

The DUNE experiment and its Far Detectors

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Contents

- Neutrino Oscillation
 - A brief Reminder
- Long Baseline Oscillation Experiments
 - The Signal
 - The State of the Art
- DUNE experiment
- DUNE detectors
 - Dual Phase 3x1x1
 - ProtoDUNE Dual-Phase



Illustration by Sandbox Studio/Symmetry Magazine

Neutrino Oscillation

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

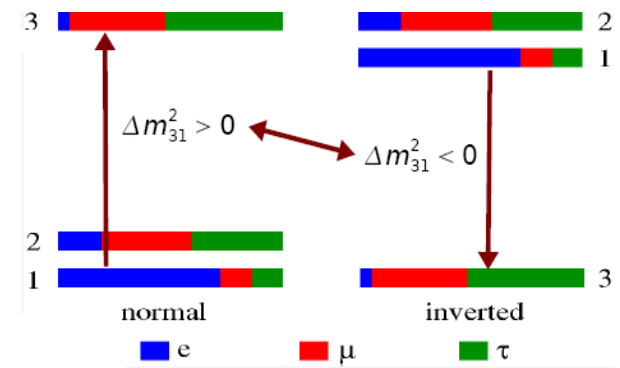
Pontecorvo – Maki – Nakagawa – Sakata (PMNS) matrix

- 3 mixing angles
- 1 CP phase

$$U = \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & \sin\theta_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin\theta_{13}e^{i\delta} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}$$

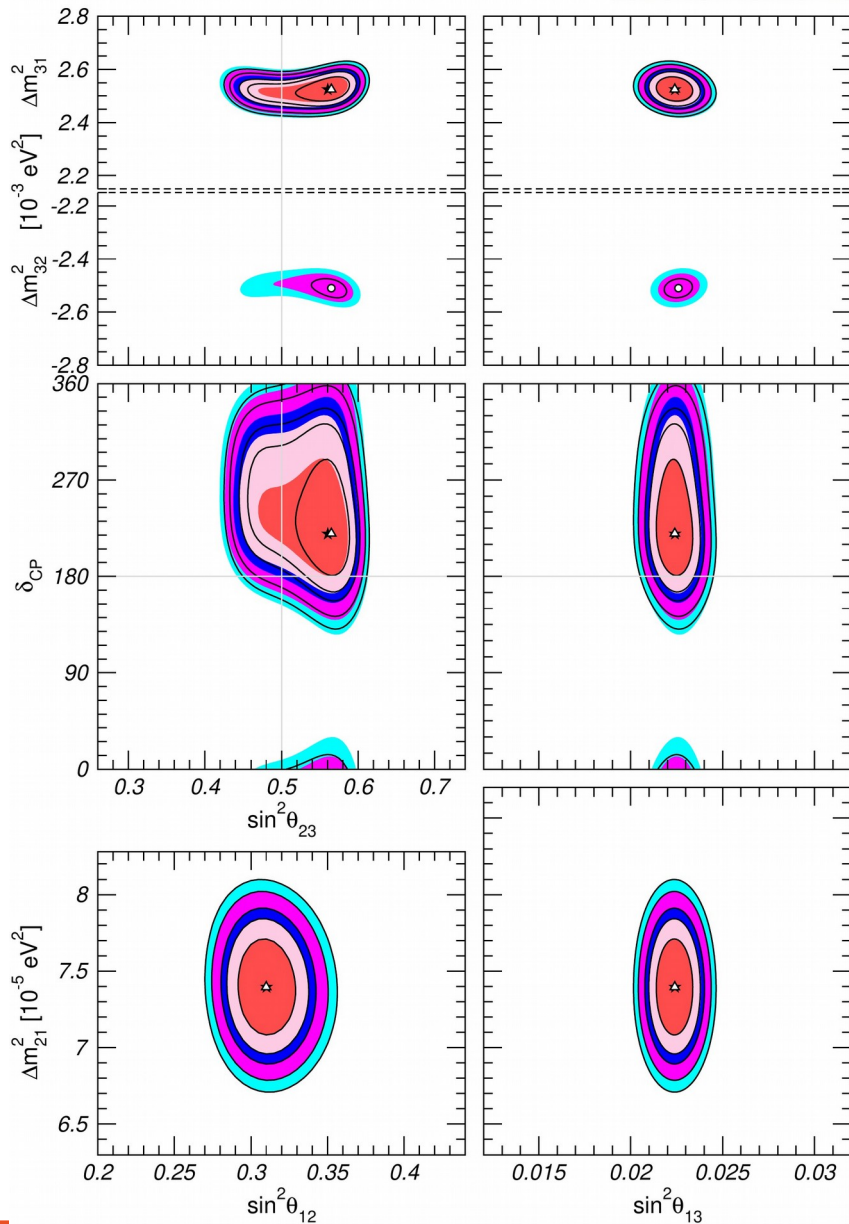
solar atmospheric

- 2 mass splittings Δm^2_{ij}
- $\theta_{13} \rightarrow \delta_{cp}$ and sign of Δm^2_{31}



Global Fits

NuFIT 4.1 (2019)



	Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 10.4$)	
	bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range
with SK atmospheric data				
$\sin^2 \theta_{12}$	$0.310^{+0.013}_{-0.012}$	$0.275 \rightarrow 0.350$	$0.310^{+0.013}_{-0.012}$	$0.275 \rightarrow 0.350$
$\theta_{12}/^\circ$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.75}$	$31.61 \rightarrow 36.27$
$\sin^2 \theta_{23}$	$0.563^{+0.018}_{-0.024}$	$0.433 \rightarrow 0.609$	$0.565^{+0.017}_{-0.022}$	$0.436 \rightarrow 0.610$
$\theta_{23}/^\circ$	$48.6^{+1.0}_{-1.4}$	$41.1 \rightarrow 51.3$	$48.8^{+1.0}_{-1.2}$	$41.4 \rightarrow 51.3$
$\sin^2 \theta_{13}$	$0.02237^{+0.00066}_{-0.00065}$	$0.02044 \rightarrow 0.02435$	$0.02259^{+0.00065}_{-0.00065}$	$0.02064 \rightarrow 0.02457$
$\theta_{13}/^\circ$	$8.60^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.98$	$8.64^{+0.12}_{-0.13}$	$8.26 \rightarrow 9.02$
$\delta_{CP}/^\circ$	221^{+39}_{-28}	$144 \rightarrow 357$	282^{+23}_{-25}	$205 \rightarrow 348$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$
$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.528^{+0.029}_{-0.031}$	$+2.436 \rightarrow +2.618$	$-2.510^{+0.030}_{-0.031}$	$-2.601 \rightarrow -2.419$

JHEP 01 (2019) 106 arXiv:1811.05487

NuFIT 4.1 (2019), www.nu-fit.org

$$|U_{\text{PMNS}}| \sim \begin{pmatrix} 0.8 & 0.5 & 0.1 \\ 0.5 & 0.6 & 0.7 \\ 0.3 & 0.6 & 0.7 \end{pmatrix}$$

Long Baseline Experiments

Beam of ν_μ
or $\bar{\nu}_\mu$



hundreds of km through
the Earth

Detect ν_μ ($\bar{\nu}_\mu$)
 ν_e ($\bar{\nu}_e$)
and if you can
 ν_τ ($\bar{\nu}_\tau$)

ν_e appearance

$$P(\nu_\mu \rightarrow \nu_e) \approx \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(A-1)\Delta}{(A-1)^2}$$

$$+ 2\alpha \sin \theta_{13} \cos \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \cos \Delta$$

$$- 2\alpha \sin \theta_{13} \sin \delta_{CP} \sin 2\theta_{12} \sin 2\theta_{23} \frac{\sin A\Delta}{A} \frac{\sin(A-1)\Delta}{(A-1)} \sin \Delta$$

$$\alpha = \frac{\Delta m_{21}^2}{\Delta m_{31}^2}$$

$$\Delta = \frac{\Delta m_{31}^2 L}{4E}$$

$$A = +G_f N_e \frac{L}{\sqrt{2}\Delta}$$

- ν_e appearance: mass hierarchy, δ_{CP} and octant of θ_{23}
- ν_μ disappearance: high precision $|\Delta m_{32}|$ and $\sin^2 2\theta_{23}$, constrain octant

Long Baseline Experiments

Beam of ν_μ
or $\bar{\nu}_\mu$



hundreds of km through
the Earth

Detect ν_μ ($\bar{\nu}_\mu$)

ν_e ($\bar{\nu}_e$)

and if you can

ν_τ ($\bar{\nu}_\tau$)

Compare oscillation probabilities $P(\nu_\mu \rightarrow \nu_e)$ to $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

They are not the same due to:

- 1) $\delta_{CP} \neq 0$ or π (CP violation!!)
- 2) asymmetry due to matter effects (the Earth is made of matter)

- ν_e appearance: mass hierarchy, δ_{CP} and octant of θ_{23}
- ν_μ disappearance: high precision $|\Delta m_{32}|$ and $\sin^2 2\theta_{23}$, constrain octant

Long Baseline Experiments

Beam of ν_μ
or $\bar{\nu}_\mu$



hundreds of km through
the Earth

Detect ν_μ ($\bar{\nu}_\mu$)

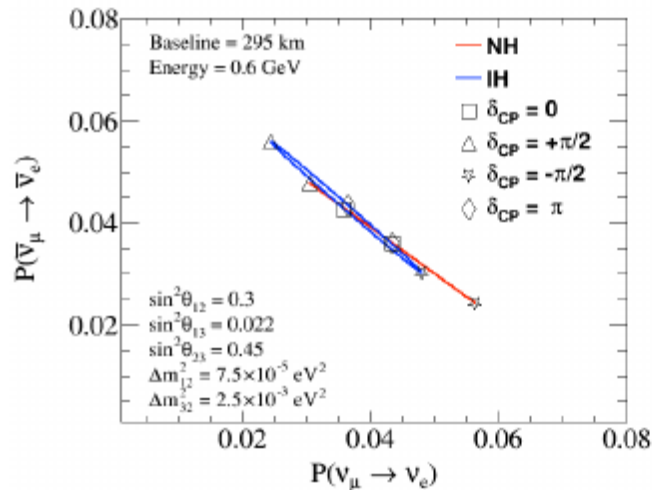
ν_e ($\bar{\nu}_e$)

and if you can

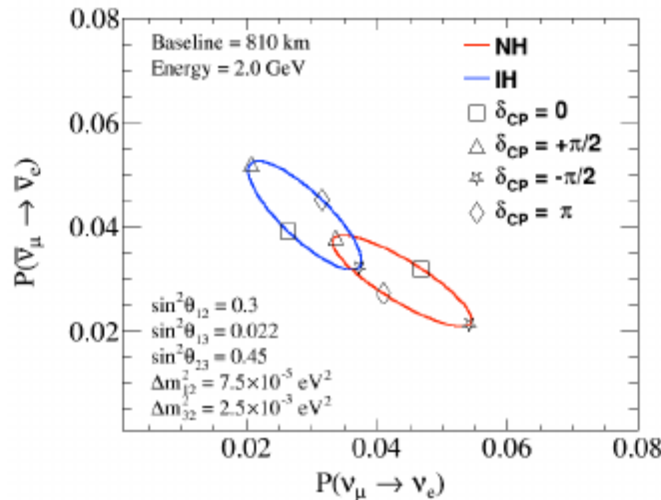
ν_τ ($\bar{\nu}_\tau$)

If the baseline is long enough, the matter effect dominates, and δCP and neutrino mass ordering disentangle.

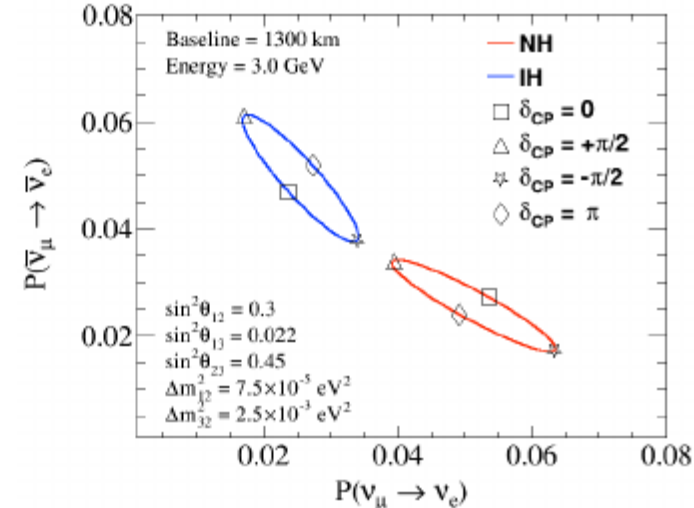
Baseline 295 km
Energy 0.6 GeV



Baseline 810 km
Energy 2.0 GeV



Baseline 1300 km
Energy 3.0 GeV



Long Baseline Experiments

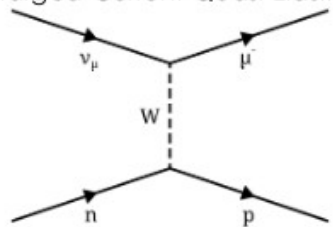
– Neutrino Signal

Real ν_μ as seen by μ BooNE

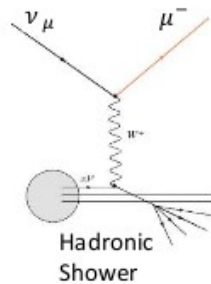
Neutrino flavour determined by outgoing lepton

Depending on neutrino energy, interactions can be quite complex (multiple products and showers)

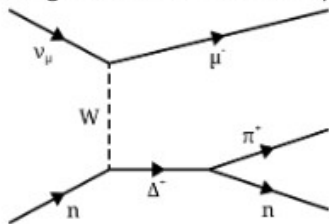
CCQE (1p1h)
(Charged-Current Quasi-Elastic)



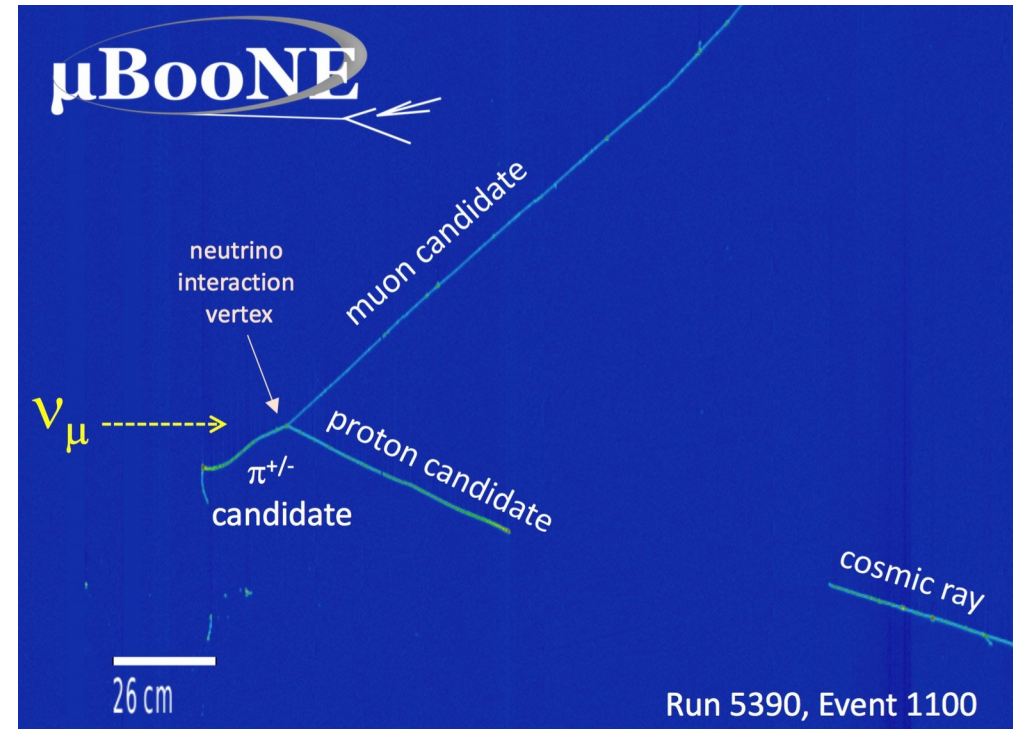
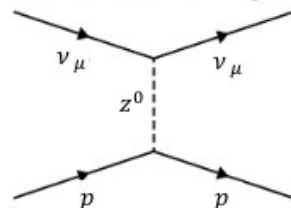
CCDIS
(Deep Inelastic Scattering)



