Hunting the missing baryon number violation and beyond at the ESS



D. Milstead Stockholm University

Outline

- Why look for neutron oscillations?
- How to look for neutron oscillations
- Nnbar and HIBEAM at the ESS

Particle physics today in one slide

- The Standard Model is complete
- Neutrino masses are not understood.
- Other outstanding problems remain e.g. baryogenesis, dark matter...
- The LHC has arguably given the most important scientific result of the century (so far).
- Discoveries of fundamental importance are no longer assured (nor are they excluded) by exploring at higher collision energy.
- A strong need for a diverse field.

Baryon and lepton number violation

- BN,LN "accidental" SM symmetries at perturbative level
 - -BNV, LNV in SM non-perturbatively (eg instantons)
 - BNV Sakharov condition for baryogenesis
- BNV,LNV generic features of SM extensions
- Need to explore the possible selection rules:

$$\Delta B \neq 0$$
, $\Delta L = 0$, $\Delta [B - L] \neq 0$
 $\Delta B = 0$, $\Delta L \neq 0$, $\Delta [B - L] \neq 0$
 $\Delta L \neq 0$, $\Delta B \neq 0$, $\Delta [B - L] = 0$

An experimentalist's view

Decay mode Partial mean life (x 10³⁰ yrs)

$N \rightarrow e^+\pi^-$	> 2000 (n), > 8200 (p)
$N ightarrow \mu^+ \pi$ $N ightarrow u \pi$	> 1000 (n), > 6600 (p)
$N \rightarrow \nu \pi$	> 1100 (n), > 390 (p)
$\rho \rightarrow e^+ \eta$ (DDD)	> 4200
$\stackrel{p \to e^+\eta}{\underset{p \to \mu^+\eta}{}}$ (RPP)	> 1300
$n ightarrow u \eta $	> 158
$N \rightarrow e^+ \rho$	> 217 (n), > 710 (p)
$N \rightarrow \mu^{+} \rho$	> 228 (n), > 160 (p)
$N \rightarrow \nu \rho$	> 19 (n). > 162 (p)
$p \mapsto e^+\omega$	> 320
$ \rho \rightarrow e^+ \omega $ $ \rho \rightarrow \mu^+ \omega $ $ \eta \rightarrow \nu \omega $	> 780
$n\mapsto u\omega $	> 108
$N \rightarrow e^+ K$	> 17 (n), > 1000 (p)
$N \rightarrow \mu^+ K$	> 26 (n), > 1600 (p)
$N \rightarrow \nu K$	> 86 (n), > 5900 (p)
$n \rightarrow \nu K_S^0$	> 260
$p \to e^+ K^* (892)^0$	> 84
$N \rightarrow \nu K^*(892)$	>78 (n), >51 (p)
$p \rightarrow e^{+}\pi^{+}\pi^{-}$ $p \rightarrow e^{+}\pi^{0}\pi^{0}$	> 82
	> 147
$ ho ightarrow ho^+ \pi^- \pi^0$	> 52
$ ho ightarrow \ \mu^+ \pi^+ \pi^+$	> 133
$ ho \mapsto \stackrel{'}{\mu}^{+} \pi^{0} \pi^{0}$	> 101
$n \rightarrow \mu^+\pi^-\pi^0$	> 74
$n ightarrow e^+ K_\perp^0 \pi^-$	> 18
$n \rightarrow : e^- \pi^+$ $n \rightarrow : \mu^- \pi^+$	> 65
	> 49
$n \mapsto e^- \rho^+$ $n \mapsto \mu^- \rho^+$	> 62 > 7
	> 32
$n \rightarrow e^- K^+$ $n \rightarrow \mu^- K^+$	> 57
$p \rightarrow e^{-}\pi^{+}\pi^{+}$	> 30
$n \rightarrow e^- \pi^+ \pi^0$	> 29
$p \rightarrow \mu^+ \pi^+ \pi^+$	> 17
$n \rightarrow \mu - \pi + \pi^0$	> 34
$p \rightarrow e^- \pi^+ K^+$	> 75
$p \rightarrow \mu^- \pi^+ K^+$	> 245

