❖ Climatic Effects
 - ❖ Ocean circulation & Water cycle
 - ❖ Glacial movements
- ❖ Earthquake monitoring
- ❖ Atmospheric Science

Fundamental Physics

Atomic clocks & Atom Interferometers:

- ❖ Tests of the Equivalence Principle
- ❖ Tests of General & Special Relativity
- ❖ Gravitational Red Shift
- ❖ Gravity Wave detection in the mid-frequency band
- ❖ Dark Mater detection

Telecommunications

Atomic clocks:

- ❖ Global time reference,
- ❖ High accuracy satellite navigation,
- ❖ Internet synchronisation.

Optical links and Quantum Memories

- ❖ Deep space ranging & communications
- ❖ Increased spectral efficiency and data rates
- ❖ Secure communications via QKD
- ❖ Entanglement Distribution



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Applications of Quantum Technologies

Earth Observation/Planetary Science/Space Weather

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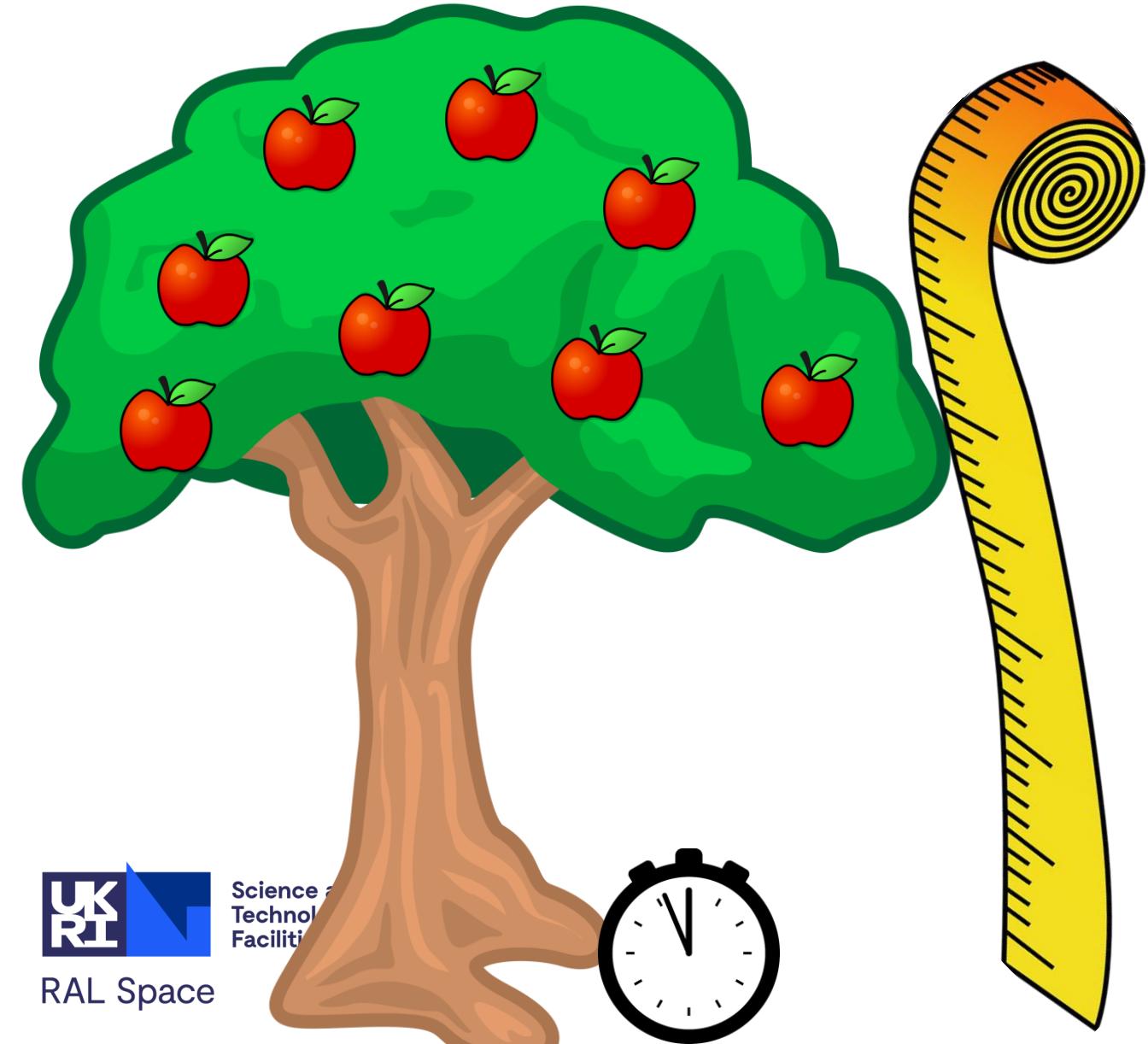


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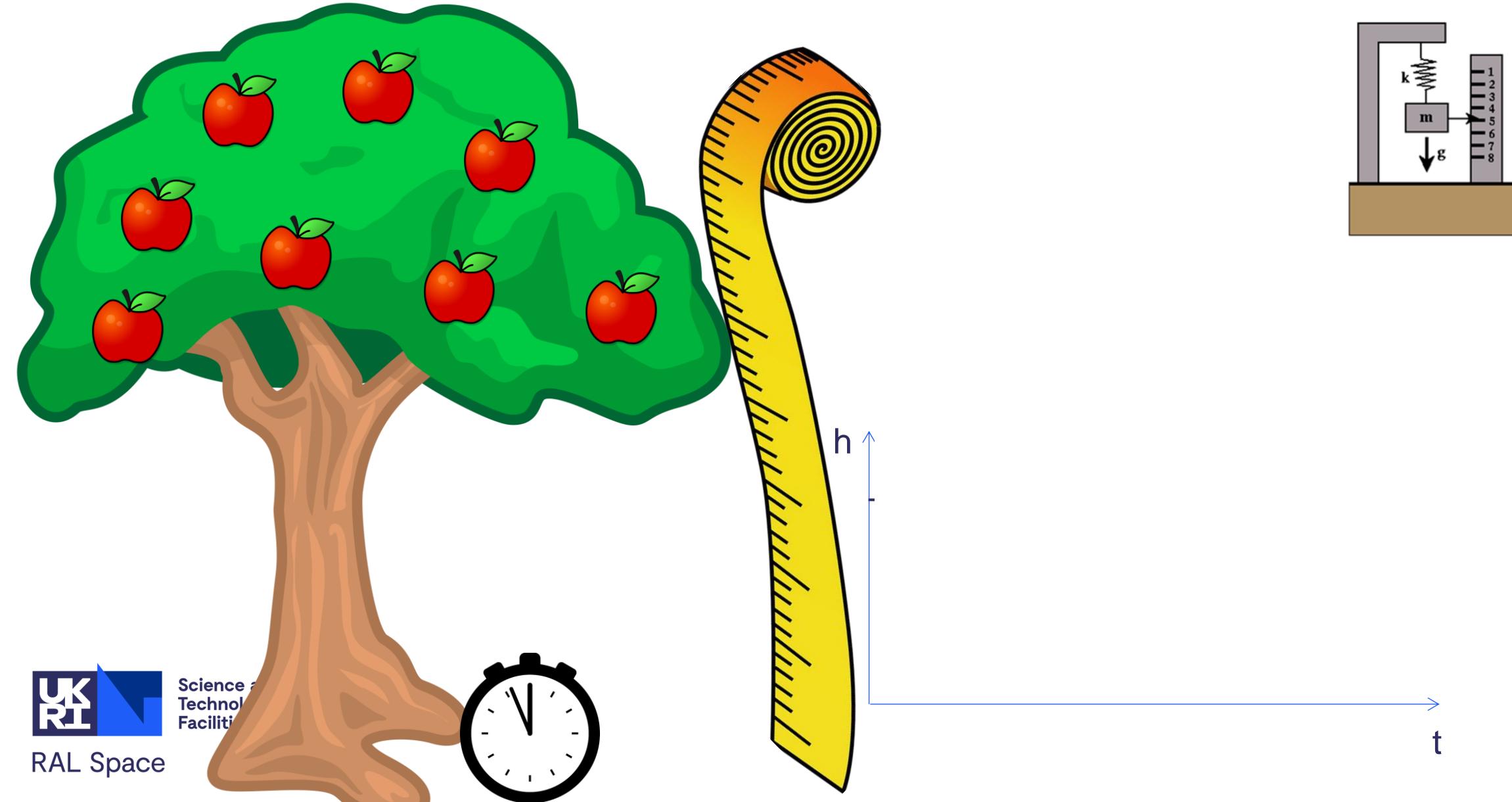
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Space Based
Ground Based

Classical accelerometry/gravimetry



Classical accelerometry/gravimetry

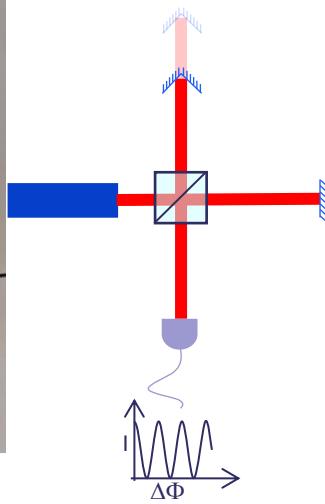


Classical accelerometry/gravimetry

Falling Corner Cube



Microg-LaCoste FG5



Spring elongation



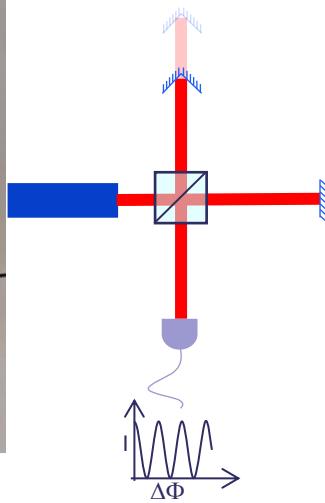
Scintrex CG5

Classical accelerometry/gravimetry

Falling Corner Cube



Microg-LaCoste FG5



Spring elongation



Scintrex CG5

Limitations:

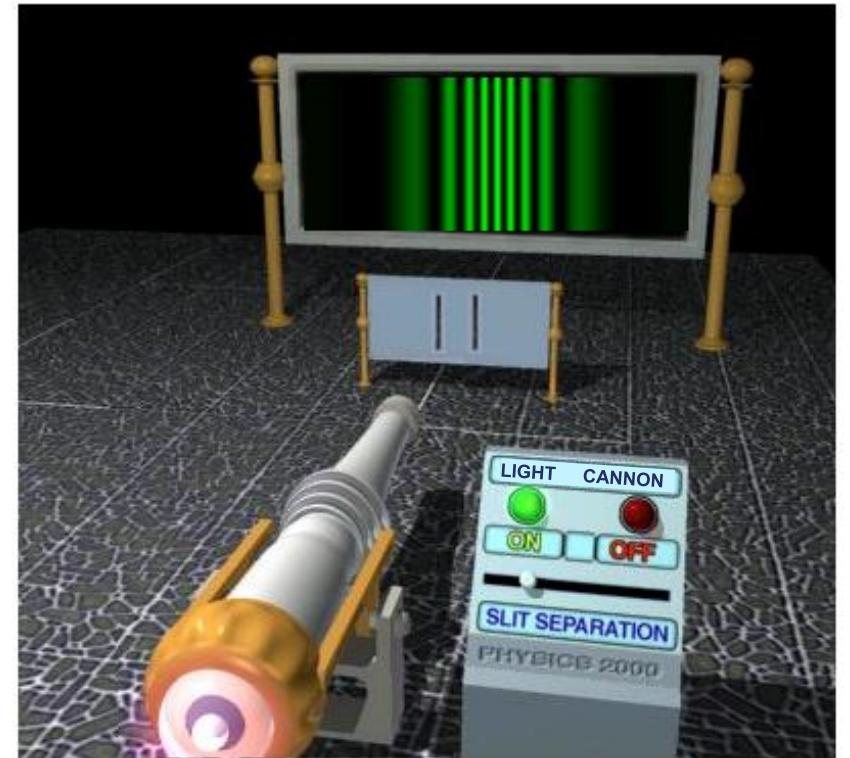
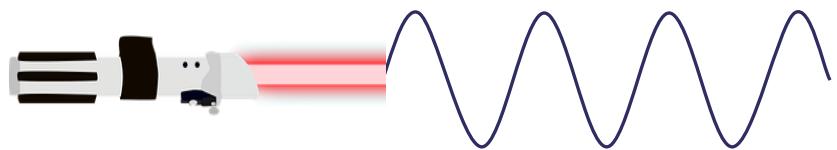
Free fall	Spring
Macroscopic mass	Ageing → Drift
Low rep rate	Readout noise
Residual vibrations	
Ground rebound	

EM Wave Interferometry

- ❖ Young's double slit

Phil. Trans. Roy. Soc. 92, 387 (1802)

DOI:10.1098/rstl.1802.0016



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Matter Wave Interferometry

- ❖ de Broglie, wavelength of matter $\Rightarrow \lambda = \frac{h}{p} = \frac{h}{mv}$

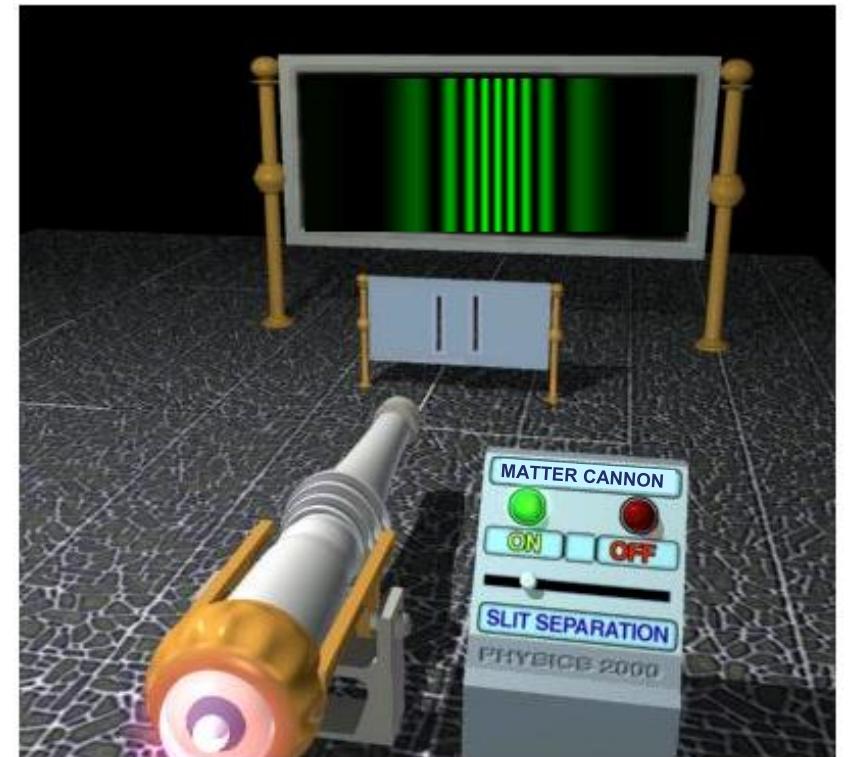
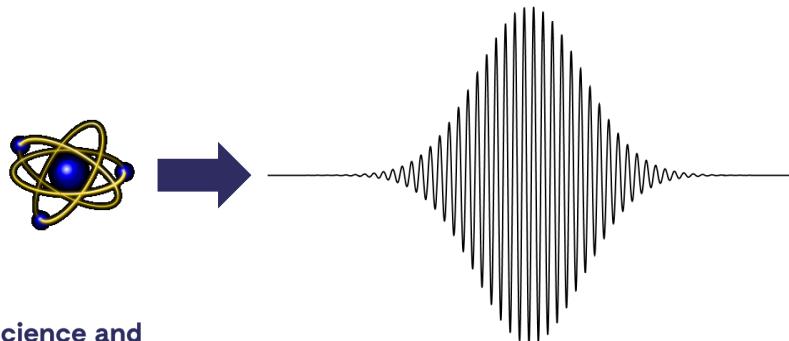
Ann. Phys., Vol. 10, N°3 (1925), pp. 22–128

DOI: 10.1051/anphys/192510030022

- ❖ Electron diffraction by GP Thomson and Reid, and Davisson and Germer

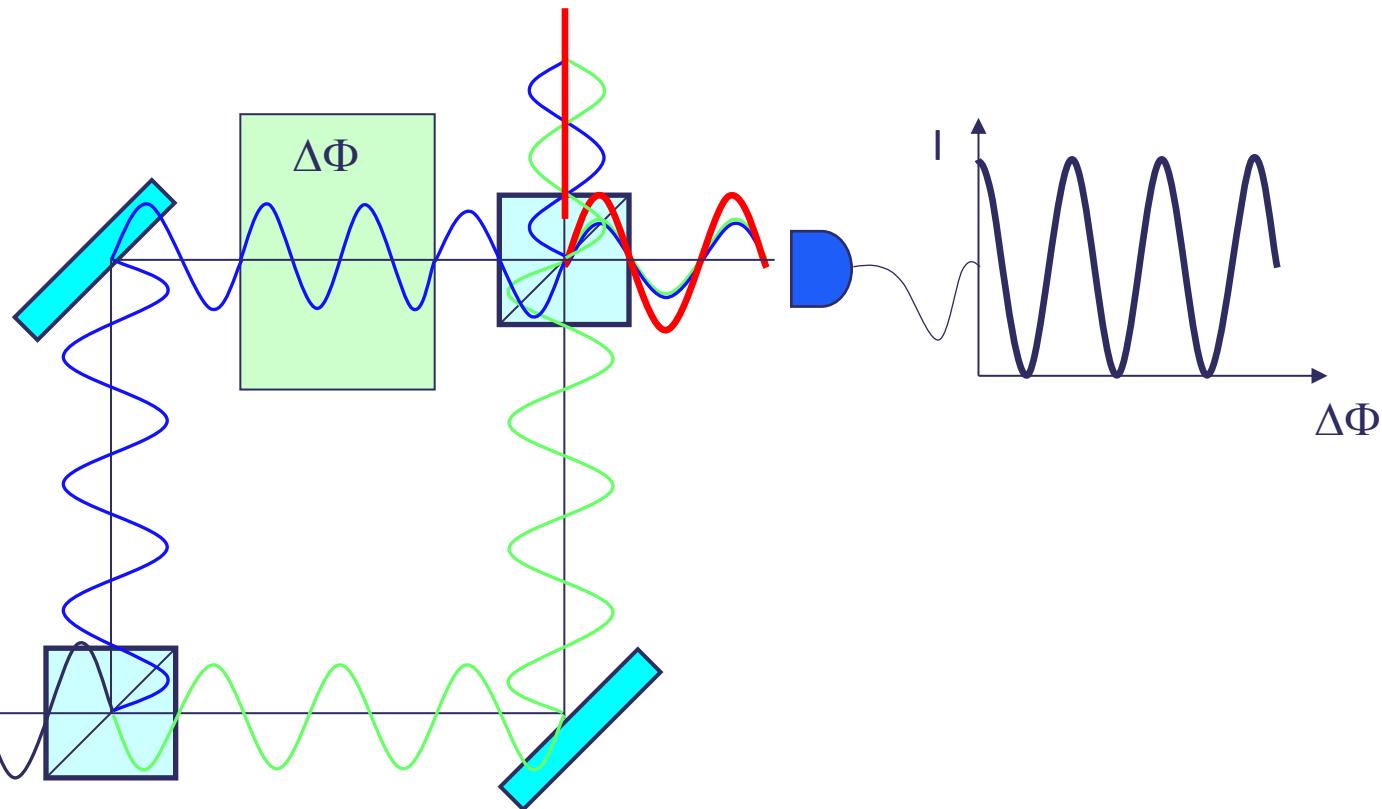
T&R: Nature **119**, 890 (1927) DOI: 10.1038/119890a0

D&G: Phys. Rev. **30**, 705 (1927) DOI: 10.1103/PhysRev.30.705



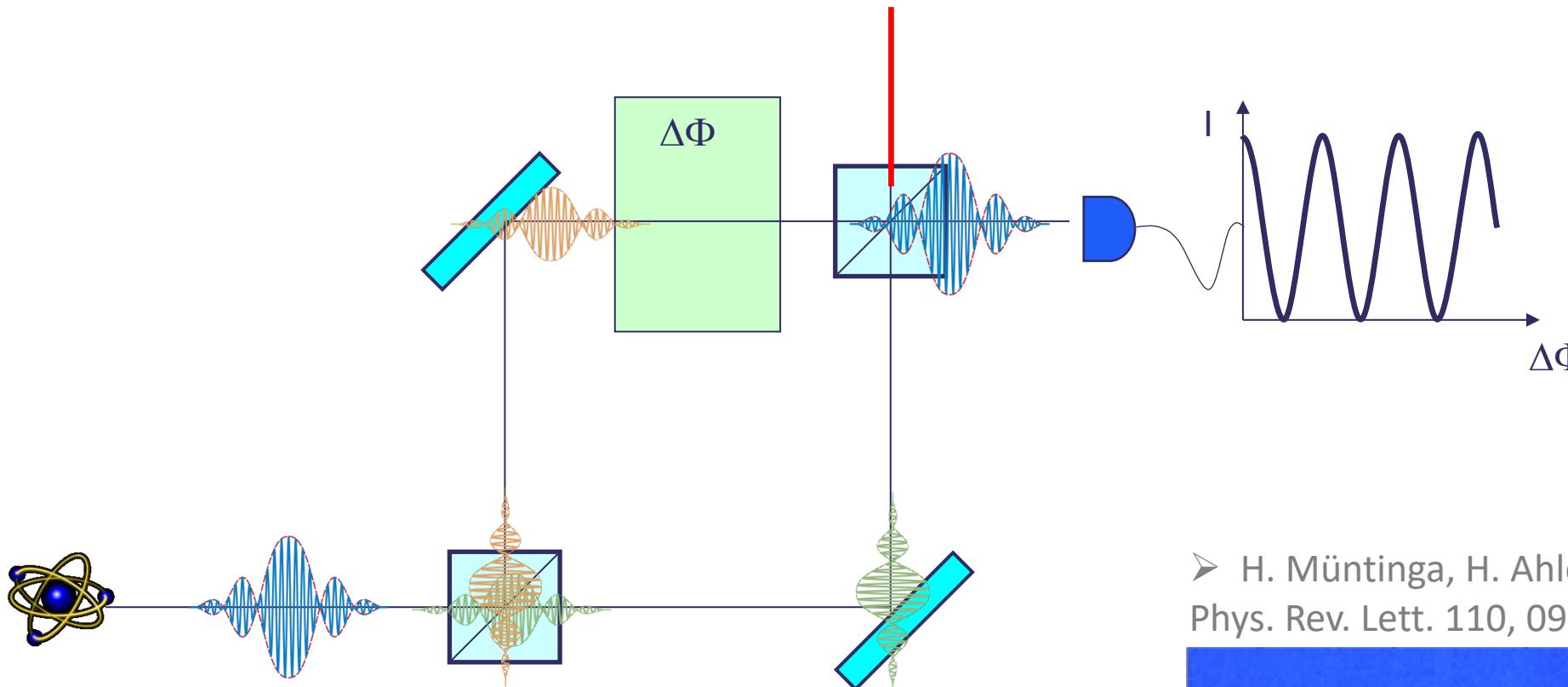
EM Wave Interferometry

Mach-Zehnder Interferometer

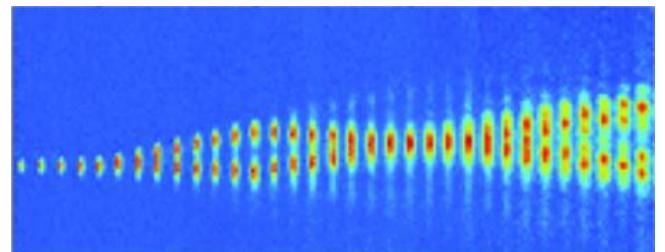


Matter Wave Interferometry

Mach-Zehnder Interferometer



➤ H. Müntinga, H. Ahlers, et al.
Phys. Rev. Lett. 110, 093602

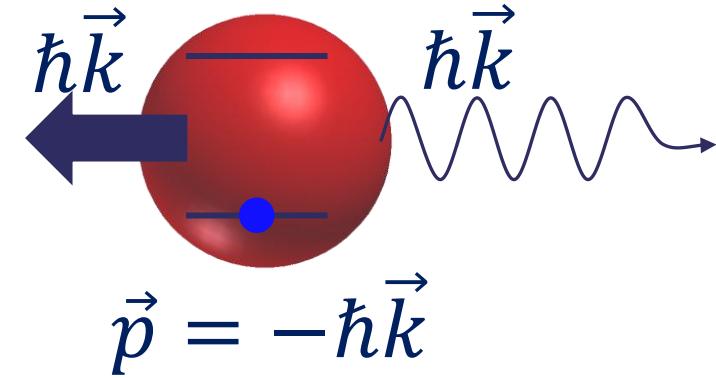
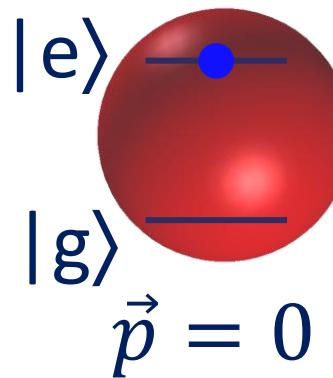
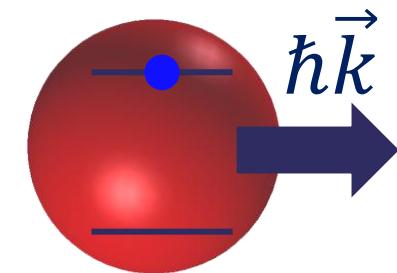
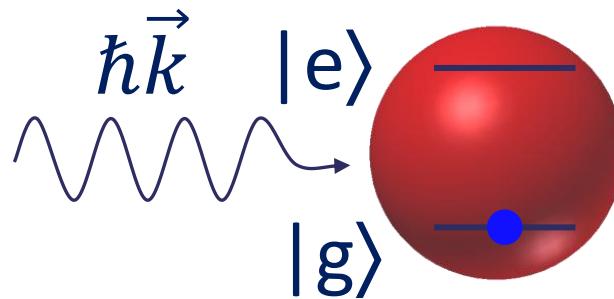


Intro to Cold Atom Interferometry

Some concepts: Energy-Momentum Conservation

$$\gamma \rightarrow \begin{cases} E = \hbar\omega \\ \vec{p} = \hbar\vec{k} \end{cases}$$