CCRES
(Charged-Current Resonant)



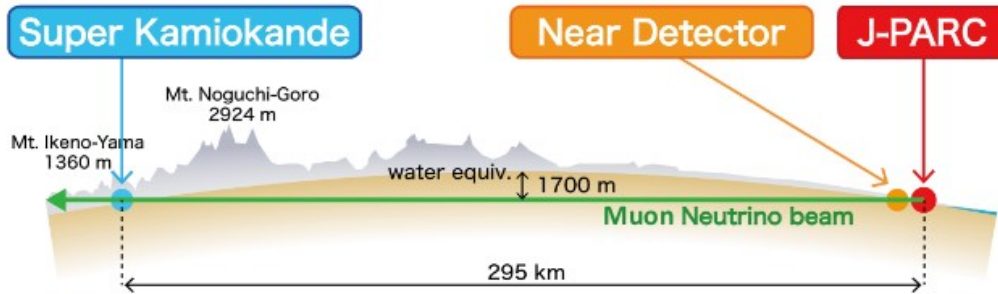
Neutral Current (NC)



Long Baseline Experiments

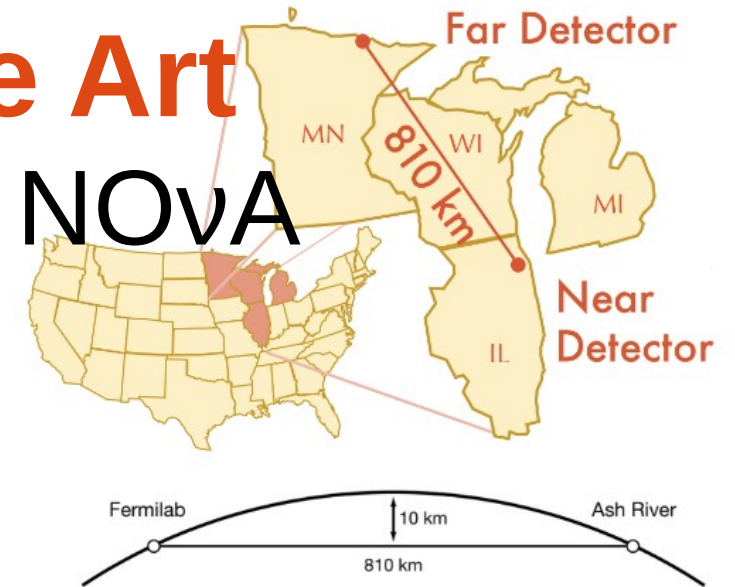
– State of the Art

T2K

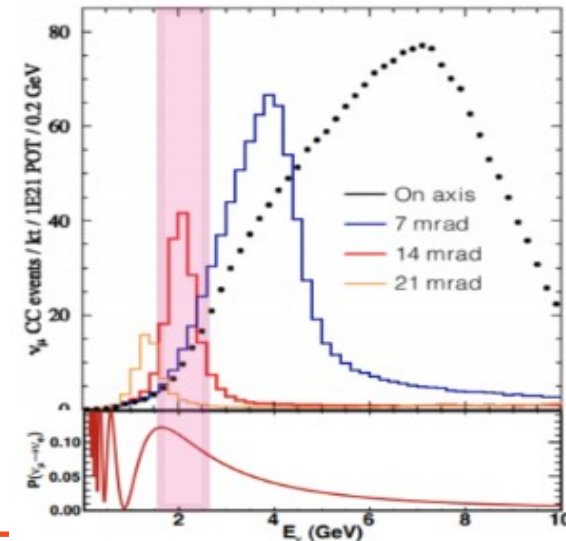


- T2K Far Detector is Water Cherenkov (SuperK 50 ktons)
- Baseline is 295 km
- Both have narrow-band beams (off-axis) peaked at 0.6 GeV (T2K) and 1.9 GeV (NOVA)
- Most events are CCQE

NOVA



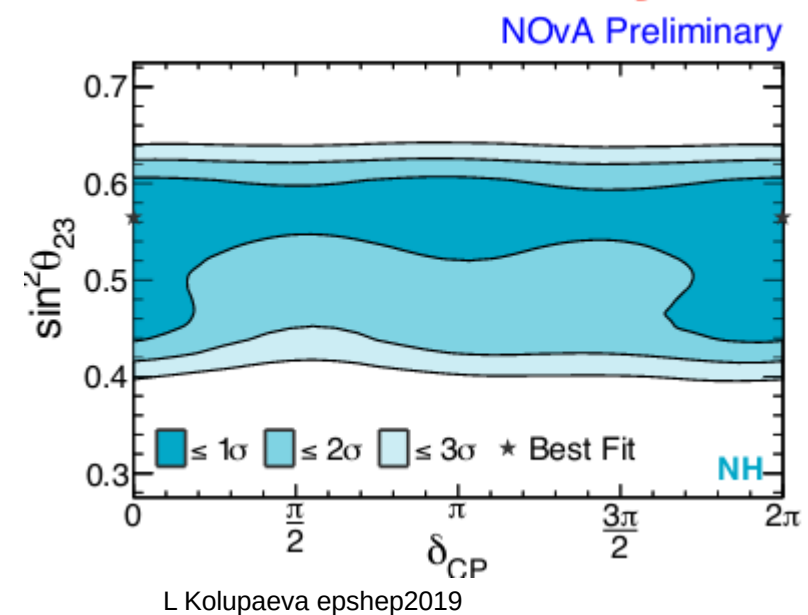
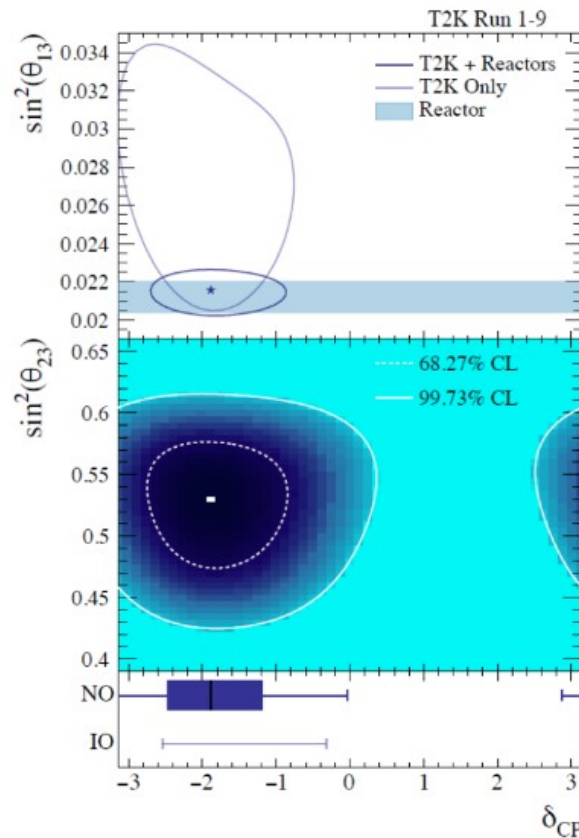
- NOVA has functionally identical Near and Far detectors (finely grained liquid scintillator; 14kton far)
- Baseline is 810km
- Higher neutrino energy
 - DIS occurs



Long Baseline Experiments

T2K – State of the Art NOvA

- Preference for maximal CP violation (confidence intervals at 3σ) and Normal Ordering
- Disfavours Inverted Ordering (at 1.9σ)



arXiv:1910.03887

Collaborations are working towards a joint analysis → 2021

Long Baseline Experiments

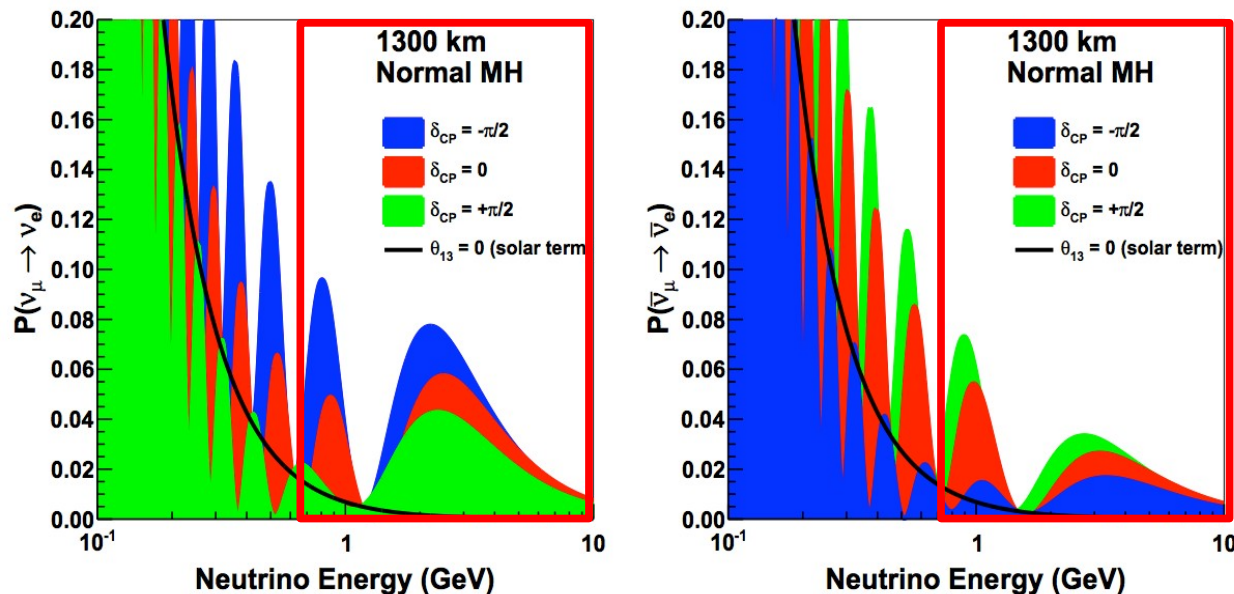
– Future

Higher power beam (more stats)

Make a spectral measurement – use a wide band beam

And a longer baseline (completely disentangle mass ordering and CP violation)

Make a spectral measurement
use a wide band beam
highly performing detectors

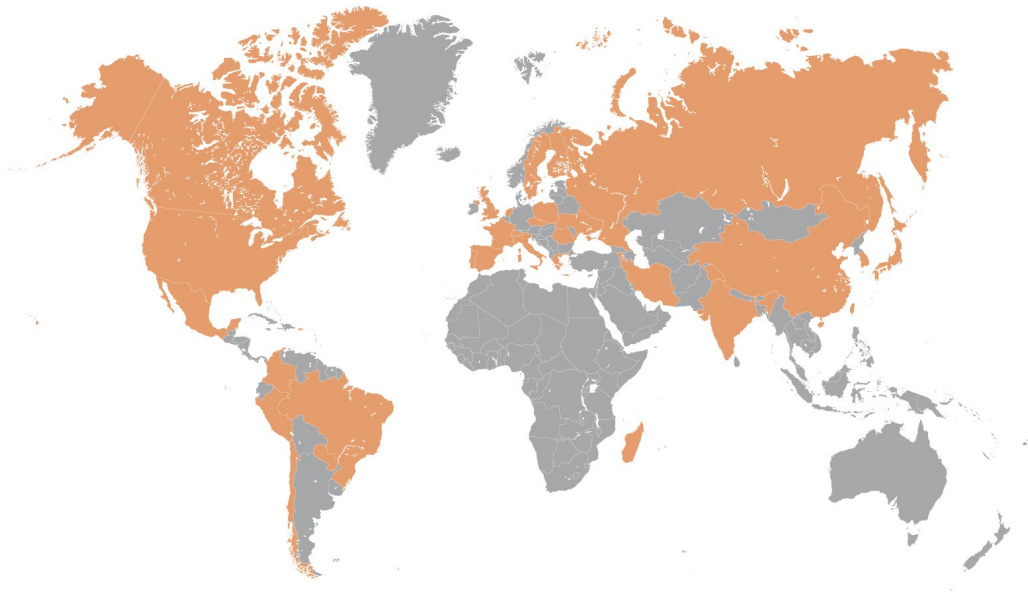


Measurement range spans
2 oscillation peaks
Gain additional power on
deltaCP

Downside – hit all forms of
neutrino interactions.
Need highly performing
Near Detectors...

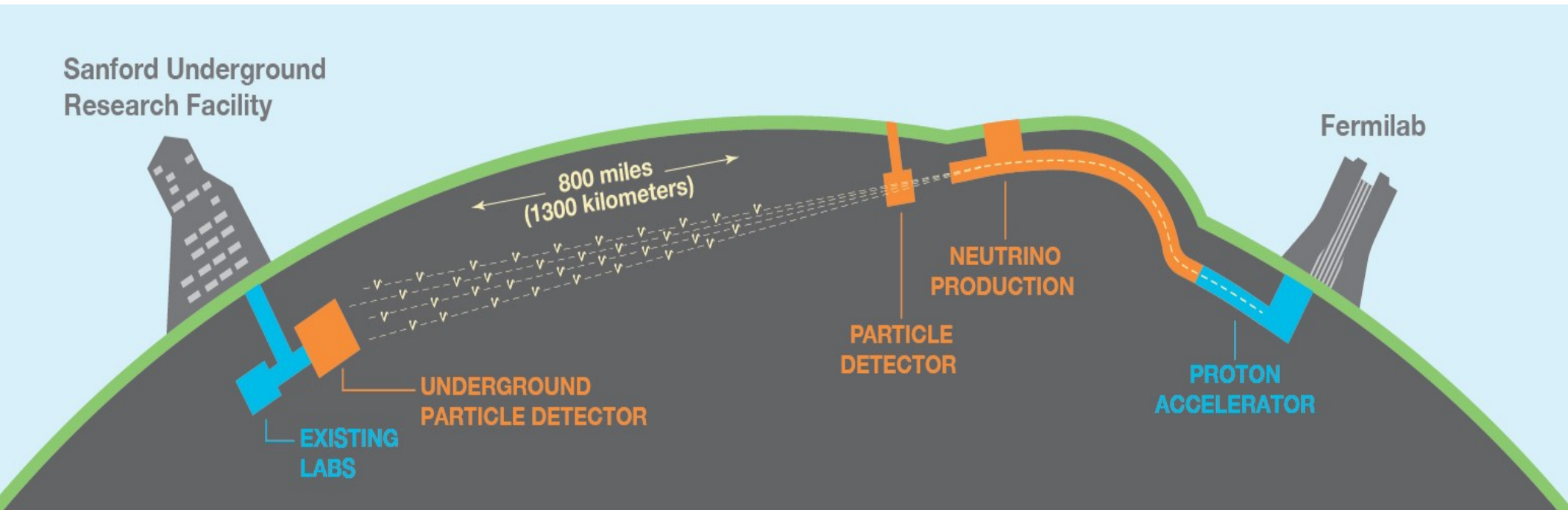
DUNE: international collaboration

- 1069 collaborators from 177 institutions in 31 countries
- 578 faculty, 184 postdocs, 109 engineers, 198 Ph.D. students



Armenia (3), Brazil (31), Canada (1), CERN (37), Chile (3), China (2), Colombia (8), Czech Republic (11), Spain (35), Finland (4), France (38), Greece (5), India (44), Iran (2), Italy (66), Japan (7), Madagascar (4), Mexico (10), The Netherlands (6), Paraguay (4), Peru (7), Poland (6), Portugal (6), Romania (7), Russia (10), South Korea (5), Sweden (1), Switzerland (30), UK (146), Ukraine (4), USA (528)