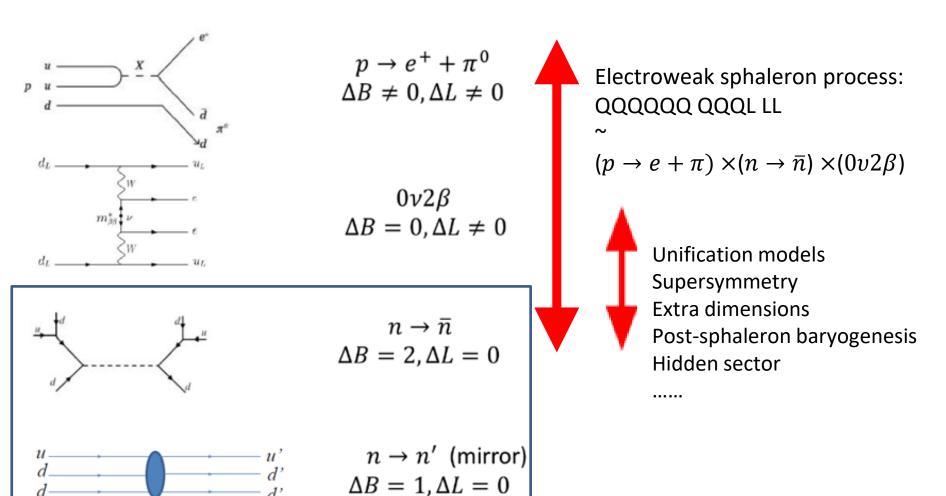
$p \rightarrow 1$	et γ	> 670
$p \mapsto 1$	$\mu^{+}\gamma$	> 478
	$ u\gamma$	> 28
$\square p \mapsto$	$e^+ \gamma \gamma$	> 100
$n \rightarrow 1$	$ u\gamma\gamma$	> 219
p	e+e+e+	> 793
$p \rightarrow 1$	$e^+\mu^+\mu^-$	> 359
$p \mapsto 1$	$e^+ u u$	> 170
$n \mapsto 1$	$e^+e^- u$	> 257
$n \rightarrow 1$	$\mu^+ \dot{e}^+ u $	> 83
$n \rightarrow 1$	$\mu^+\mu^- u$	> 79
$p \rightarrow$	μ+e+e	529
p	$\mu^+\mu^+\mu^-$	> 675
	$\mu^+ u u$	> 220
	$e^{-}\mu^{+}\mu^{+}$	> 6
$ \stackrel{n}{N} $	3 u	> 0.0005
· · · N · → ·	e ⁺ anything	> 0.6 (n, p)
$ N \rightarrow $	μ^+ anything	$ \cdot $ > 12 (n, p)
:N →:	e ⁺ π ⁰ anything	\Rightarrow 0.6 (h, p)
$pp \rightarrow$	π ⁺ π ⁺ _+ _0	> 0.7
$pn \rightarrow$	$\pi^{+}\pi^{0}$ $\pi^{+}\pi^{-}$	> 2 > 0.7
	$\pi^0 \pi^0$	> 0.7
nn →	K+K+	> 3.4
ייי טע		> 5.8
	e+μ+:	> 3.6
	μ^{+}	> 1.7
	$e^+\overline{ u}$	> 2.8
	$\mu^{+}\overline{\nu}$	> 1.6
pn →		> 1.0
	$ au^+\overline{ u}_{ au}$	> 1.0
nn →	$ u_e \overline{ u}_e$	

$$\Delta B \neq 0, \Delta L \neq 0$$

$$\Delta B \neq 0, \Delta L = 0$$

Few searches for $\Delta B \neq 0$, $\Delta L = 0$

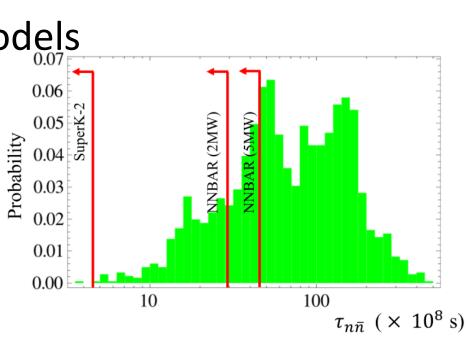
Candidate BNV,LNV processes



Neutron oscillations are a key part of the landscape of new physics Symbiosis with other processes

Neutron-antineutron oscillations

- R-parity violating supersymmetry
- Extra dimensions
- Post-sphaleron baryogenesis d
- Left-right symmetric models
- etc. etc.



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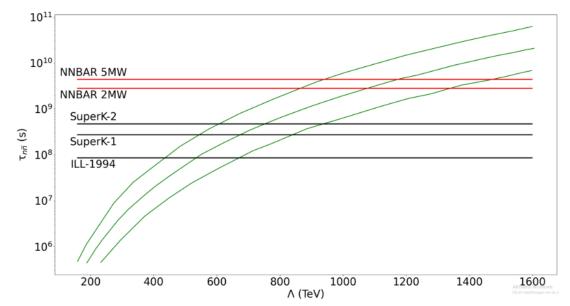
Energy scale of new physics for

$$n \to \overline{n}$$
 ?