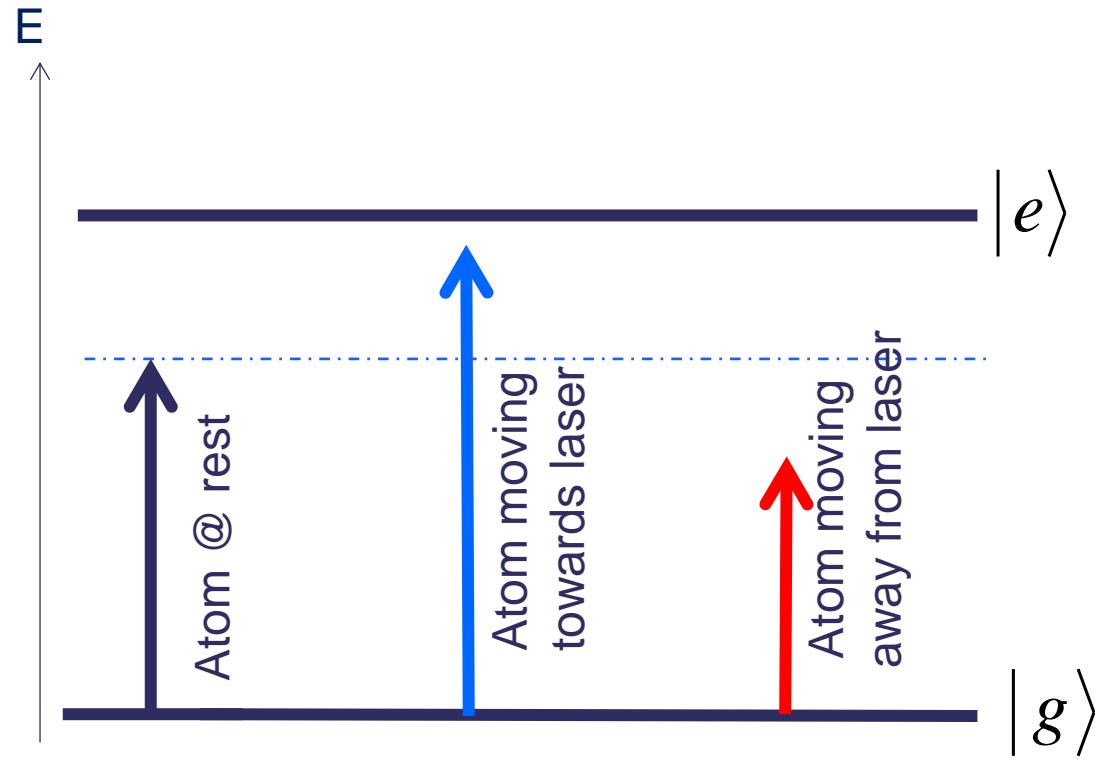
$$|\vec{p}| = \frac{E}{c}$$
$$|\vec{k}| = \frac{\omega}{c} = \frac{2\pi}{\lambda}$$



Intro to Cold Atom Interferometry

Some concepts: Laser Cooling and Trapping

Doppler Effect



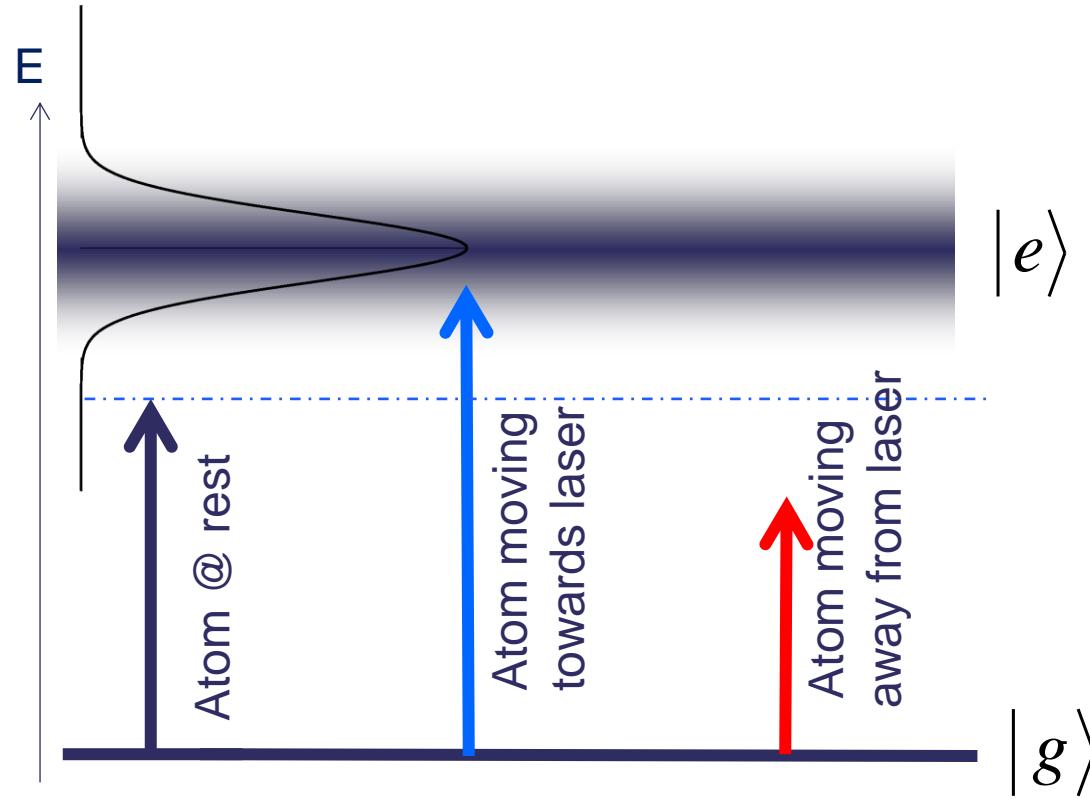
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Intro to Cold Atom Interferometry

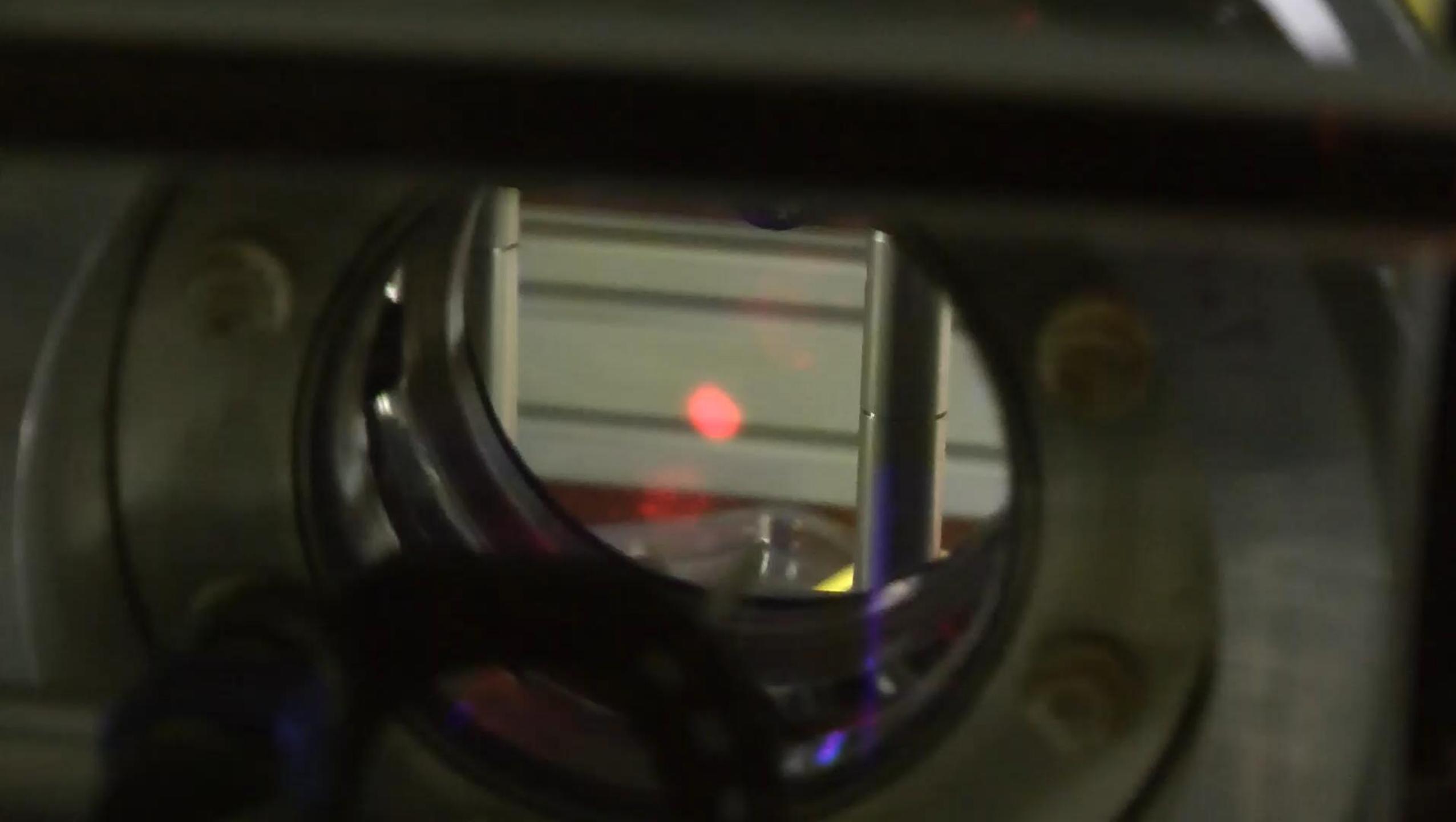
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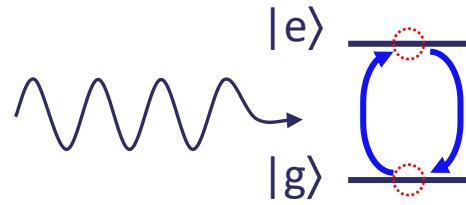
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Intro to Cold Atom Interferometry

Some concepts: Rabi Oscillations

Continuous excitation of a two level atom

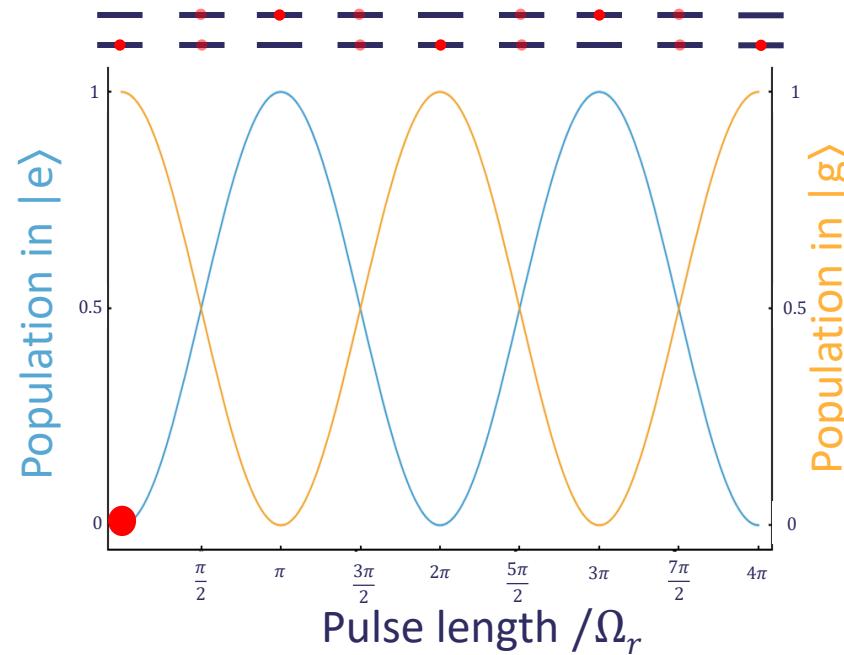


$$\text{Rabi frequency} \rightarrow \Omega_r = \sqrt{\Omega_{eg}^2 + \delta^2}$$

$$\Omega_{eg} = \frac{\langle e | \vec{d} \cdot \vec{E}_0 | g \rangle}{\hbar}$$

$$\begin{aligned}\delta &= \omega_L - \omega_{eg} \\ \omega_{eg} &= E_{eg}/\hbar\end{aligned}$$

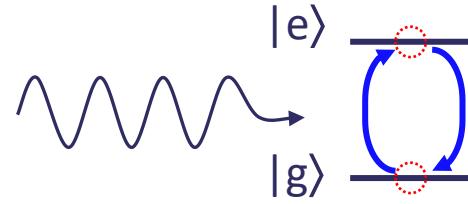
$$\text{Population in } |e\rangle \propto \sin^2\left(\frac{\Omega_r \cdot t}{2}\right)$$



Intro to Cold Atom Interferometry

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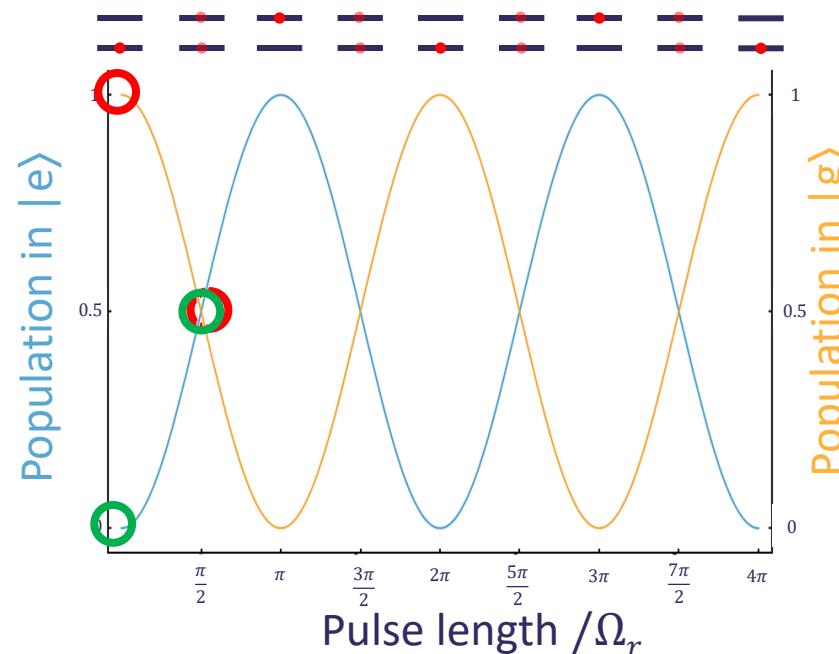


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$$P_{ex} = \frac{1}{2}$$

$$|\psi_0\rangle = |g\rangle$$

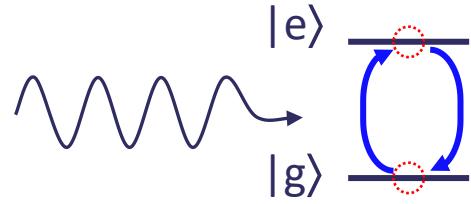


$$|\psi\rangle = \frac{1}{\sqrt{2}}(|g\rangle + |e\rangle)$$

Intro to Cold Atom Interferometry

Some concepts: Rabi Oscillations

Continuous excitation of a two level atom

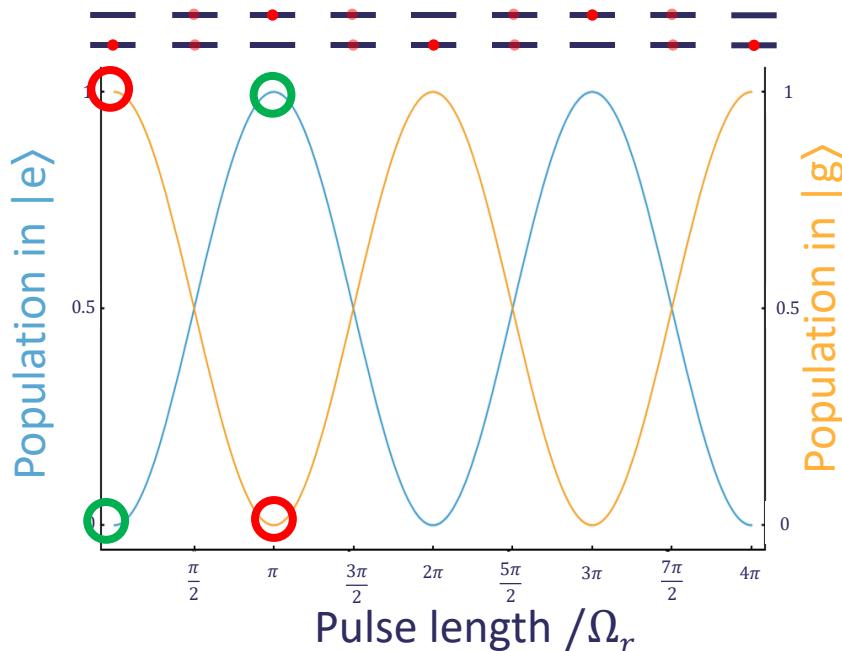


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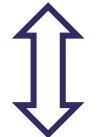
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$$P_{ex} = 1$$

$$|\psi_0\rangle = |g\rangle$$



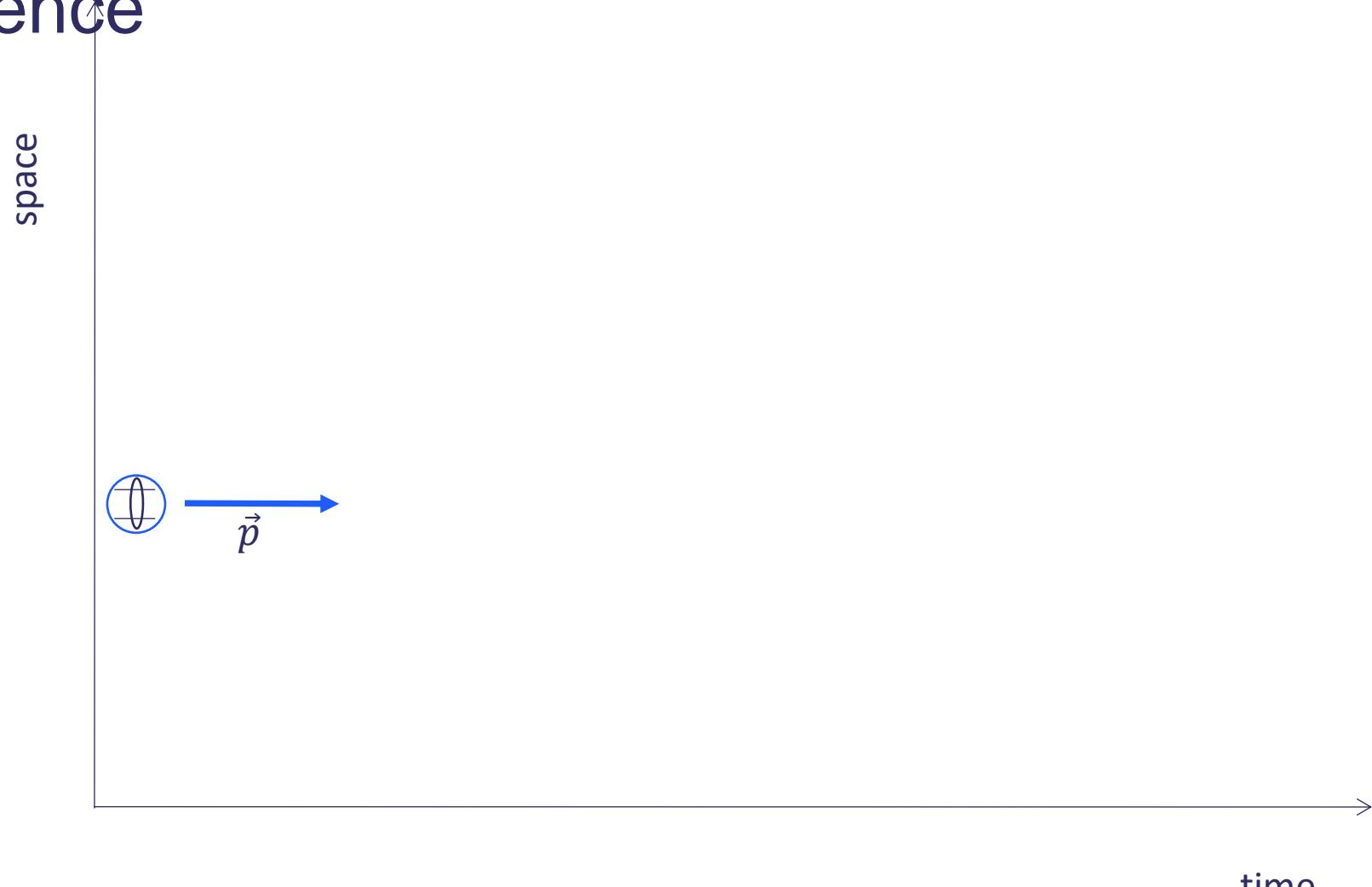
$$|\psi_0\rangle = |e\rangle$$

Intro to Cold Atom Interferometry

Experimental sequence

Intro to Cold Atom Interferometry

Experimental sequence



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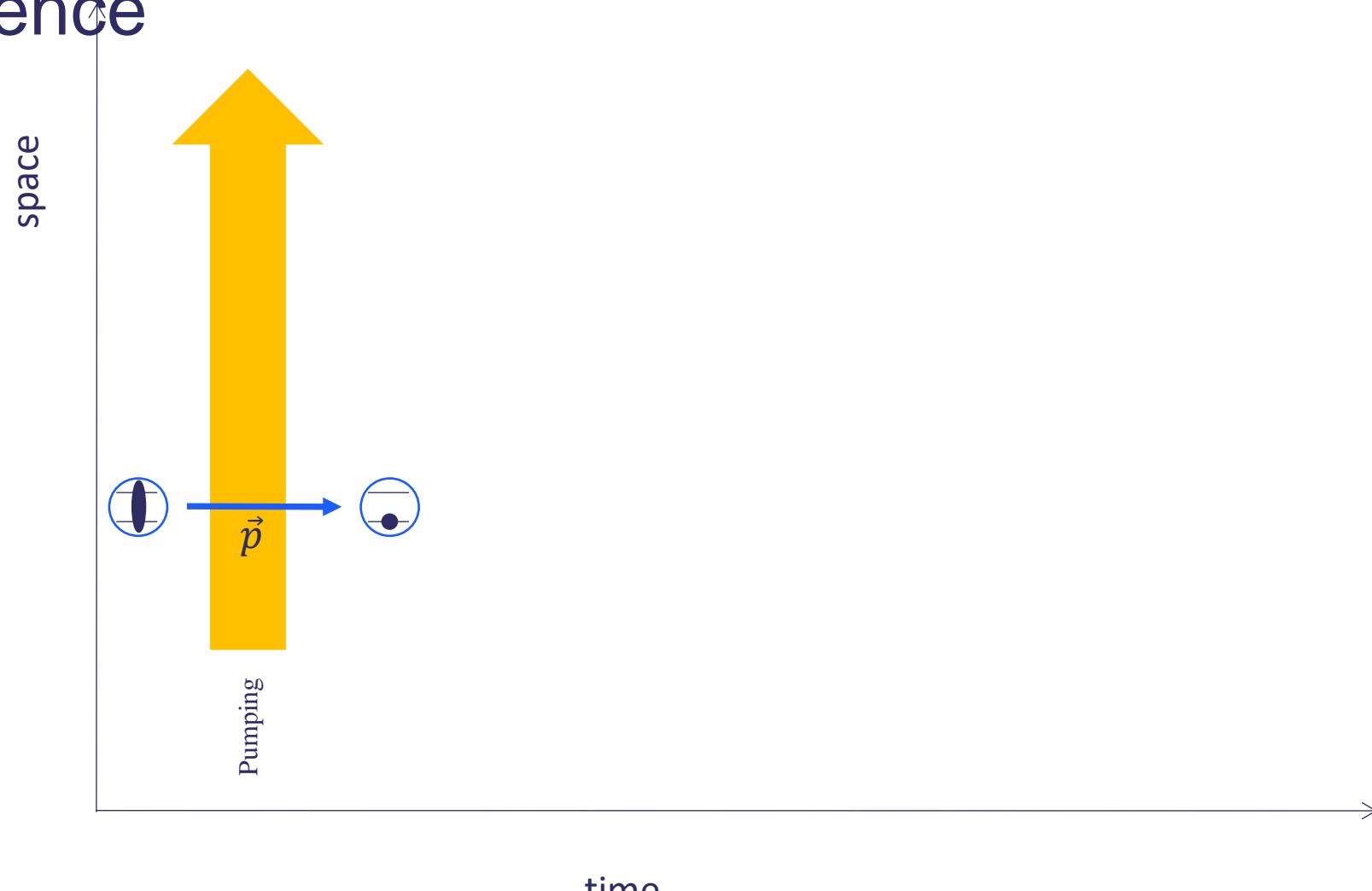
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Intro to Cold Atom Interferometry

Experimental sequence

State preparation and release

$$|\psi\rangle = |g, \vec{p}_0\rangle$$

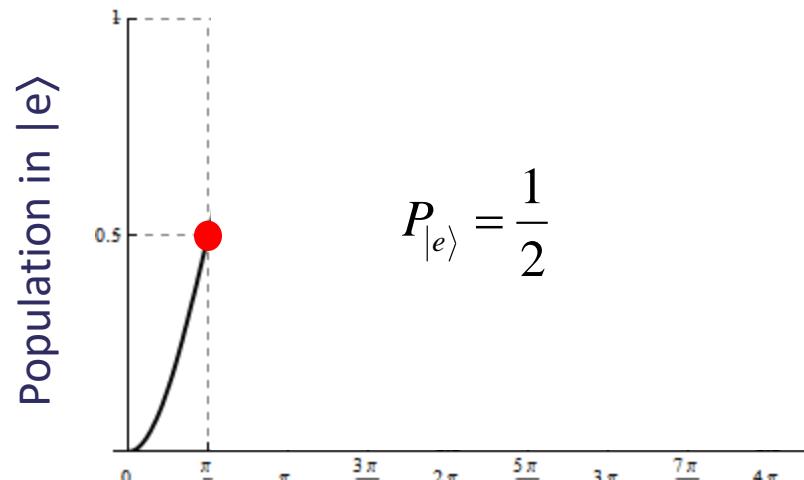


Intro to Cold Atom Interferometry

Experimental sequence

Pulse length = $\pi/2 / \Omega_r$

$$|\psi\rangle = \frac{1}{\sqrt{2}}(|g, \vec{p}_0\rangle + |e, \vec{p}_0 + \hbar\vec{k}\rangle)$$

