# Project 6: Development of Low-Mass Dark Matter Searches with the DarkSide-20k Experiment

### RAL/PPD PhD Open Day 19th February 2025



Science and Technology **Facilities** Council

**Particle Physics** 



### Dr Ashlea Kemp & Prof Darren Price RAL/University of Manchester



The University of Manchester





### The Project At a glance, you will...

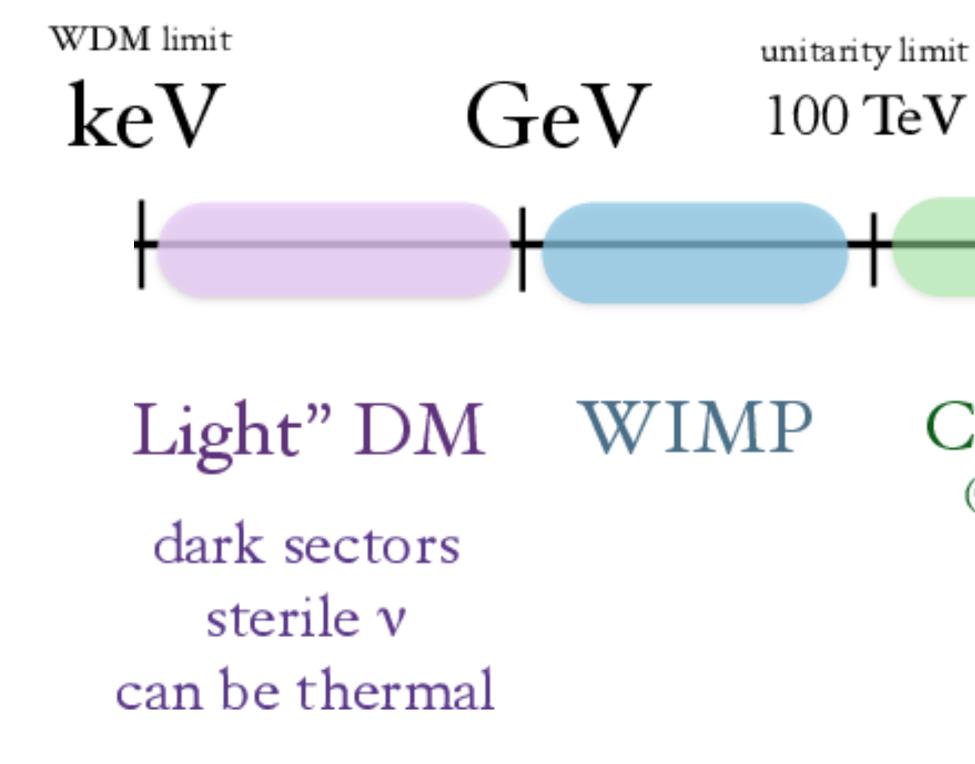
- Take part in the characterisation and calibration of the **novel** matter experiment (which are being built here at RAL!)
- Actively participate in the installation and commissioning of the detector on-site at LNGS.
- Directly contribute to the **very first** dark matter search in DarkSide-20k!
- the search for a range of **low-mass** dark matter candidates.

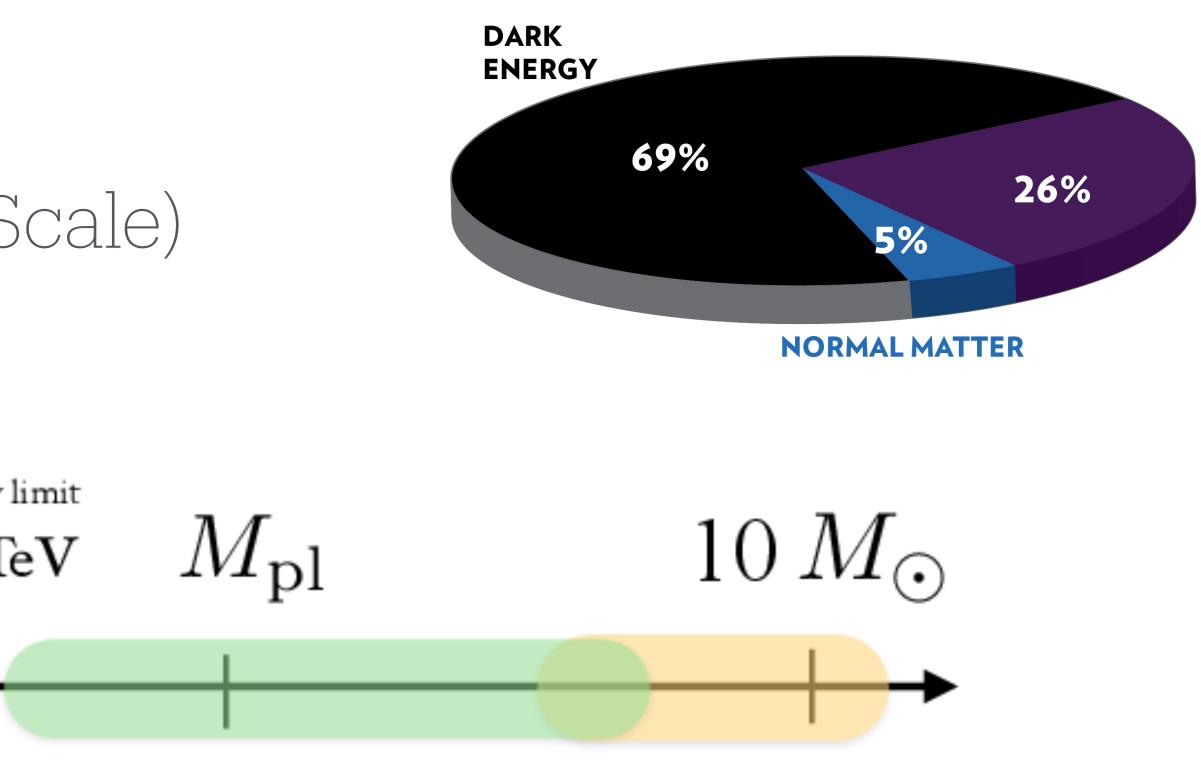
silicon photosensors custom-designed for the DarkSide-20k dark

• Play a pivotal role in developing new ideas to enhance and expand

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### The Challenge Mass Scale of Dark Matter (Not to Scale)





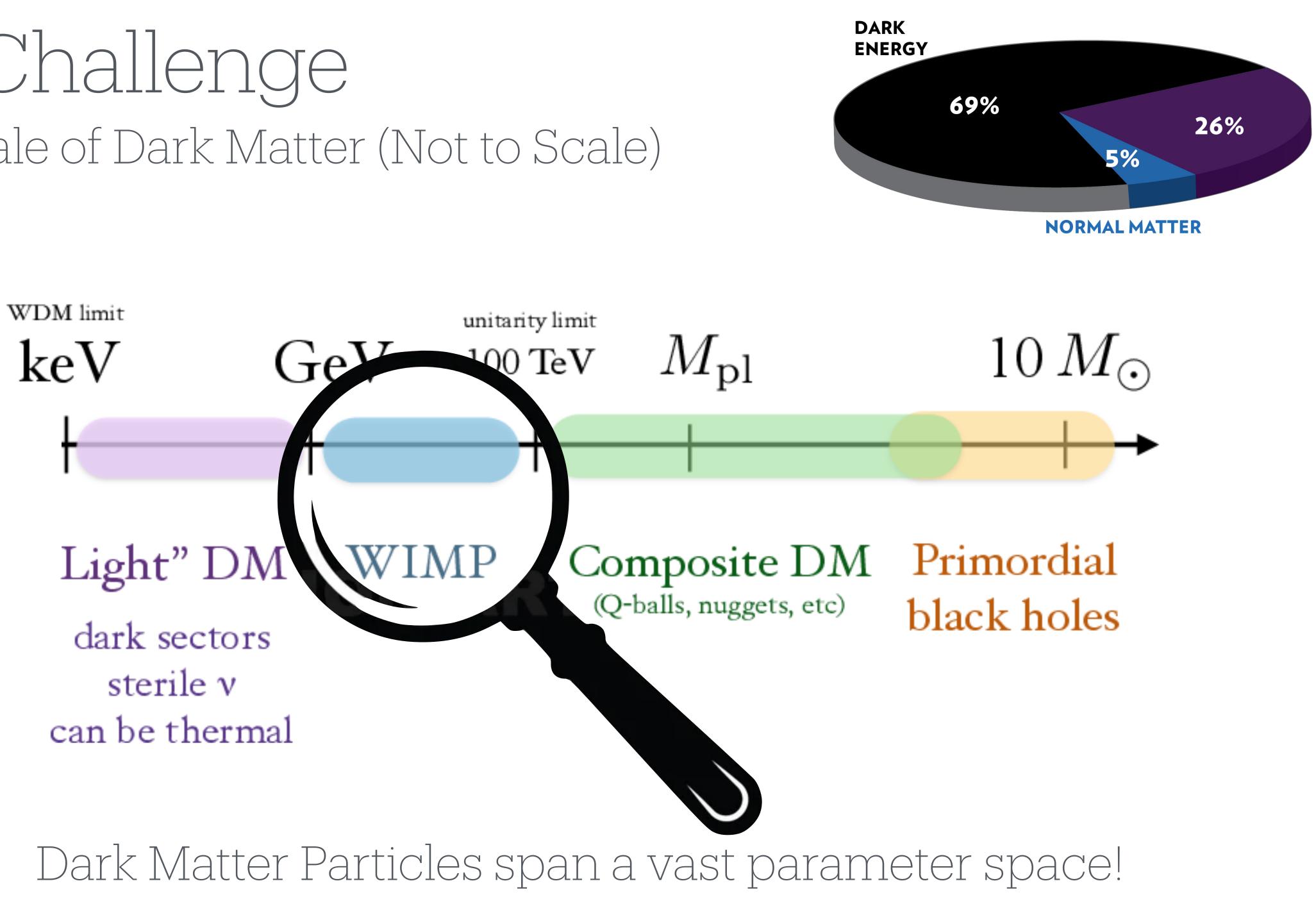
Composite DM Primordial (Q-balls, nuggets, etc) black holes