Simple EFT approach
Dimension 9 (6 quark) operator

$$\mathcal{O}_{\Delta\mathcal{B}=2} = \frac{1}{\mathcal{M}^5} (udd)^2 + \text{h.c.}$$

$$\varepsilon_{n\bar{n}} = \frac{C\Lambda_{\text{QCD}}^6}{\sqrt{5}}$$



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A search for $n \to \overline{n}$ probes the PeV scale for new physics.

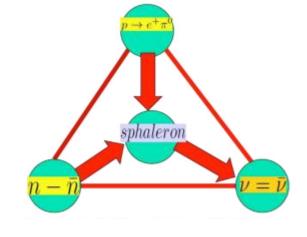
A combination of BSM signatures in the SM

Symbiosis between proton decay, neutrinoless double beta decay and $n \to \bar{n}$ in the Standard Model:

Electroweak sphaleron process:

QQQQQQ QQQL LL

$$(p \rightarrow e + \pi) \times (n \rightarrow \bar{n}) \times (0v2\beta)$$



Observation of two processes implies the existence of the other one.

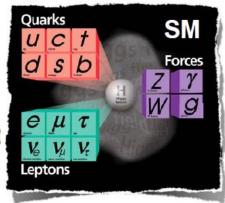
Sterile neutrons

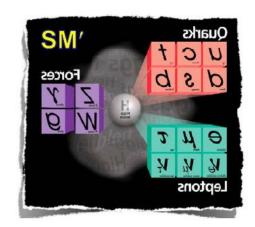
"Hidden/mirror" sector

Restores parity symmetry.

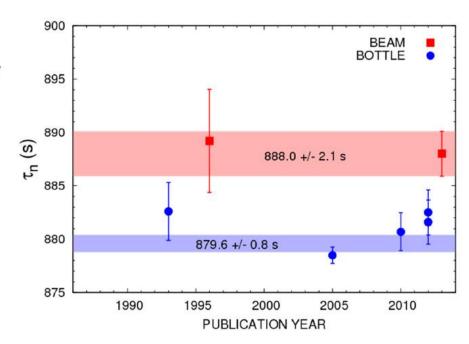
Possible mixing for Q = 0 particles, eg, $n \rightarrow n'$

Mirror matter : dark matter candidates (m< 10 GeV)





Can explain 5σ neutron lifetime discrepancy seen in bottle and beam experiments.



Outline

- Why neutron oscillations?
- How to look for neutron oscillations
- Nnbar and HIBEAM at the ESS

$n \rightarrow \bar{n}$ mixing formalism



$$i\hbar \frac{\partial}{\partial t} {n \choose \bar{n}} = {E_n \quad \delta m \choose \delta m \quad E_{\bar{n}}} {n \choose \bar{n}}$$

$$\delta m = \langle \bar{n} | H_{eff} | n \rangle < 10^{-29} \text{ MeV} = n\bar{n} \text{ mixing physics}$$

$$P_{n \to \bar{n}} = \left(\frac{\delta m}{\Delta E}\right)^2 \sin^2(\Delta E \times t) \quad ; \Delta E = E_n - E_{\bar{n}}$$

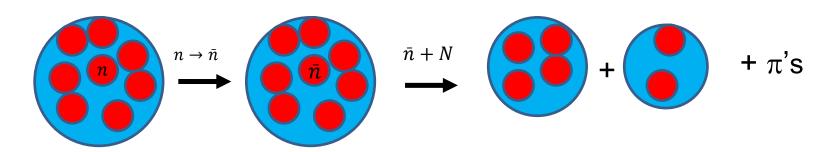
Two interesting cases:

- Free neutron oscillation: $\Delta E \times t \ll 1 \Rightarrow P \sim (\delta m \times t)^2$
- Bound neutron oscillation: $\Delta E \times t \gg 1$

Quasi-free limit : $\Delta Et \sim 1 \Rightarrow P \sim (\delta m \times t)^2$

Searching with bound neutrons

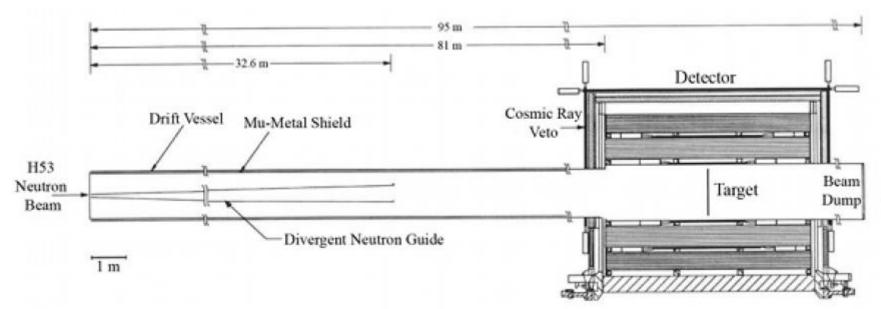
Nuclear disintegration after neutron oscillation



$$P_{n \to \bar{n}} = \left(\frac{\delta m}{\Delta E}\right)^2 \sin^2(\Delta E \times t)$$
,
 $\Delta E \sim 10 - 100 \text{ MeV}$.
 $\Rightarrow \text{Suppression: } \left(\frac{\delta m}{\Delta E}\right)^2 < 10^{-60}$