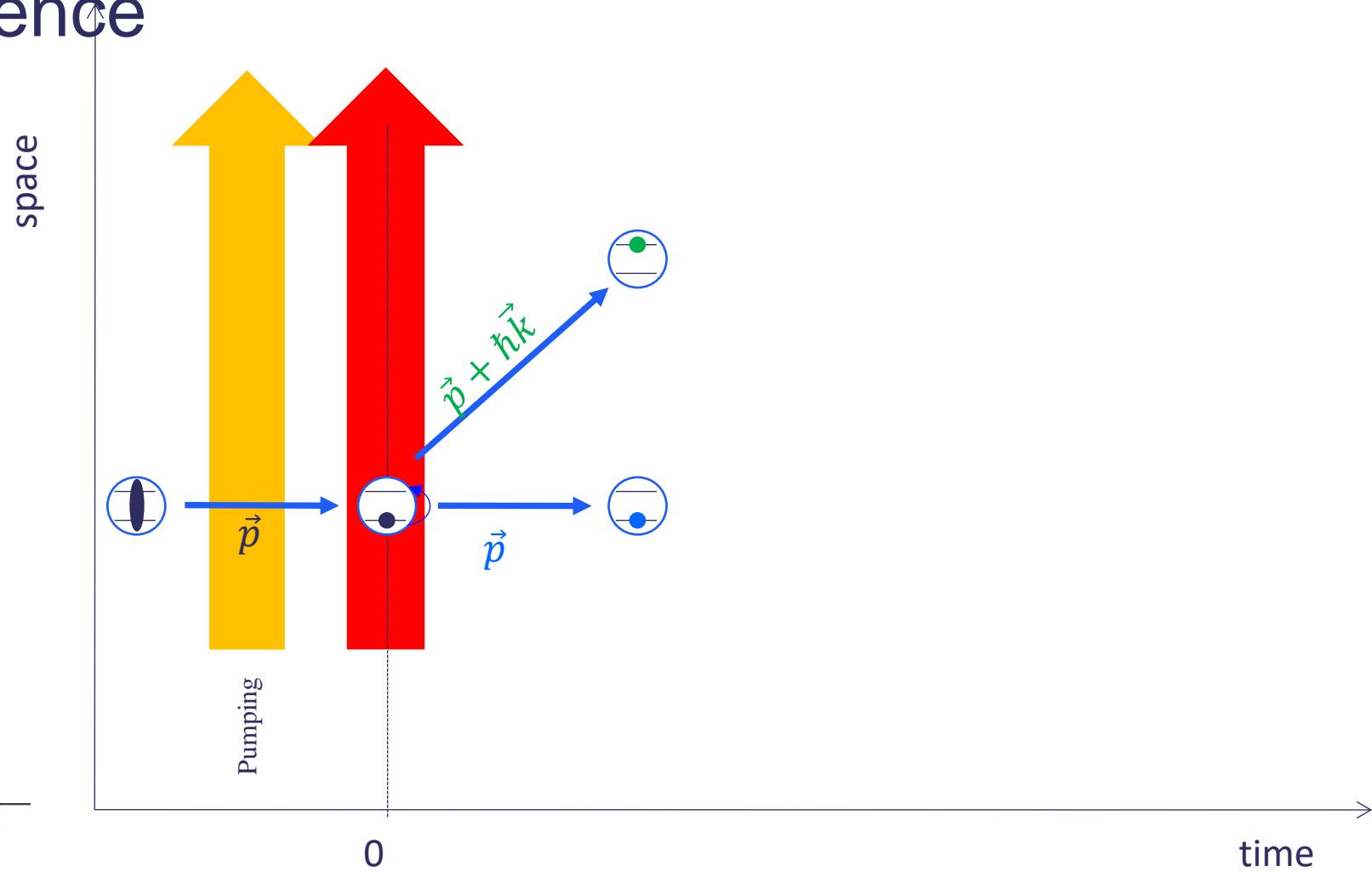


$$P_{|e\rangle} = \frac{1}{2}$$



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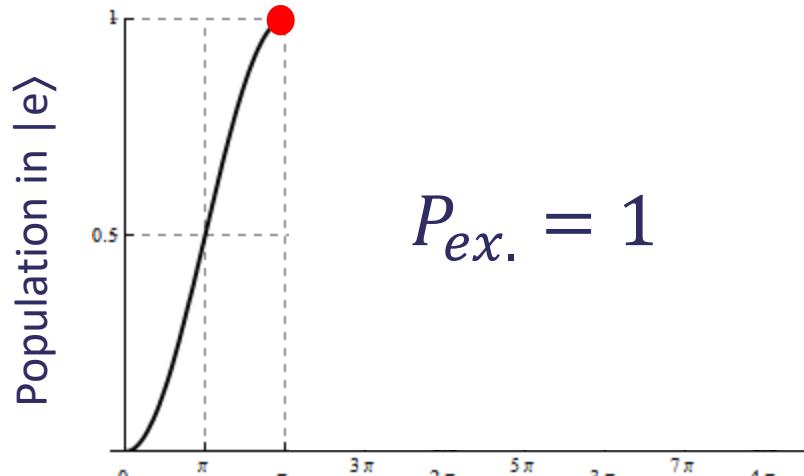


Intro to Cold Atom Interferometry

Experimental sequence

Pulse length = $\pi \cdot \Omega_r$

$$|g\rangle \Rightarrow |e, \vec{p}_0 + \hbar\vec{k}\rangle$$
$$|e\rangle \Rightarrow |g, \vec{p}_0\rangle$$

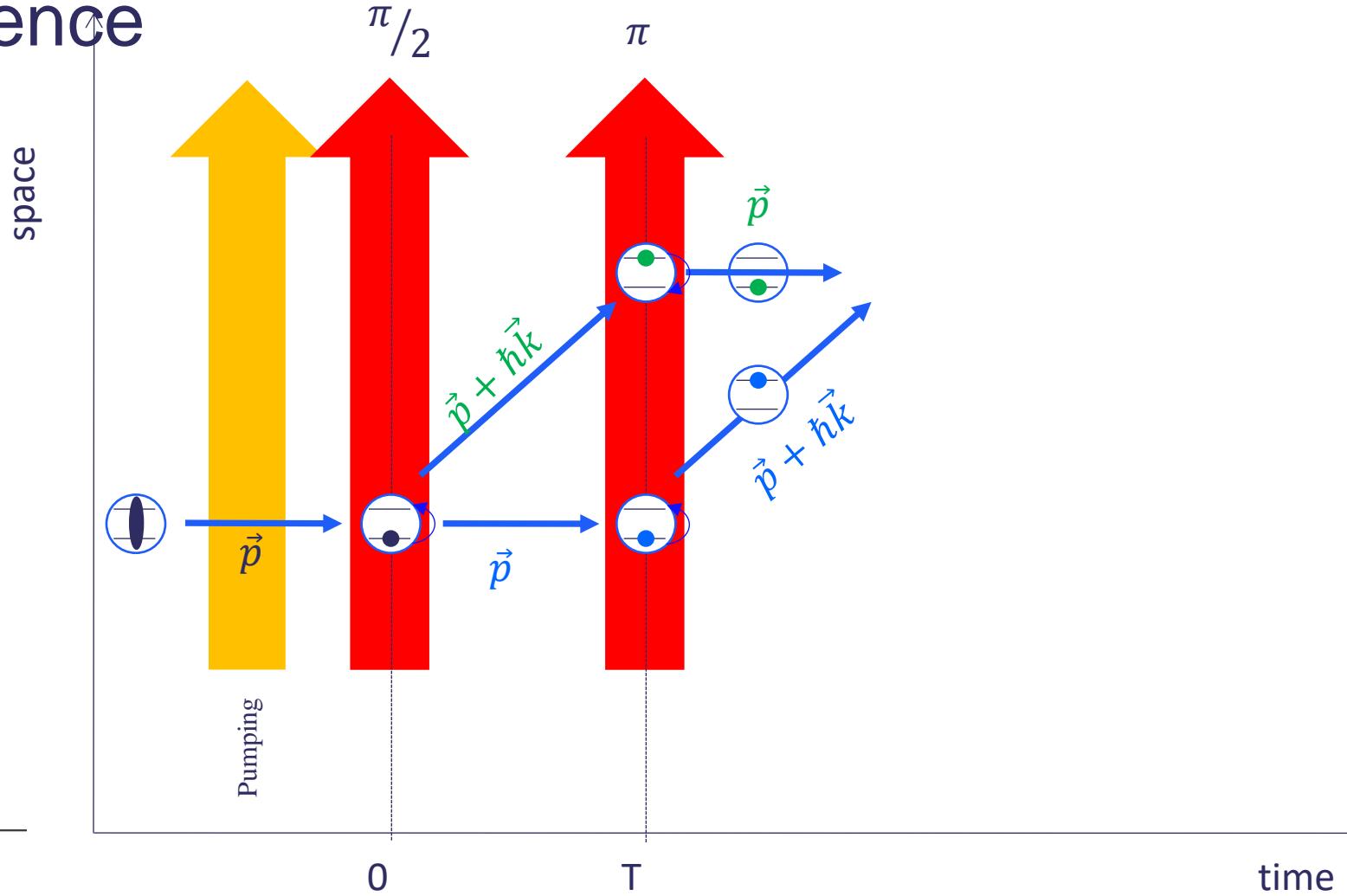


$$P_{ex.} = 1$$



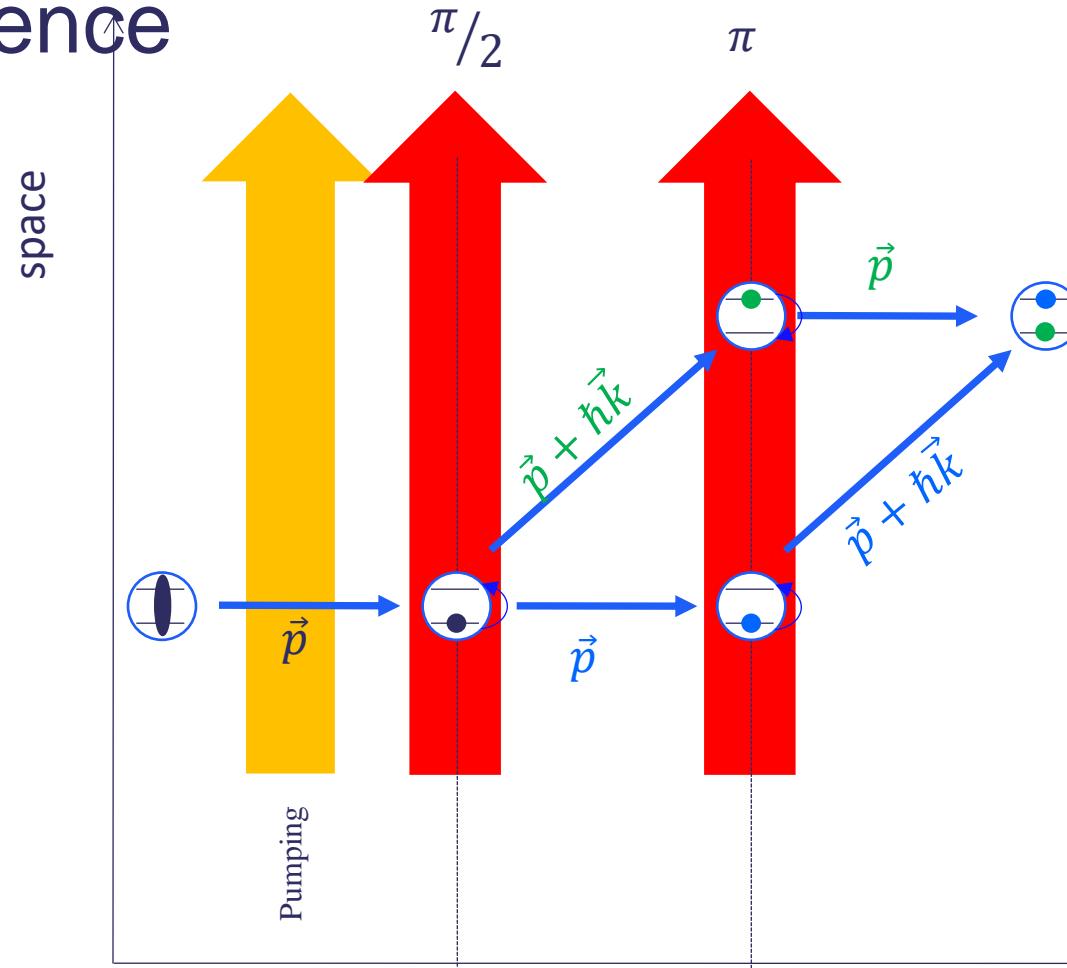
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Intro to Cold Atom Interferometry

Experimental sequence



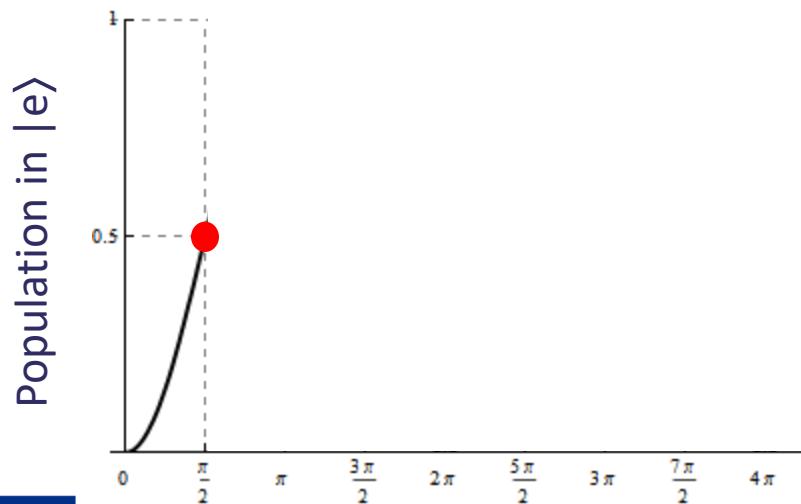
Intro to Cold Atom Interferometry

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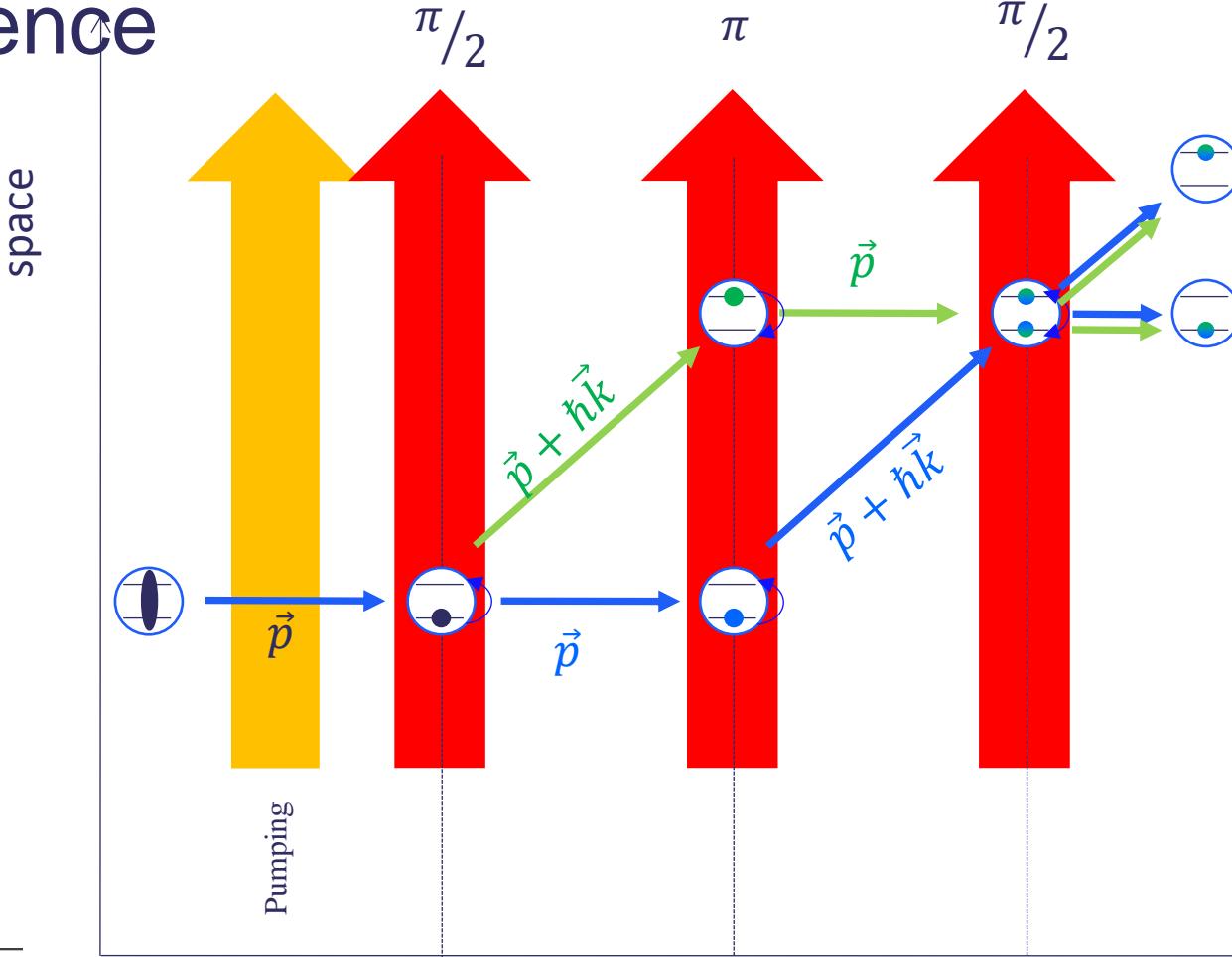
$$|g\rangle \Rightarrow \frac{1}{\sqrt{2}} (|g, \vec{p}_0\rangle + |e, \vec{p}_0 + \hbar\vec{k}\rangle)$$

$$|e\rangle \Rightarrow \frac{1}{\sqrt{2}} (|e, \vec{p}_0 + \hbar\vec{k}\rangle + |g, \vec{p}_0\rangle)$$



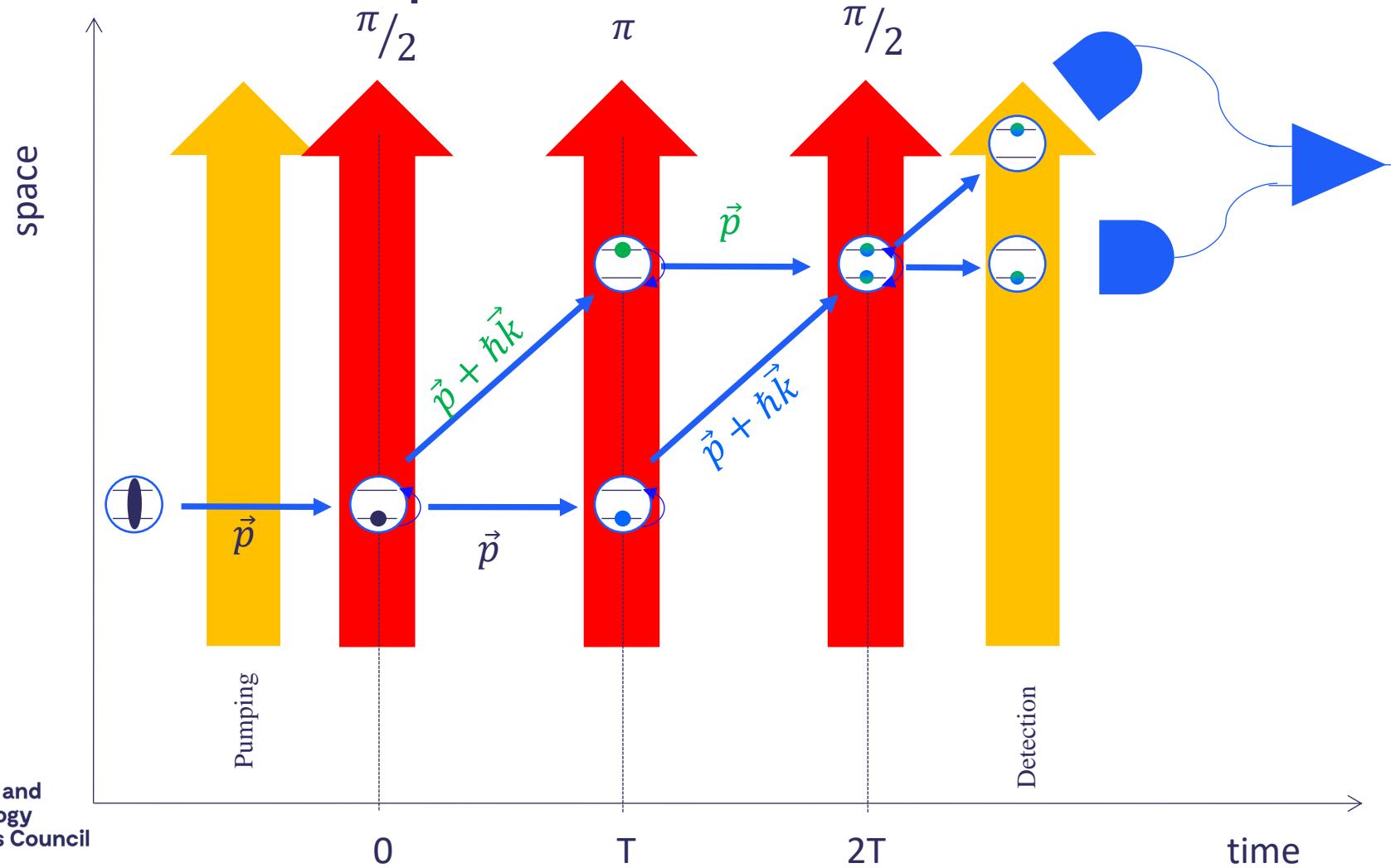
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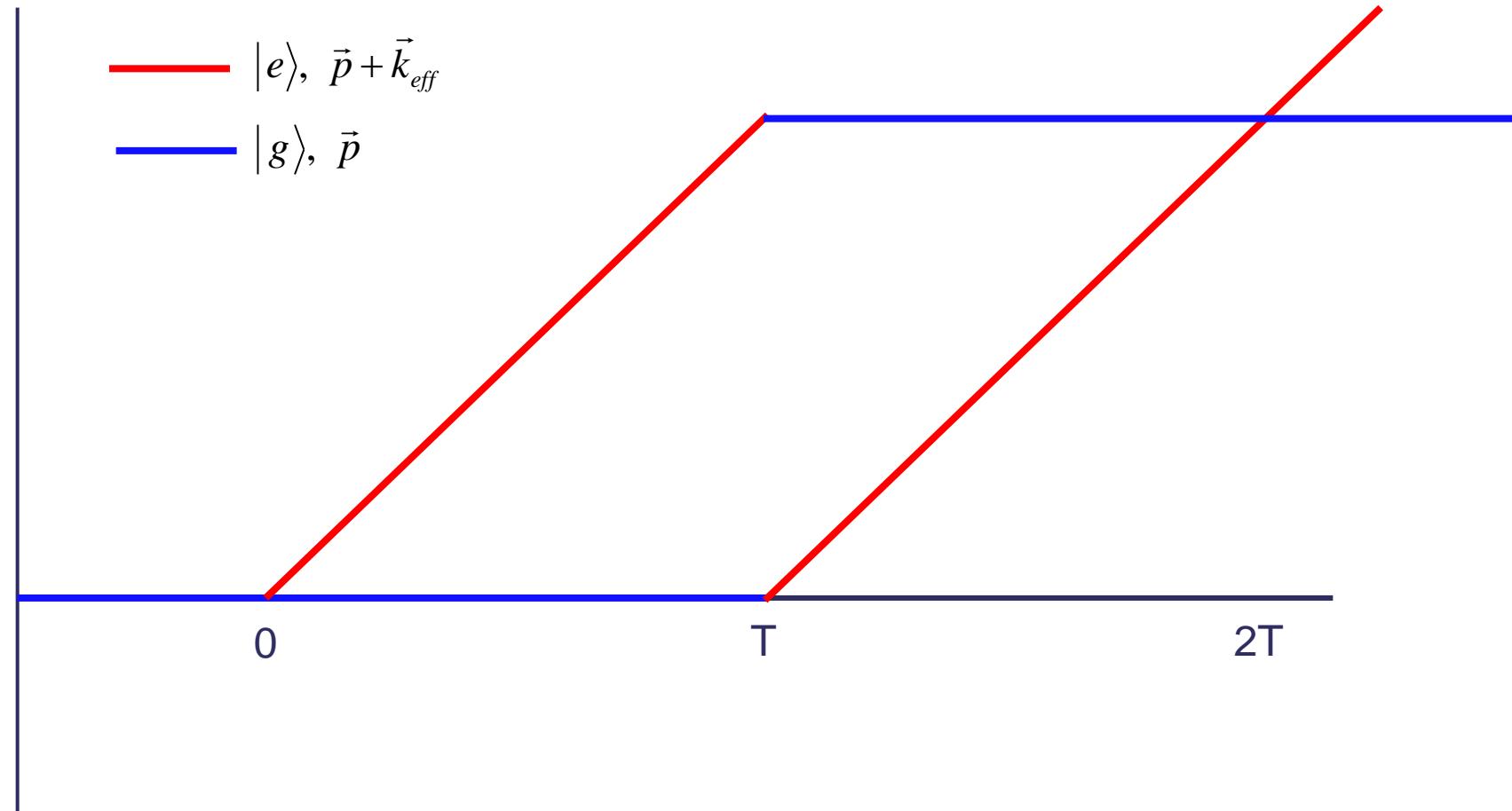
Intro to Cold Atom Interferometry

Experimental sequence



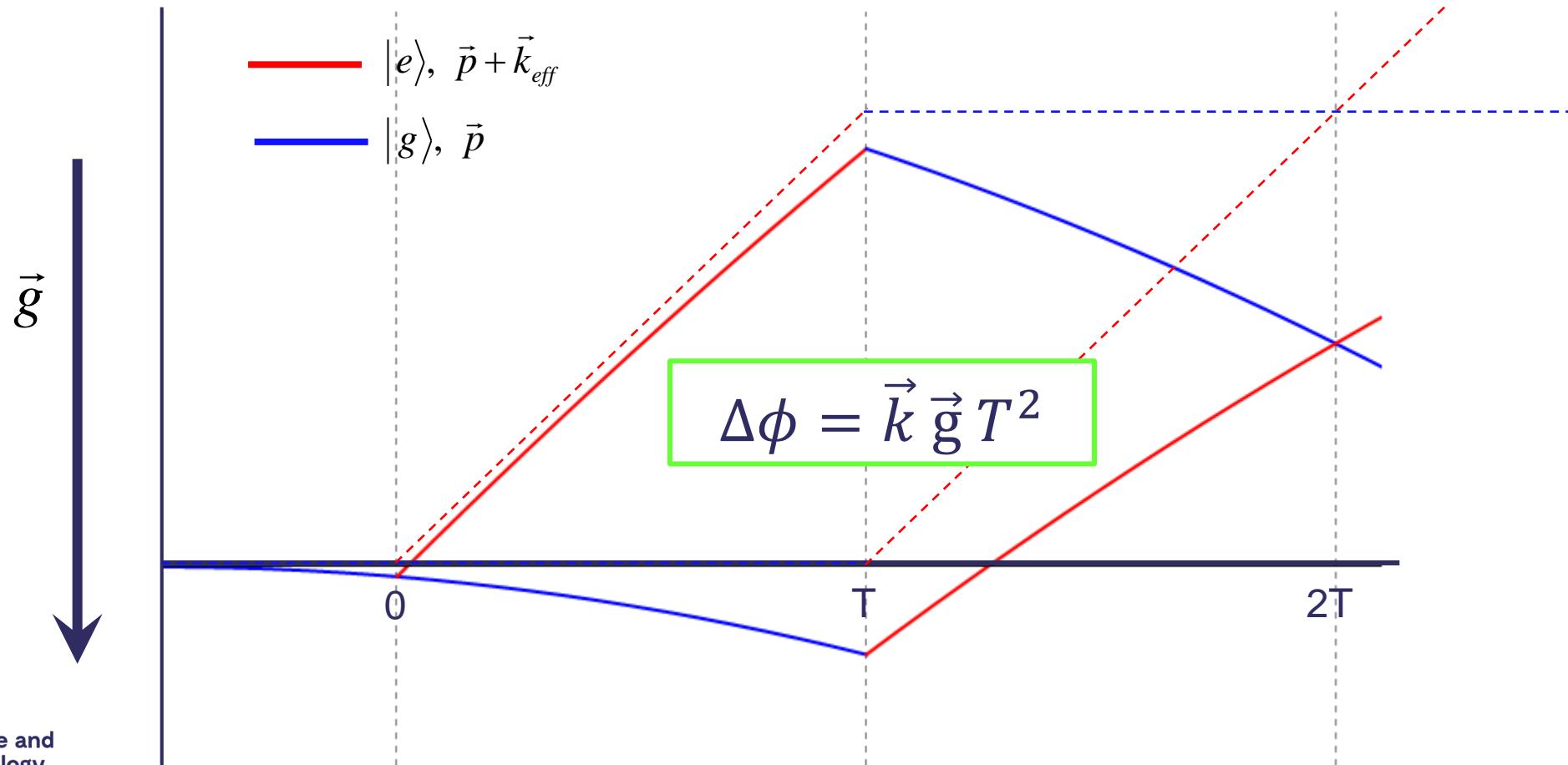
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Experimental sequence



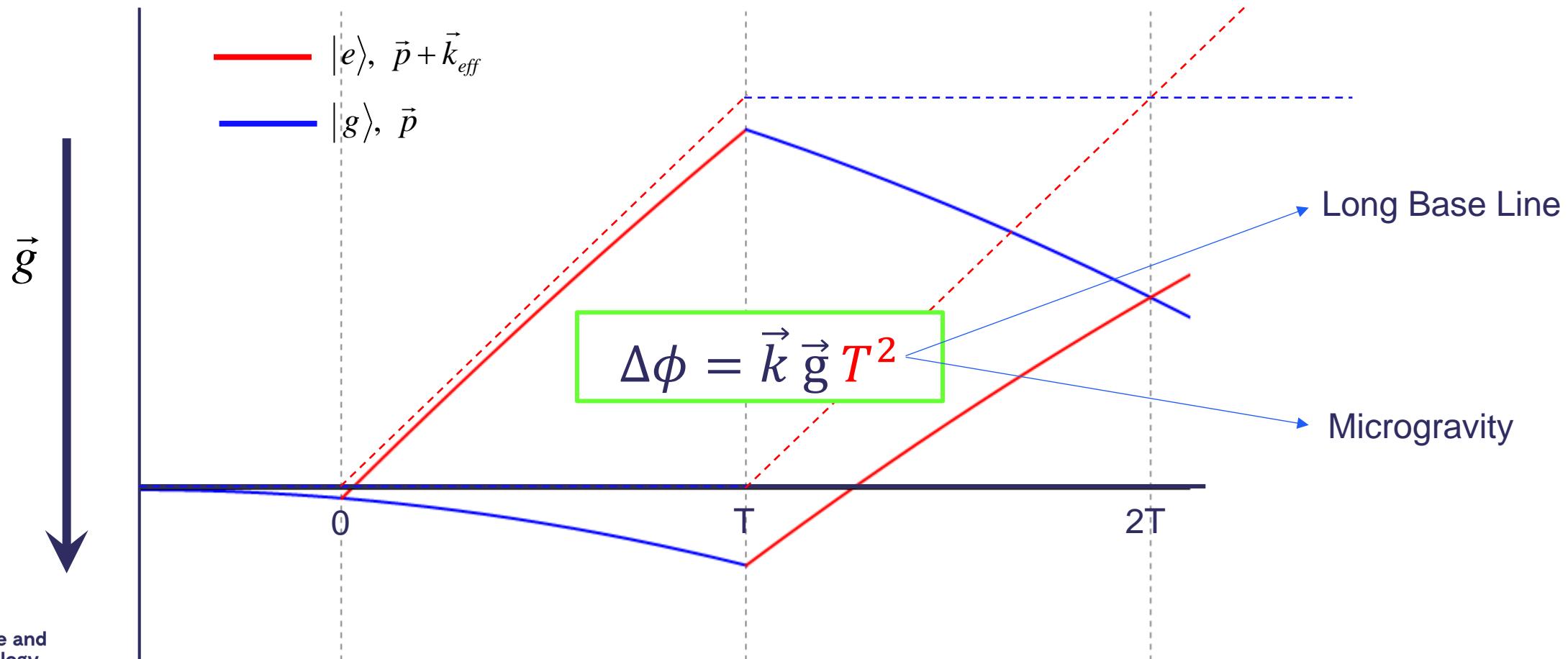
Intro to Cold Atom Interferometry

Experimental sequence



Intro to Cold Atom Interferometry

Experimental sequence



Gravity Mapping and ...