DUNE is next generation neutrino oscillation experiment



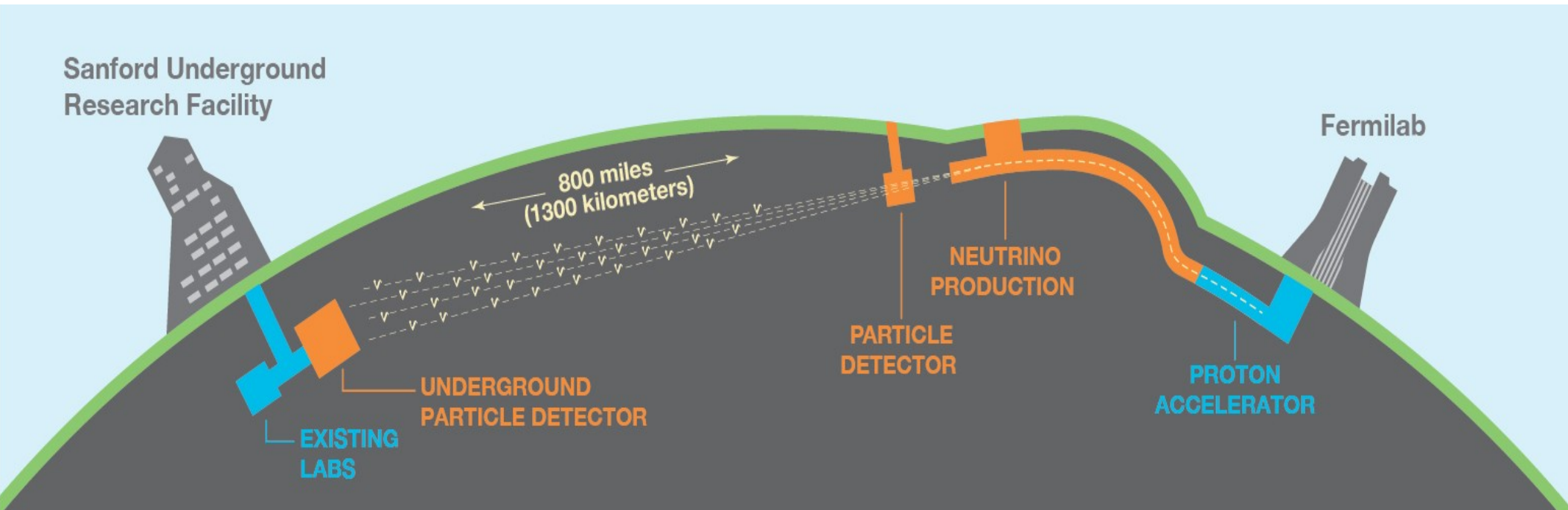
Physics goals :

Neutrino oscillations : measure ν_μ disappearance + ν_e and ν_τ appearance (both neutrino and anti-neutrino modes)

Mass Ordering, leptonic CP Violation discovery, θ_{23} octant and more in a single experiment + physics beyond the Standard Model

Large underground detectors : Nucleon Decay searches, SuperNovae core collapse etc

DUNE is next generation neutrino oscillation experiment

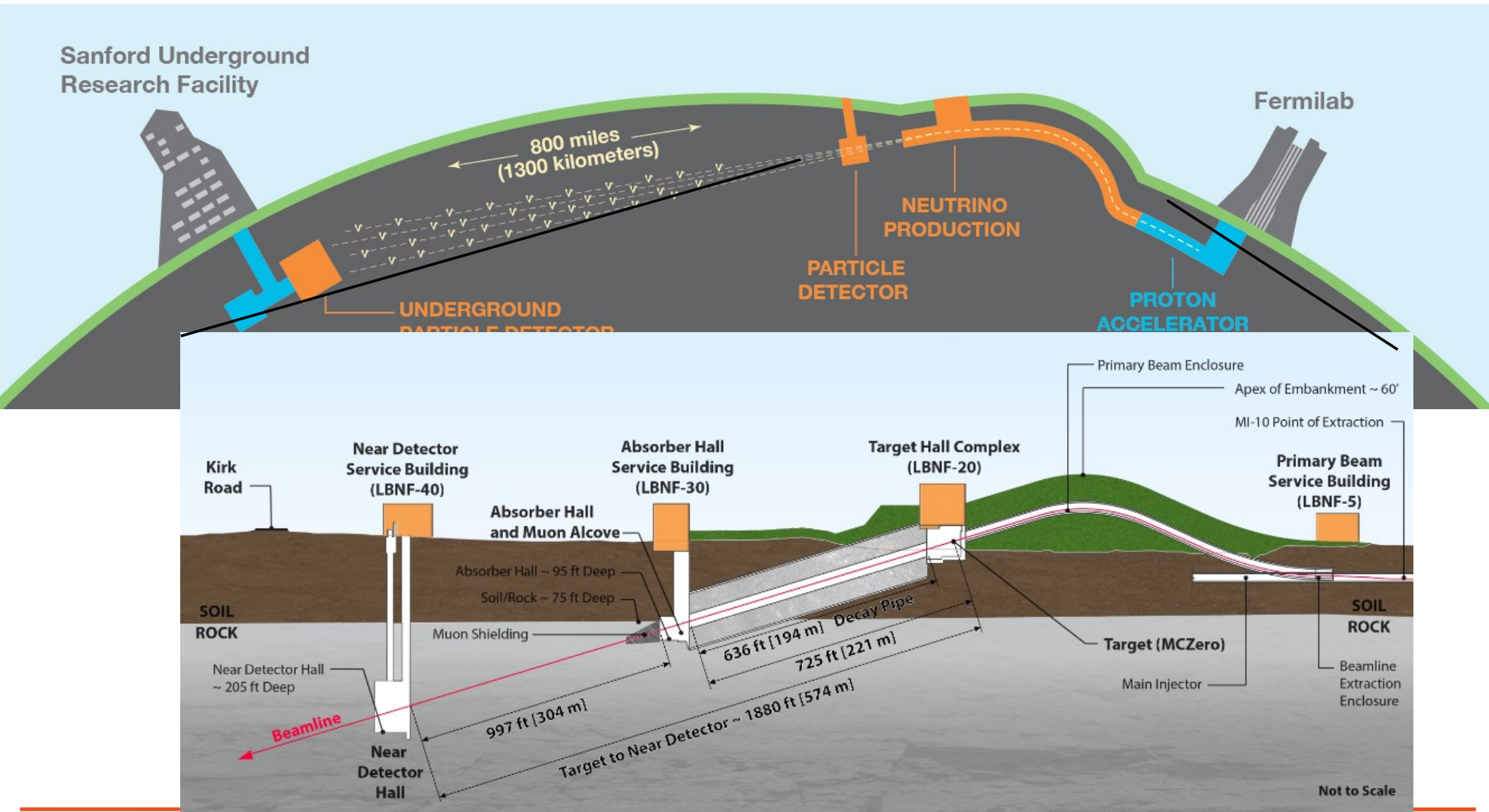


Far detectors at SURF:
4 x 10 kt (fiducial) Liquid Argon TPCs
1.5 km underground

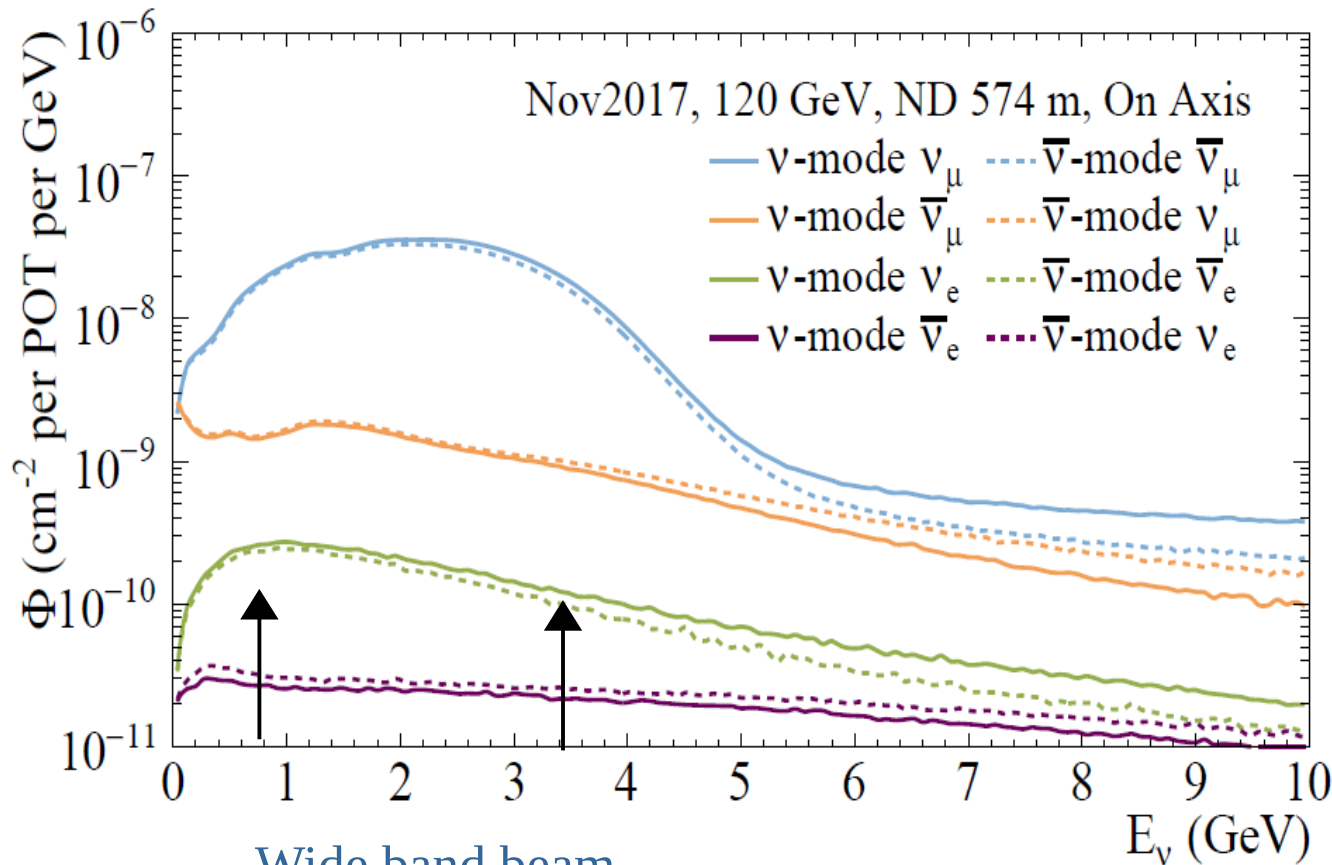
1.2 MW wide-band beam from
Fermilab (upgradable to 2.4 MW)

Near Detector to measure initial
composition

Long Baseline Neutrino Facility



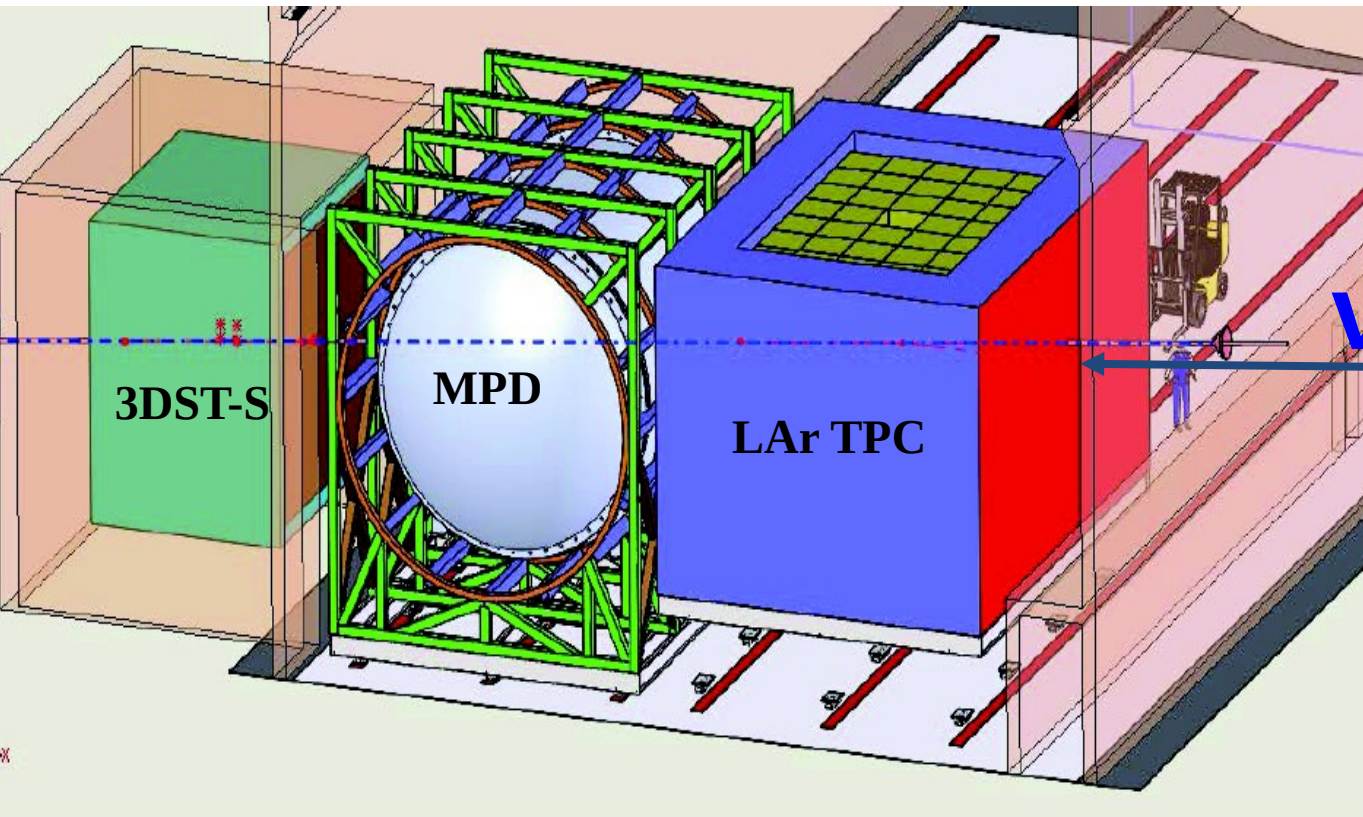
LBNF beam



Wide band beam
2 oscillation maxima

- 120 GeV Main Injector proton beam
- 1.2 MW initial beam power, upgradeable to 2.4 MW
- Beamline and focusing system optimized for CP violation sensitivity