Best current limits (SuperKamiokande) $\Rightarrow \tau_{free} > 4.7 \times 10^8 \text{ s}$ Irreducible bg's prevent large improvements. Model—dependent (nuclear interactions).

Free neutron search at ILL



Institute Laue—Langevin (Early 1990's). Cold neutron beam from 58MW reactor.

 $\sim 130 \mu m$ thick carbon target 100m propagation in field-free region

Signal of at least two tracks with E > 850 MeV 0 candidate events, 0 background. $\Rightarrow \tau_{n \to \bar{n}} > 0.86 \times 10^8 \text{s}.$

Outline

Why neutron oscillations?



How to look for neutron oscillations



NNbar and HIBEAM at the ESS

The European Spallation Source

High intensity spallation neutron source

Multidisplinary research centre with 17 European nations participating.

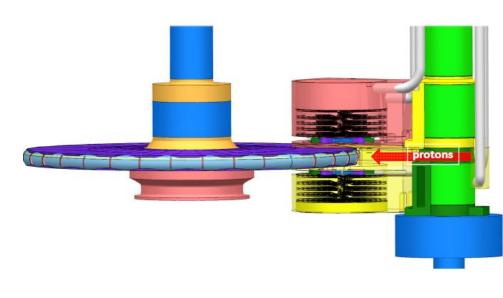
Lund, Sweden. Start operations in 2027/2028.

Up to 2 GeV protons (3ms long pulse, 14 Hz) hit rotating tungsten target.

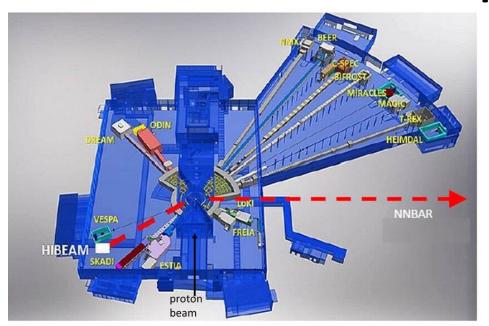
Cold neutrons after interaction with moderators.

15 beamlines/instruments – none are Swedish-led





Beamlines and program





R&D Annihilation detector prototype Conceptual design reports for HIBEAM/NNBAR

TDRs and small scale experiment

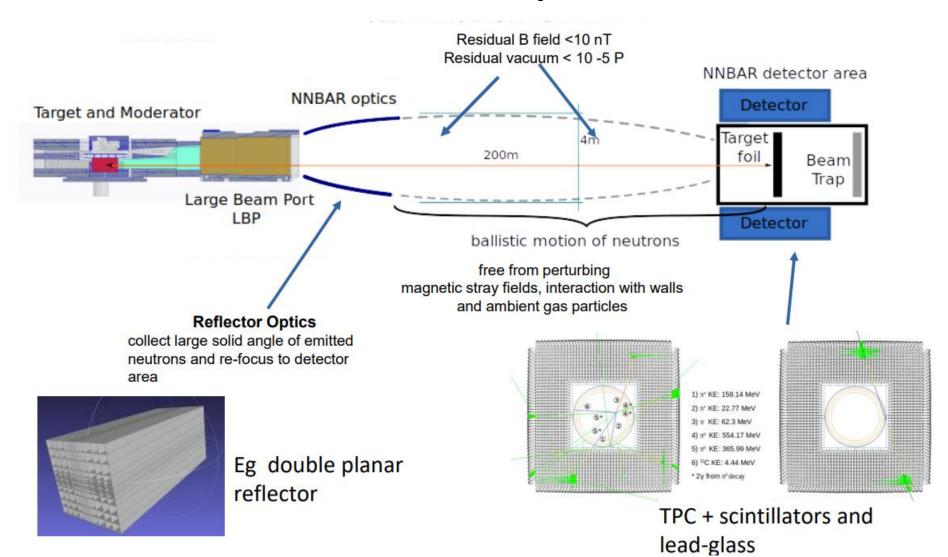
HIBEAM

High precision induced: $n \to n', \ n \to \overline{n} \ \ (\text{x10 improvement})$ First search for free $n \to \overline{n}$ at a spallation source Eg at upgraded test beamline