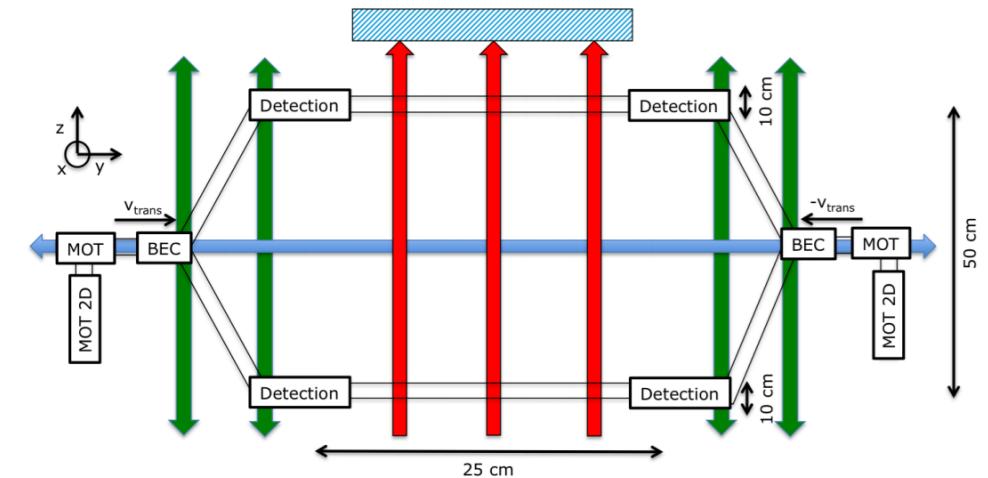
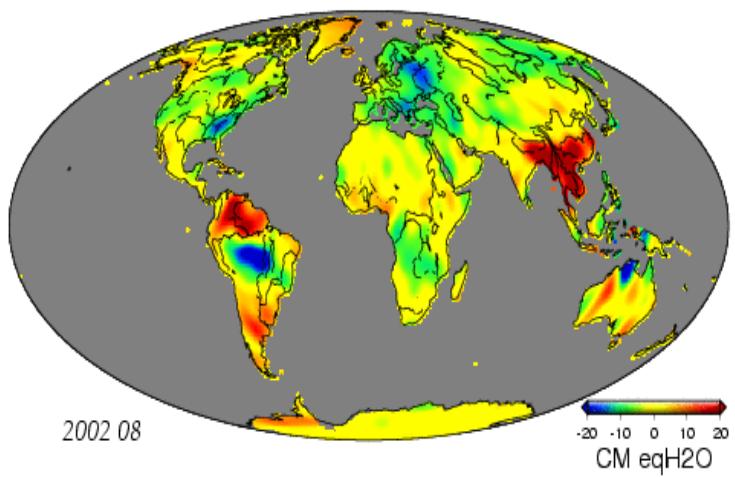
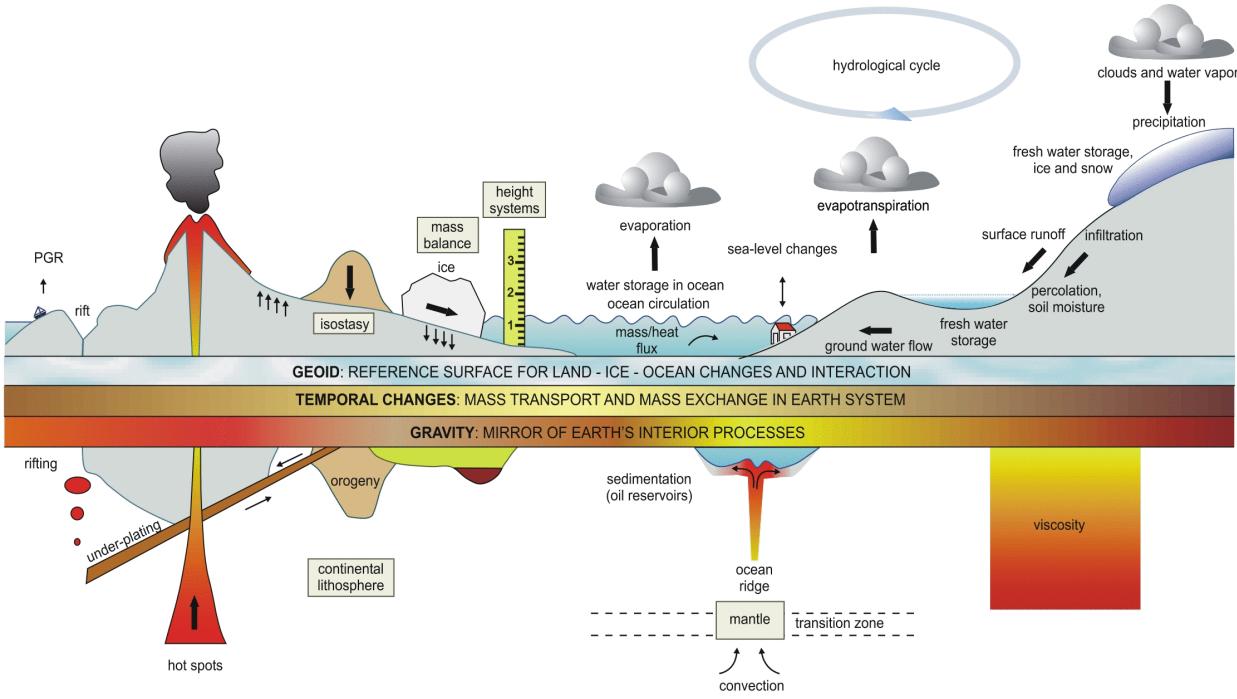
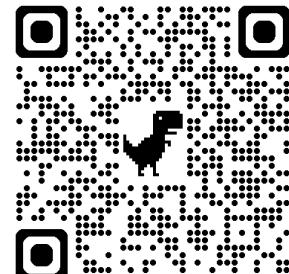
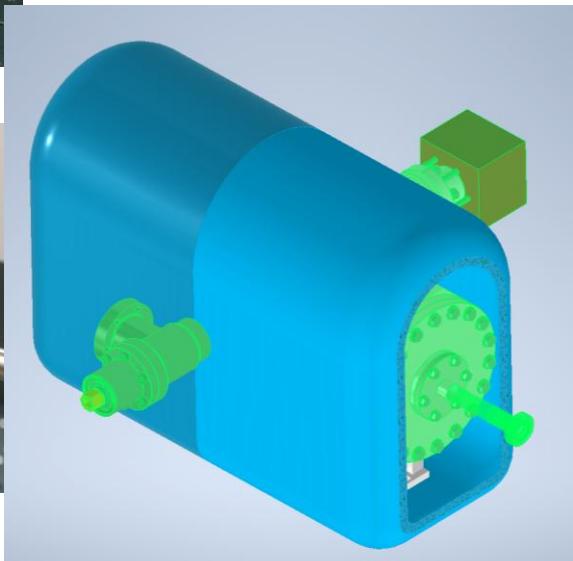
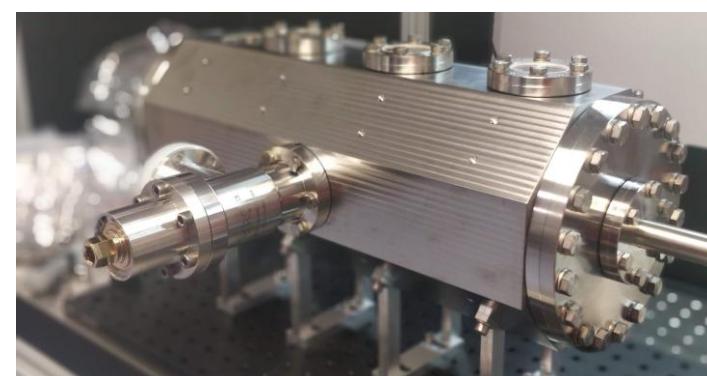
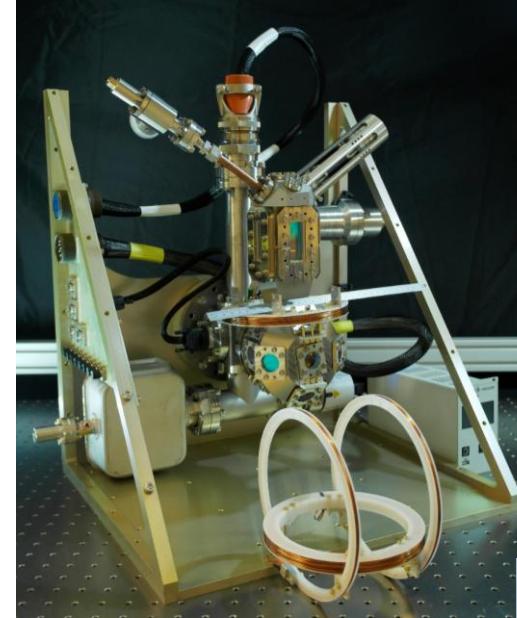
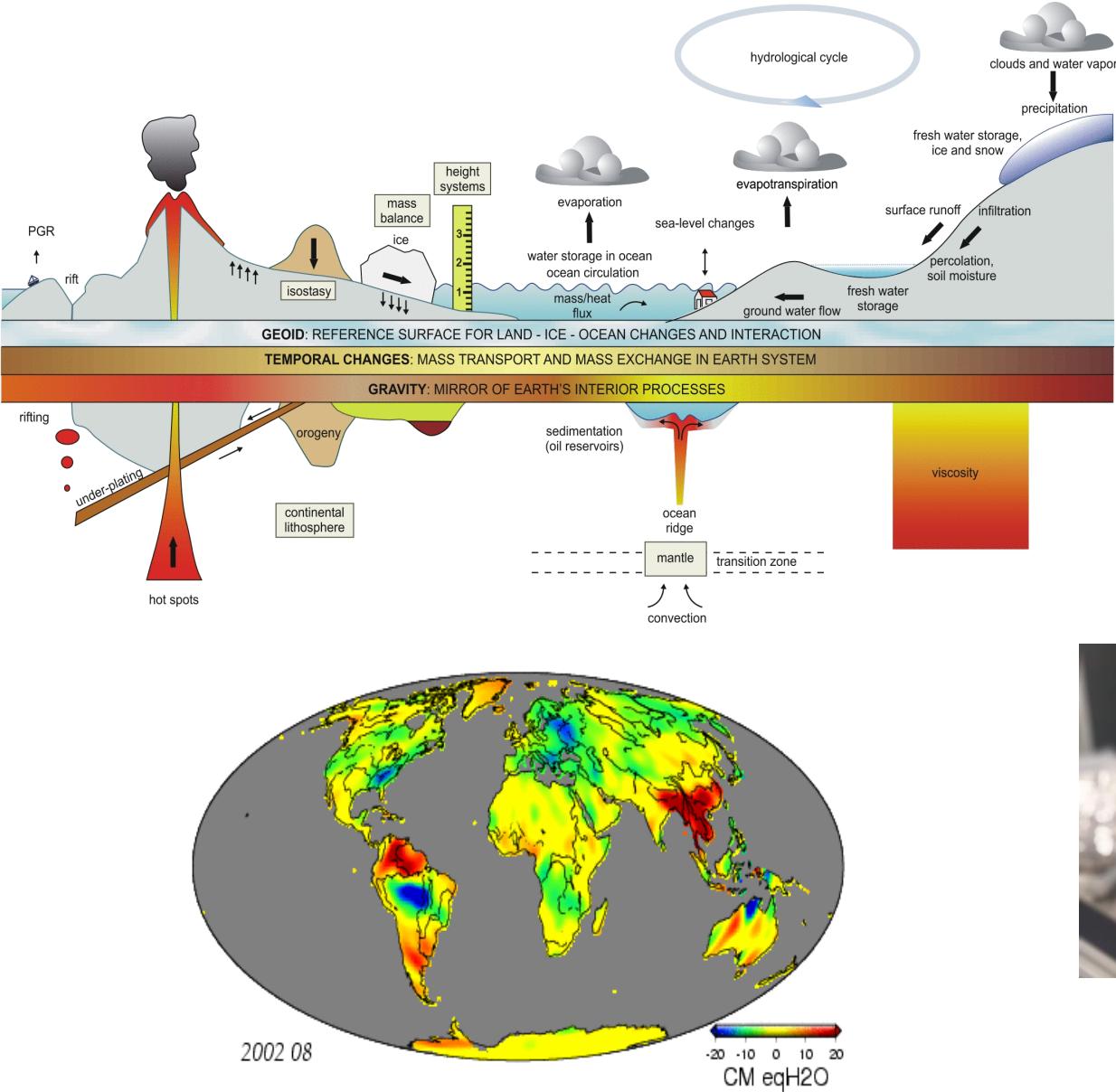


FIG. 3: (Color online) Scheme of the vacuum chamber (not drawn to scale; orientative sizing). The red arrows represent the Raman lasers for the interferometry part; the blue arrows represent the light pulse giving the transverse velocity v_{trans} ; the green arrows represent the light pulses guiding the atoms from the cooling room to the interferometer room; the blue rectangle is the mirror, which will be the inertial reference.

Based on O. Carraz et al, Microgravity Science and Technology v26, pp 139-145 (2014)



Gravity Mapping and ...



... Atmospheric Science

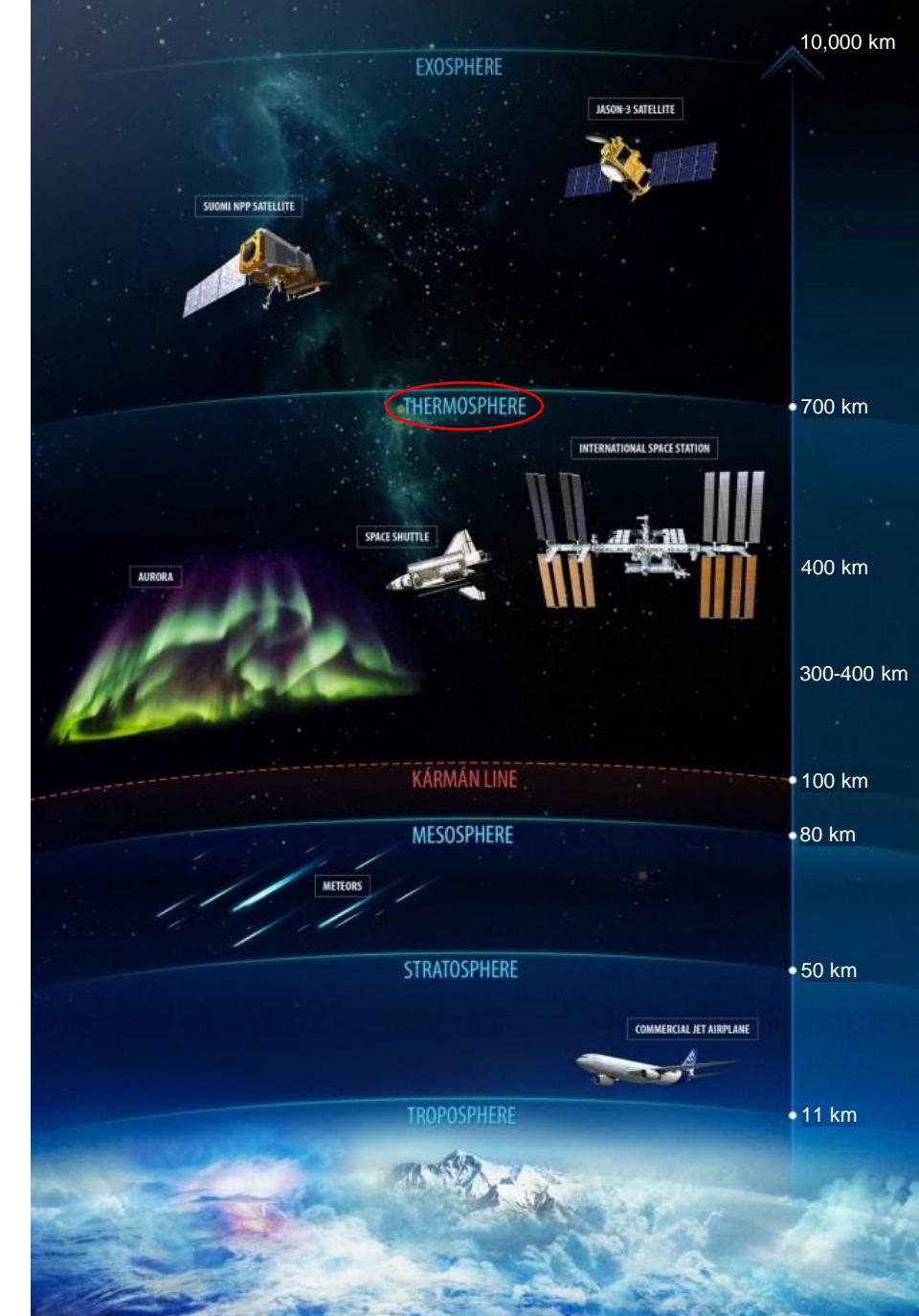
The Thermosphere

....is the interface between Earth and outer space

- important in the Earth system energy balance
- very limited knowledge → energy transport models impacted
- improved data sets → improved Climate modelling

....is where most of the New Space satellites fly

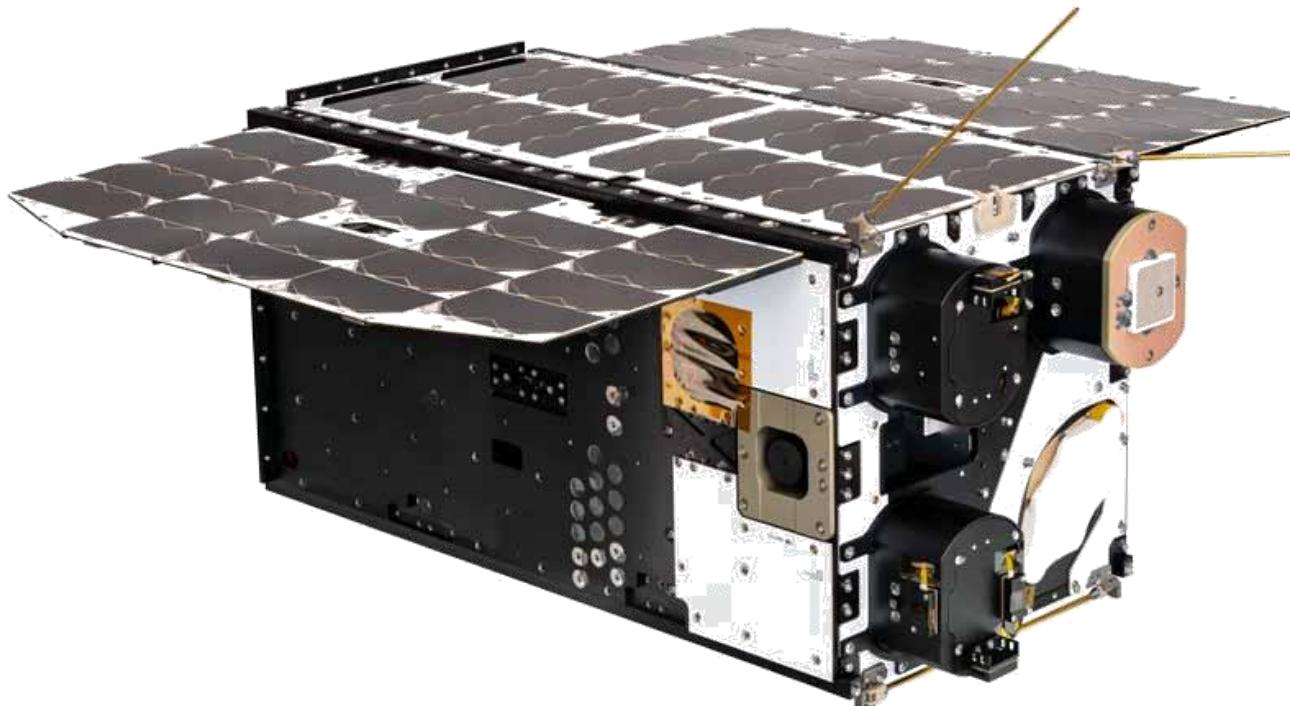
commercial interest on density data to improve predictions on orbit decay



... Atmospheric Science

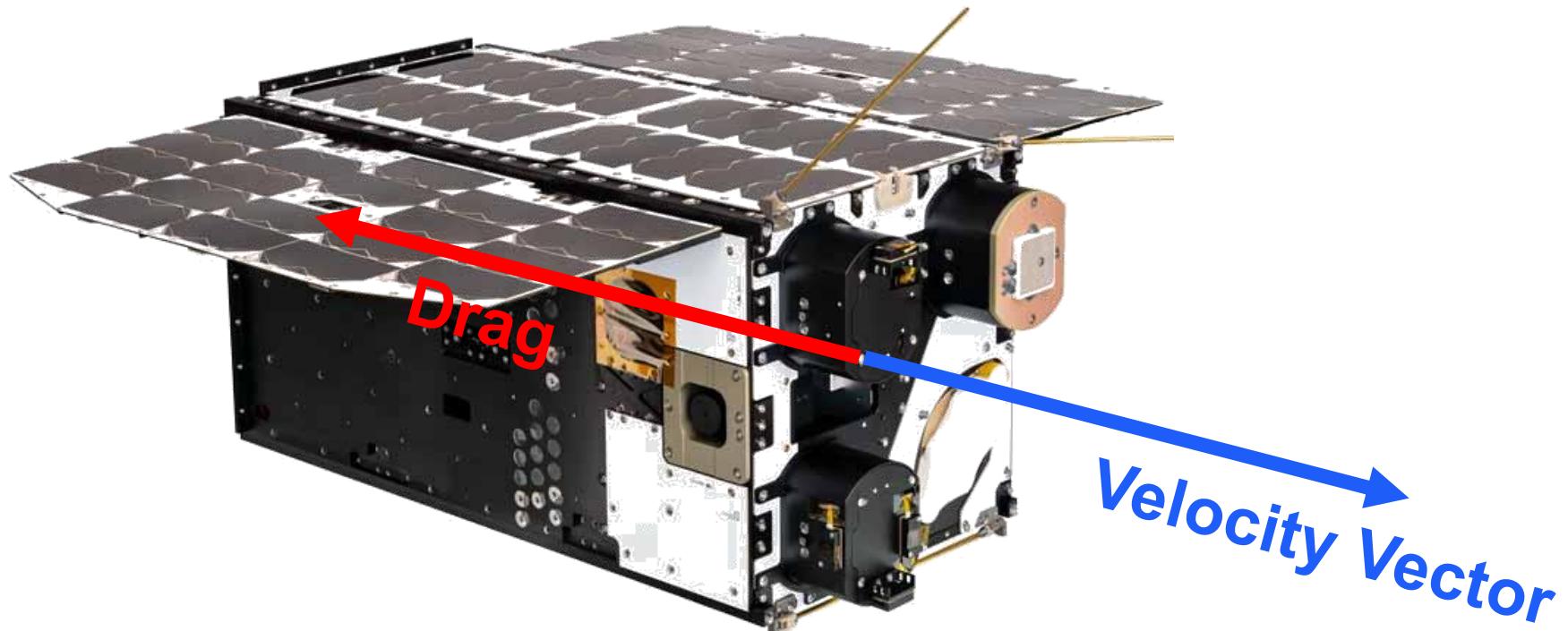
CAITDM Cold Atom Interferometric Thermosphere Drag Measurement