### Dark Matter Particles span a vast parameter space!



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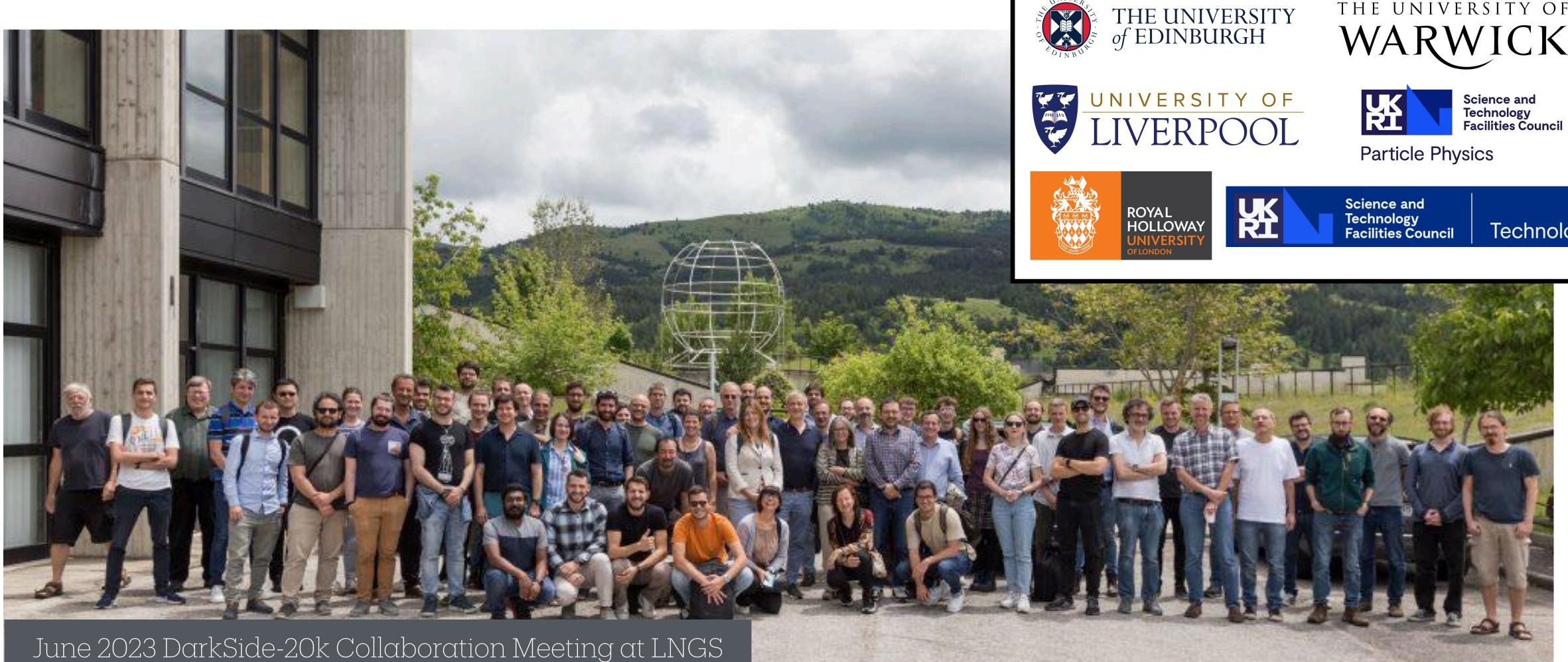
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## DarkSide-20k: Overview

Global Argon Dark Matter Collaboration (GADMC) comprised of 400+ people across 14 countries.









MANCHESTER 1824

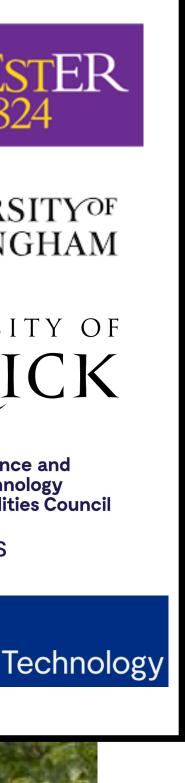












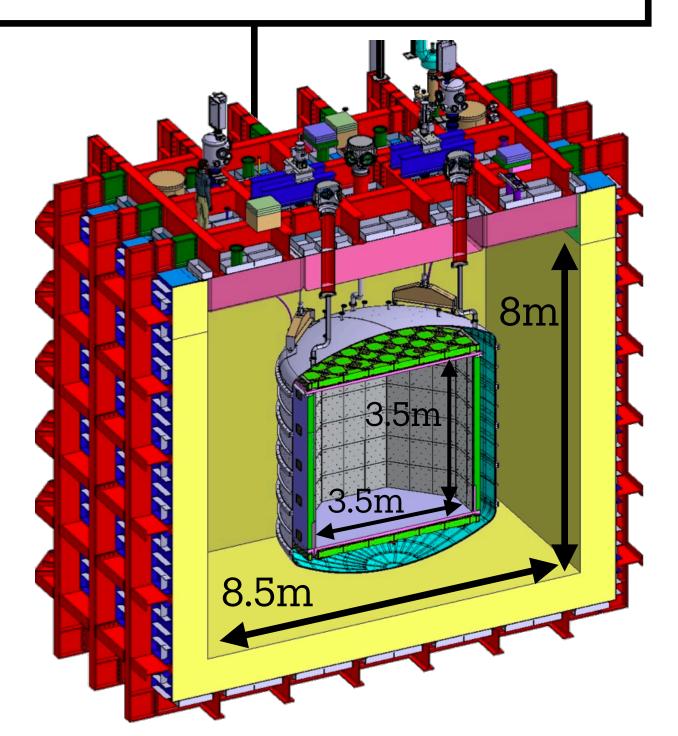


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# DarkSide-20k: Detector

Cosmogenic (Outer) Veto:

- 650t Atmospheric Argon.
- Instrumented with Silicon Photomultipliers (SiPMs) with sparse coverage.