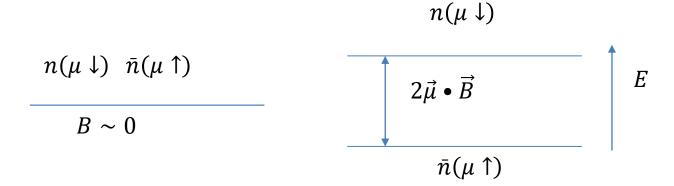
NNBAR High sensitivity free $n \to \bar{n}$ (x1000 improvement) At the Large Beam Port

NNBAR

The NNBAR Experiment



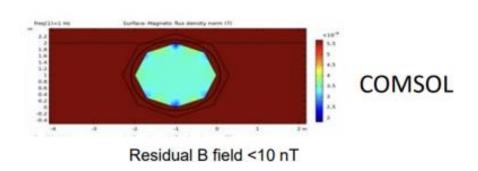
The need for magnetic shielding



Degeneracy of n, \bar{n} broken in B—field due to dipole interactions: $\Delta E = 2\vec{\mu} \cdot \vec{B}$

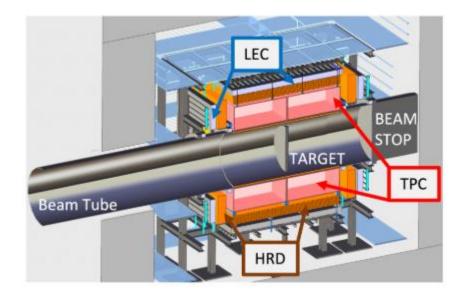
Flight time ≤ 1 s For quasi-free condition $\Delta E \times t \ll 1$ $\Rightarrow B \leq 10$ nT and vacuum $\leq 10^{-5}$ Pa.

Outer and inner octagon-shaped passive shield of 1-2 mm thick sheets of mumetal.

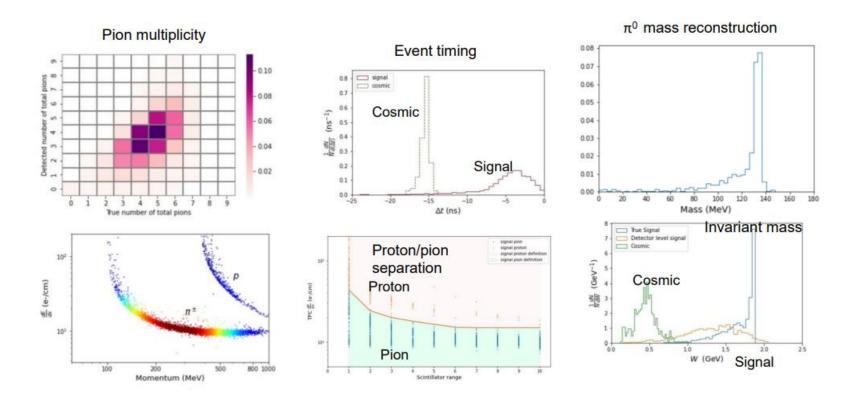


NNBAR detector

- ~2 GeV invariant mass pionic final state
- Lead-glass em calo
- Scintillator staves
- TPC
- Cosmic veto (scintillators)



Geant-4 detector simulation



A Computing and Detector Simulation Framework for the HIBEAM/NNBAR Experimental Program at the ESS

Ioshua Barrow^{10,11}, Gustaaf Brooijmans², José Ignacio Marquez Damian³, Douglas DiJulio³, Katherine Dunne⁴, Elena Golubeva⁵, Yuri Kamyshkov¹, Thomas Kittelmann³, Esben Klinkby⁸, Zsófi Kókai³, Jan Makkinje², Bernhard Meirose^{4,6,*}, David Milstead⁴, André Nepomuceno⁷, Anders Oskarsson⁶, Kemal Ramic³, Nicola Rizzi⁸, Valentina Santoro³, Samuel Silverstein⁴, Alan Takibayev³, Richard Wagner⁹, Sze-Chun Yiu⁴, Luca Zanini³, and



Symmetry 14 (2022) 1, 76

Backgrounds

- Cosmic rays (neutral and charged dominant at ILL)
- Thermal neutrons, beta-delayed neutrons
- Low energy photons from the activation of the target + beamline. While these are low energy (1 MeV), pile-up happens.
- Spallation bg -high energy, can be removed with timing
- Nuclear fragments
- Geant4 and MCNP study for different beamline configurations and neutron poisons

Mechanical assembly/design

Design study by Julich engineers/Sam S./Anders Oskarssson



The given weight loads are just estimated based on the draft design and assumptions on each component's individual weight. The values can deviate later in the design process upwards or downwards, when a more detailed design of all components is available.