Near detector system at Fermilab



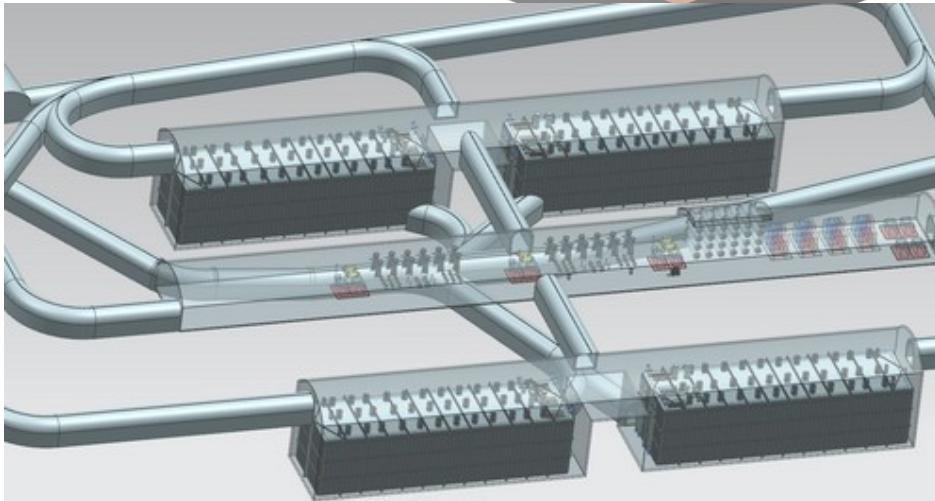
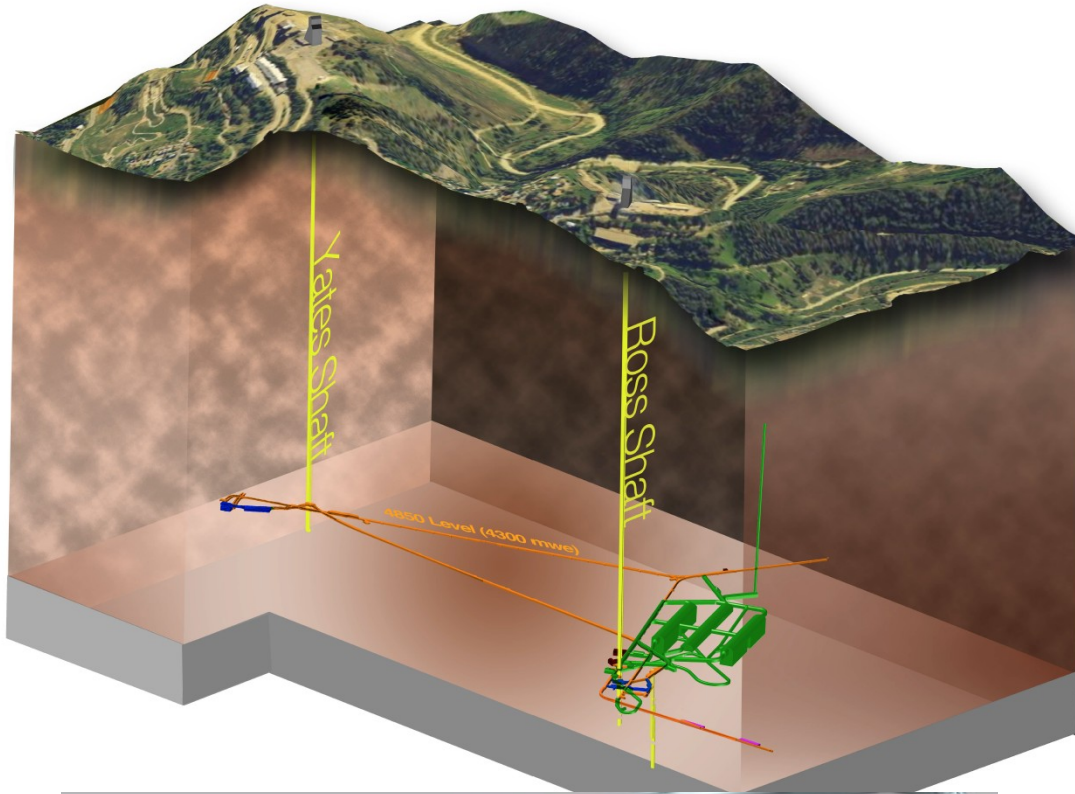
- Predict the neutrino spectrum at the FD
- Measure interactions on Ar
- Measure neutrino energy
- Constrain x-section model
- Measure neutrino flux
- Obtain data with different fluxes
- Monitor the neutrino beam

Liquid Argon TPC

Multi-Purpose Detector - HP gaseous Argon TPC + ECAL + Magnet

3D Scintillating Tracking Spectrometer

Far detector at SURF



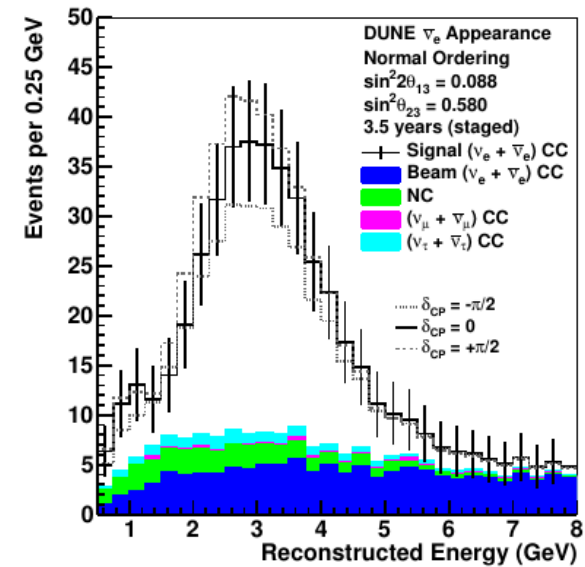
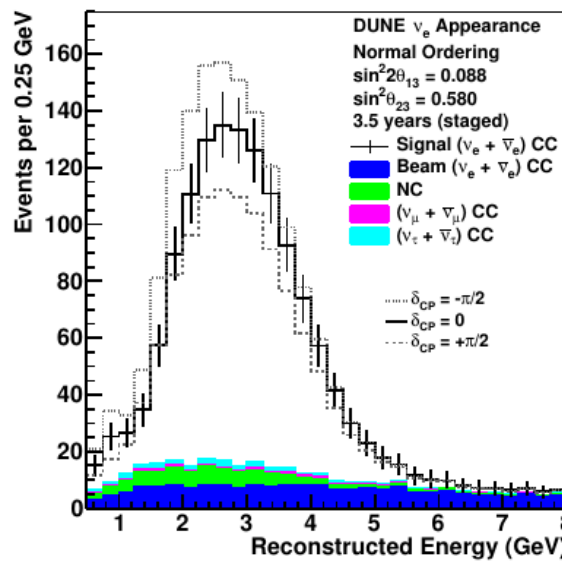
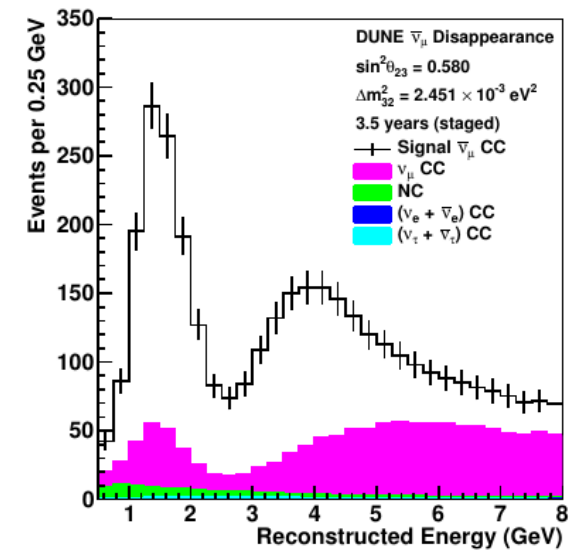
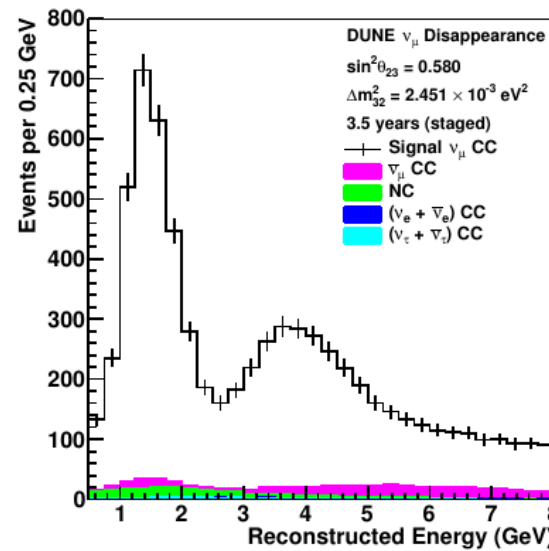
- Sanford Underground Research Facility in Lead, South Dakota
- Four 10-kt Fiducial LAr TPC modules, located 1.48 km underground
- Excavation
- first module operational in 2024
- Start of run: 2 FD modules (20 kt), 1.2 MW beam power, with ND
- +1 year: 3 FD modules (30 kt)
- +3 years: 4 FD modules (40 kt)
- +6 years: upgrade to 2.4 MW beam

Neutrino Oscillations

Measure appearance and disappearance for both neutrino and anti-neutrinos

disentangle Mass Ordering and CP effects

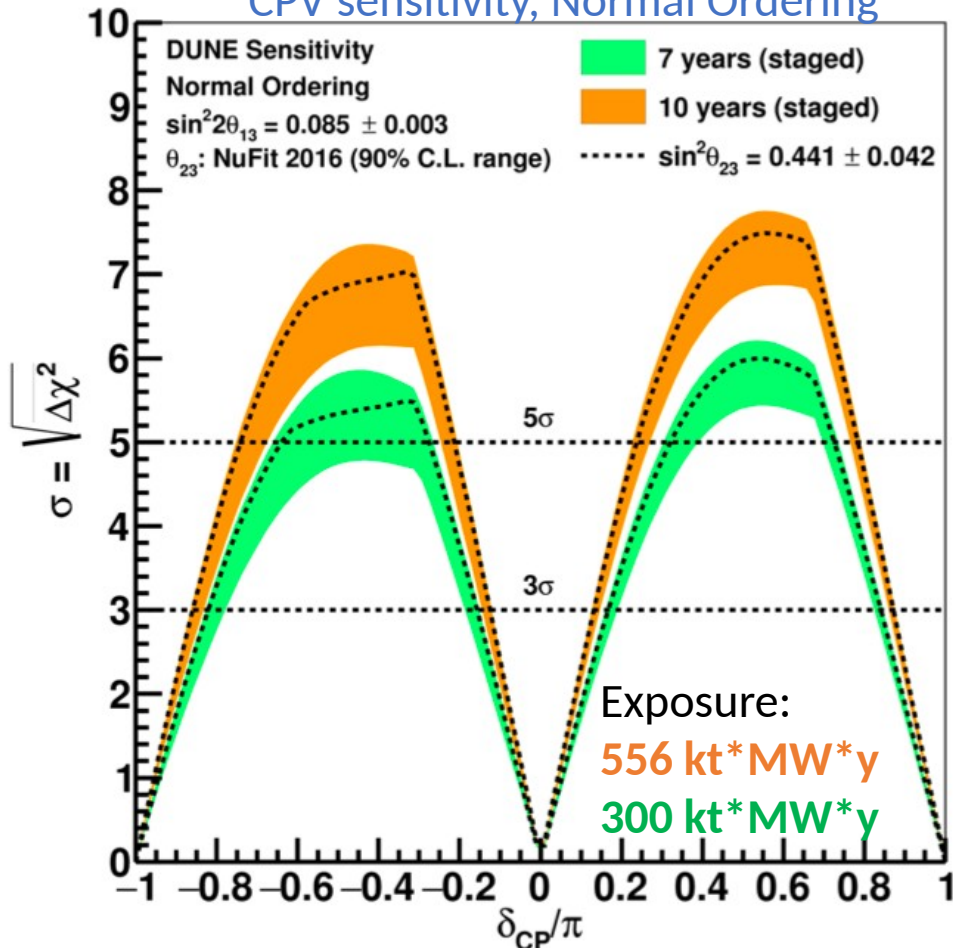
Spectral measurement - 1st and 2nd maxima



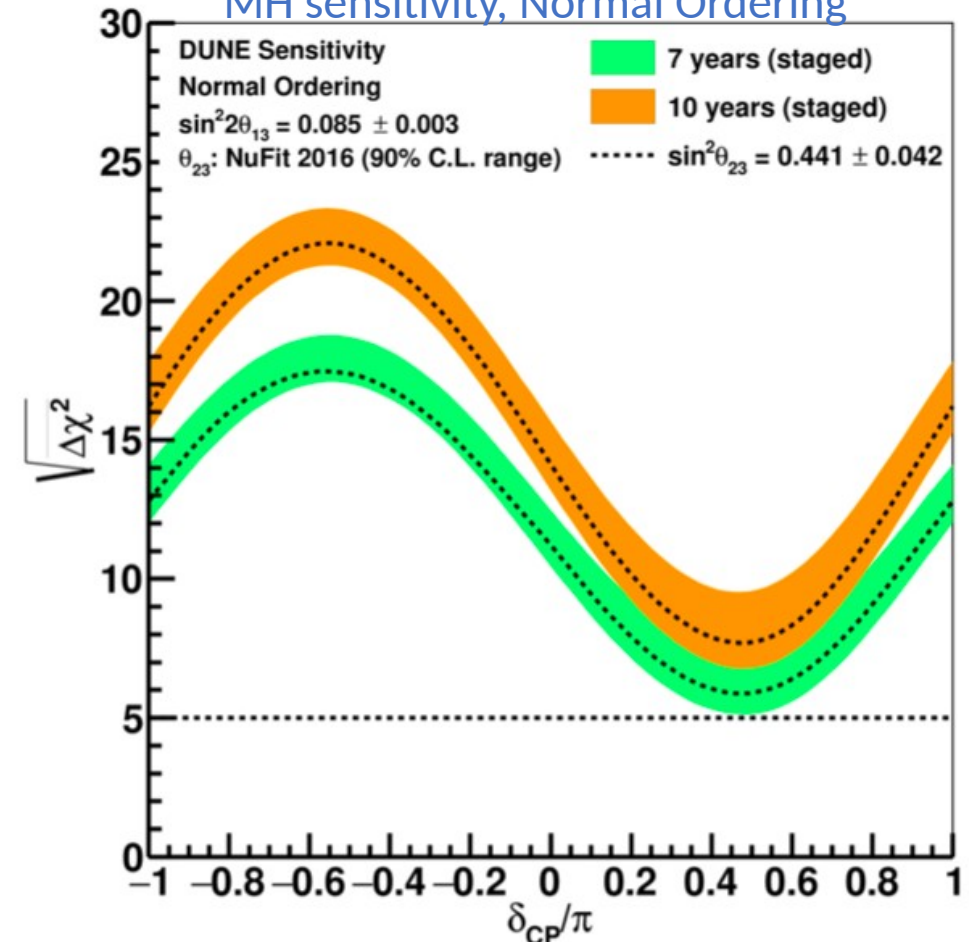
DUNE Technical Design Report (TDR) arXiv:2002.03005