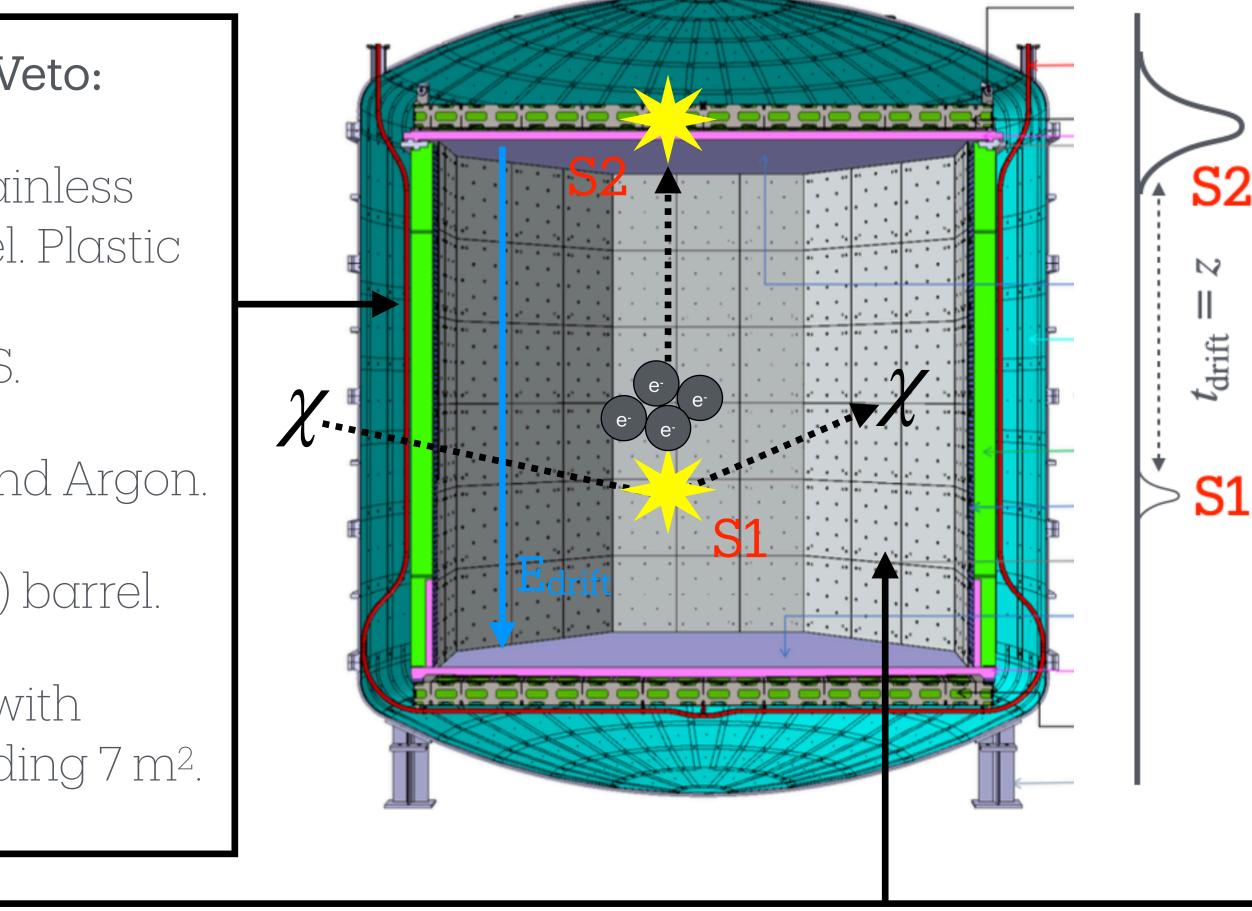


#### Neutron (Inner) Veto:

- Enclosed in Stainless Steel (SS) vessel. Plastic neutron shield surrounding SS.
- 35t Underground Argon.
- PMMA (acrylic) barrel.
- Instrumented with SiPMs; UK building  $7 \text{ m}^2$ .

#### Dual-Phase TPC:

• 50t Underground Argon. Instrumented with 2x Optical Plates of SiPM arrays with  $21 \,\mathrm{m}^2$  coverage.



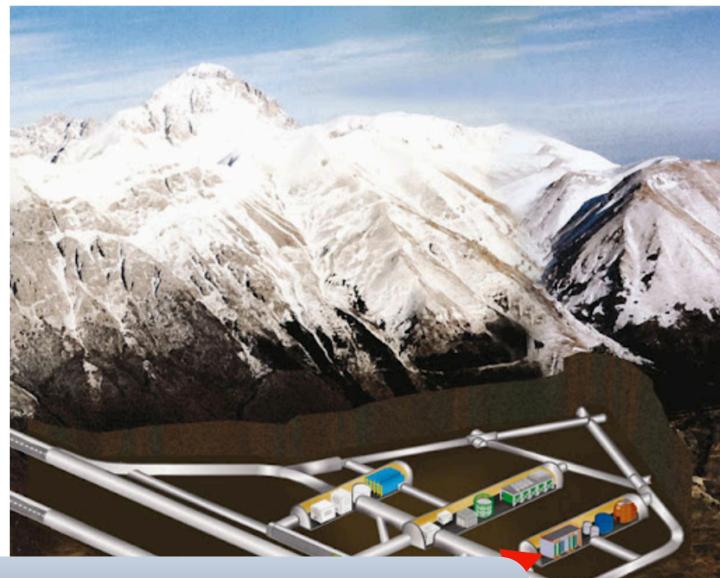




# DarkSide-20k: Status

Construction at LNGS well underway:

- Cryostat and infrastructures in LNGS Hall C complete.
- Cryogenics system operating in Hall C.
- ✓ TPC components in production.
- ✓ Installation of UK photodetectors starting late 2025.
- ✓ Construction complete 2026: data-taking from 2027.



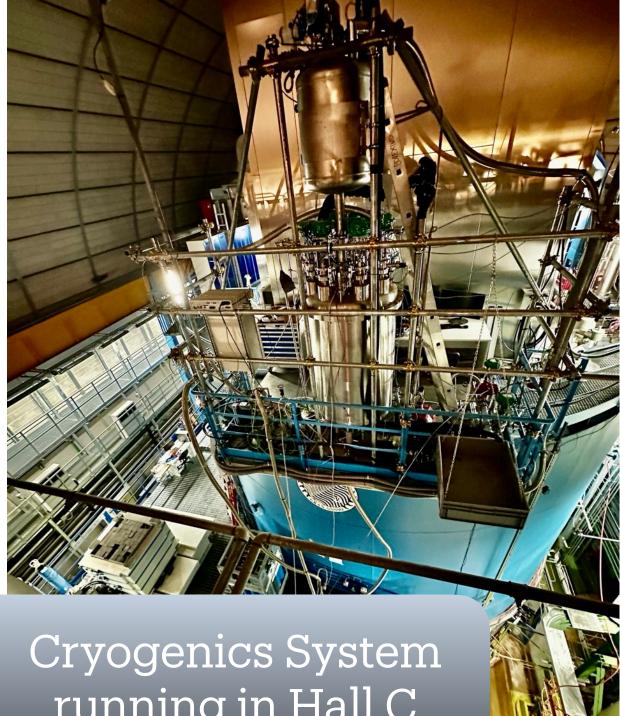


Cryostat complete in Hall C (LNGS)

DarkSide-20k located in Hall C at LNGS, Italy (3400 m.w.e)

UK groups building 7m<sup>2</sup> of Silicon Detector Readout, Production, and Installation, including here at RAL!

Do go and see the Cleanroom tour today if you can!



running in Hall C



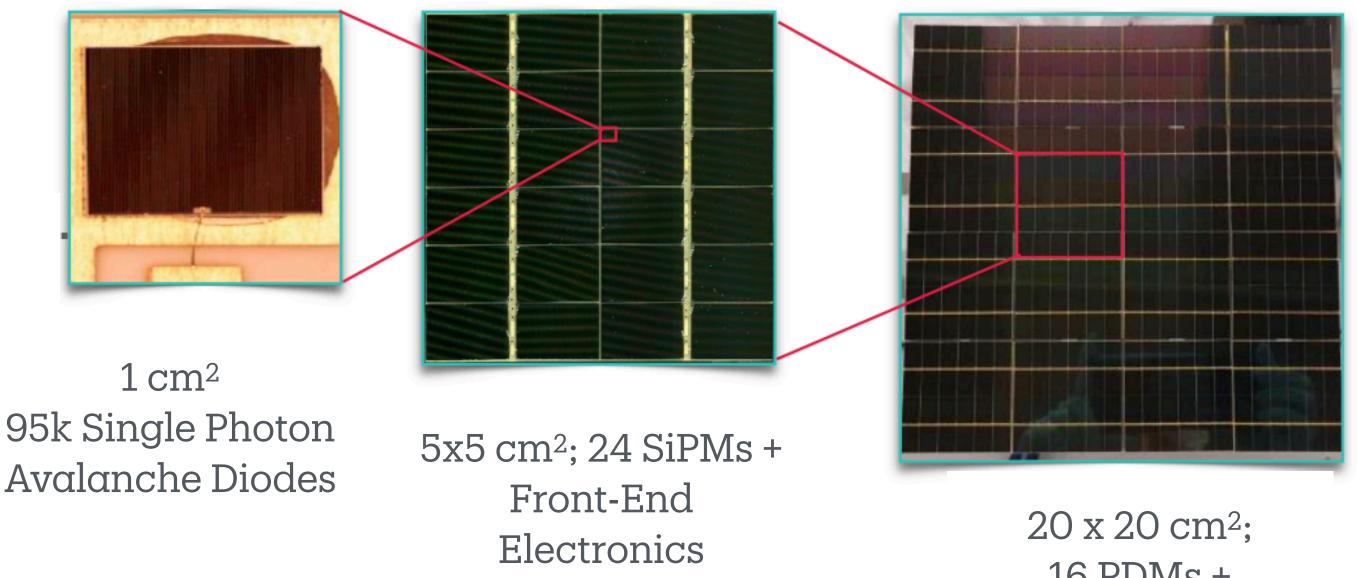


# DarkSide-20k: Light Readout with SiPMs

Silicon Photomultiplier (SiPM)