- Lasers + PL Electronics
- Physics Package
- Bus +Avionics



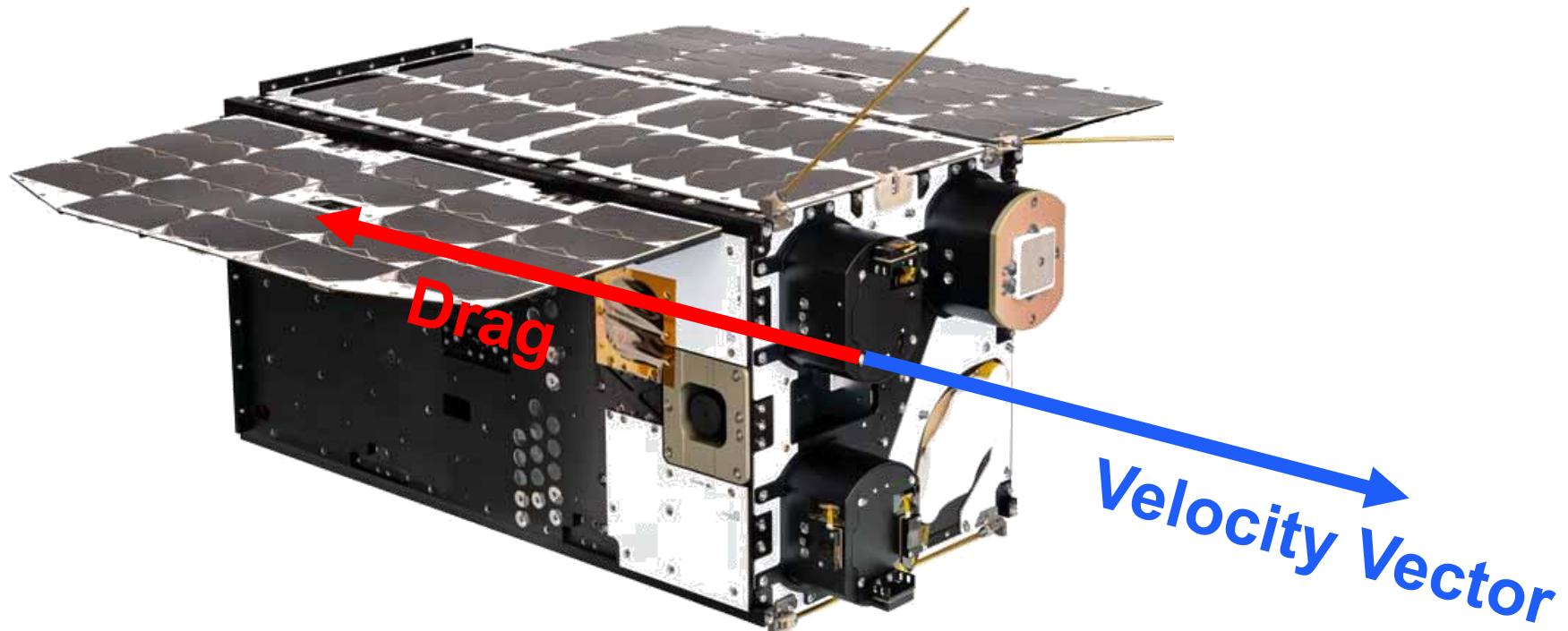
CAITDM

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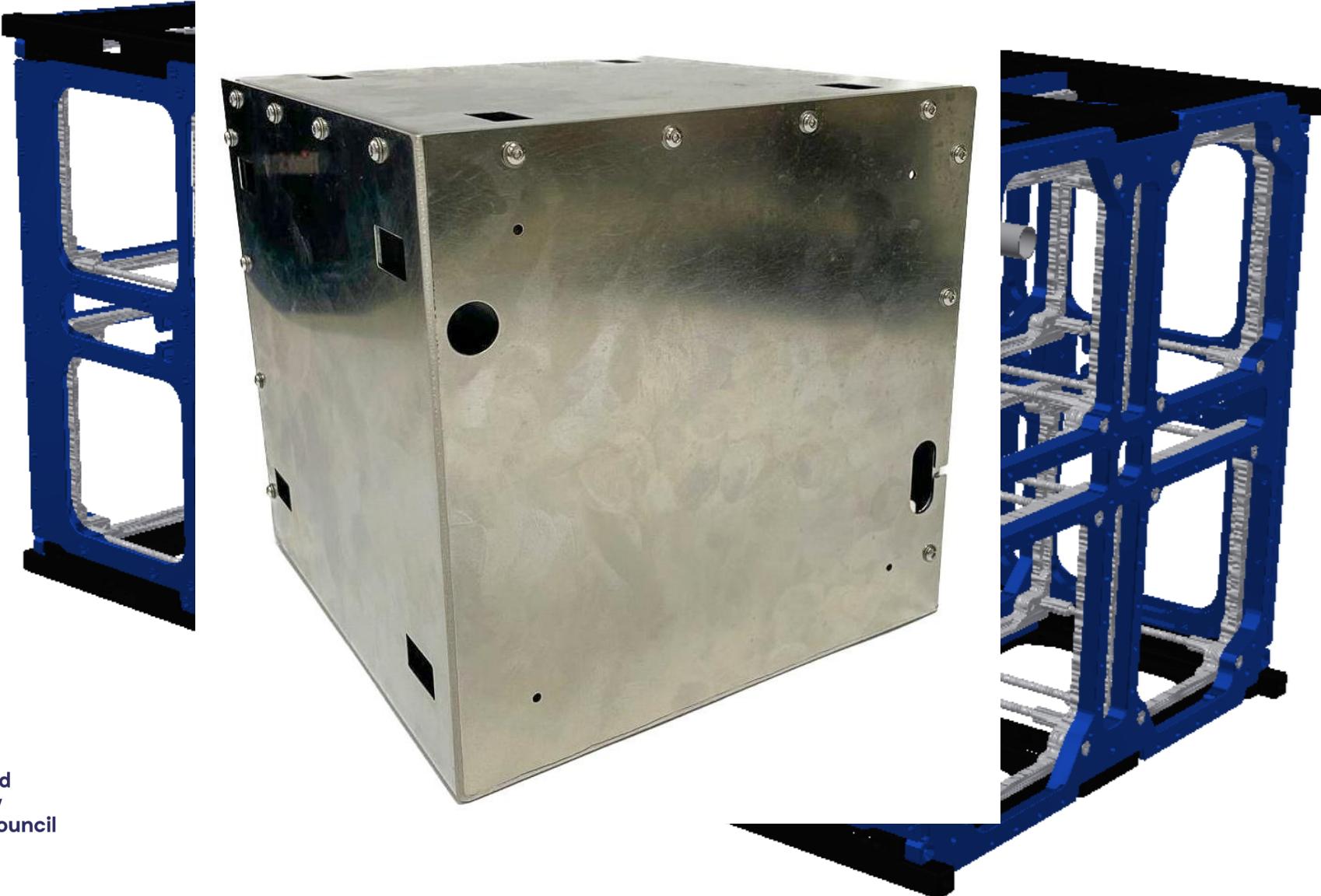


CAITDM

- Lasers + PL Electronics
- Physics Package
- Bus +Avionics

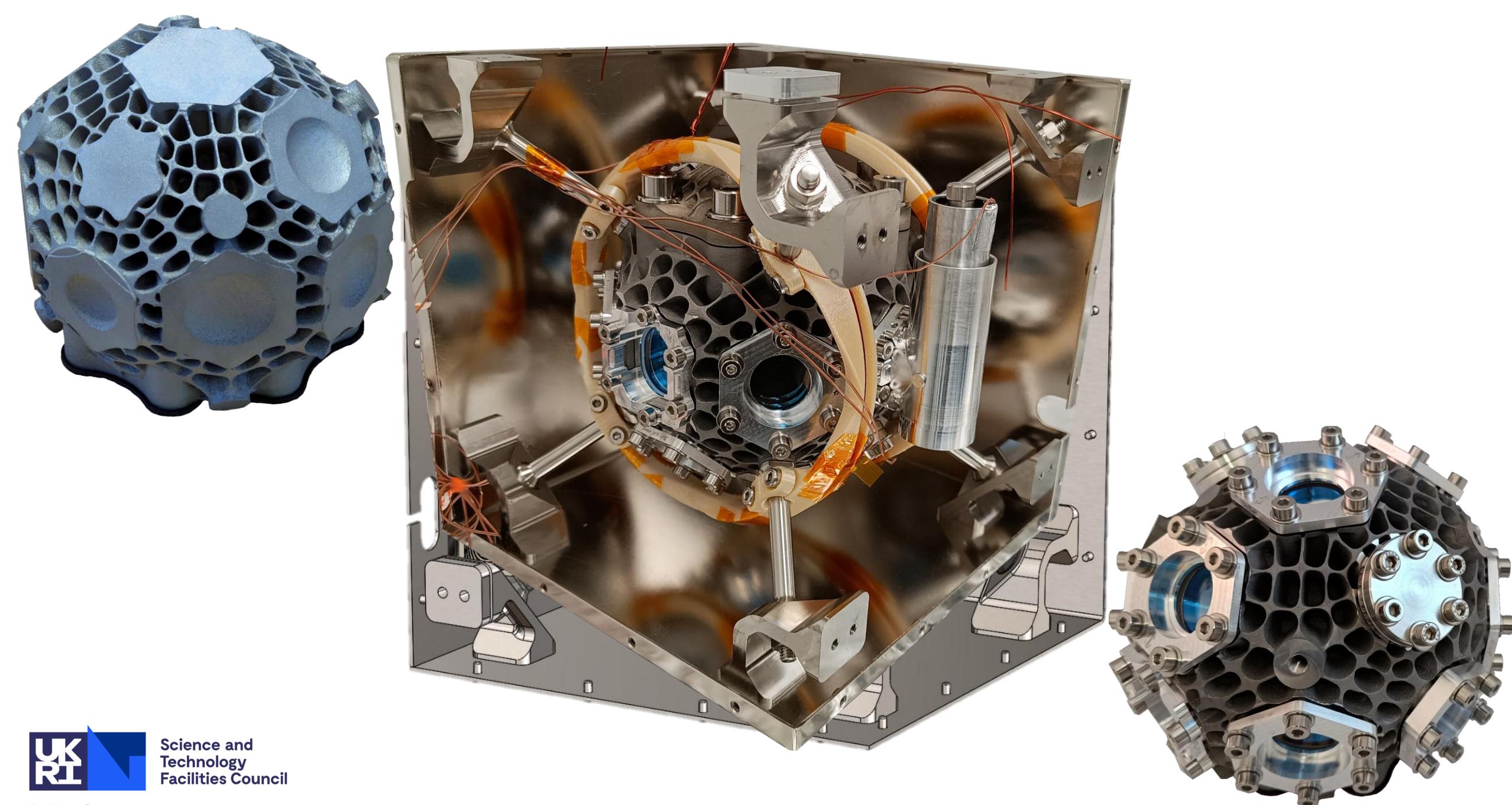


Physics Package



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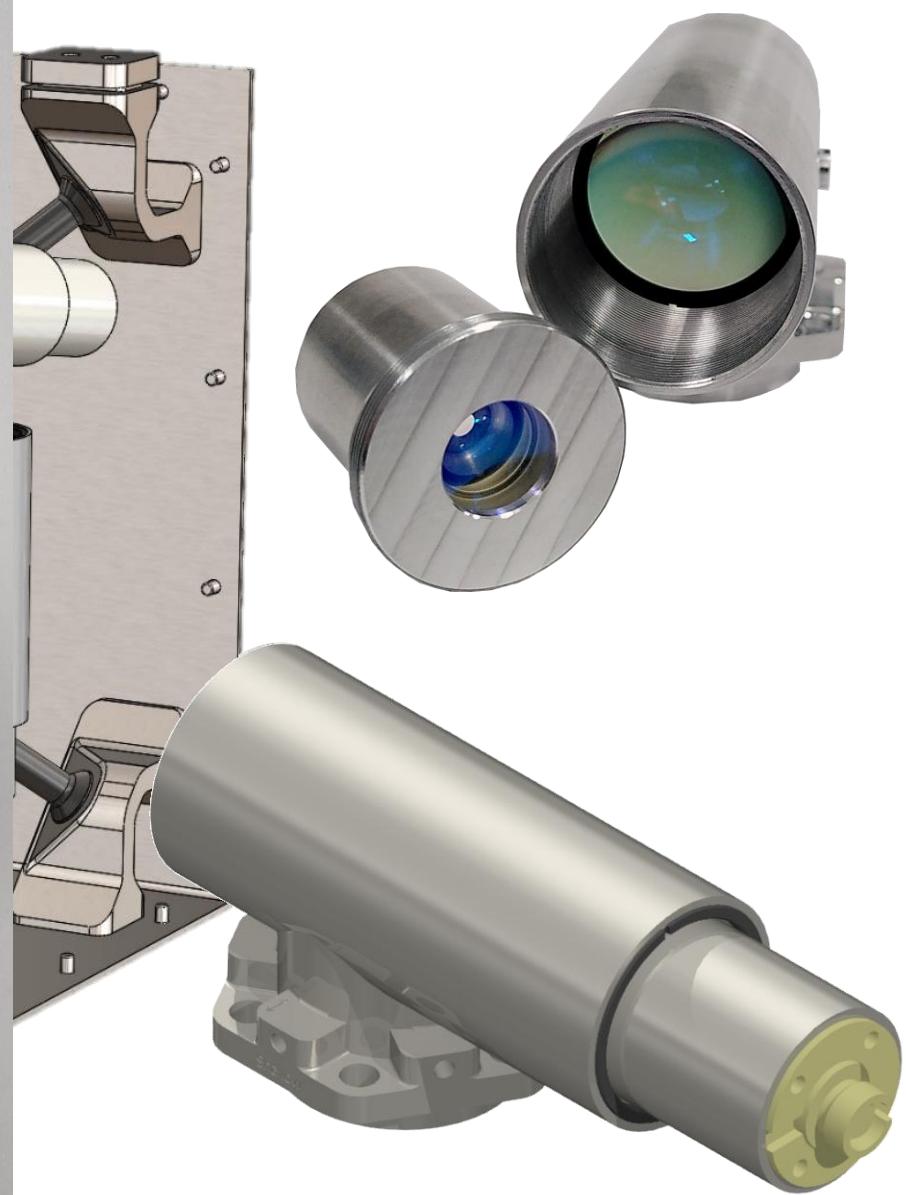
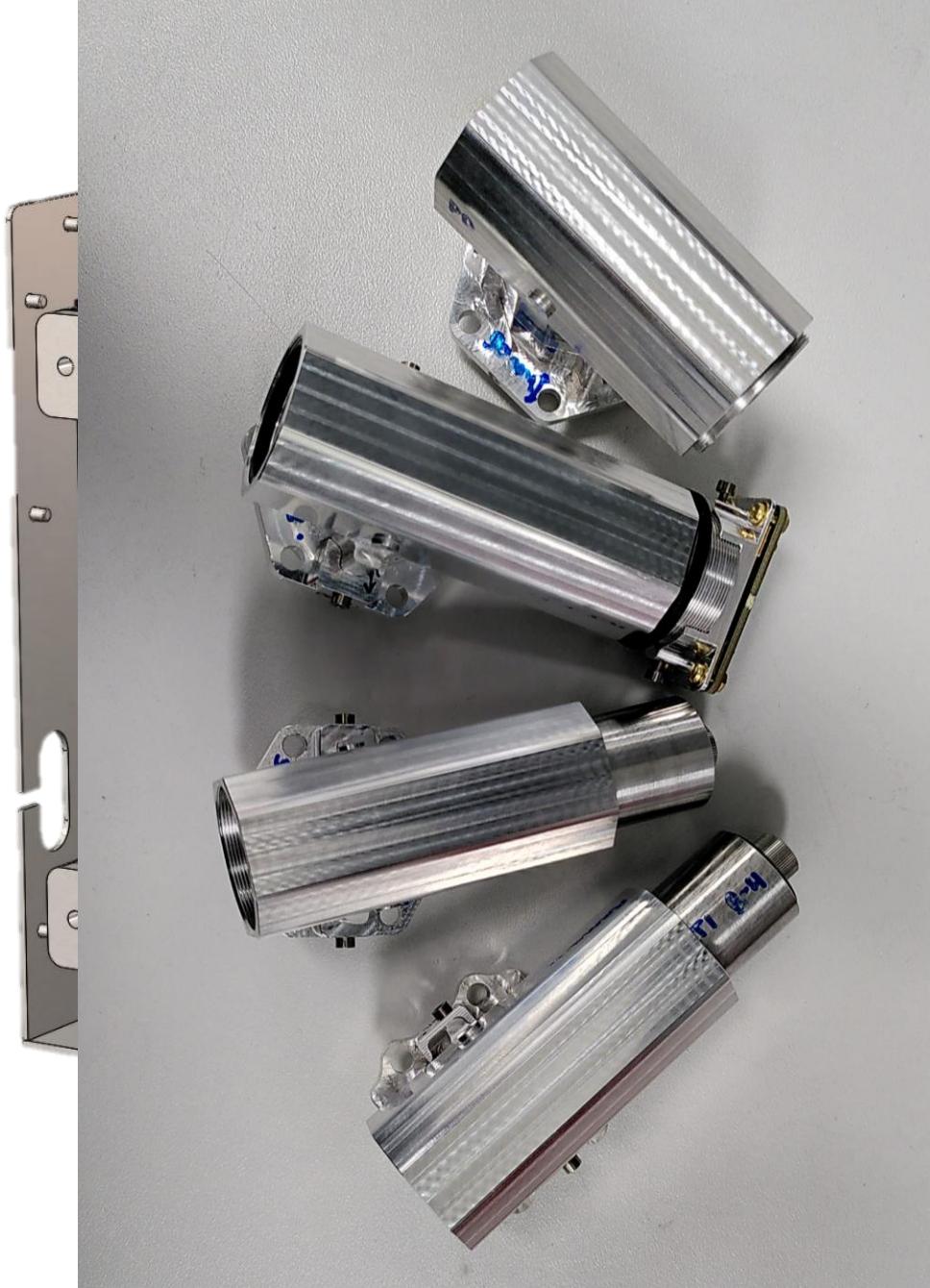
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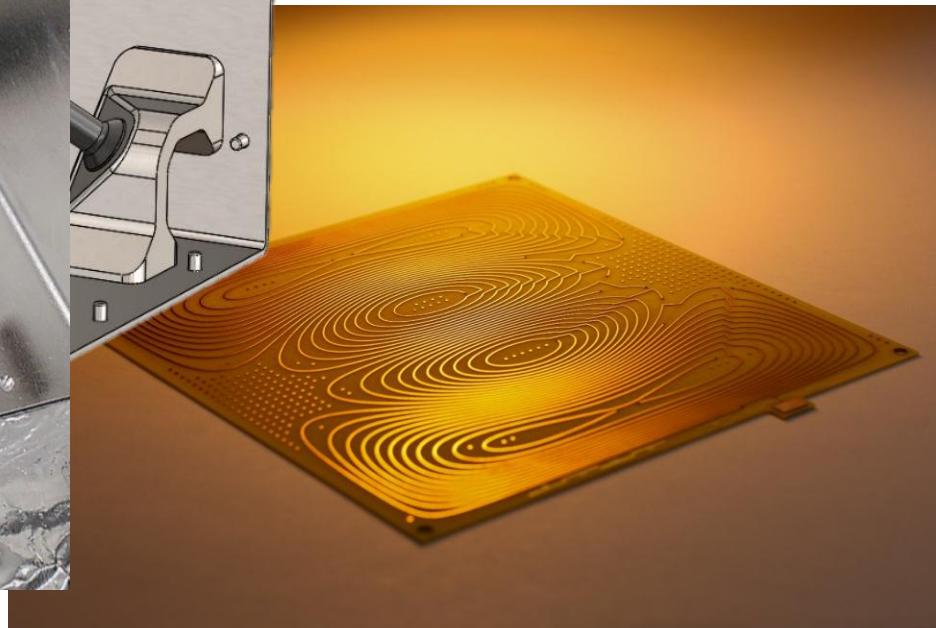
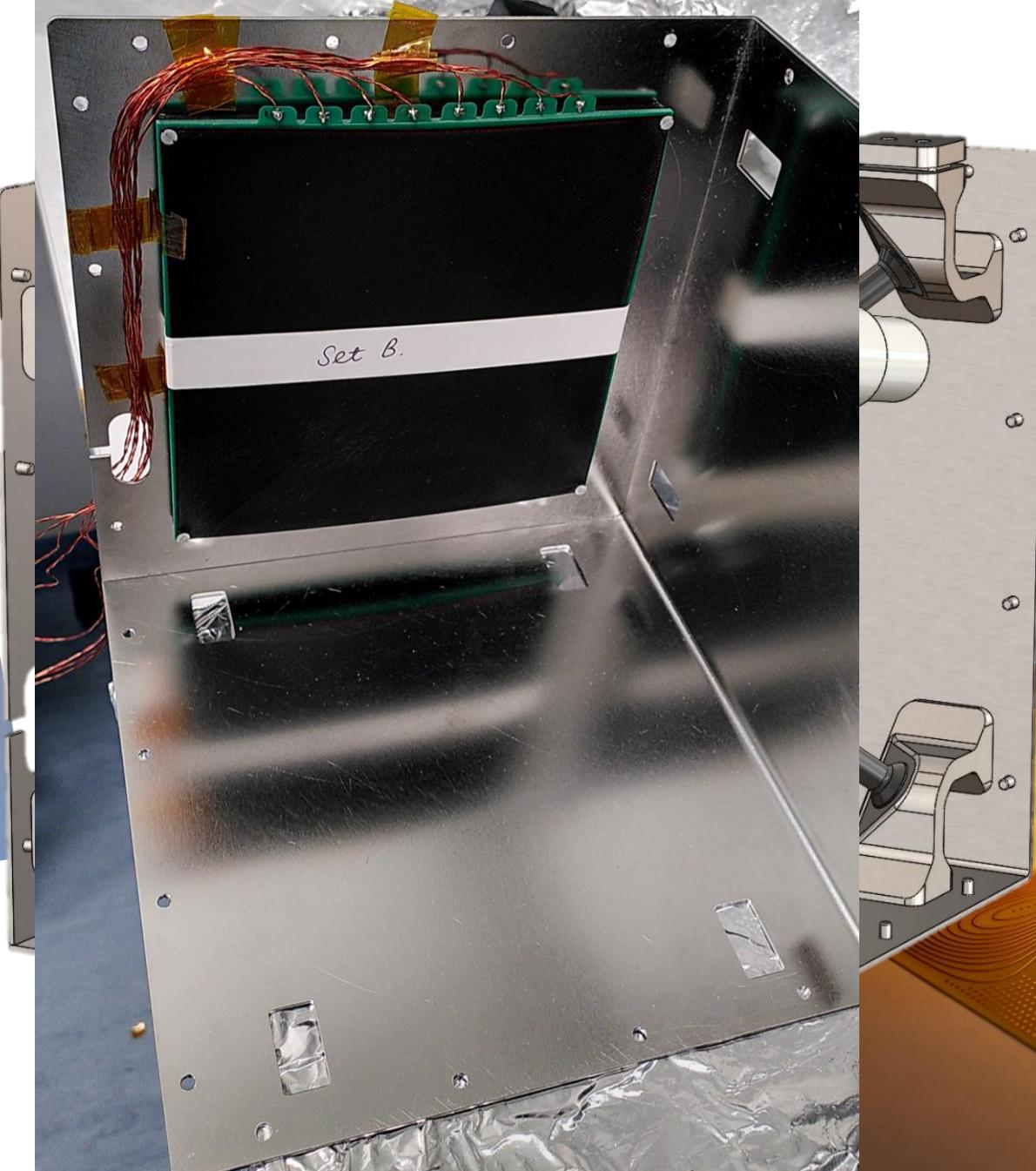
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Environmental Testing

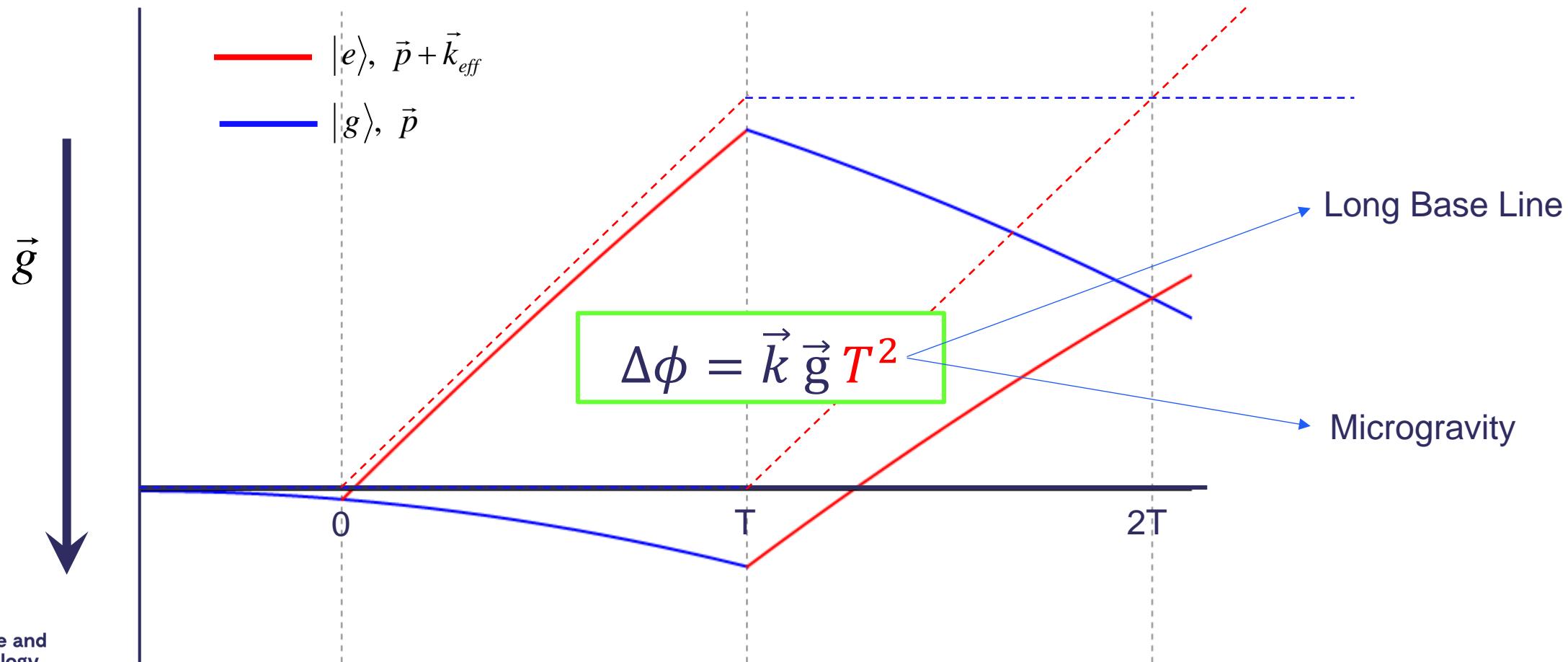
Thermal Cycling: 4 cycles -30°C to +60°C