sensitivity to CPV and Mass Ordering

CPV sensitivity, Normal Ordering



MH sensitivity, Normal Ordering



Width of bands indicates variation in possible central values of θ_{23}

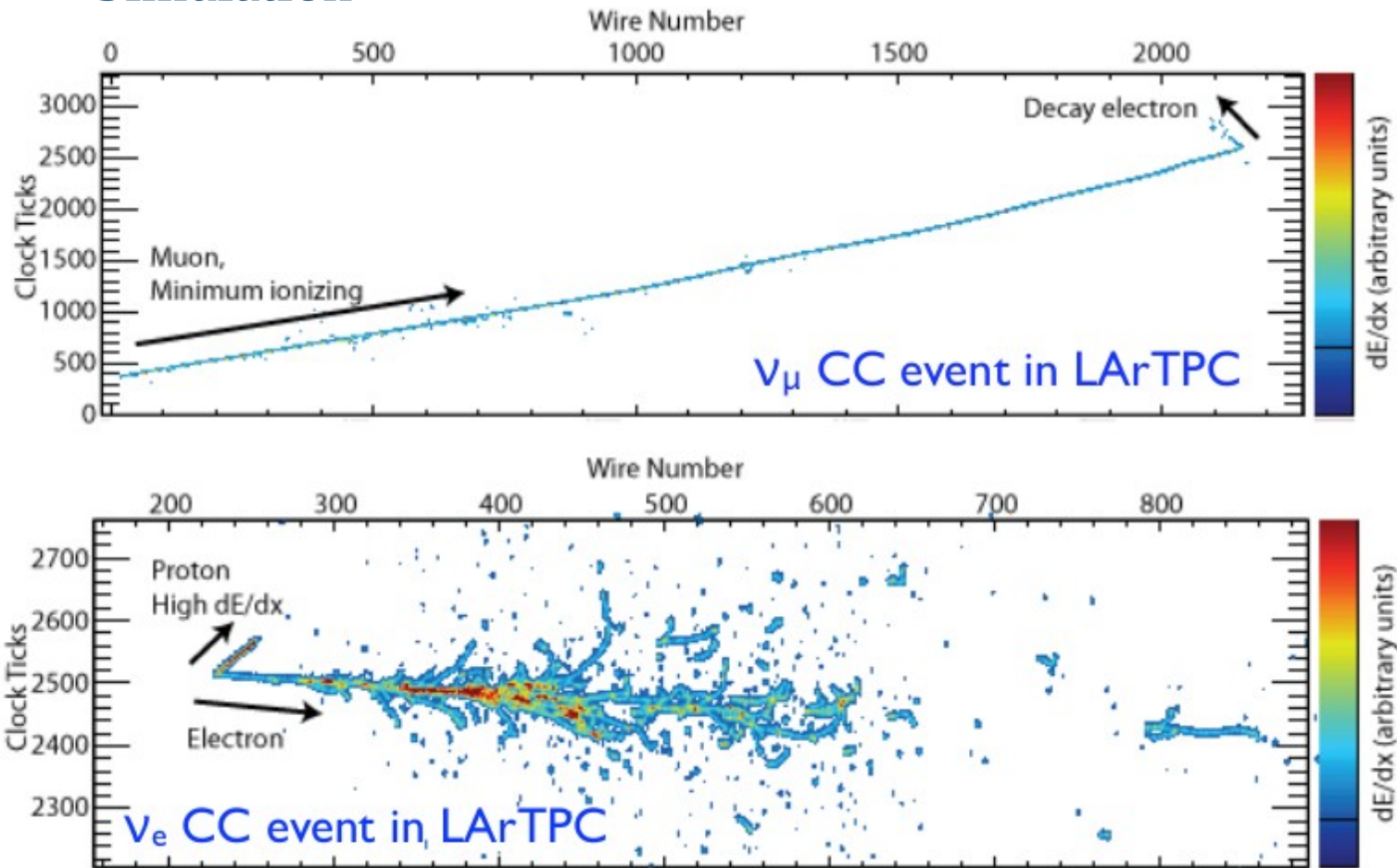
>5 σ sensitivity for both orderings and the full range of δ_{CP}

Updated analysis with full FD simulation & reconstruction, detailed systematic uncertainties, and ND, in TDR

LAr TPCs

- Gigantic detectors deep underground with Excellent calorimetric and spatial resolution
- Wide off-beam physics program

Simulation



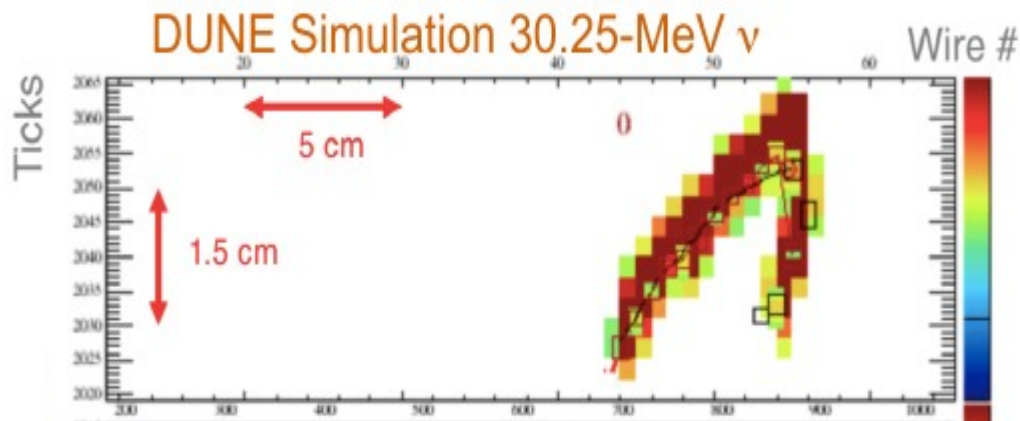
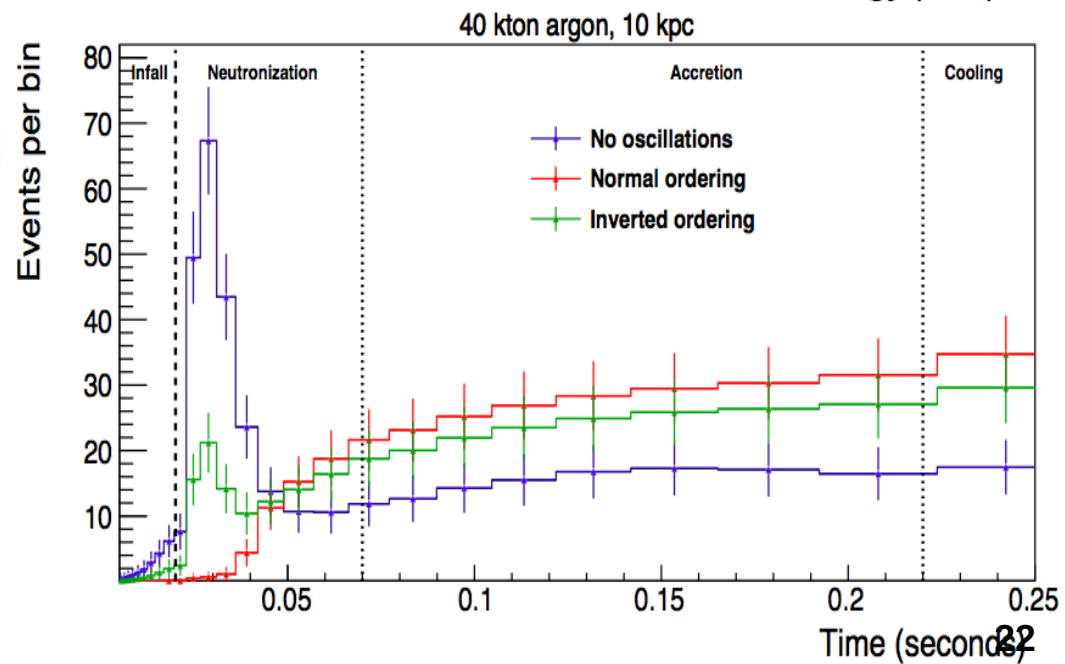
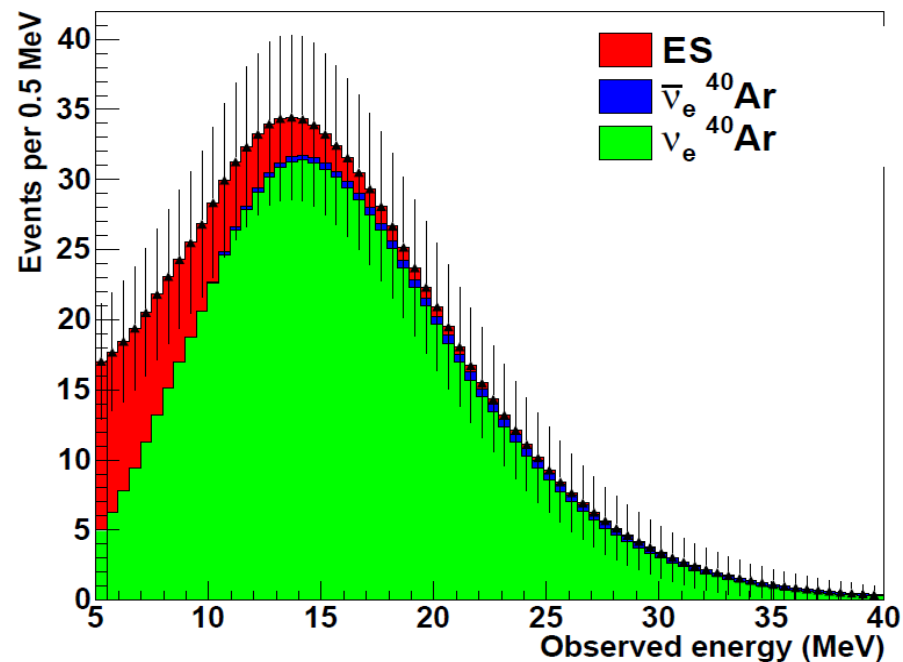
non-beam physics: SuperNova burst

Uniquely DUNE is sensitive to ν_e



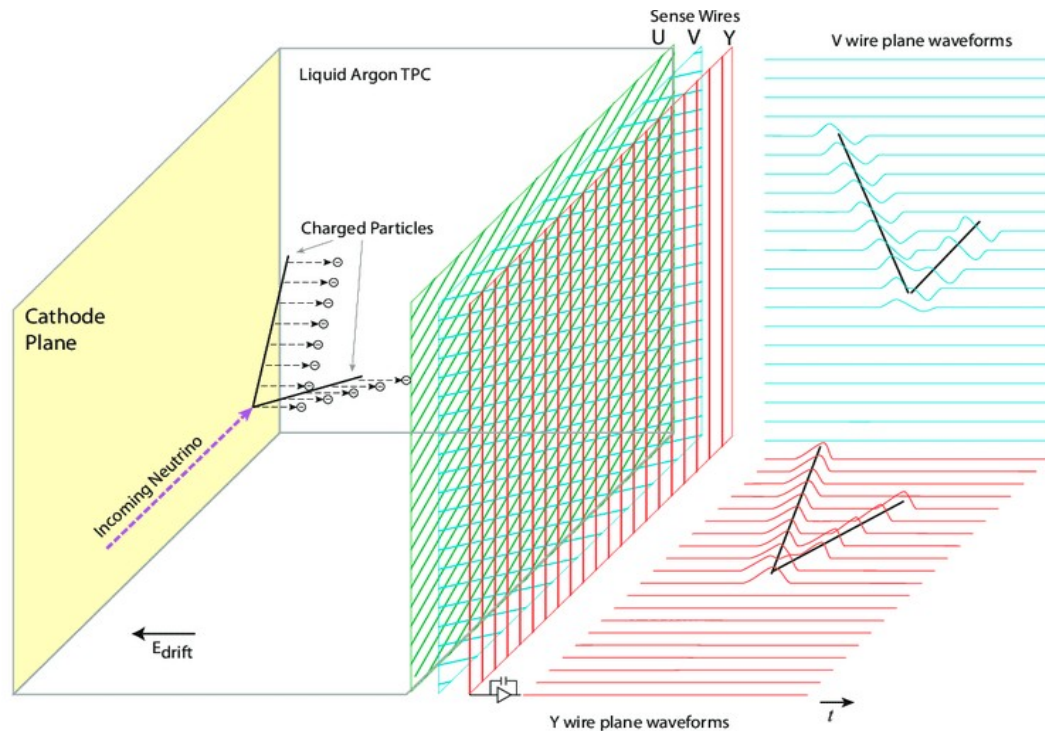
5-50 MeV signals: efficient triggering + continuous data collection

Monte Carlo studies ongoing with Marley generator and full simulation

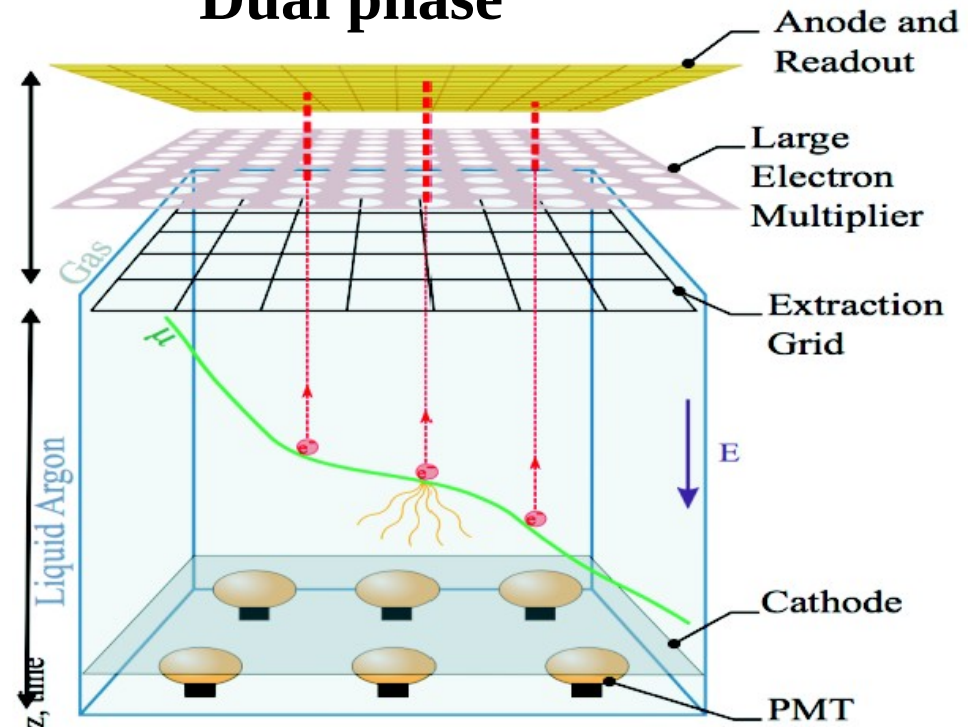


Two detector technologies

Single phase



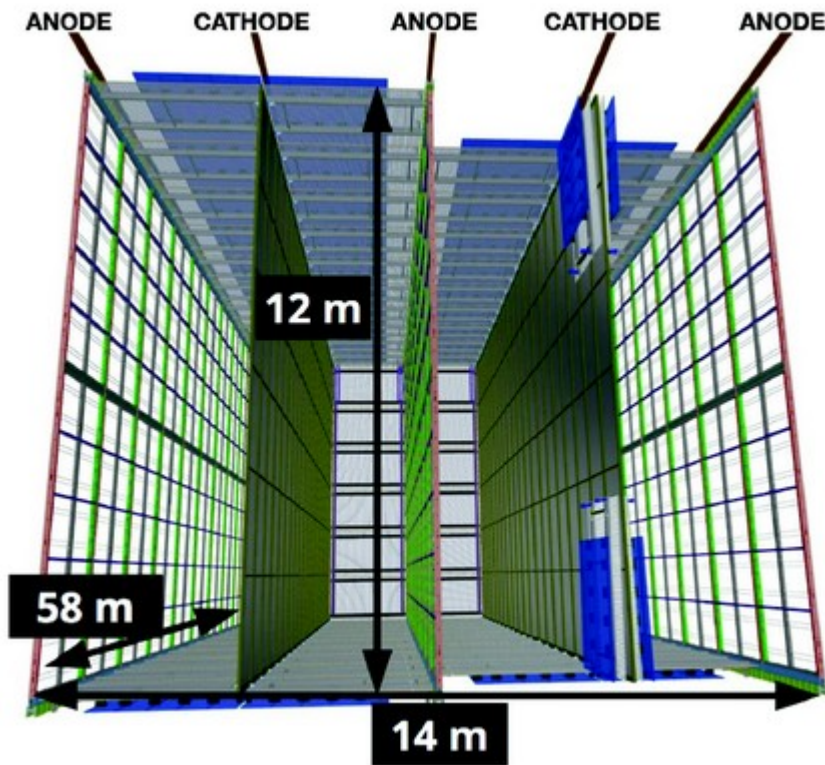
Dual phase



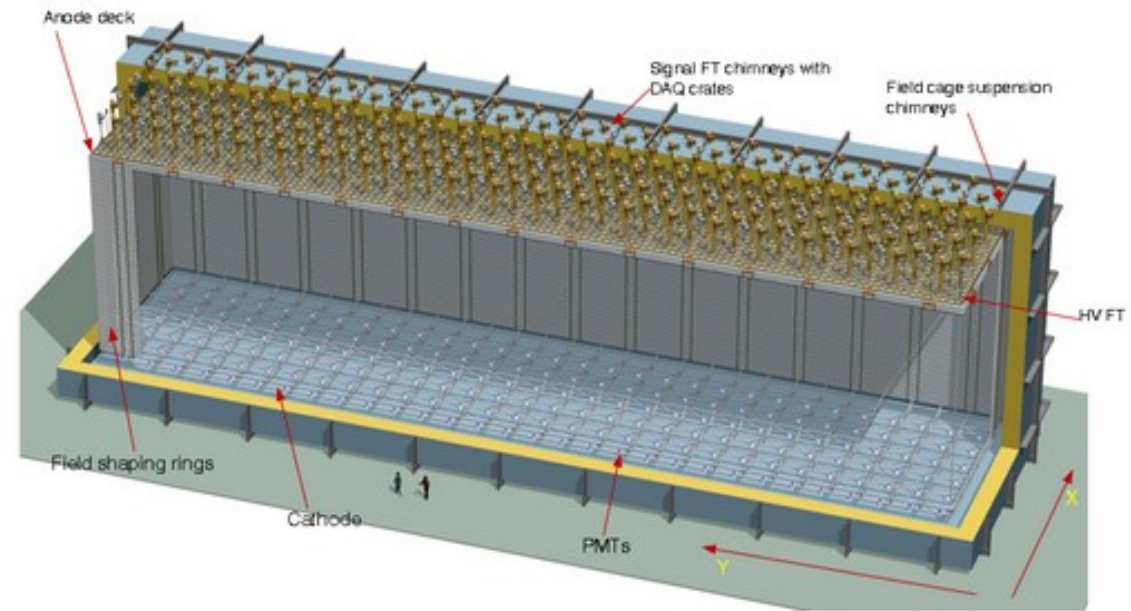
Horizontal drift, distance 3.6 m
 Anode wires immersed in LAr
 Vertical Anode and Cathode Plane Assemblies (APA, CPA)
 1 collection, 2 induction wires at 37.7° , wire pitch 5 mm
 Photon detectors: light guides + SiPM, embedded in APAs