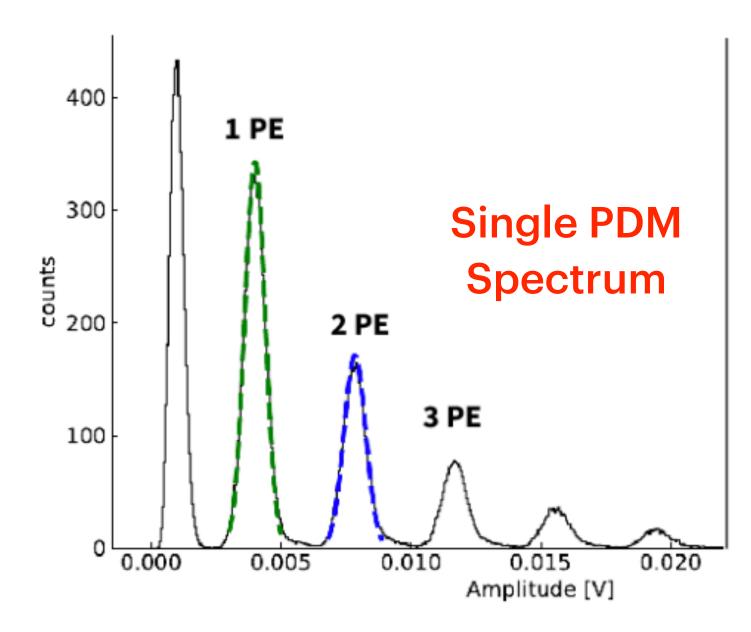
Photodetector Module (PDM)

Photodetector Unit (PDU)



First dark matter experiment to be fully instrumented with SiPMs!

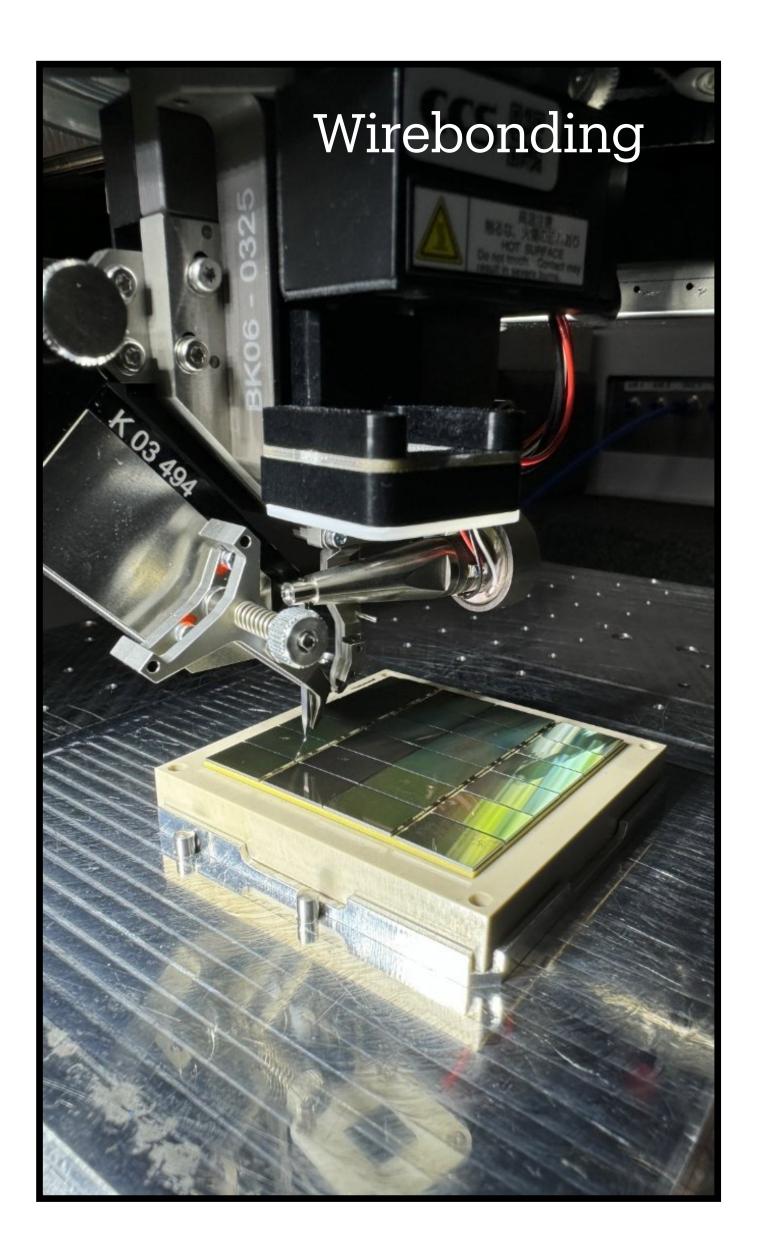
Why SiPMs? Lower noise; lower radioactivity; higher photon detection efficiency; and excellent single photon resolution compared to traditional photomultiplier tubes (PMTs).