Vibration to NASA GEVS levels



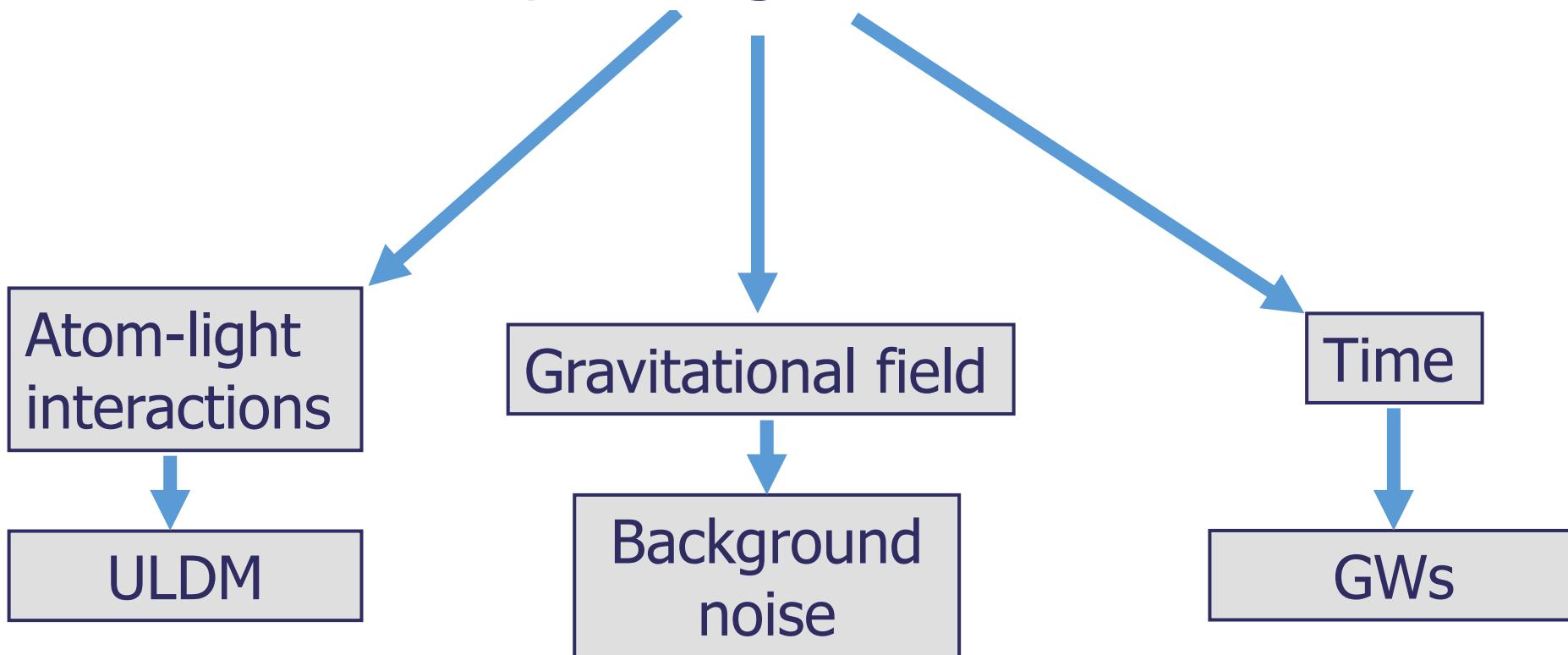
Intro to Cold Atom Interferometry

Experimental sequence



Gravitational Waves and UL Dark Matter

What we measure: $\Delta\phi = \vec{k} \cdot \vec{g} T^2$



Fundamental Physics: Worldwide

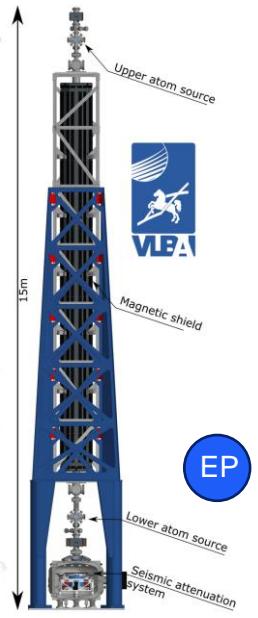
10m fountain, Stanford, US



10m AI, Wuhan, CN



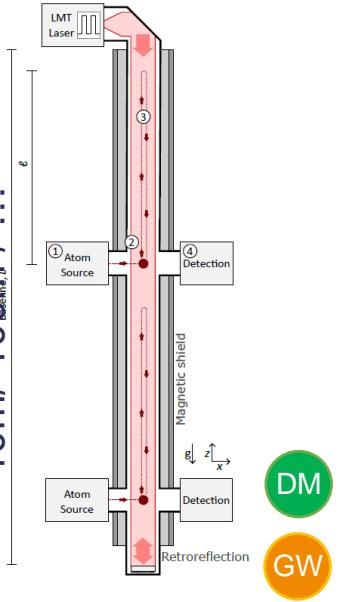
15m VLBAI, Hannover,
DE



MAGIS, Stanford/Fermilab, US
10m / 100m / ...



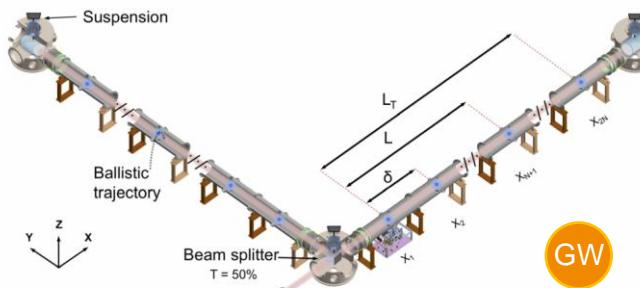
ALION, UK
10m/ 100m / ...



MIGA, FR



ELGAR, FR



ZAIGA, Wuhan, CN



EP Equivalence
Principle

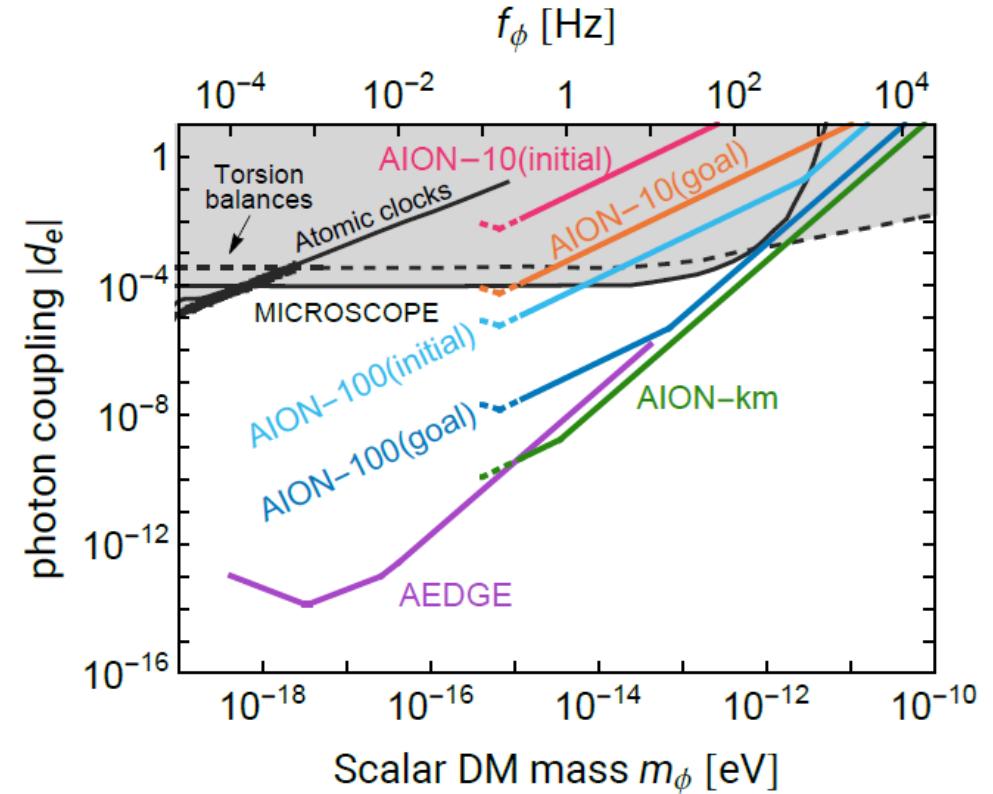
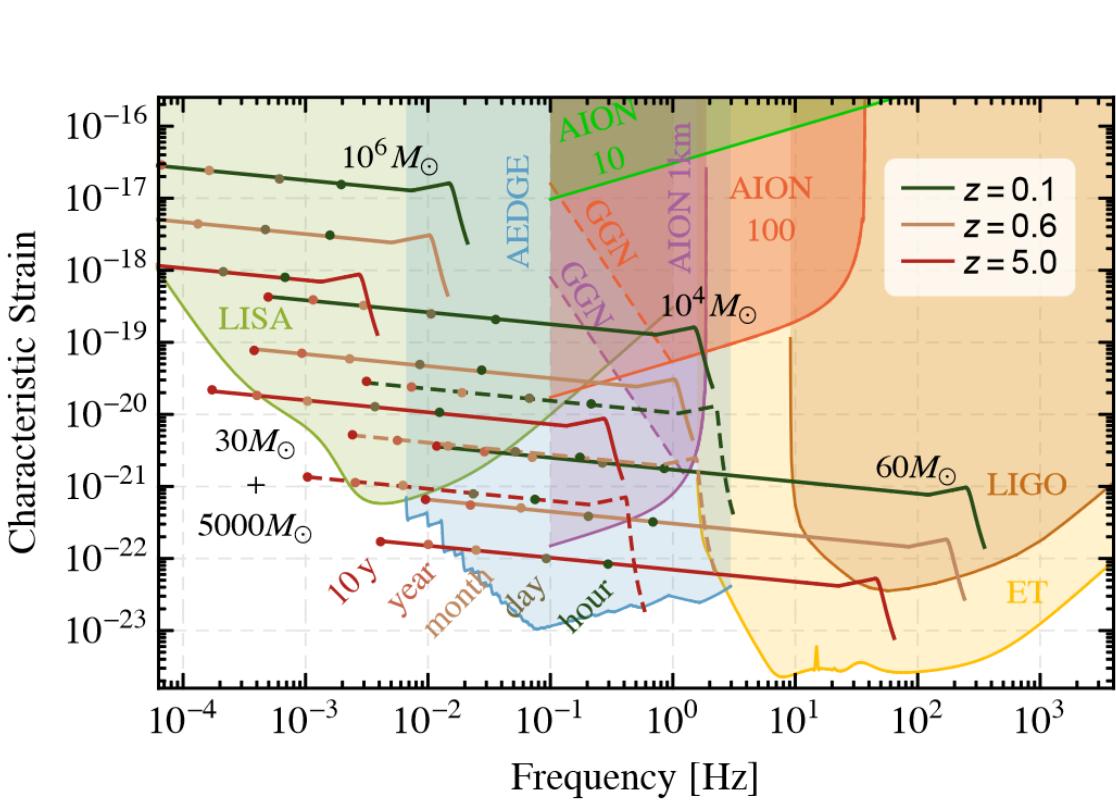
GW Gravitational
Waves

DM UL Dark
Matter



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Gravitational Waves and UL Dark Matter



- ❖ Explore mid-frequency band GW
- ❖ Precision sensor for ultra-light DM



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Particle Physics

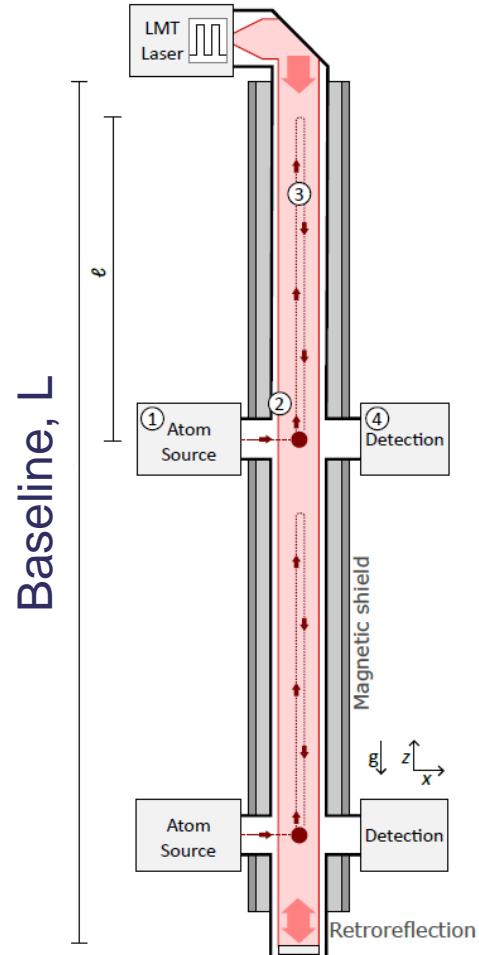


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Atom Interferometry Observatory & Network



Model Philosophy

AION-10 – testing and characterisation

AION-100 – 10^4 solar-mass black hole mergers

AION-km – neutron star binaries or black hole mergers, multi-messenger astronomy

L. Badurina *et al.*, AION: an atom interferometer observatory and network, JCAP05(2020)011



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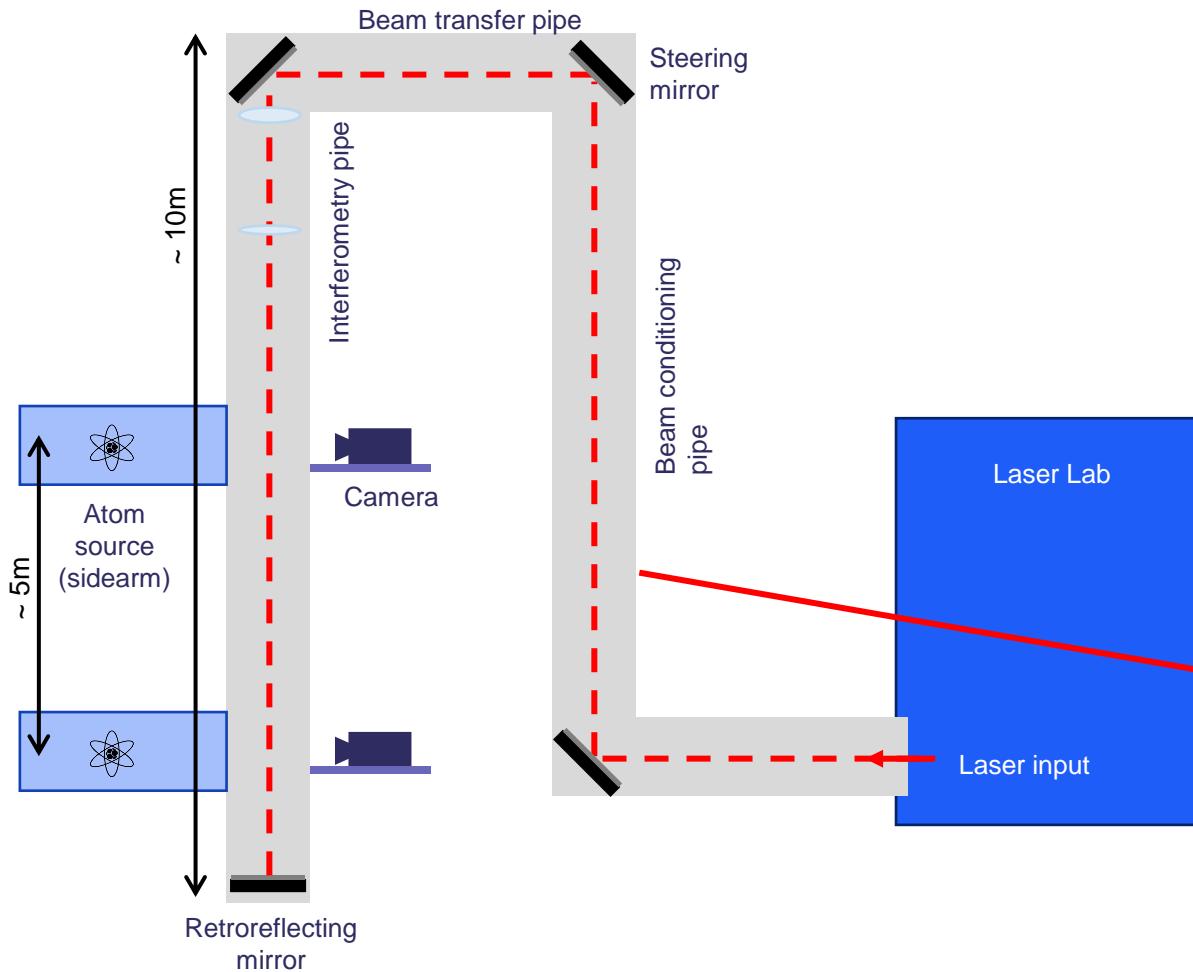
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Beecroft building, Oxford



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Gravitational Waves

Phase shift accumulated consists of:

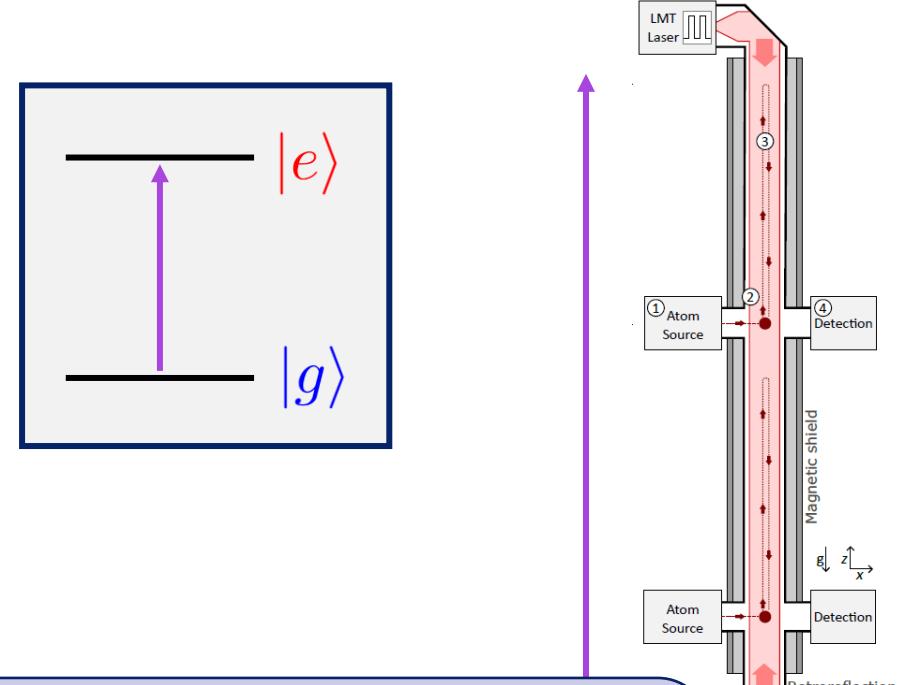
- **Propagation** phase along each path
- **Separation** phase at the end of the interferometer
- **Laser** phase at each node

$$\Delta\phi_{\text{total}} = \Delta\phi_{\text{propagation}} + \phi_{\text{separation}} + \phi_{\text{laser}}$$

- Each laser pulse imprints phase on the wavefunction
- Position of the atom (at the time of the pulse) is encoded in its wavefunction

$$\phi_{\text{laser}} = k \cdot x(t_0) - \omega t_0 + \phi$$

↑
Atom position



For a differential measurement:
Use the same laser pulse for both clouds... cancel the laser noise



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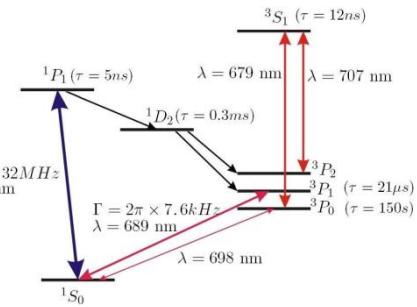
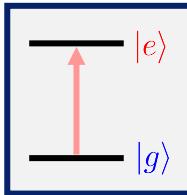
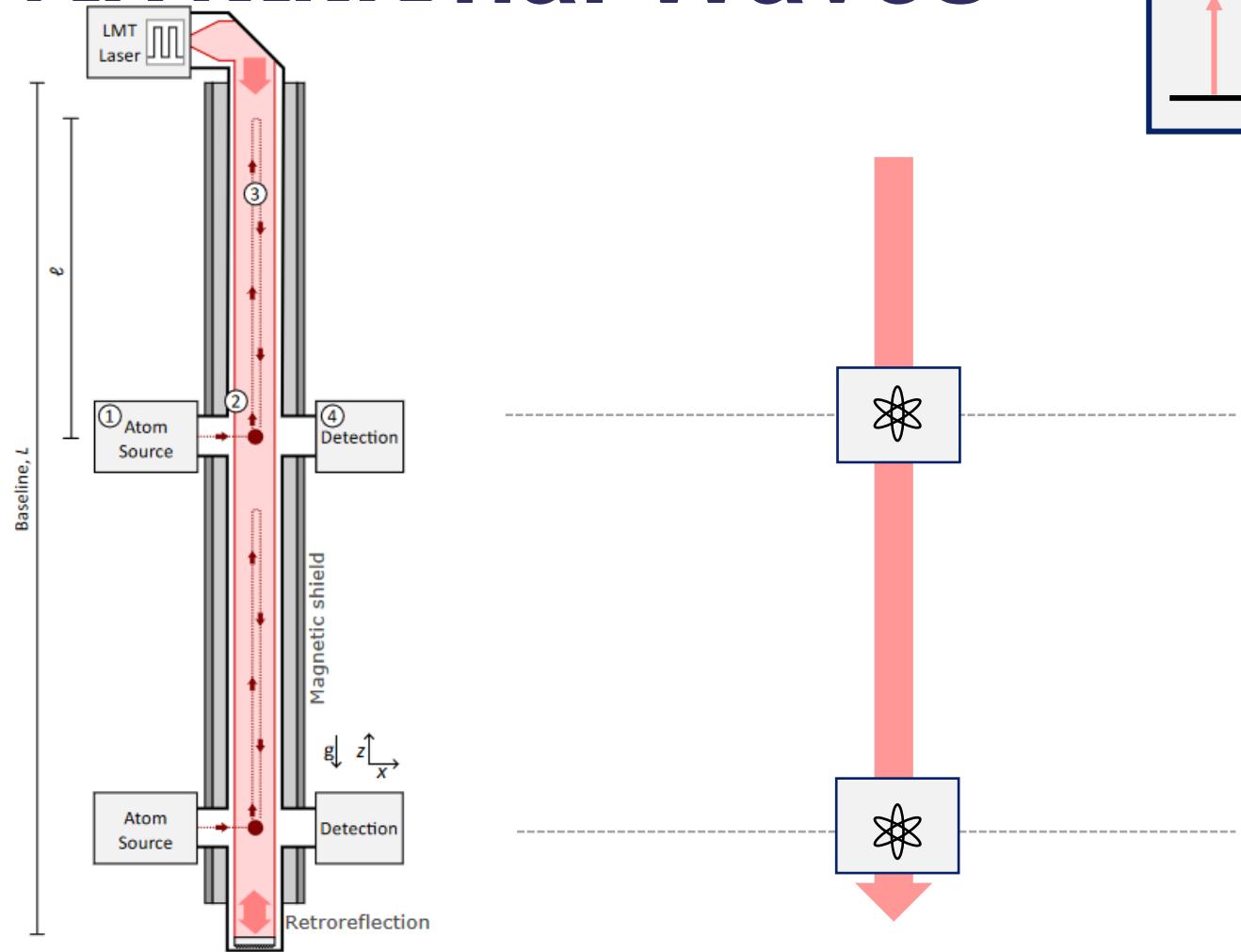
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Hogan, Johnson, & Kasevich (2008). Light-pulse atom interferometry. Proc. Int. School Phys. Enrico Fermi.

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Gravitational Waves



Sr atoms have a long-lived excited state or '*clock transition*'

✓ Can use single colour photons

✓ Using a common laser beam for both interferometers allows noise rejection



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Particle Physics



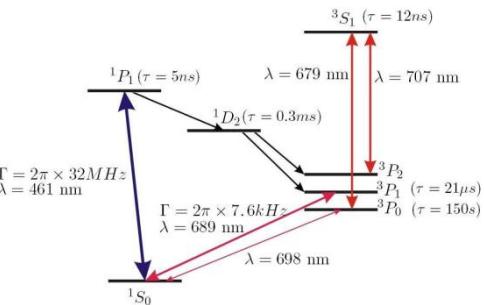
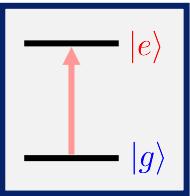
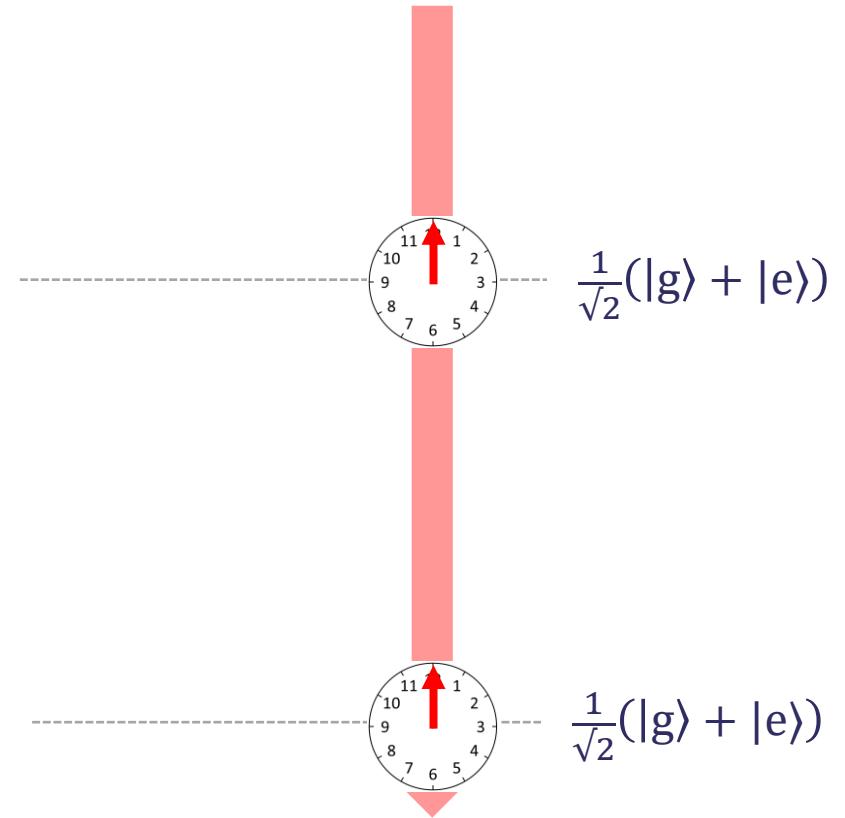
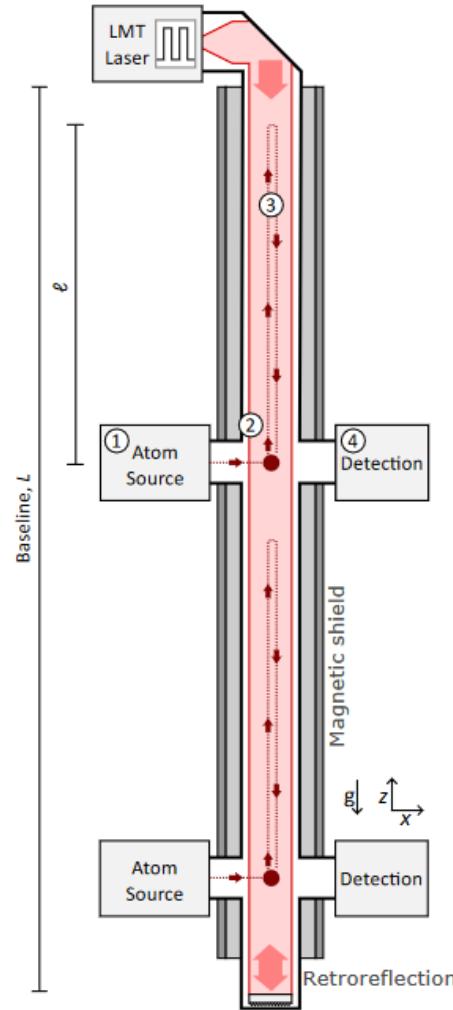
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Hogan, Johnson, & Kasevich (2008). Light-pulse atom interferometry. Proc. Int. School Phys. Enrico Fermi.