Vertical drift, distance 12 m
 Ionization electrons extracted from LAr to gas
 Signal amplified in GAr by Large Electron Multiplier (LEM)
 Charge collected on 2 orthogonal views, 3mm pitch
 Photon detectors: PMTs below the cathode

Far Detectors



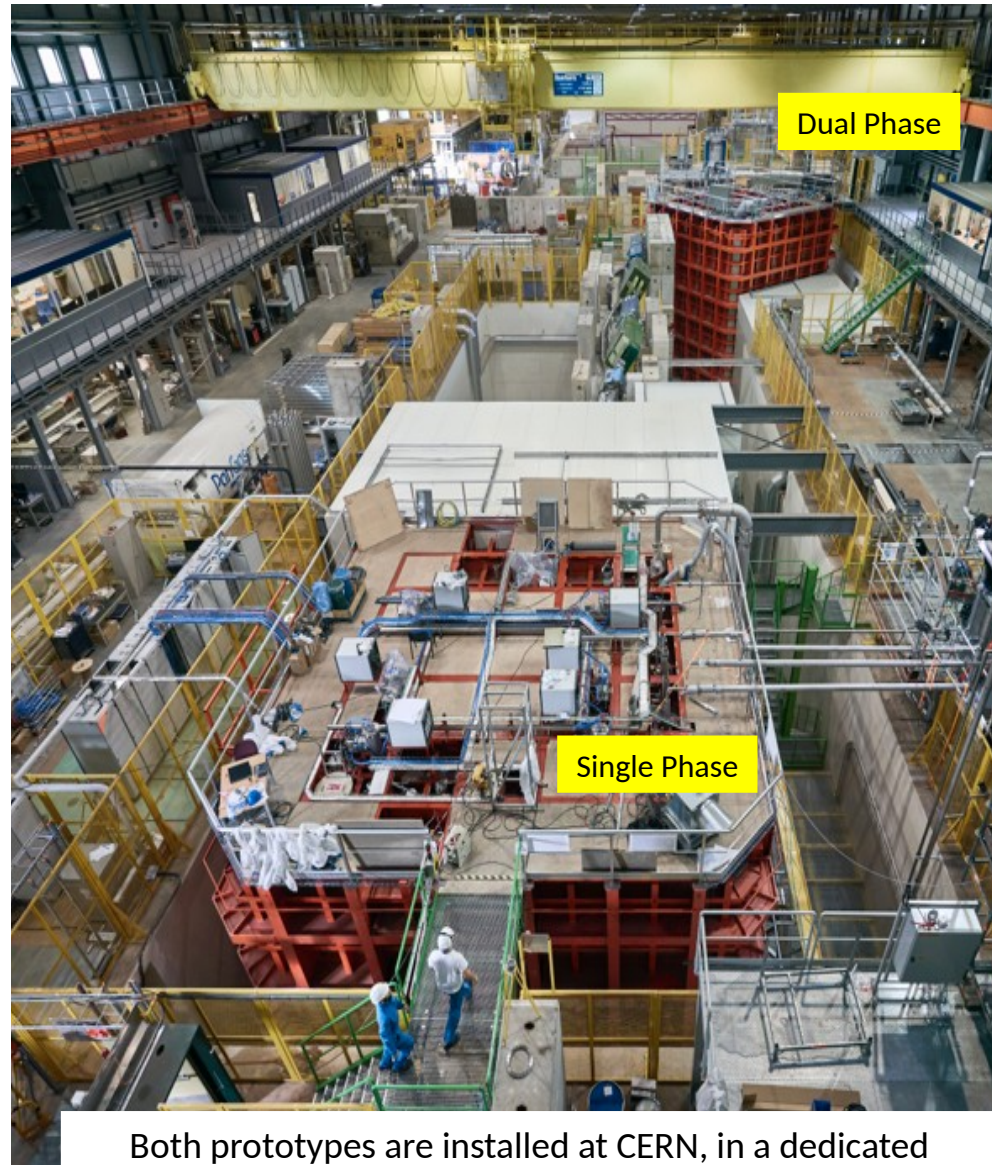
- Single Phase is modular
- Shorter drift lengths (and distance to photodetectors)
- Immersed cold electronics (inaccessible)



- Dual Phase is homogenous
- Drift is longer (12m) (and distance to photodetectors)
- Cold Electronics at the top of the detector (accessible)

protoDUNEs

- Giant Liquid Argon TPCs (LArTPC) – one Single-Phase, one Dual-Phase hosted at the CERN neutrino platform
- Necessary R&D step towards the DUNE Far Detectors
 - Tests of all engineering solutions and installation procedures
 - Use full-size components identical to those planned for DUNE FD
- 300t fiducial mass of LAr
 - Technology demonstrators
 - Demonstrate long term performance and stability
- Charged particle test beams to characterise detector response over the energy range of interest for DUNE (~ 0.5 GeV to 8 GeV)



Both prototypes are installed at CERN, in a dedicated extension of the North Area

A history of Single Phase LAr TPCs

ICARUS T-600 @ CNGS (2010-2012, 760 tons LAr)



Successfully reconstructed neutrino events from CNGS beam (~ 17 GeV)

Argoneut @ FNAL (2009-2010, 240 kg LAr)



Small TPC, precise measurements of cross-sections and neutrino interactions

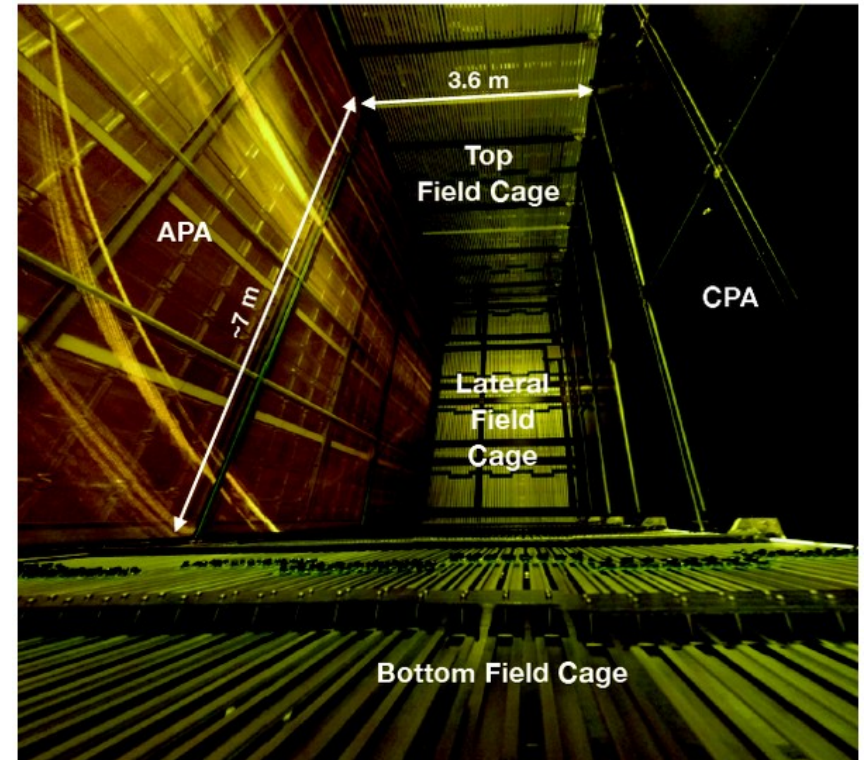
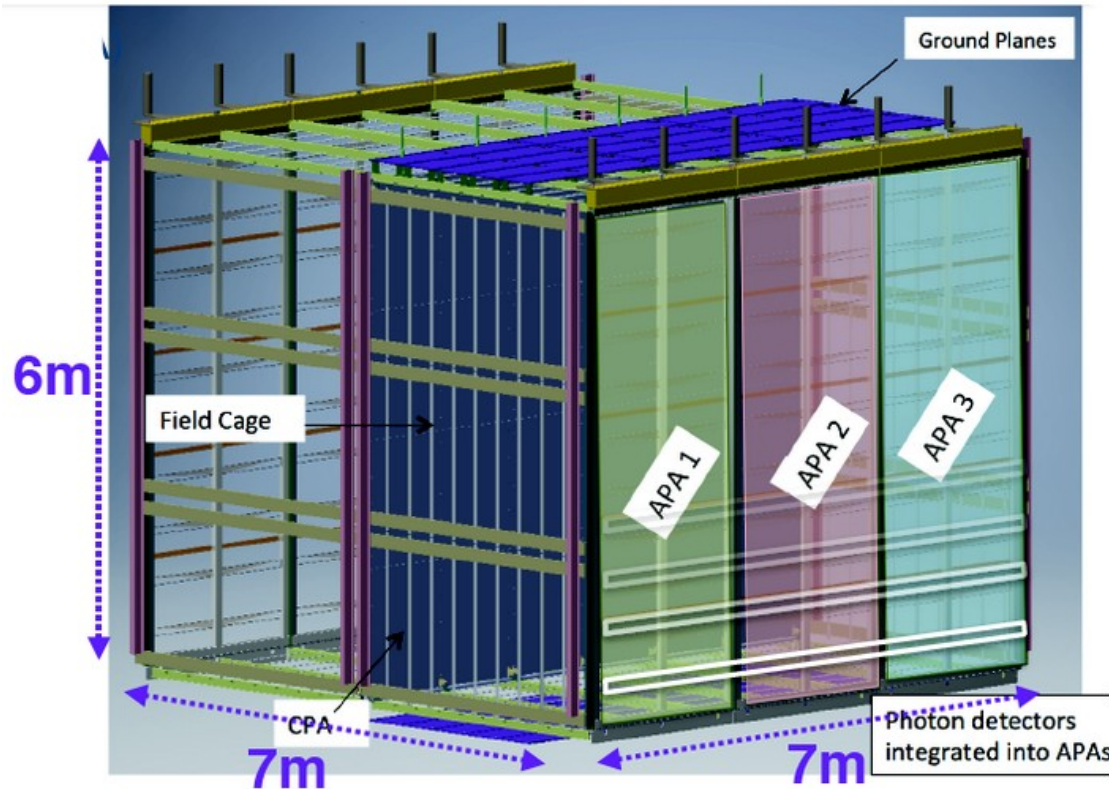
MicroBooNE @ FNAL (2015-ongoing , 170 tons LAr)



Sterile neutrino search. Neutrino event selection and reconstruction. Leads to protoDUNE Single Phase

taken from A. Chatterjee

Single Phase protoDUNE

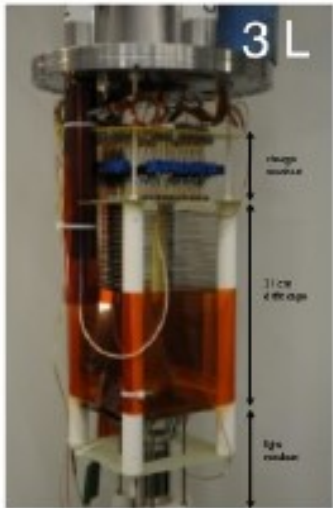


Successful Beam test in 2018 (Sept-Nov) – pions, protons, kaons from 0.3-7GeV (~4M triggers)

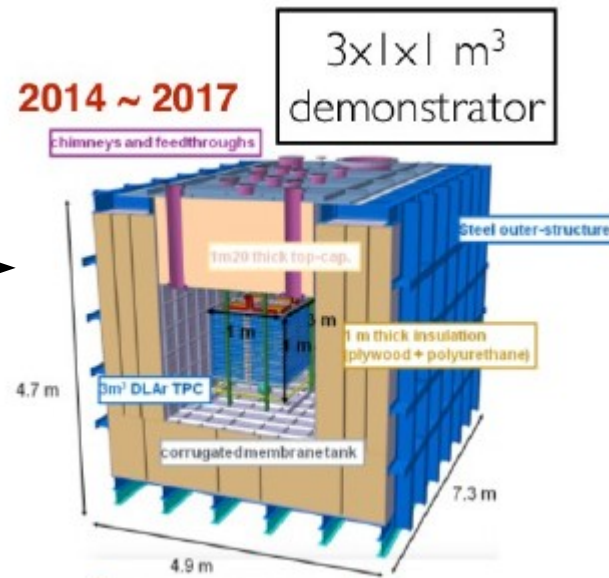
Achieved stable running, >5ms electron lifetime, S/N > 20