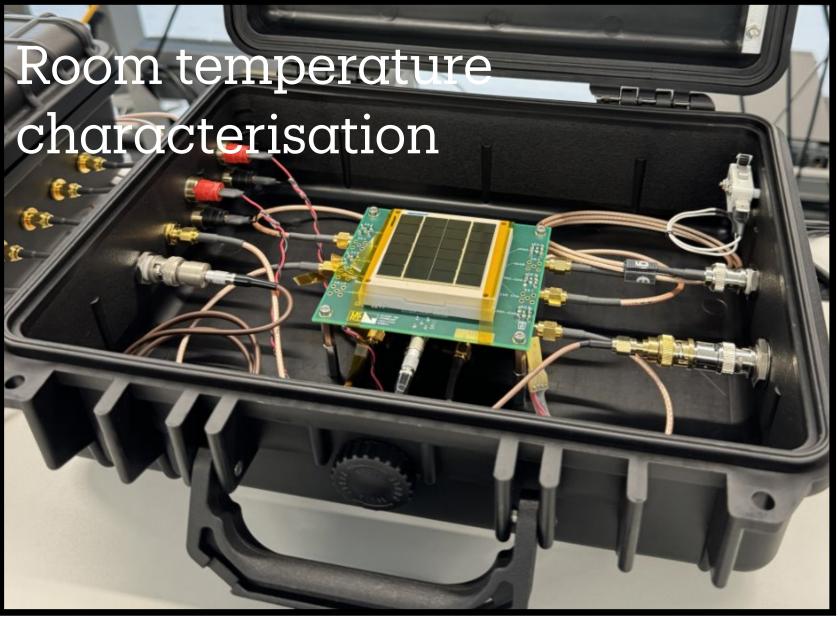


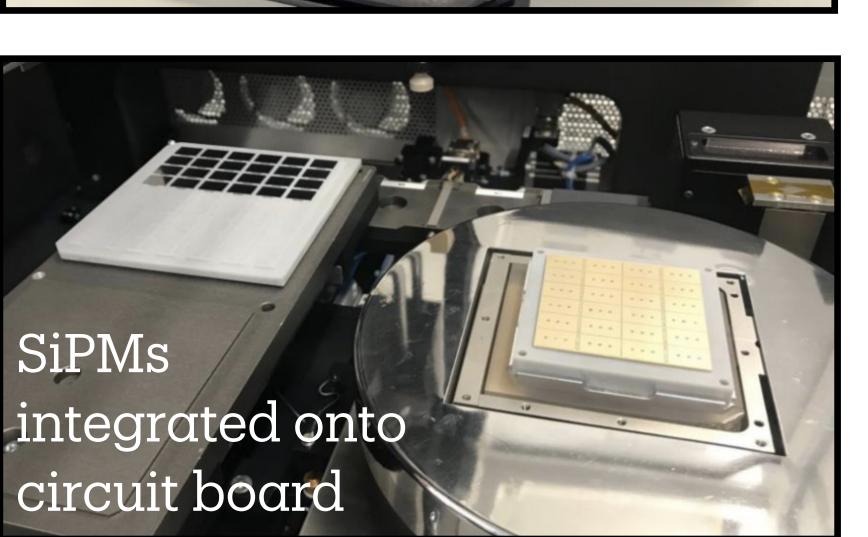
16 PDMs + Motherboard

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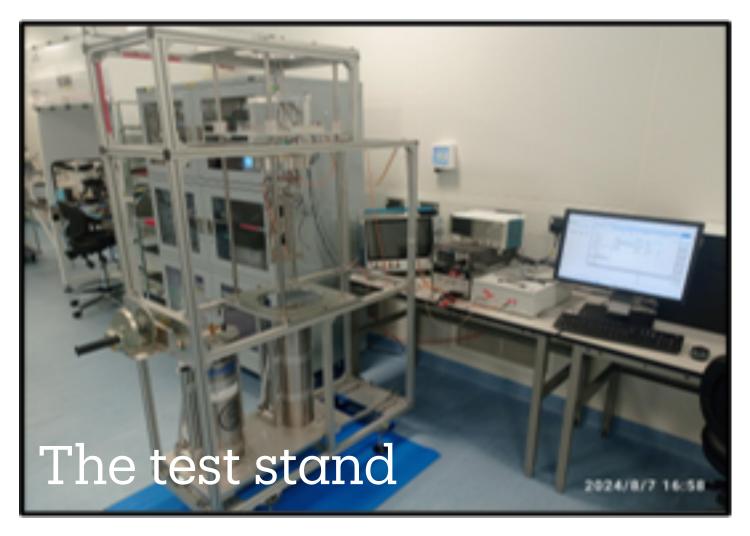
## DarkSide-20k: PDM Assembly & Testing @ RAL!











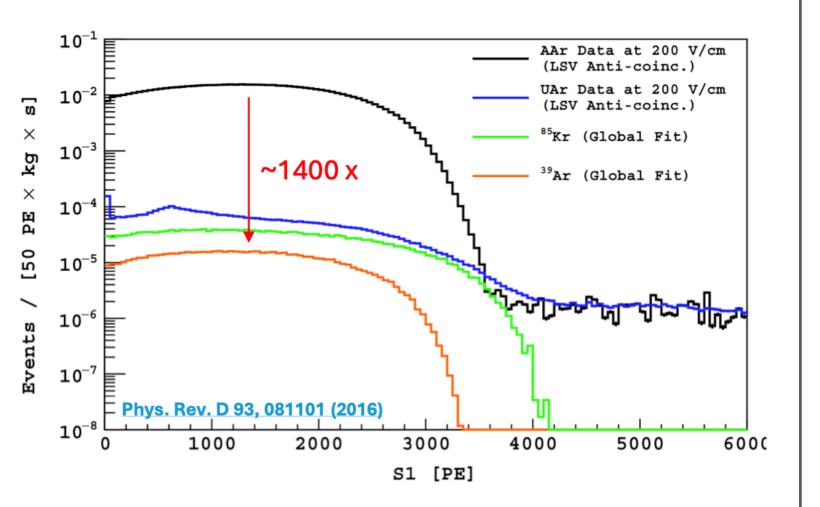






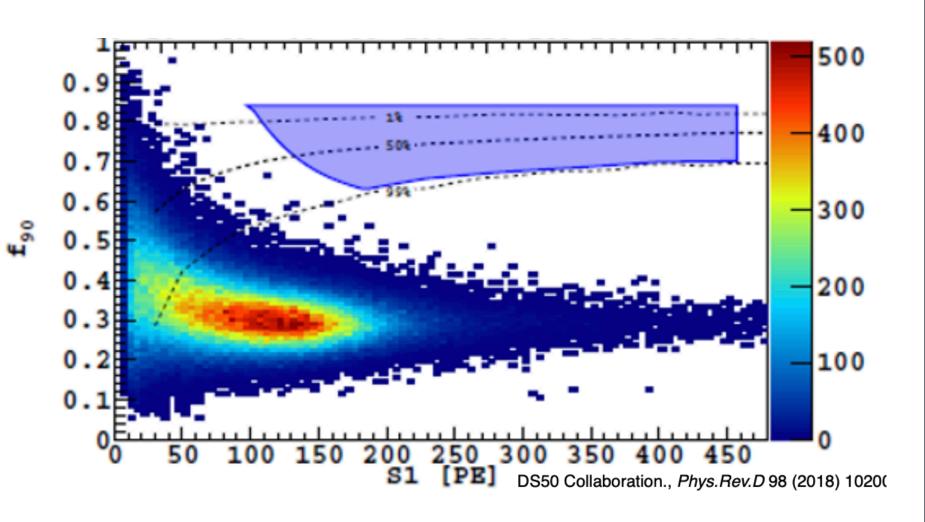
### DarkSide-20k: How we will know if we found Dark Matter?

1) Underground Argon



Atmospheric Argon has <sup>39</sup>Ar radioactive isotope with high activity of 1 Bq/kg: high electron recoil background rate.

<sup>39</sup>Ar Depletion factor ~1400!



discrimination.

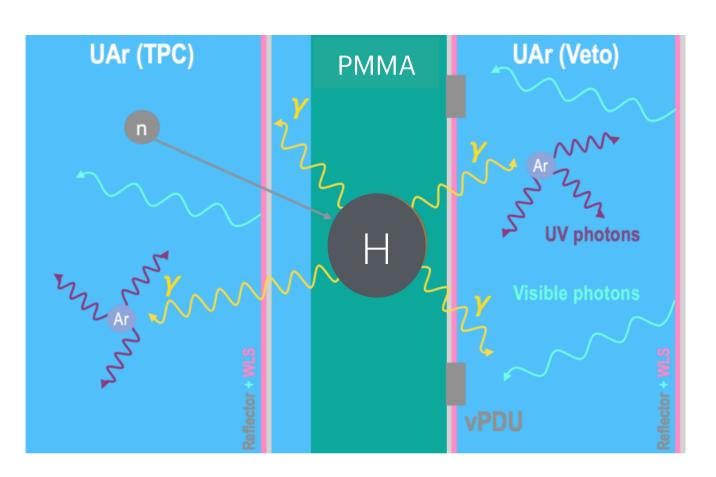
Electron recoils (background) and nuclear recoils (signal) produce very different pulse shapes. Can use pulse shapes to discriminate between them.

World-leading PSD demonstrated ~10<sup>10</sup> electron recoil rejection power.

#### 2) Pulse-Shape Discrimination (PSD)

#### Major Argon advantage: Strong background

### 3) Neutron Tagging



Neutrons most dangerous background for WIMP search.

However, we expect WIMPs to interact only once, unlike neutrons which will interact several times.

Thermal neutron capture produces high energy gamma: use as veto signal!.

<0.1 neutron WIMP-like event in 200 tonne-years.