Steel support structure ca. 40000 kg

Al support structure ca. 4000 kg

Aluminum tube without Target ca. 3700 kg

Time Projection Chamber (TPC) 12 x 50 kg = 600 kg

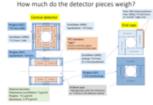
Scintillator 4x 5000 kg (top, bottom, side B) 2x 3000 kg (side A) + Al-structure ca. 3000kg

Al profile structure and lead glass blocks 97000 kg



Working platforms and stairs ca. 12000 kg

ca. 40000 kg Steel Support structure Al Profile structure ca. 4000 kg Aluminum tube ca. 3700 kg Time Projection Chamber 600 kg Al Profile structure and glass blocks ca. 97000 kg ca. 29000 kg Working platforms and stairs: ca. 12000 kg +Cables: X kg +Target:



Sam Silverstein, Stockholm University

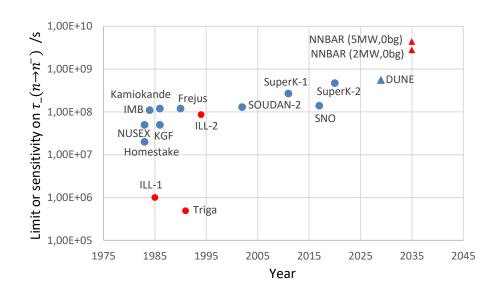


Capability of the experiment

Background suppression selections.

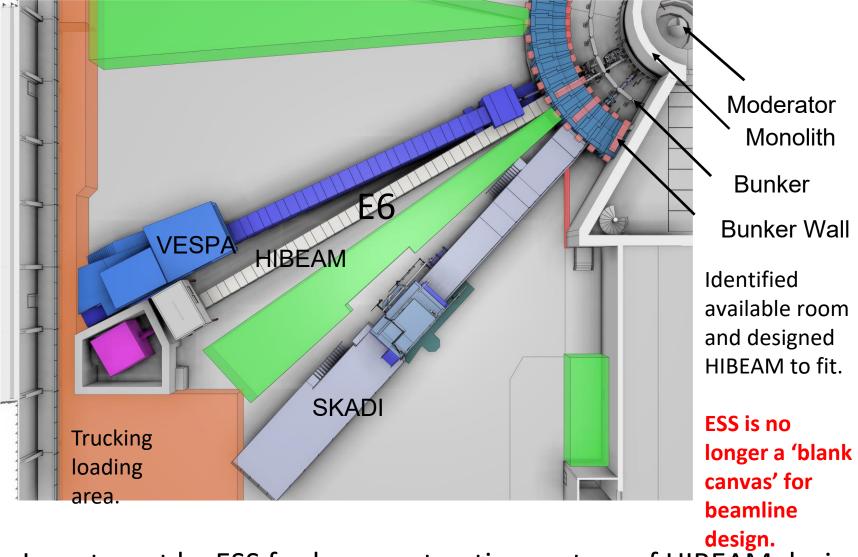
Selection	Signal	Non-muon background	Muon background
Scintillator energy loss ∈ [20, 2000] MeV	0.89	0.008	0.3
TPC track cut	0.87	2.3×10^{-3}	9.0×10^{-3}
Pion count $\geqslant 1$	0.82	7.8×10^{-9}	5.9×10^{-4}
Invariant mass $W \geqslant 0.5 \mathrm{GeV}$	0.8	7.8×10^{-9}	1.5×10^{-4}
Sphericity ≥ 0.2	0.71	1.8×10^{-11}	7.8×10^{-9}
$E_{\text{scint, y} > 0, \text{ filtered}} \leq 320 \text{ MeV & } E_{\text{scint, y} < 0, \text{ filtered}} \leq 930 \text{ MeV}$	0.68	-	-

10³ increase in discovery potential compared to previous experiment



HIBEAM

HIBEAM optimal beamline: E6



Investment by ESS for beam extraction system of HIBEAM design (1.1MEuro). Without this, impossible for any new instrument to operate before 2030's.