A history of Dual Phase LAr TPCs

Small R&D TPCs



2007 ~ 2014

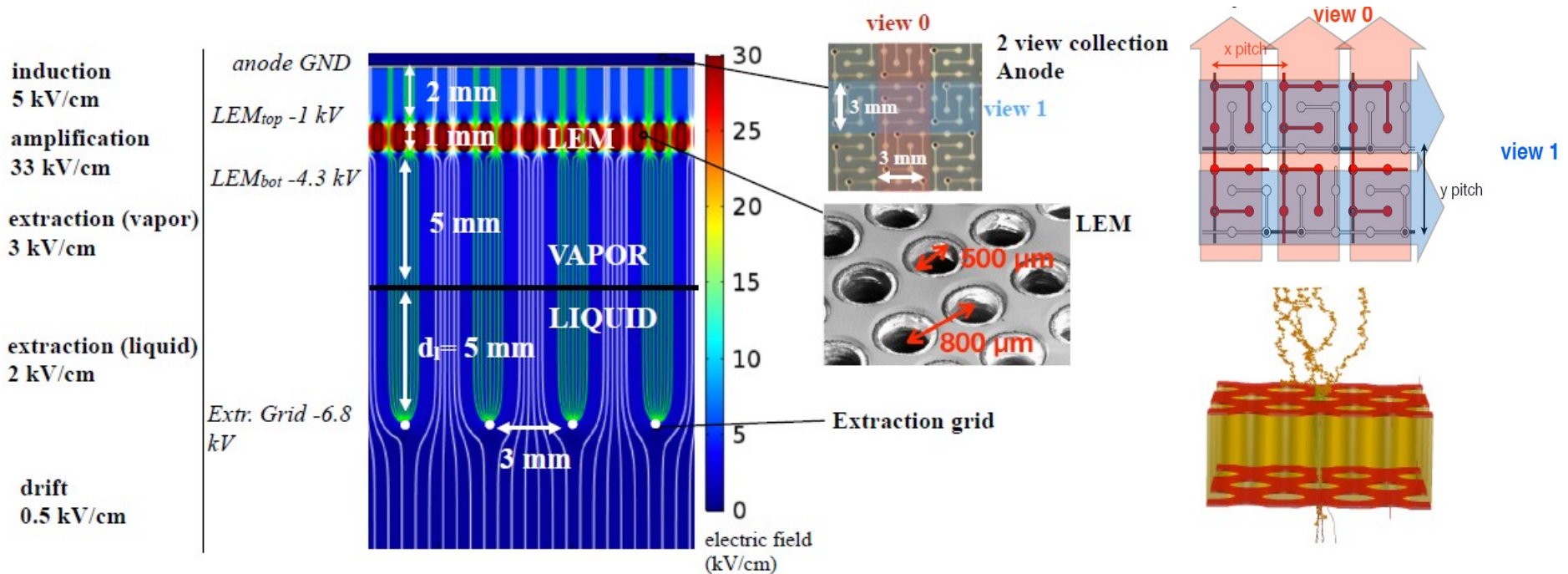


ProtoDUNE DP
6m x 6m x 6m

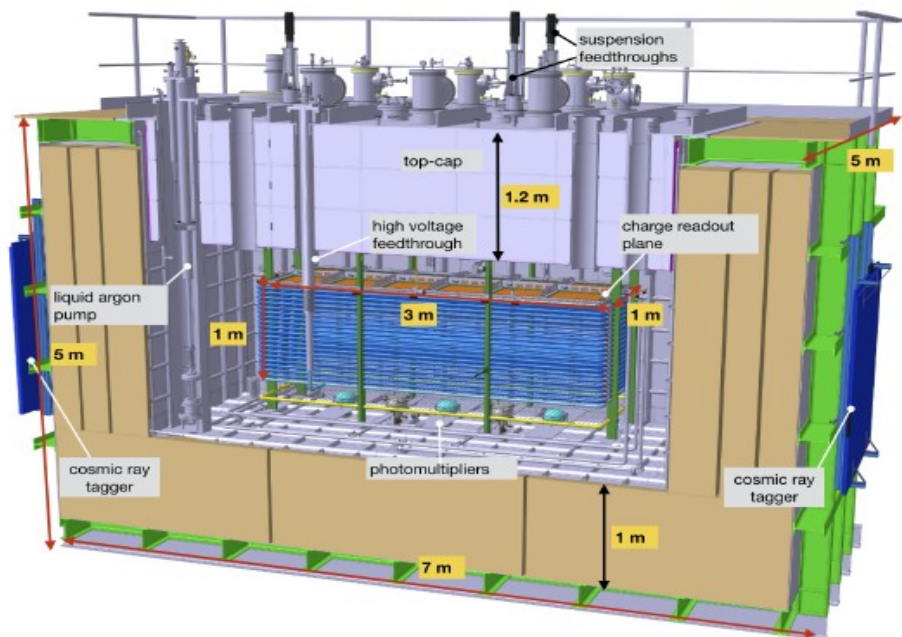
Charge Read Out Planes

Each CRP is 3m x 3m and contains:

- 1 x Extraction Grid
- 36 x Large Electron Multipliers (0.5m x 0.5m)
- 36 x Anodes



Dual-Phase 4t demonstrator



3x1x1 m³ (4t) fiducial volume,
installed at CERN

500k cosmic events in summer-fall 2017

First results on

- charge amplification
- light detection
- LAr purity

[arXiv:1806.03317 \[ins-det\]](https://arxiv.org/abs/1806.03317)

B. Aimard et al, 2018, JINST 13, P11003

Technical problems encountered with

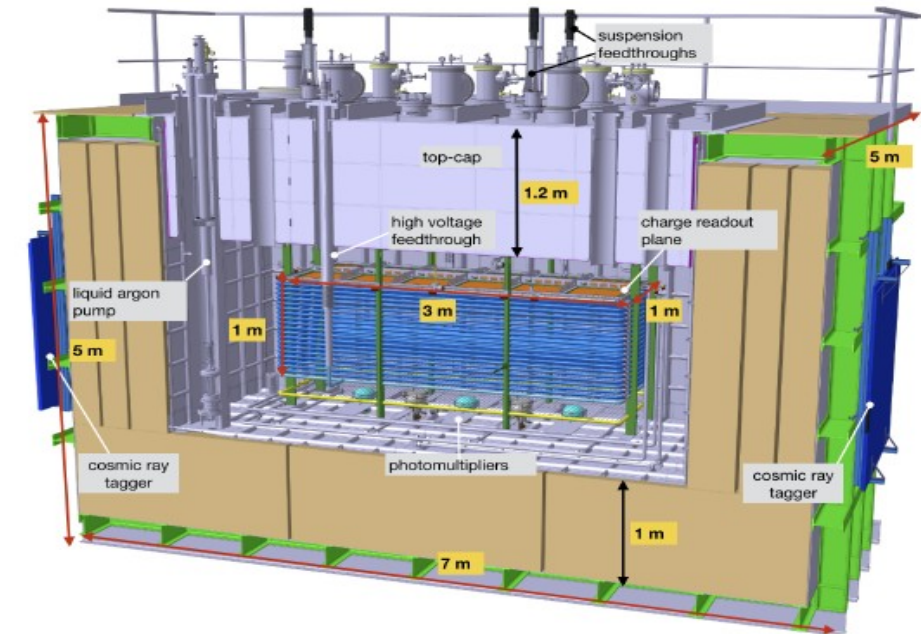
- Extraction Grid
- LEM HV

Gain needed for DUNE could not be demonstrated

Issues addressed for protoDUNE



Dual-Phase 4t demonstrator

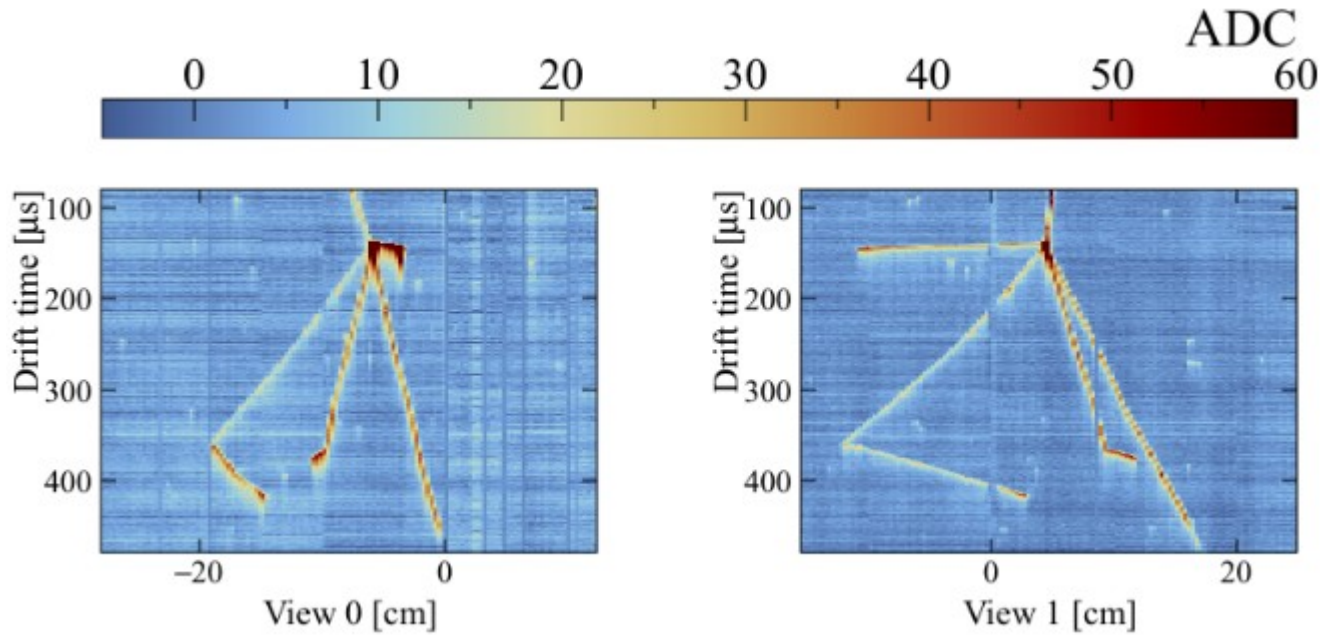


Several milestones achieved

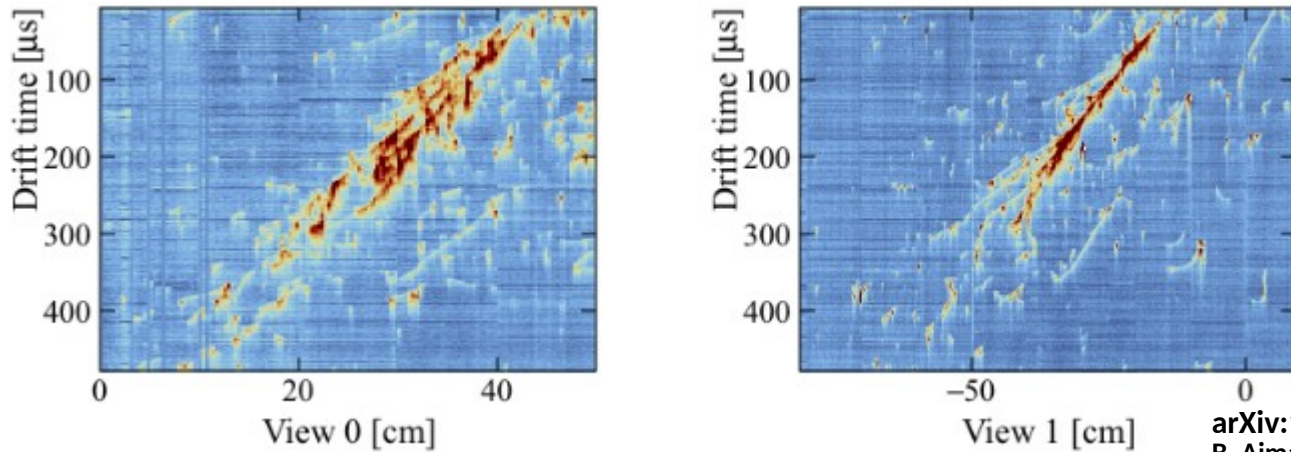
- LAr level stability over time
- Stable drift field for entire operation
- Equal charge splitting at the anode
- Purity compatible with a 4ms electron lifetime



Dual-Phase 4t demonstrator



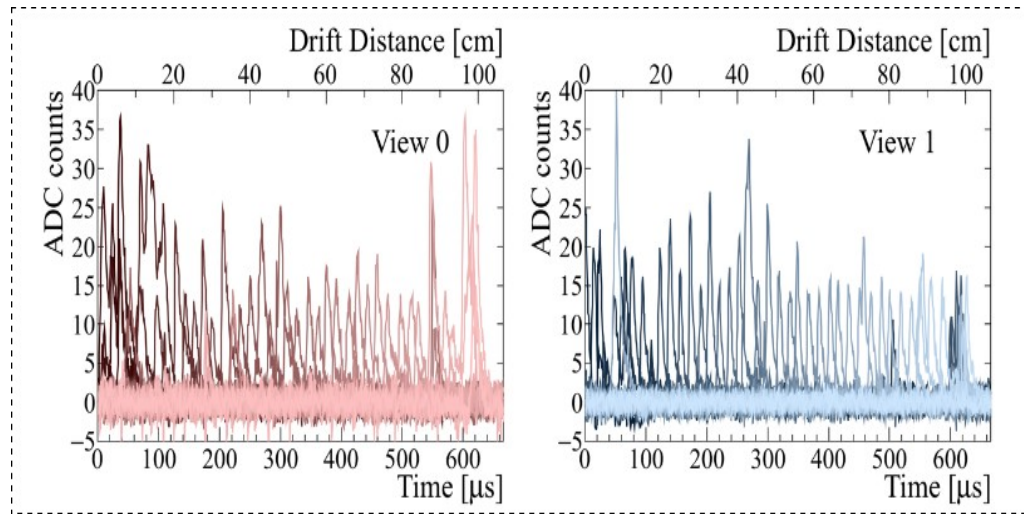
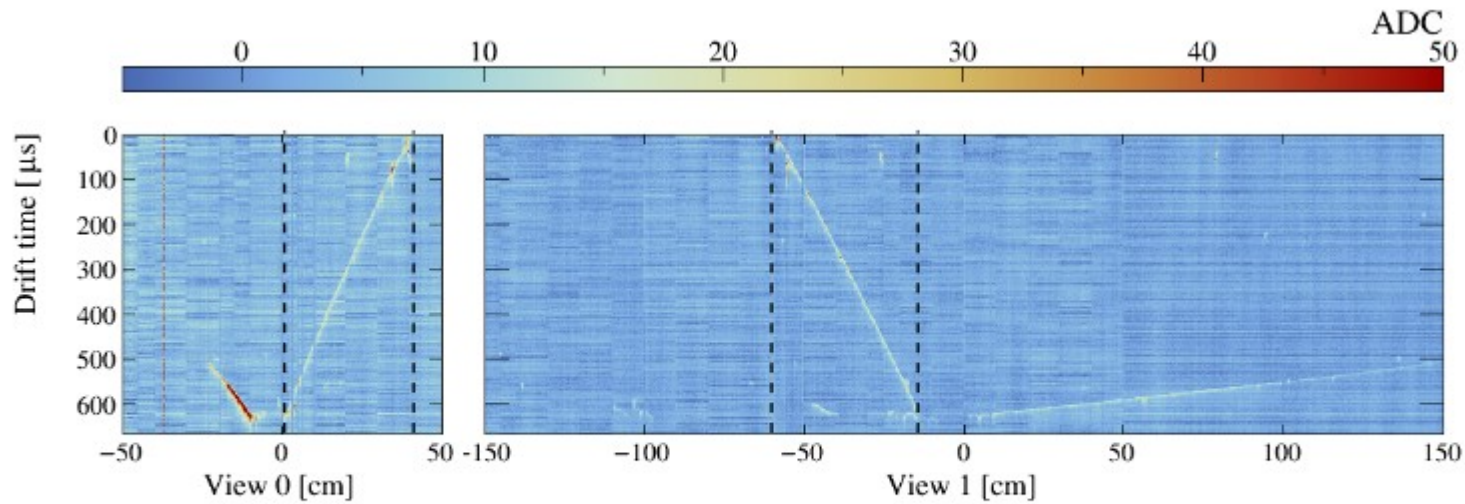
Hadronic shower