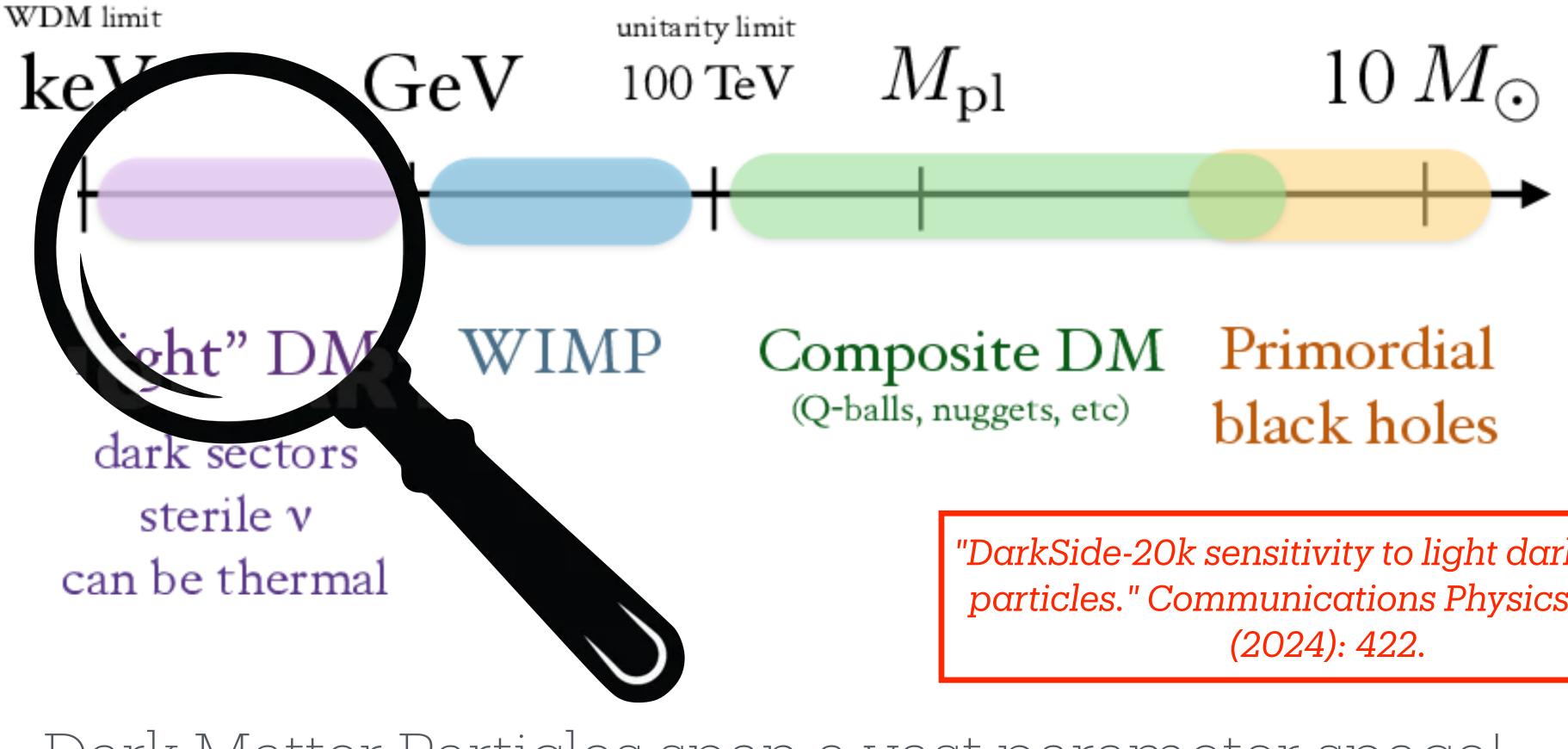








### It's not all about WIMPs... DarkSide-20k has potential to search for **both** lighter and heavier dark matter candidates!





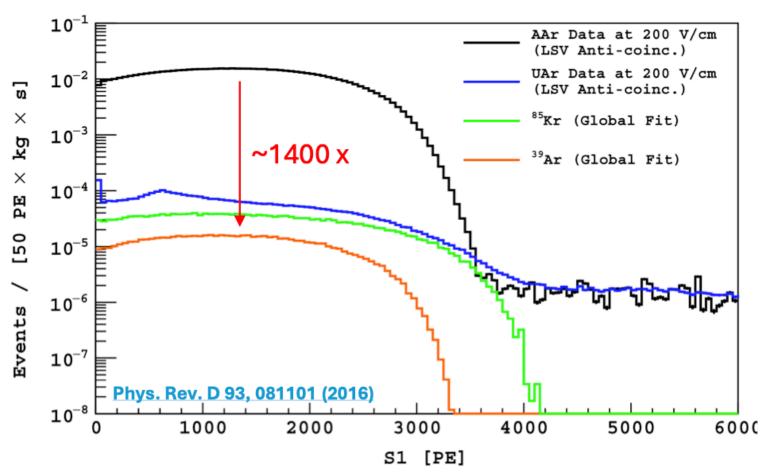
"DarkSide-20k sensitivity to light dark matter particles." Communications Physics 7, no. 1

### Dark Matter Particles span a vast parameter space!

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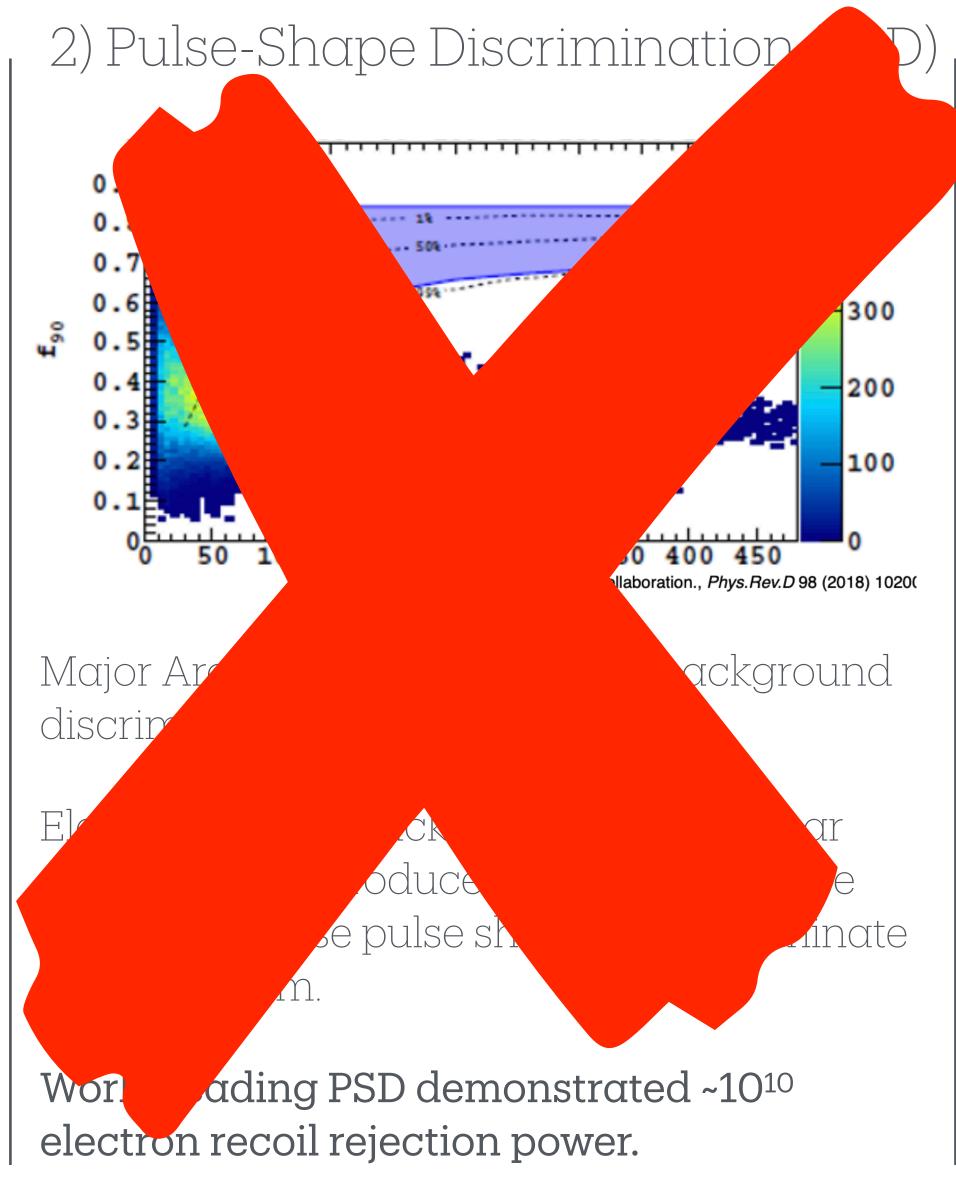
# Low-Mass Searches have their own challenges!