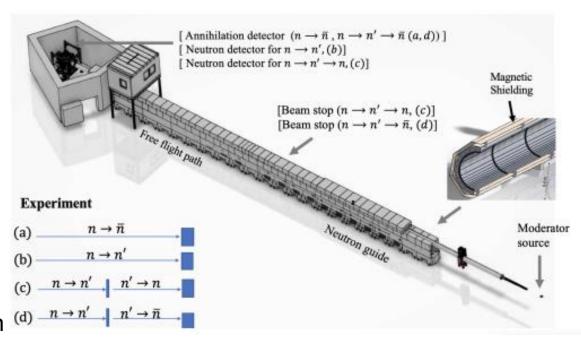
HIBEAM neutron conversions searches

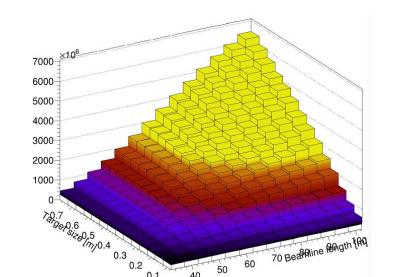
$$\hat{\mathcal{H}} = \left(\begin{array}{cccc} m_n + \vec{\mu}_n \vec{B} & \varepsilon_{n\bar{n}} & \alpha_{nn'} & \alpha_{n\bar{n}'} \\ \varepsilon_{n\bar{n}} & m_n - \vec{\mu}_n \vec{B} & \alpha_{n\bar{n}'} & \alpha_{nn'} \\ \alpha_{nn'} & \alpha_{n\bar{n}'} & m_{n'} + \vec{\mu}_{n'} \vec{B}' & \varepsilon_{n'\bar{n}'} \\ \alpha_{n\bar{n}'} & \alpha_{nn'} & \varepsilon_{n'\bar{n}'} & m_{n'} - \vec{\mu}_{n'} \vec{B}' \end{array} \right)$$

Sensitive to the full mixing Hamiltonian for $n, \bar{n}, n', \bar{n'}$

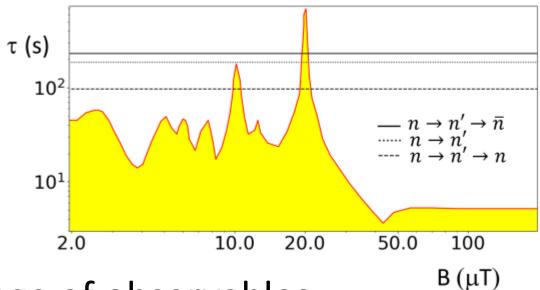
Can use bespoke annihilation detector or WASA (CsI) crystal calorimeter co-owned by UU.

Can exceed ILL experiment by factor 10





Sterile neutron searches



- Range of observables
- Order of magnitude improvement in sensitivity.

Getting to HIBEAM

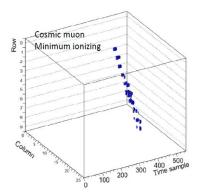
VR RFI

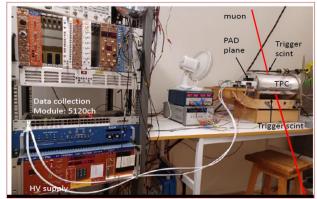
Stockholm, Lund, Chalmers, ESS

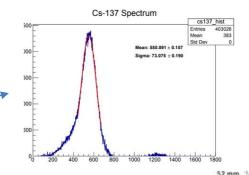
Prototype development

- TPC
- WASA crystal calorimeter
- Scintillator/lead-glass calorimeter

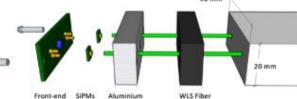
Annihilation detector Neutron detector Beamline design



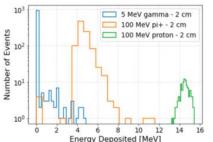


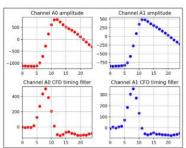












Axions@HIBEAM

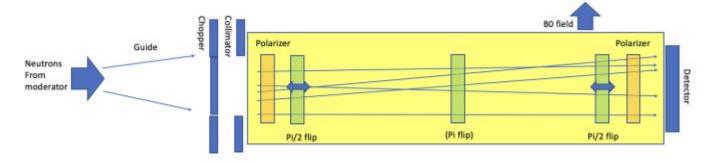
Dark matter candidate - axion. Coupling of axions to a nucleon Axions act as a pseudomagnetic field

$$H_{\mathrm{int}}(t) pprox rac{C_N a_0}{2f_a} \sin(m_a t) \, oldsymbol{\sigma}_N \cdot oldsymbol{p}_a$$

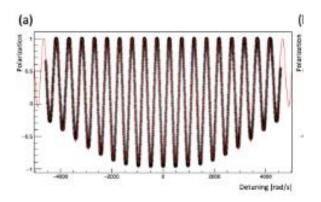
Change in Larmor frequency due to axions