Gravitational Waves



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Particle Physics

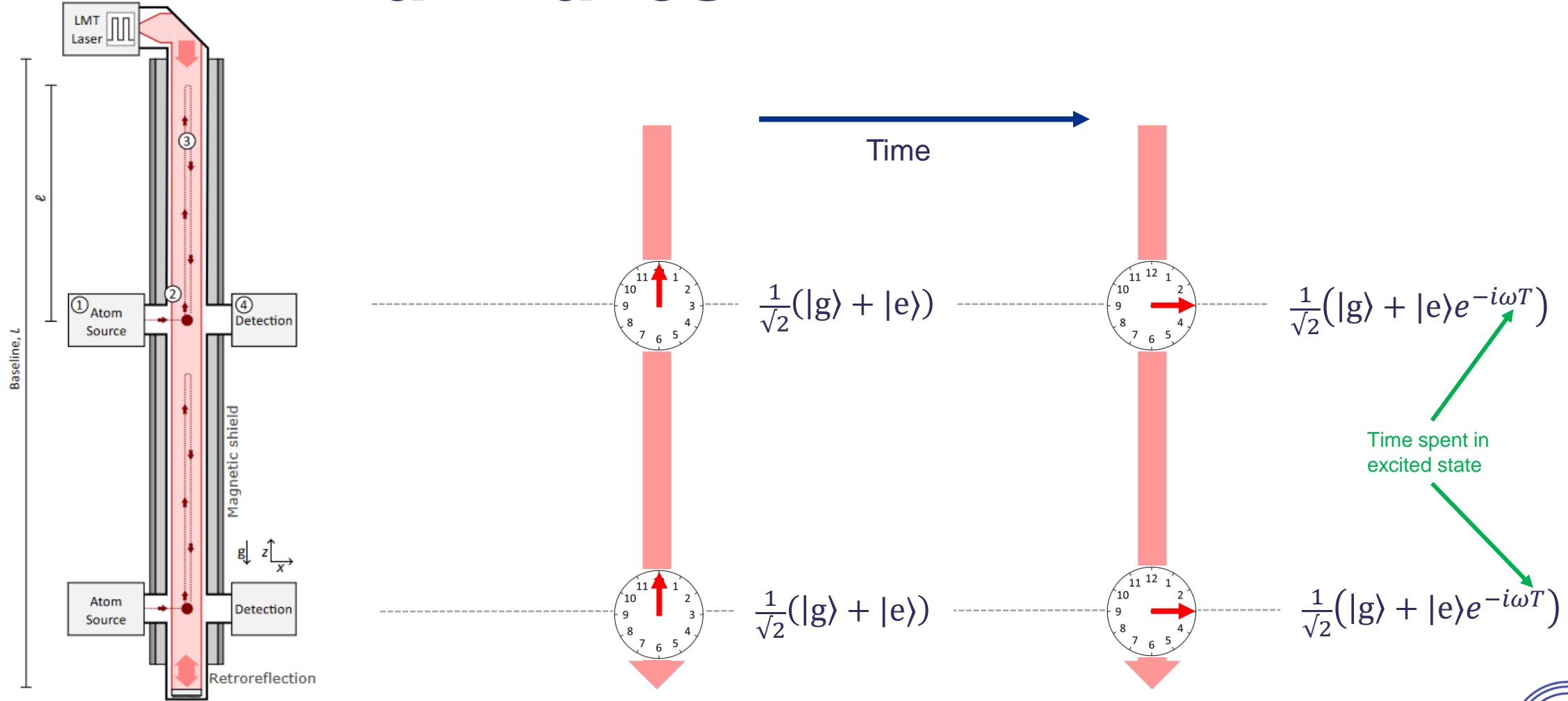


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Hogan, Johnson, & Kasevich (2008). Light-pulse atom interferometry. Proc. Int. School Phys. Enrico Fermi.

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Gravitational Waves



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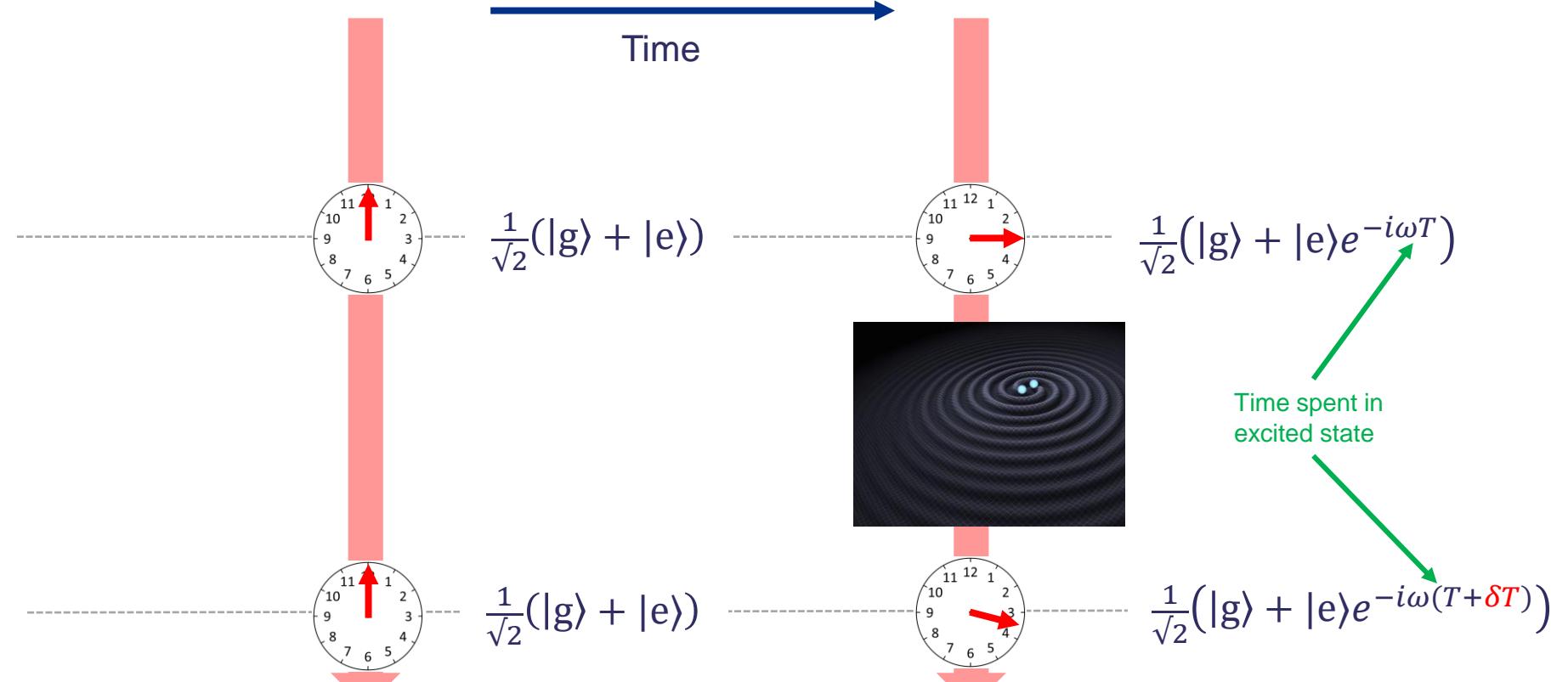
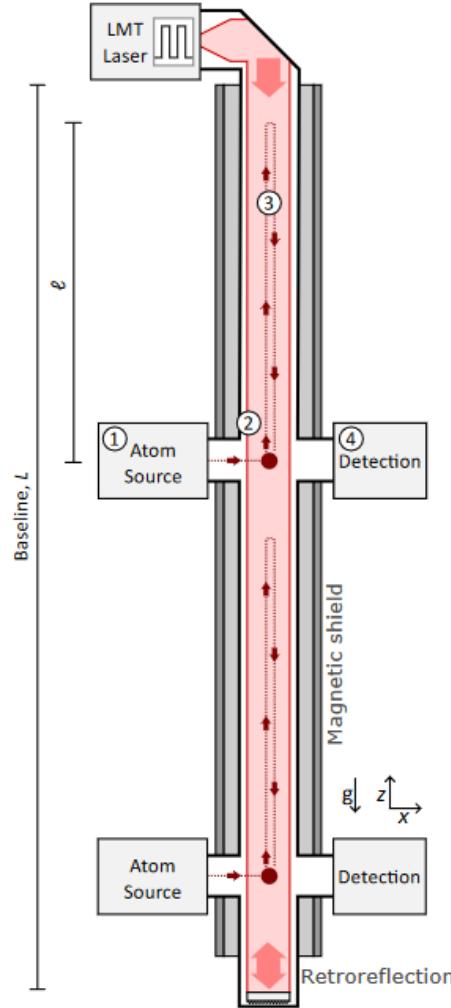
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Hogan, Johnson, & Kasevich (2008). Light-pulse atom interferometry. Proc. Int. School Phys. Enrico Fermi.

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Gravitational Waves



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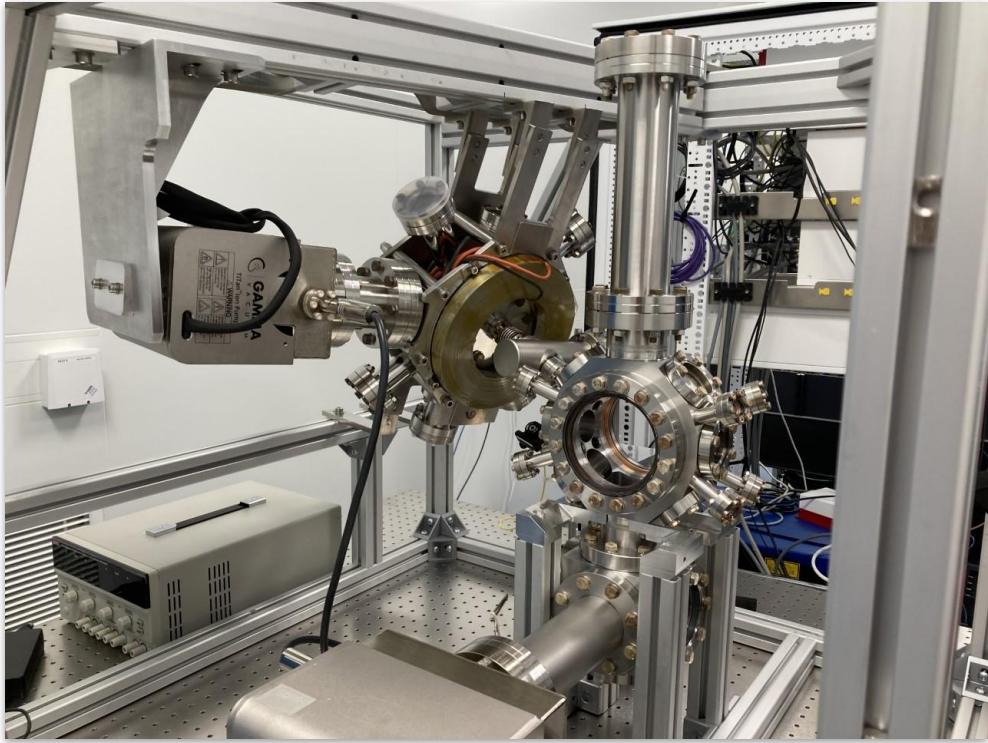


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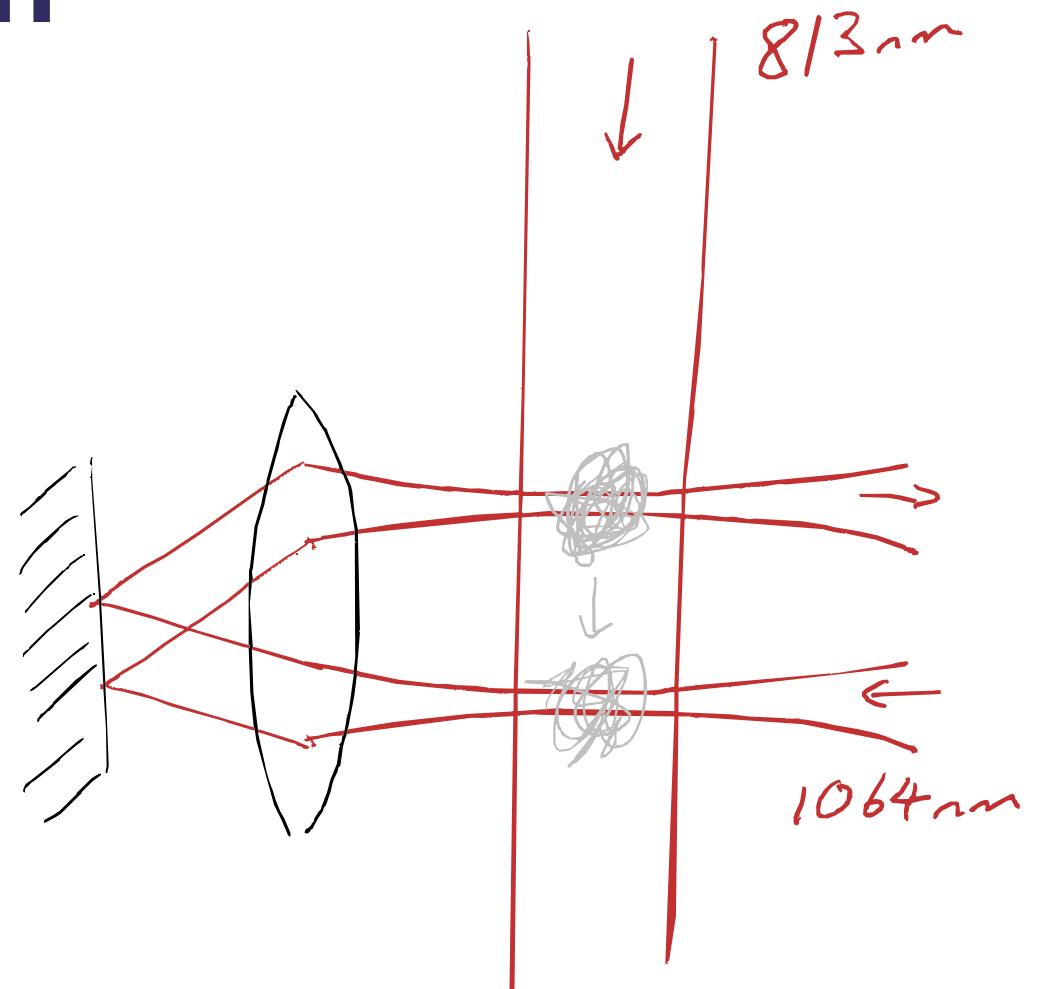
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Hogan, Johnson, & Kasevich (2008). Light-pulse atom
interferometry. Proc. Int. School Phys. Enrico Fermi.

Concept Demonstration



B. Stray et al., AVS Quantum Sci. 6, 014409 (2024)



Imperial College
London



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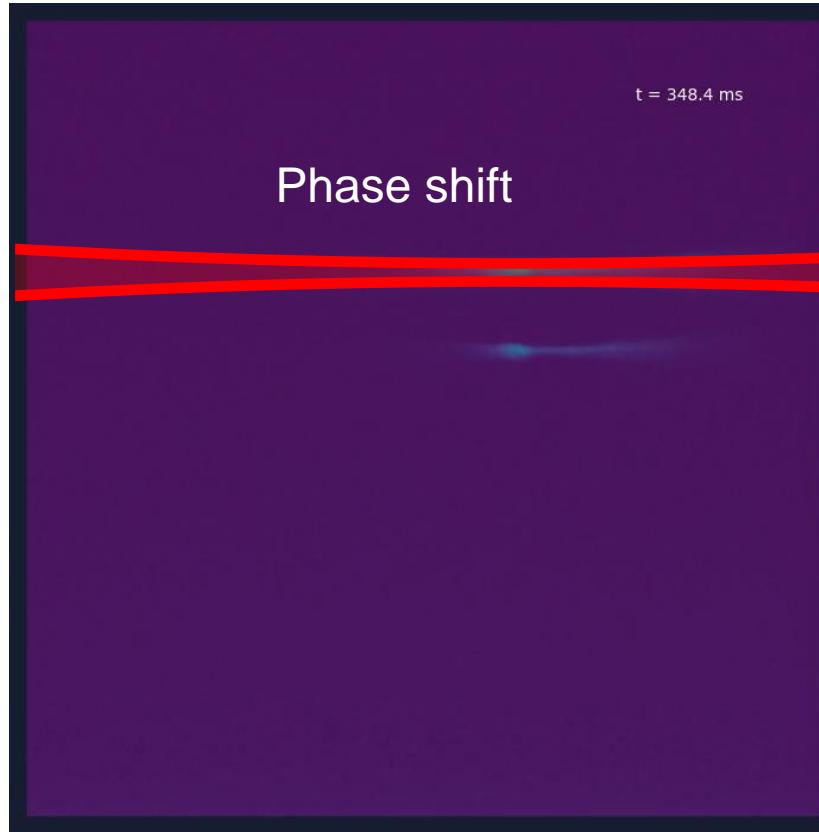
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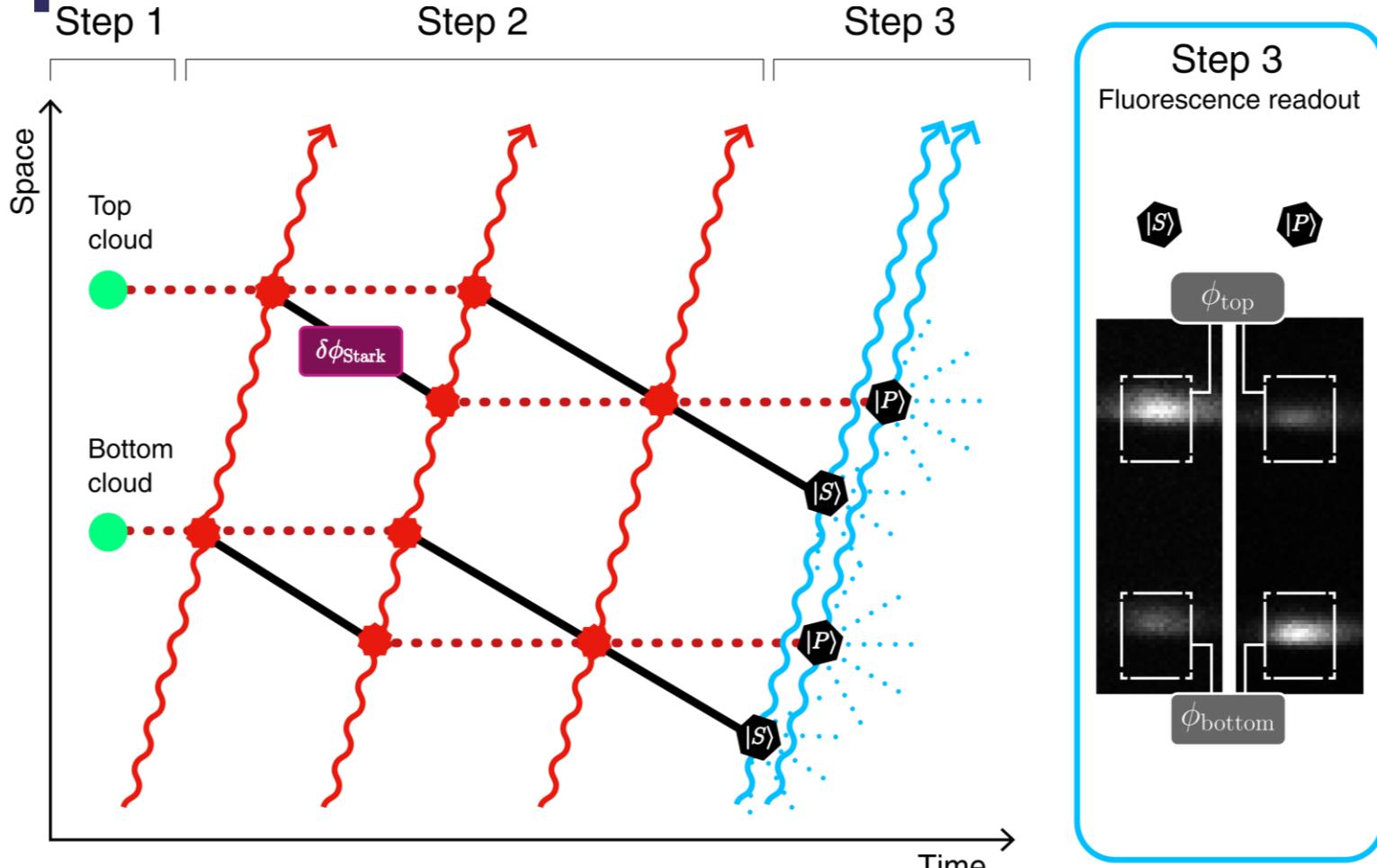
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Concept Demonstration



- 689nm Stark shifting pulse
- -80 MHz detuned from $F=9/2 \rightarrow F=11/2$
- 1mW in a 500um beam
- 30 us pulse creates a phase shift
 $\Delta\phi \approx 62^\circ$

Concept Demonstration



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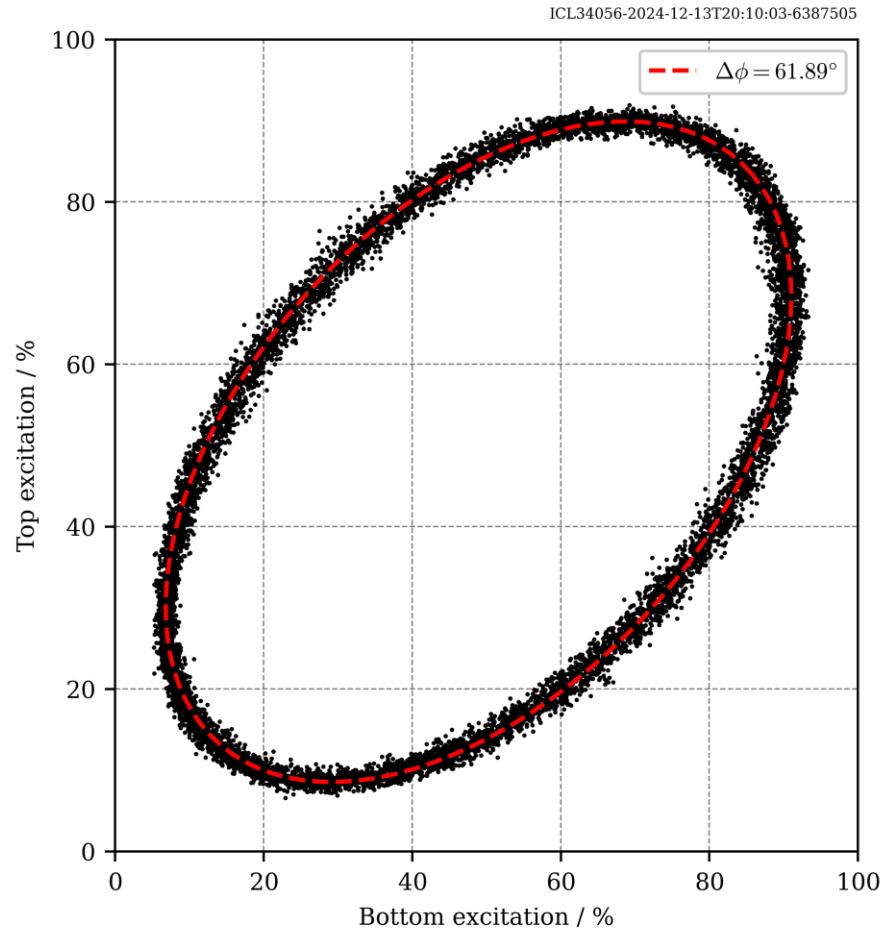
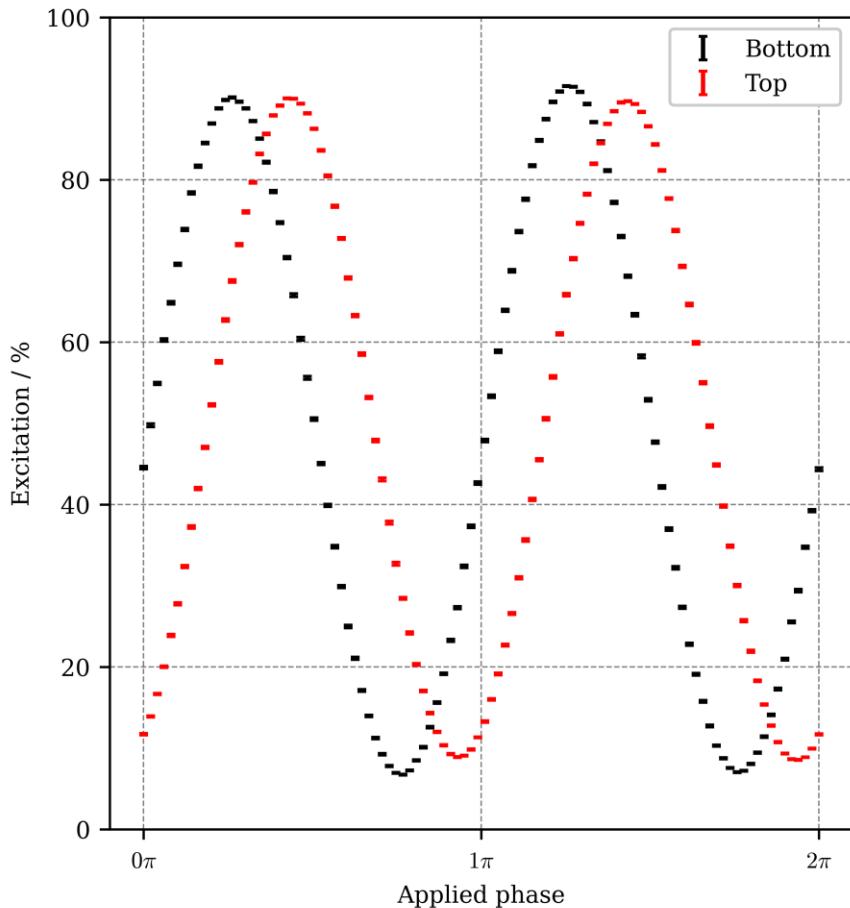
Technology