1) Underground Argon

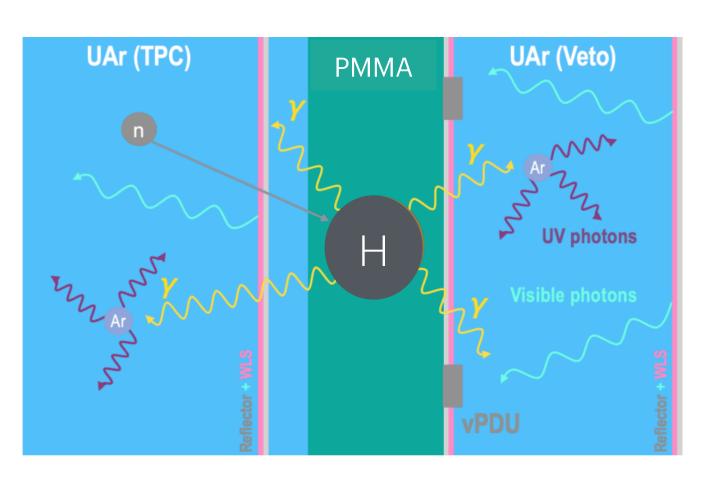


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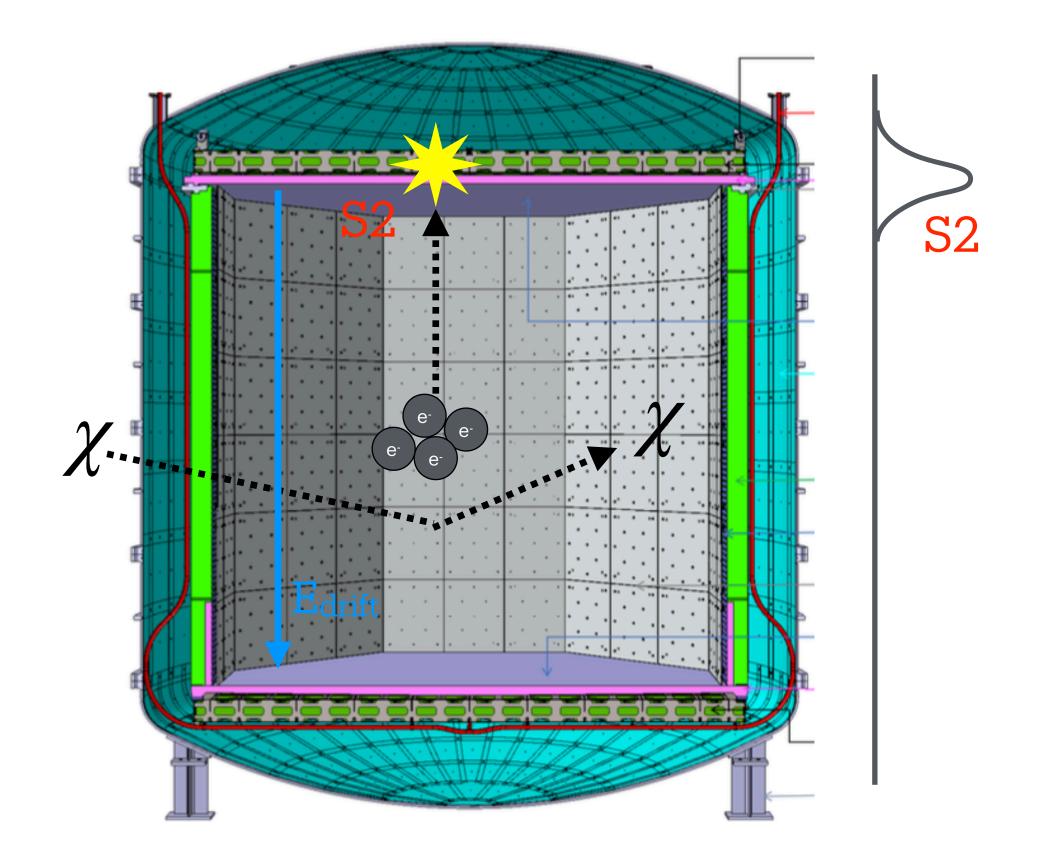


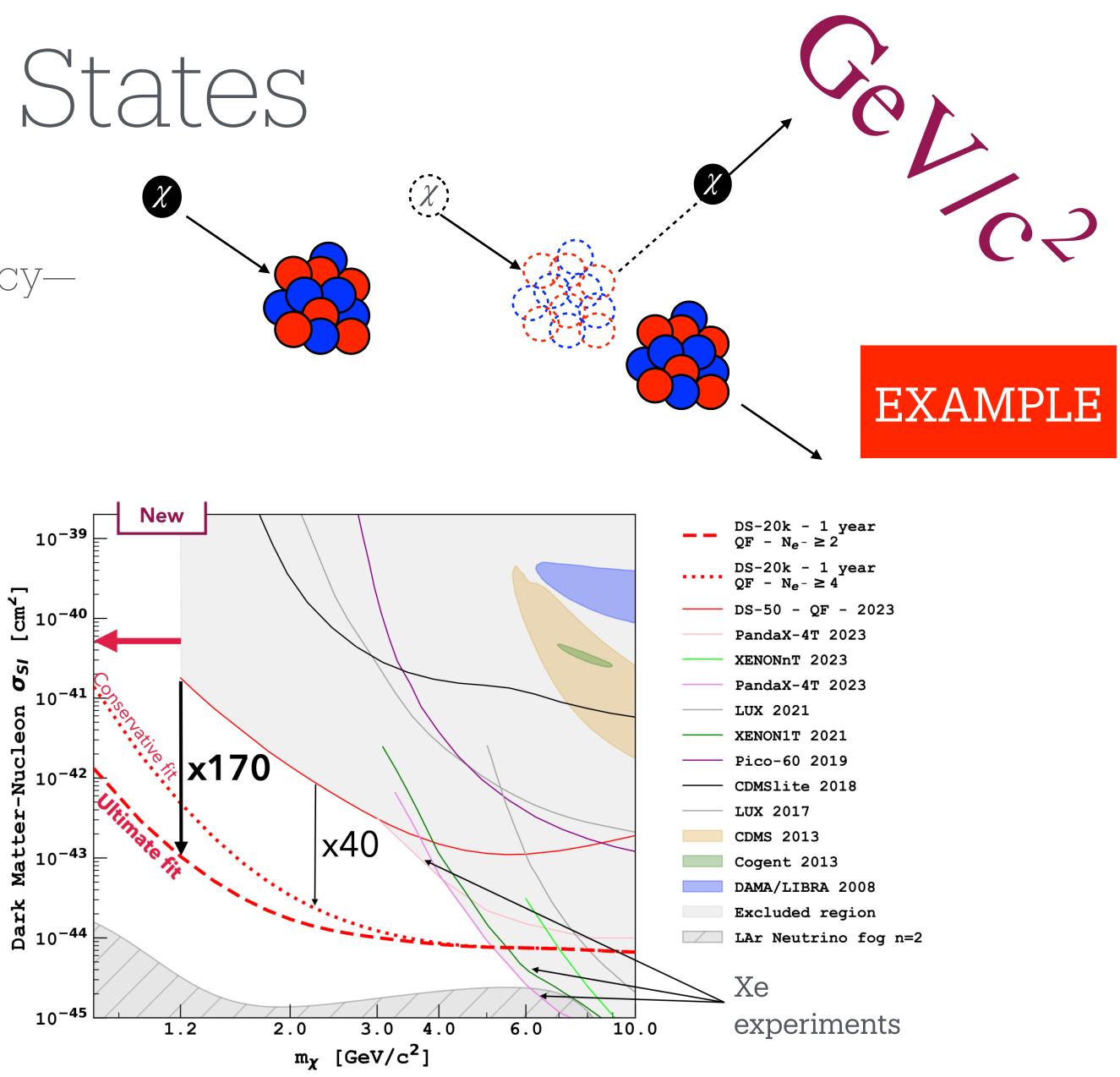




# Nuclear Recoil Final States

Dual-phase TPC design drifts and extracts single ionisation electrons in gas with near-100% efficiency signal amplified a further x20 exploiting electroluminescence in the gas phase.