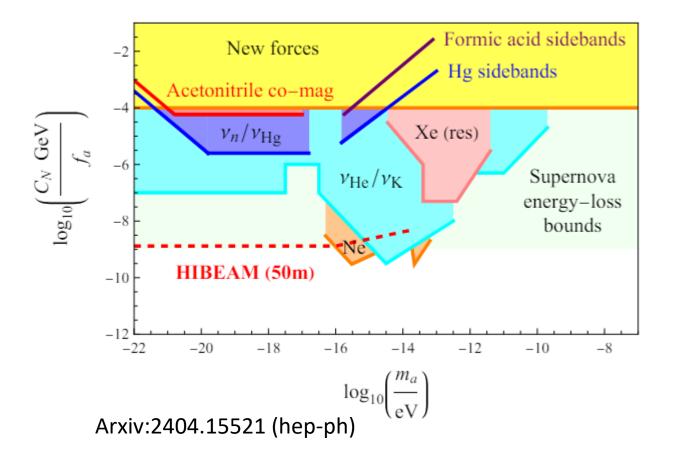
$$\hbar\omega_L = -\gamma \boldsymbol{\sigma}_N \boldsymbol{B} = H_{\text{int}}(t).$$

Ramsey set up for Lar



Fringe shifts





A small-scale pilot experiment at the ESS is funded.

Many uses for HIBEAM beyond neutron-antineutron – neutron charge, nEDM

HIBEAM/NNBAR

Started as an Expression of Interest for a neutron-antineutron search at the ESS (2015)

Signatories from 26 institutes, 8 countries.

Developed into multi-stage HIBEAM/NNBAR project

Co-spokespersons: G. Brooijmans (Columbia), D. Milstead (SU)

Lead scientist: Y. Kamyshkov (Tennesee)

Technical Coordinator (V. Santoro)

Prototype coordinator (M. Holl)

Many active institutes: SU,CTU,UU,LU (SV), TMU (DE), Tennessee, Columbia, ORNL (US), Krakow (PL), Brazil (Rio), Poland (Krakow)....

HIBEAM is supported by the Swedish Research Council (1.4MEuro (completed), 0.48Euro), the Swedish Foundation for Research Strategy (1.5MEuros), Olle Engkvist Foundation (0.4MEuro) + VR grant for collaborating with Italian institutes

NNBAR was supported as part of a 3MEuro H2020 grant for an upgraded ESS with a new lower moderator.

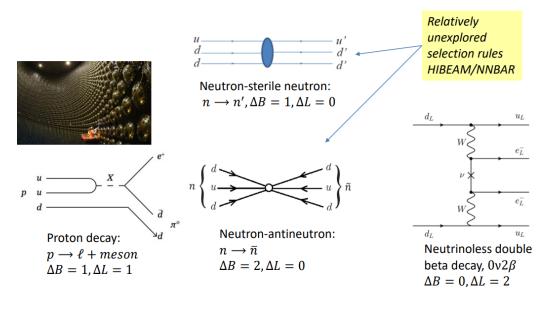
STINT award for collaboration with Brazilian institutes (B. Meirose).

Synergy Grant (SU, LU, TMU, Indiana) for ~construction of HIBEAM and first nnbar, nn' searches.

Selection of HIBEAM/NNBAR publications



Fitting into the European landscape



Plug the "observable gap" for B,L tests

+ sensitivity to many theories of physics beyond the SM (eg hidden sector (dark matter), SUSY, unification models, neutrino mass models etc.)

The 2020 Update to the European Particle Physics Strategy (Essential activities)

A. The quest for dark matter and the exploration of flavour and symmetries are crucial components of the search for new physics. This search can be done in many ways, for example through precision measurements of flavour physics and electric or magnetic dipole moments, and searches for axions dark sector candidates and feebly interacting particles. There are many options to address such physics topics including energy-frontier colliders, accelerator and non-accelerator experiments. A diverse programme that is complementary to the energy frontier is an essential part of the European particle physics Strategy. Experiments in such diverse areas that offer potential high-impact particle physics programmes at laboratories in Europe should be supported, as well as participation in such experiments in other regions of the world.

Summary

- Neutron oscillations are a key but rarely explored portal for new physics
 - baryogenesis, BNV physics, dark matter
- The ESS is opening a new discovery window
- HIBEAM/NNBAR is a multi-stage program to increase sensitivty by ~1000
 - From prototype development to physics
- HIBEAM offers a wide range of applications (neutron oscillations, axions, rare decays etc.).
- Always on the look-out for new collaborators!!

Neutron-antineutron oscillations

- R-parity violating supersymmetry
- Extra dimensions
- Post-sphaleron baryogenesis
- Left-right symmetric models
- etc. etc.