<http://arxiv.org/abs/2504.09158>

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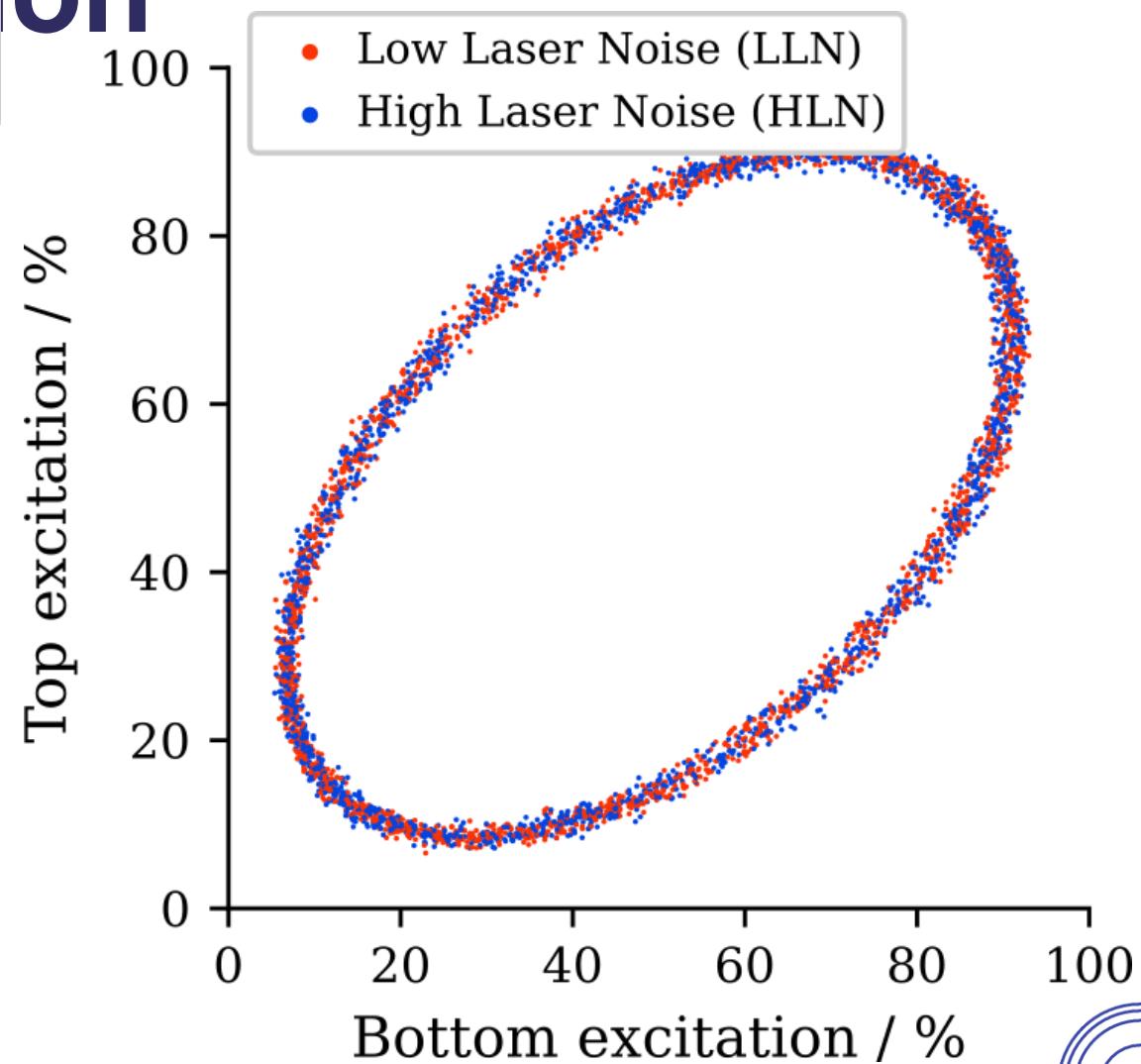
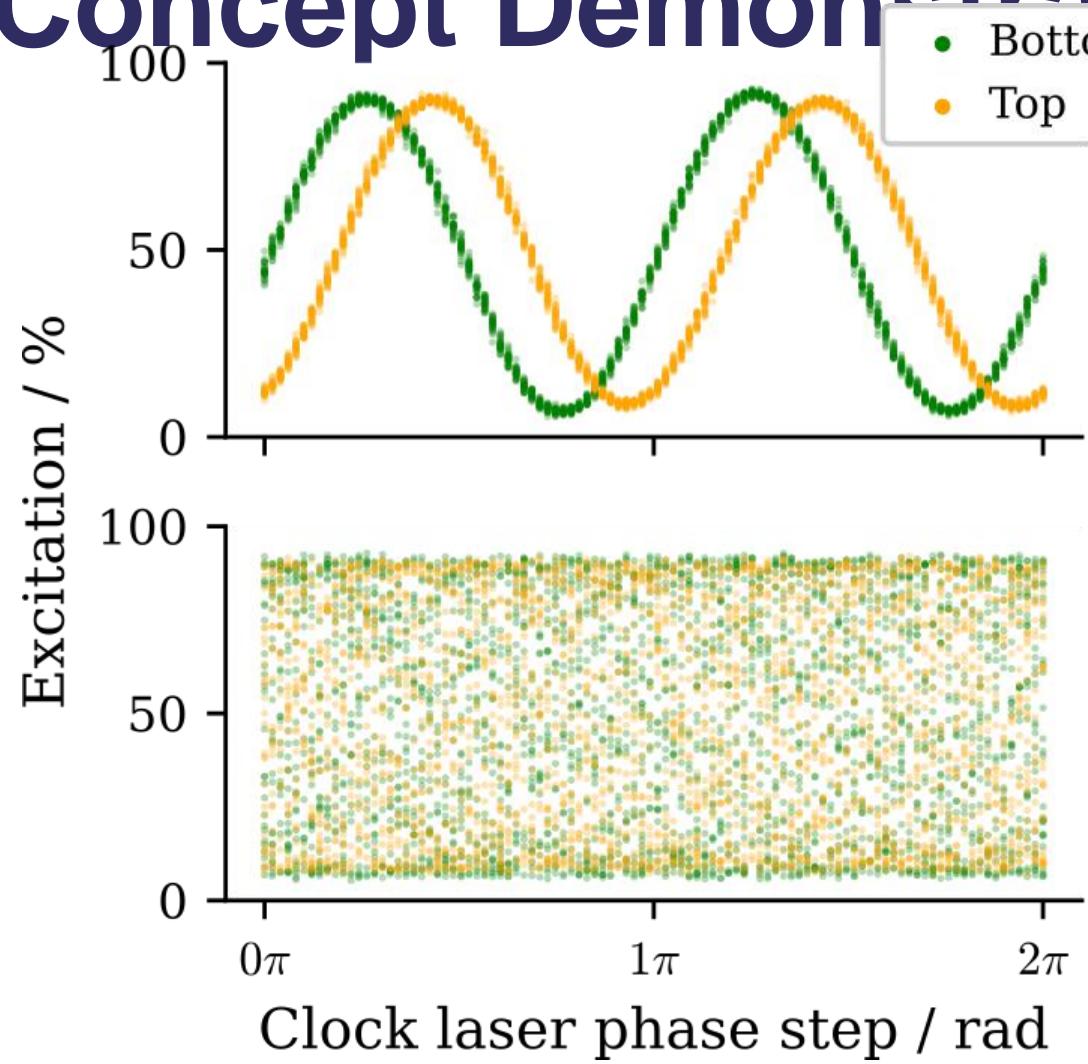
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Concept Demonstration



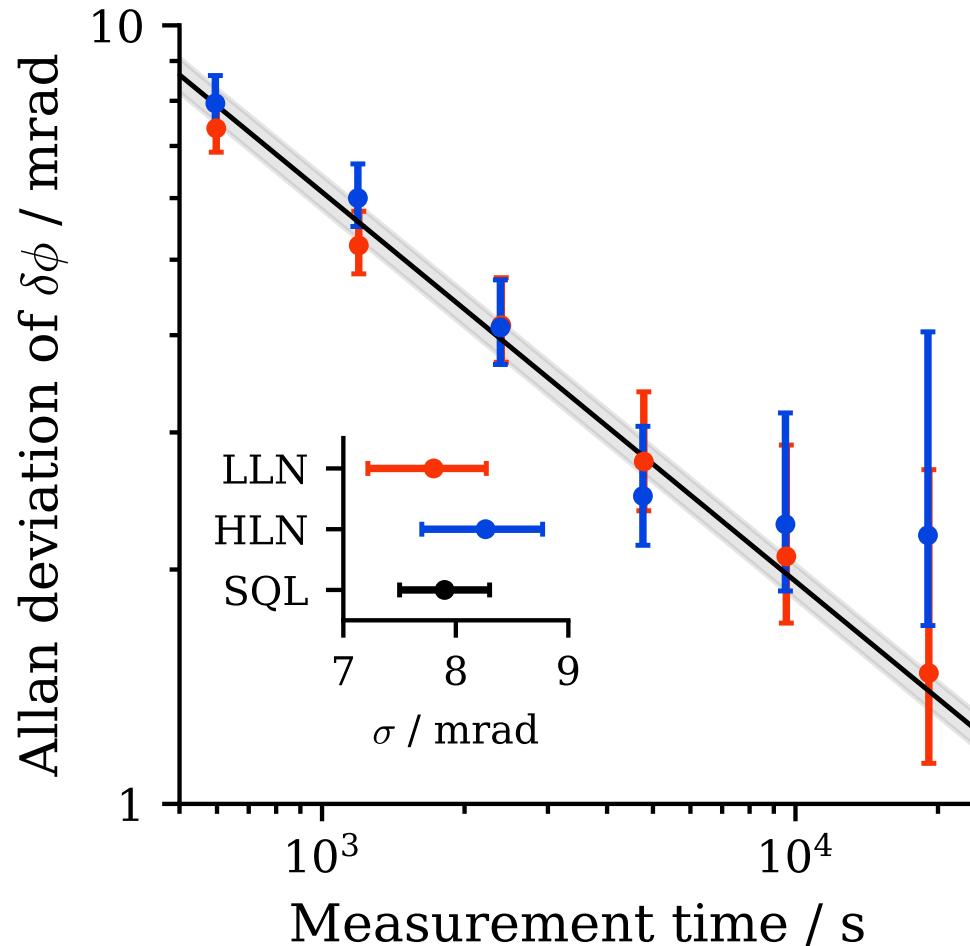
Concept Demonstration

<http://arxiv.org/abs/2504.09158>



Concept Demonstration

<http://arxiv.org/abs/2504.09158>



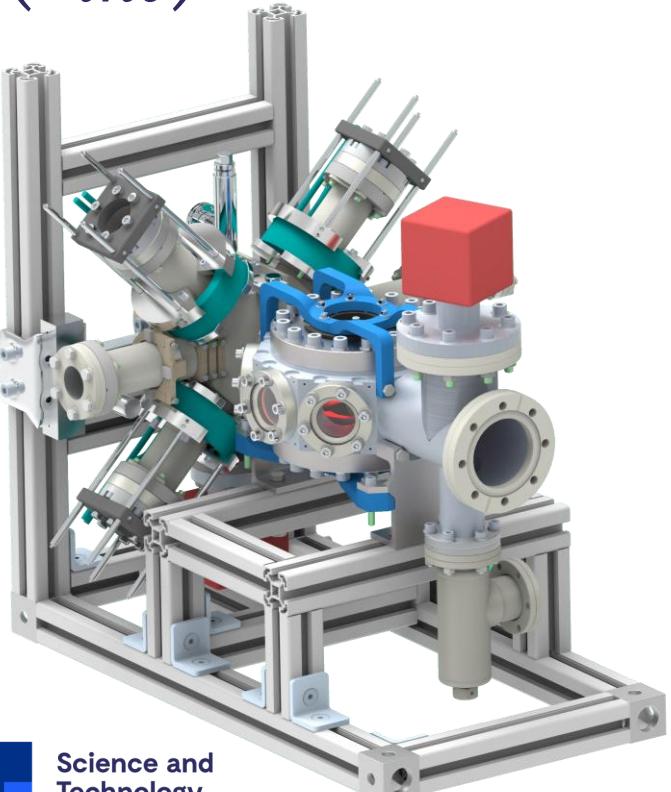
- 11200 samples
- Binned to 100 shots each
- Allan deviation
- ~ 1500 atoms/trap
- Standard deviations of $\delta\phi$ are consistent with the standard quantum limit in both the LLN and the HLN dataset:
 - $\sigma_{\text{LLN}} - \sigma_{\text{SQL}} = 0.10^{+0.55} - 0.66$ mrad
 - $\sigma_{\text{HLN}} - \sigma_{\text{SQL}} = 0.57^{+0.54} - 0.70$ mrad

Improving Sensitivity

$$d^{best} \sim \left(\frac{1}{T}\right)^{\frac{5}{4}} \frac{1}{Cn\Delta r} \left(\frac{\Delta t}{N_{atom}}\right)^{\frac{1}{2}} \left(\frac{1}{T_{int}}\right)^{\frac{1}{4}}$$

Badurina, Blas, McCabe, PRD 105, 023006 (2022)

- T Interferometer duration
- C Contrast
- n Number of LMT
- Δr Separation between interferometers
- Δt Time between interferometer sequences
- N_{atom} Number of atoms
- T_{int} Total integration time



-  More atoms
-  Colder atoms
-  Large momentum transfer
-  Reliability



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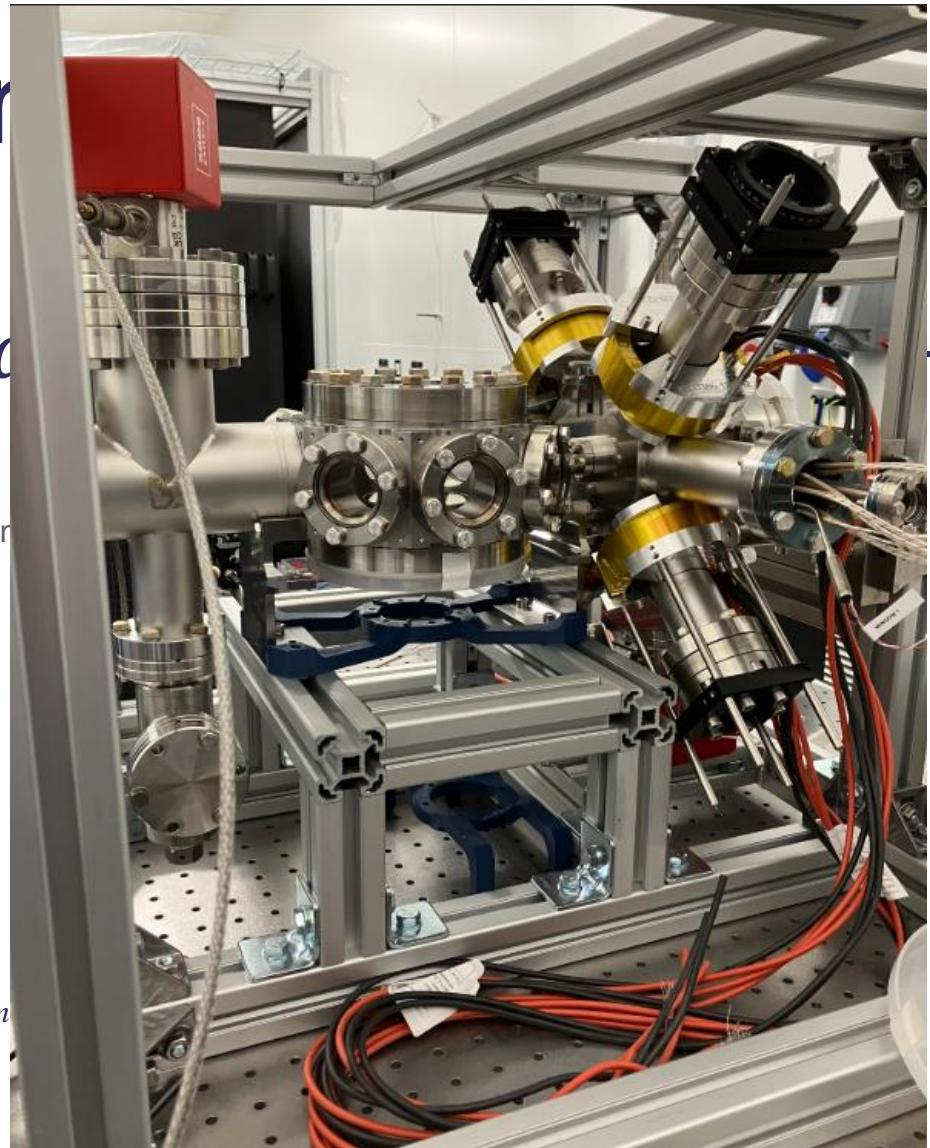
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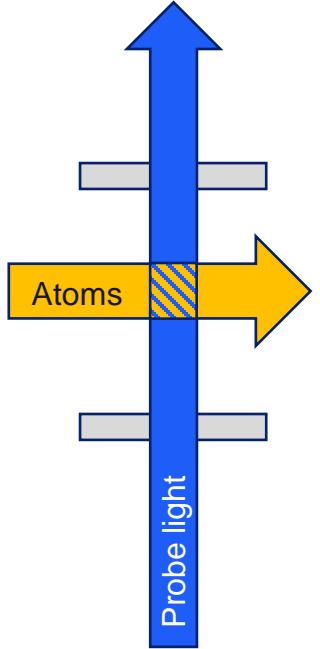
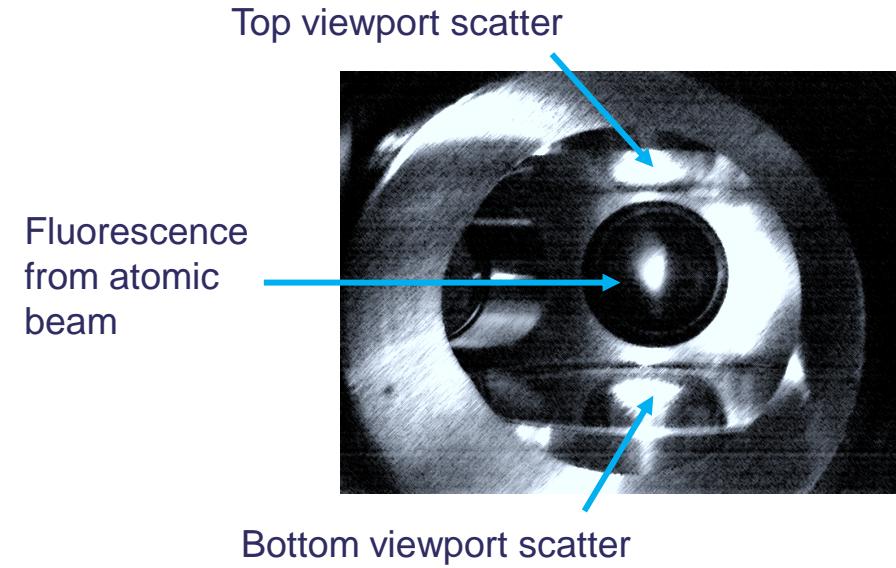
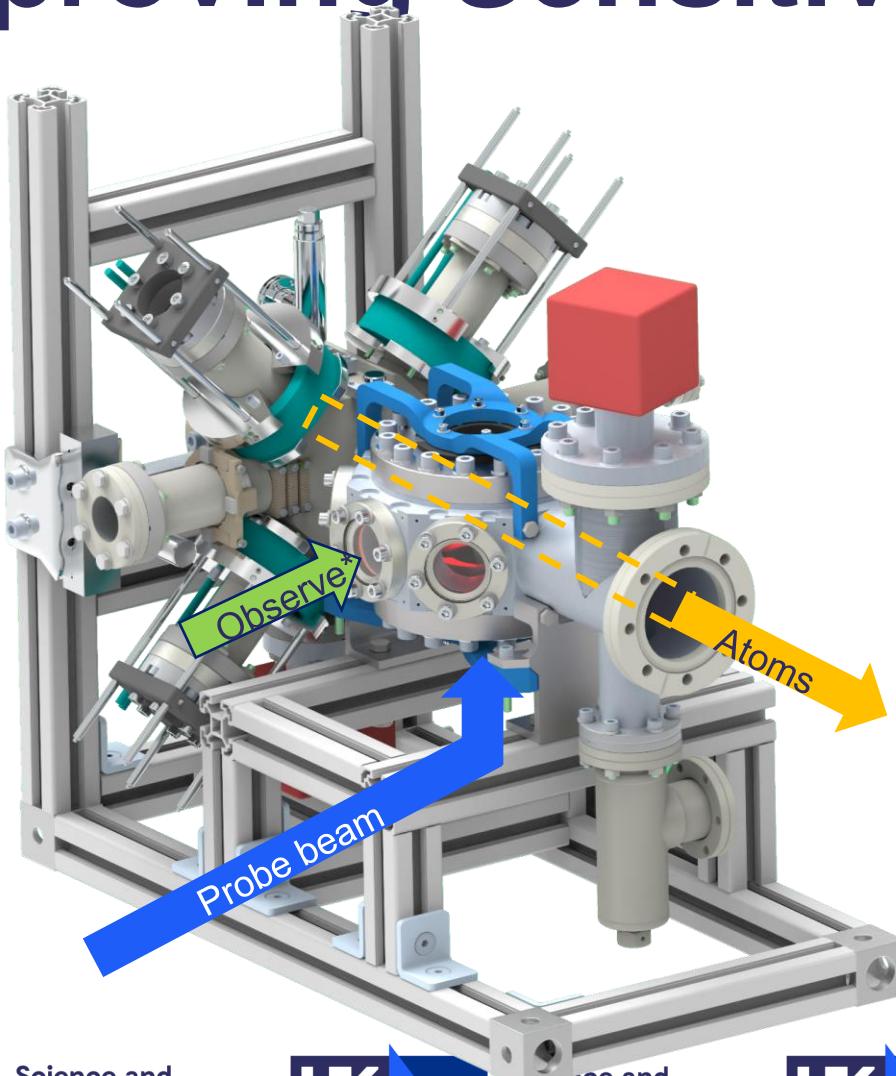
T
 C
 n
 Δr
 Δt
 N_{atom}
 T_{int}



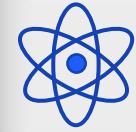
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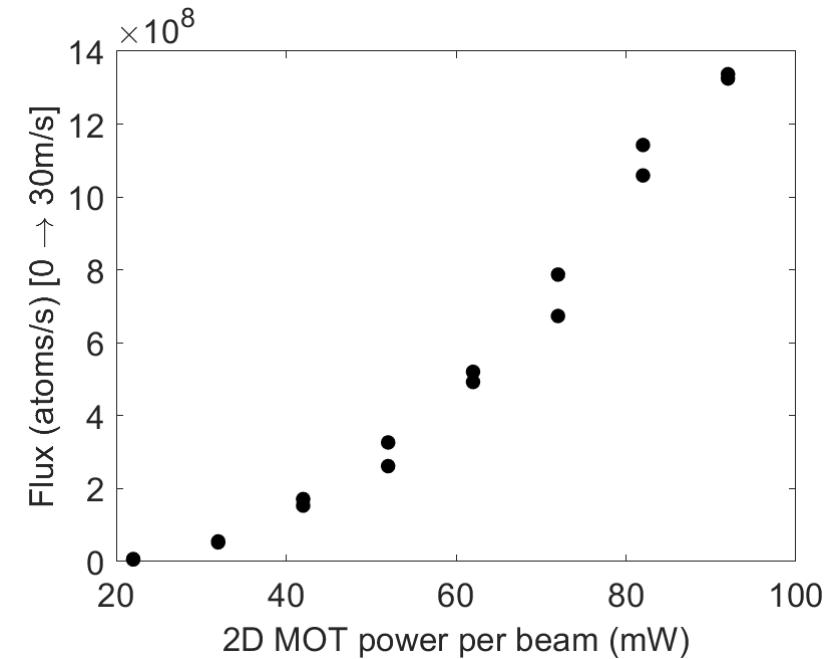
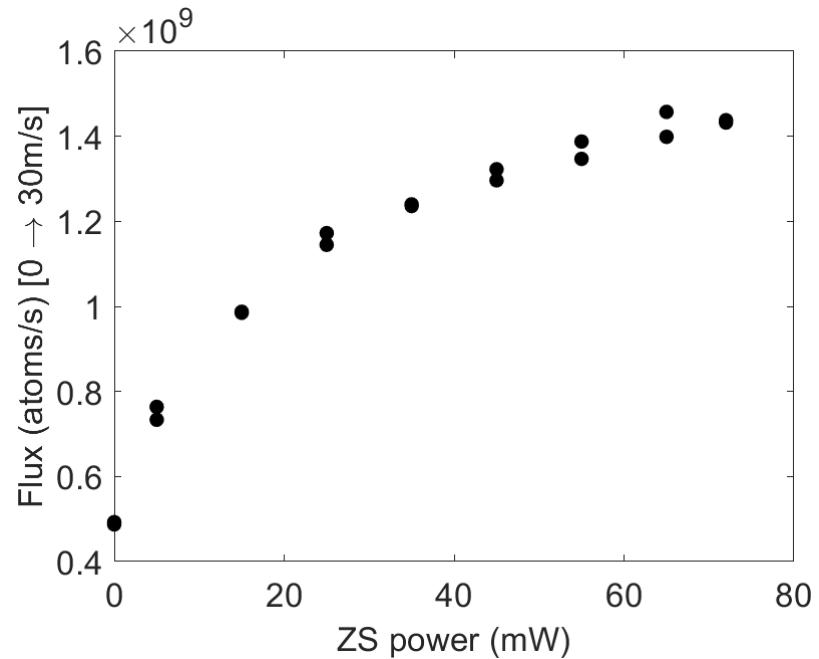
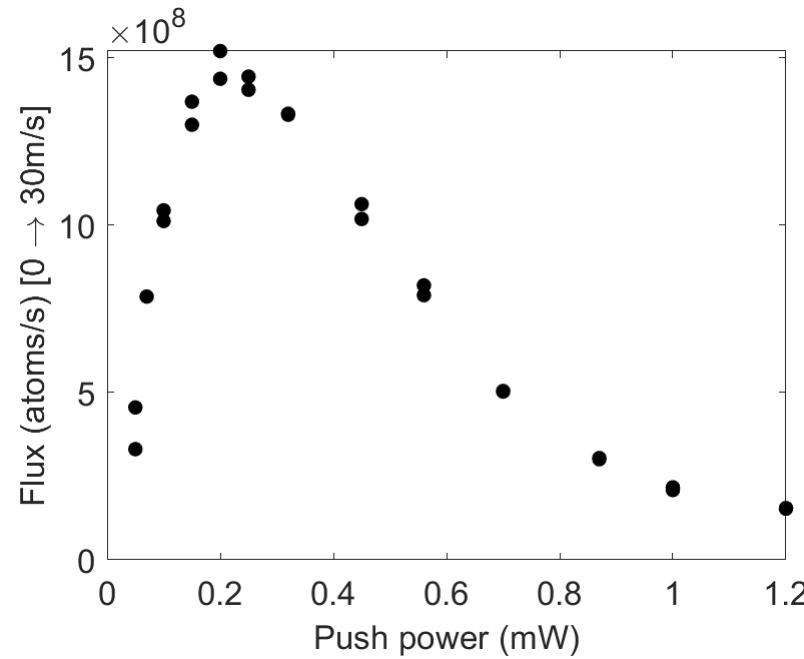
Improving Sensitivity



Improving Sensitivity



More atoms





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Thank you

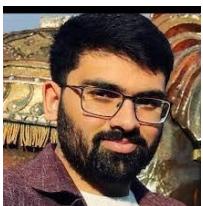
The Quantum Sensors and Tech team at RAL Space



Mark
Bason



Vicki
Henderson



Kamran
Hussain



Hamza
Labiad



Anna
Marchant

The R&D Electronics Engineering team at RAL Space



Adam
Filip



Mike
Salter



Jorn
Voegtli

Main Collaborators



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Other Collaborators

metamorphic
Imperial College
London

UNIVERSITY OF
BIRMINGHAM

UNIVERSITY OF
CAMBRIDGE

UNIVERSITY OF
OXFORD

UNIVERSITY OF
LIVERPOOL

University of
Nottingham
UK | CHINA | MALAYSIA

KING'S
COLLEGE
LONDON