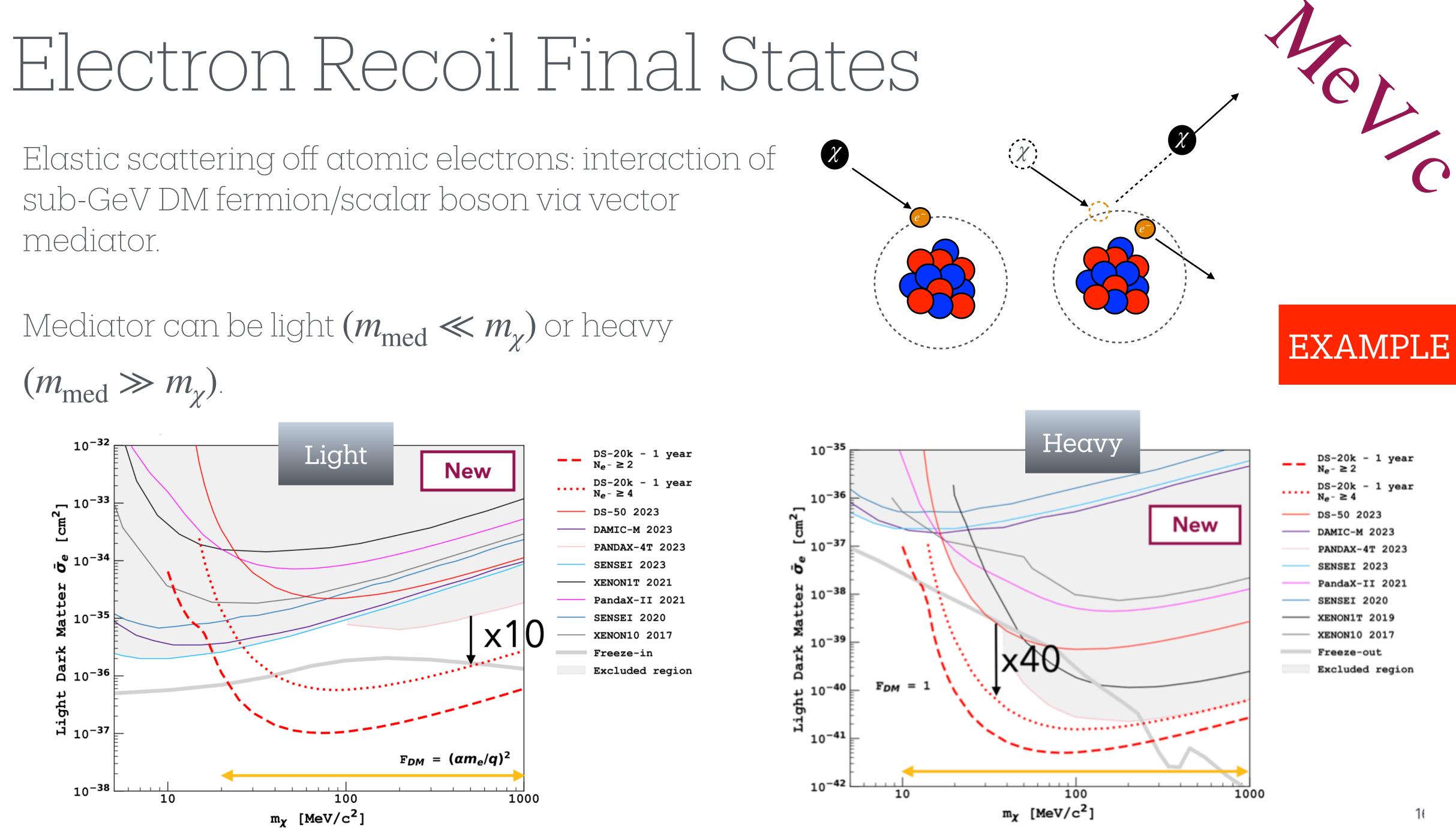


By exploiting ionisation signal (S2) only, DarkSide-20k can reach sub-keV recoil energy thresholds.



sub-GeV DM fermion/scalar boson via vector mediator.

Mediator can be light  $(m_{\rm med} \ll m_{\chi})$  or heavy







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# Summary

DarkSide-20k is a **hugely ambitious** project with **vast physics potential**: you will be joining the collaboration at undoubtedly the **most exciting time**!

Time will be spent between RAL, Manchester, and LNGS: **ample opportunities** for networking and training.

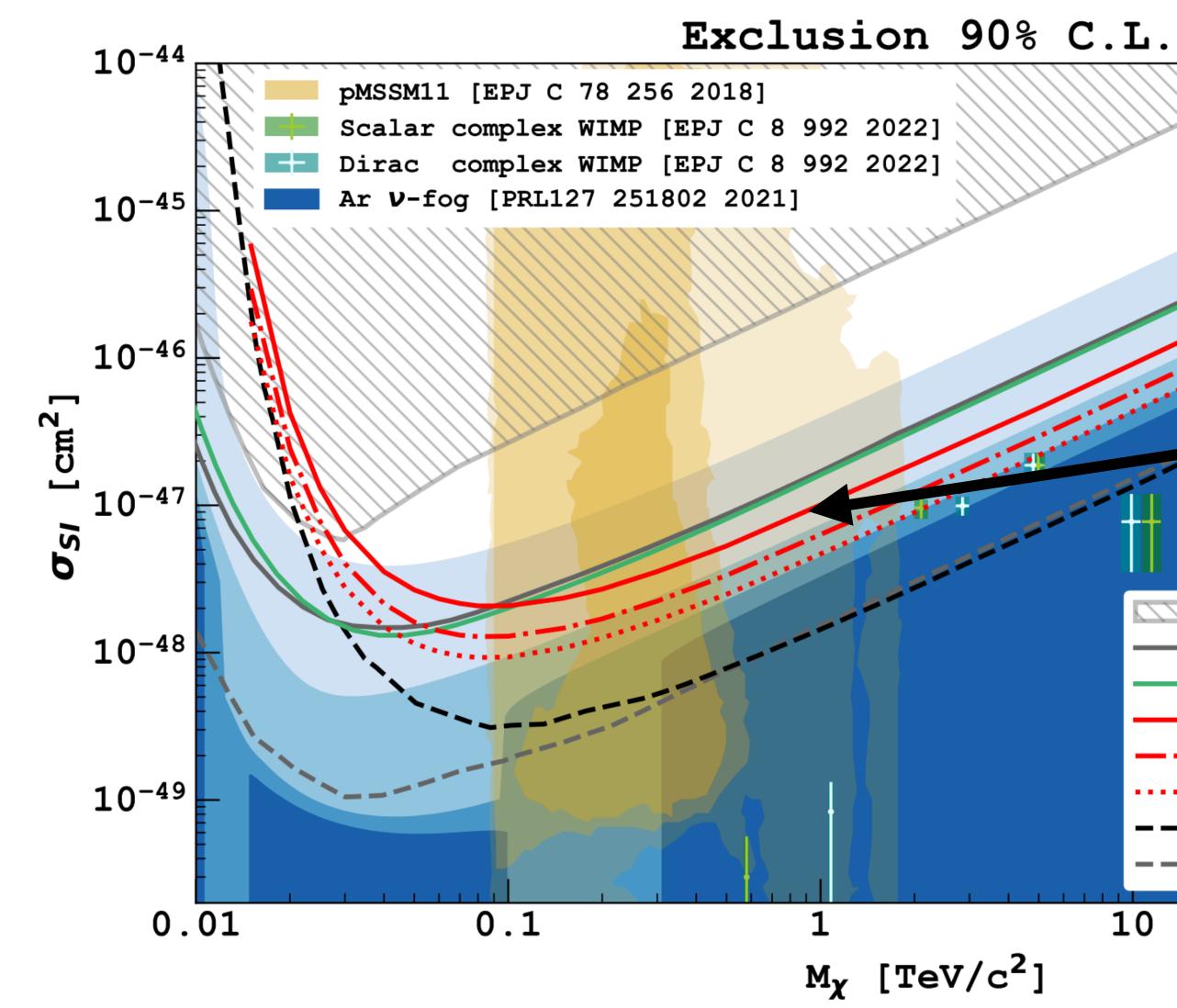
A good **mix of hardware and software** experience, with **plenty of outputs** (papers, conference opportunities).

The project is not set in stone: we want to capitalise on opportunities for new ideas, based on what **you** find interesting!

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# Back Up

## DarkSide-20k: High-Mass WIMP Sensitivity



LZ 90%CL excl [2207.03764] LZ 2.7 y (15.3 t yr) XENONNT 5 y (20.2 t yr) DS-20k Fid. 5 y (100 t yr) DS-20k Fid. 10 y (200 t yr) DS-20k Ext. 10 y (460 t yr) --- ARGO Fid. (3000 t yr) --- XLZD (1000 t yr) 10 100 500 Projected sensitivity to spin-independent WIMP-nucleon scattering cross section: 7.4 x 10<sup>-48</sup> cm<sup>2</sup> for α 1 TeV/c<sup>2</sup> WIMP.

200 tonne-years exposure.