Electronic shower

arXiv:1806.03317 [ins-det]
B. Aimard et al, 2018, JINST 13, P11003

Dual-Phase 4t demonstrator

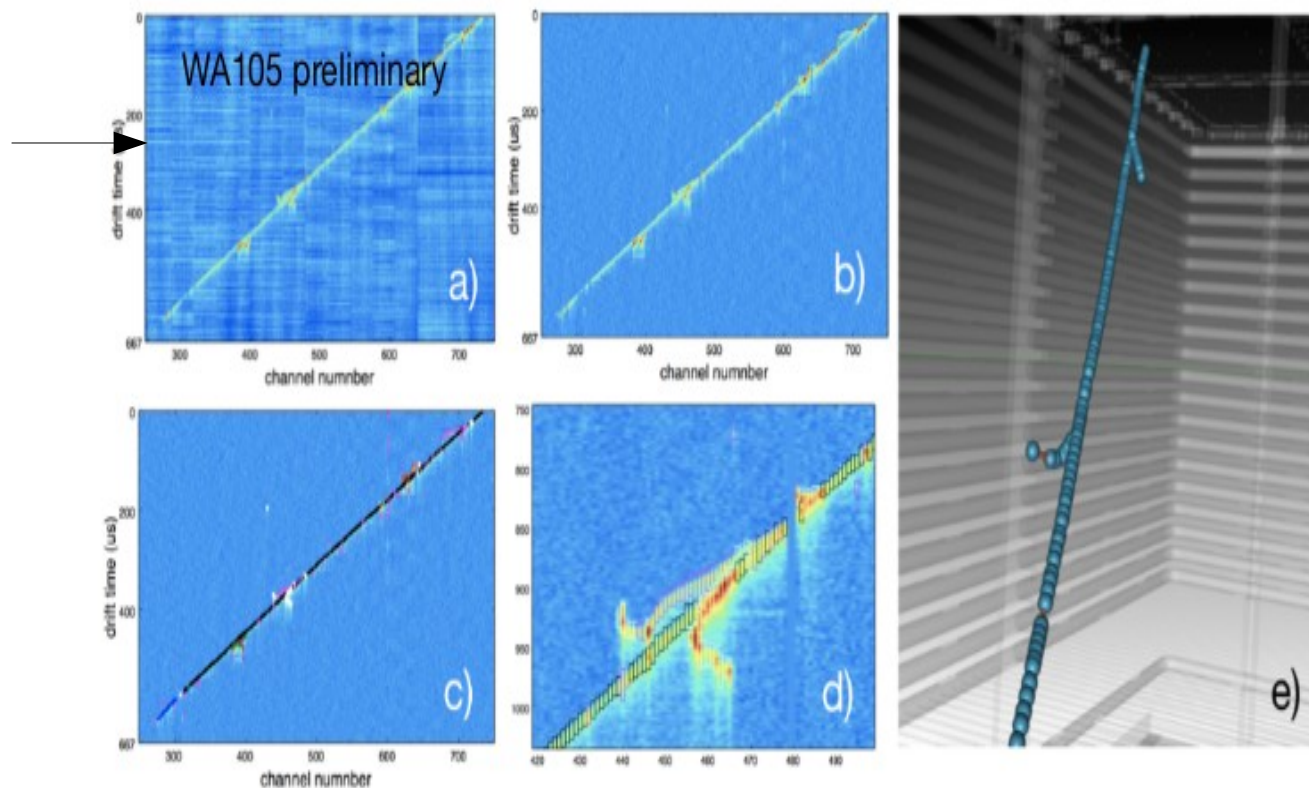


Through-going muon

arXiv:1806.03317 [ins-det]
B. Aimard et al, 2018, JINST 13, P11003

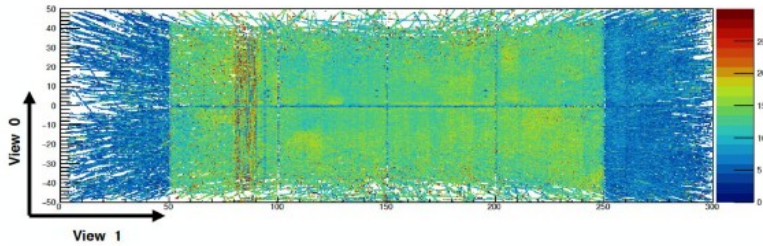
Dual-Phase 4t demonstrator

Noise filtering
3D reconstruction
Track selection

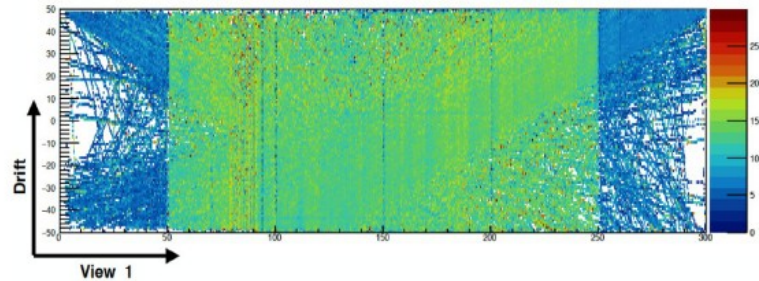


Dual-Phase 4t demonstrator

$\langle dQ/ds \rangle$ uniformity across the CRP



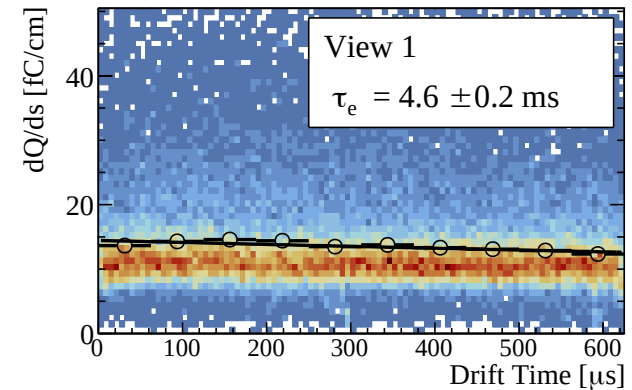
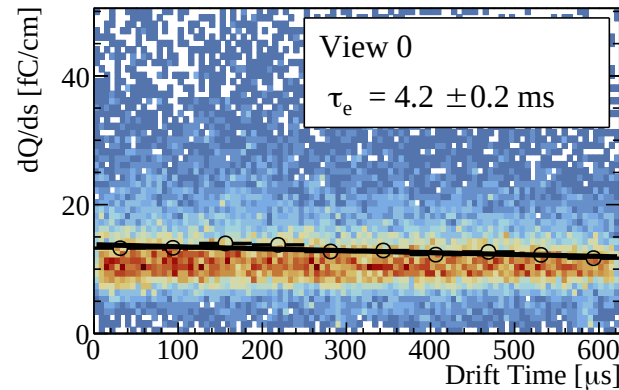
$\langle dQ/ds \rangle$ uniformity along the drift



500k cosmic events in summer-fall 2017
example crossing muons

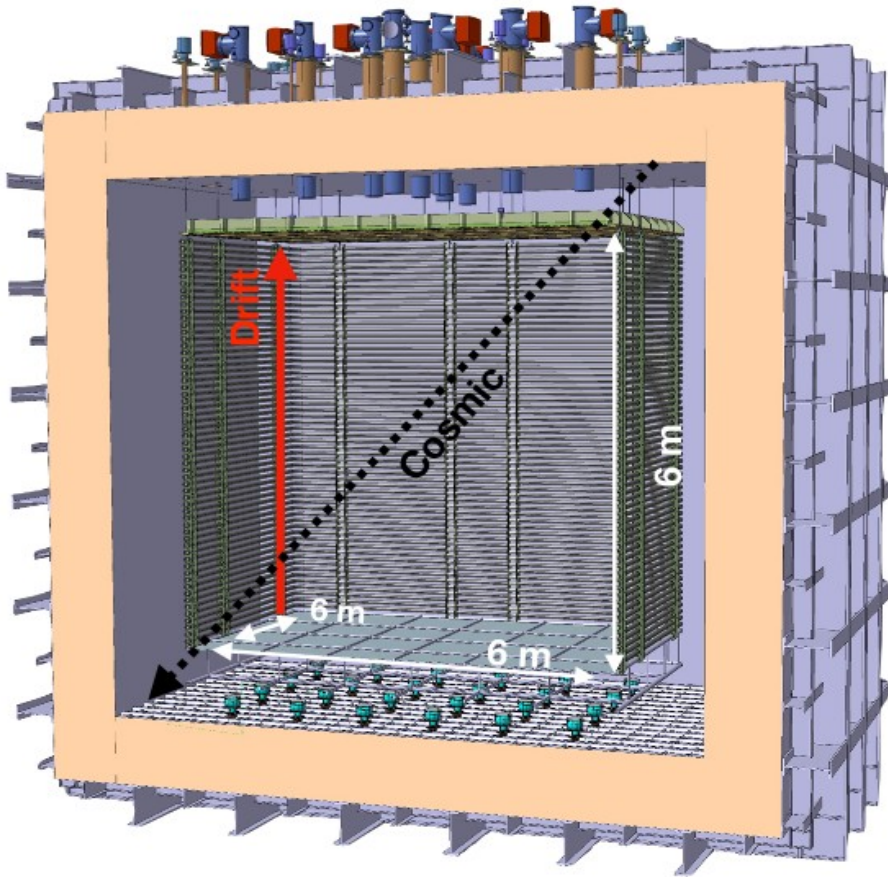
Using through-going muon tracks (top to bottom)
And 3D reconstruction
Calculate dQ/ds for each hit of track

Indicates electron lifetime



ProtoDUNE Dual-Phase





Fiducial mass is 300t

Half the drift length of DUNE FD

Drift field 0.5 kV/cm \rightarrow -300 kV on cathode

Expected S/N $>$ 20

36 photomultipliers

DUNE Interim Design Report
Dual Phase - arXiv:1807.10340

Drift Field



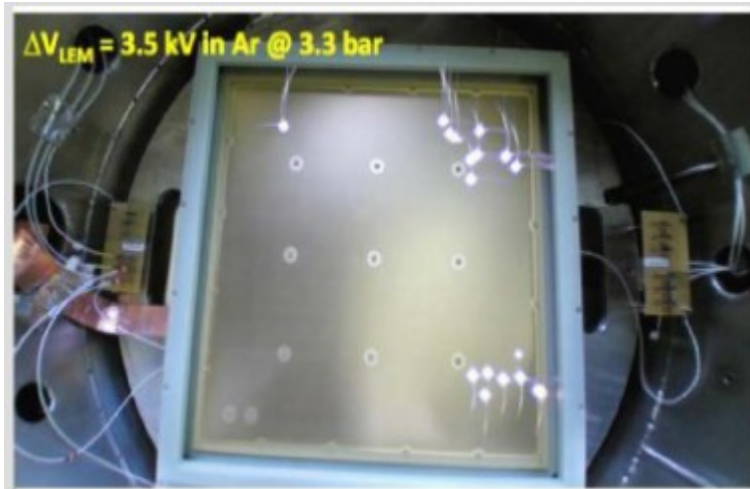
Drift field objective is 0.5 kV/cm
-300 kV on Dual-Phase cathode
Cathode design is modular (3m x 3m)
60% optical transparency
Maximum local field requirement
<30kV/cm

High Voltage feedthrough based on
ICARUS design same from Single/Dual-
phase

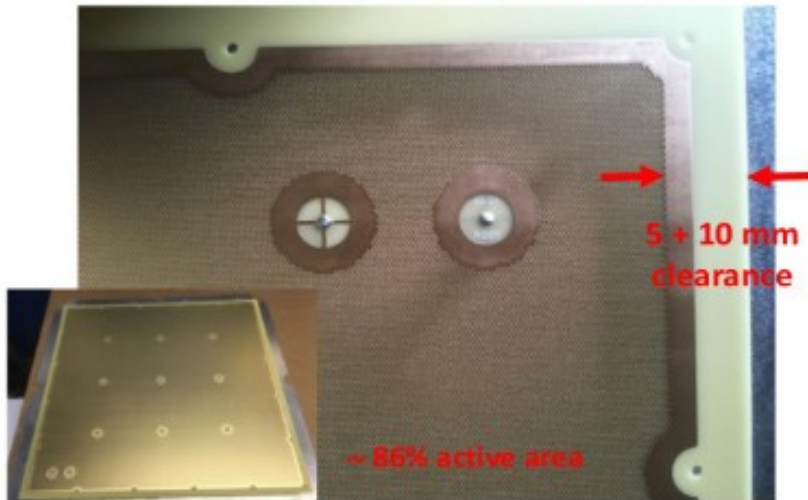
HV extender-degrader delivers HV to field
rings and cathode

- April 2018 : field cage completed, stable operation at 150 kV in middle of cathode, ground on top and bottom
- Demonstrated OK for 0.5 kV/cm

Large Electron Multiplier



CFR-35 – NP02



LEMs are tested in gaseous argon
3 types have been tested (CFR-34, 35, 36)

Arrived at a LEM design:
No trips > 64h
No sparks
Effective gain > 20

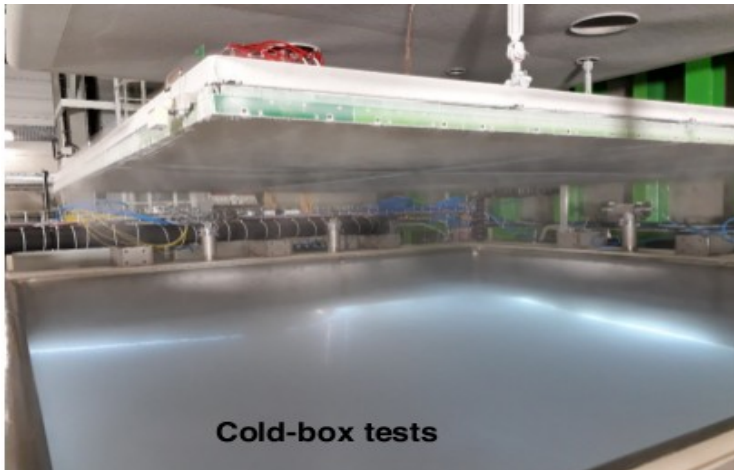
Further optimisation of active area

Charge Readout Planes



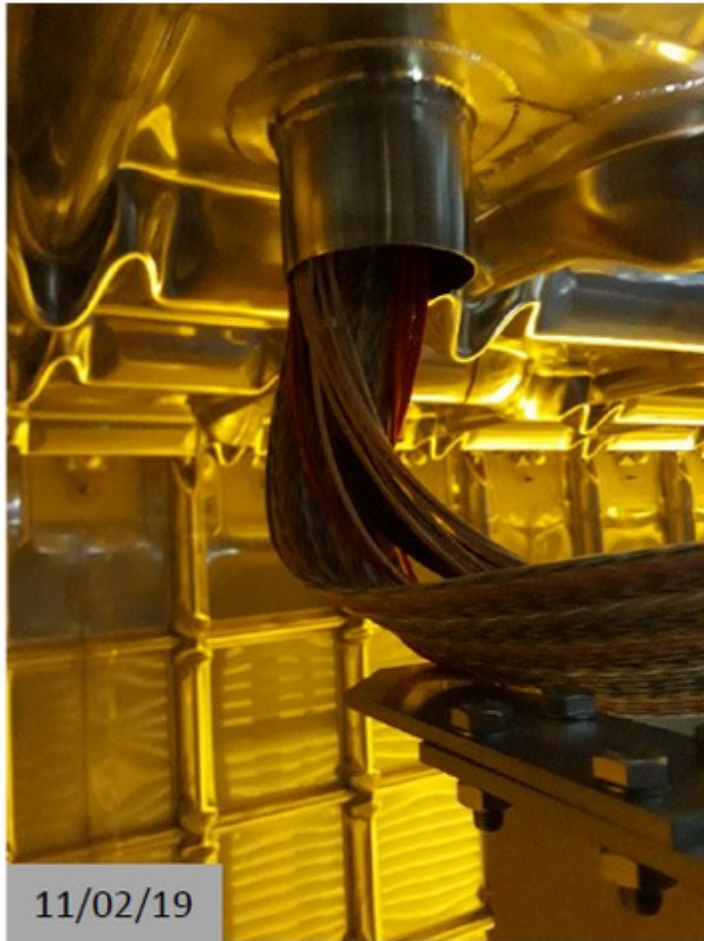
CRP assembly in clean room

- June 2018 : first Charge Readout Plane (CRP) assembly completed
- August 2018 : first CRP cold-box test completed



Cold-box tests





Assembly completed March 2019
Purging, cooling, filling completed August 2019

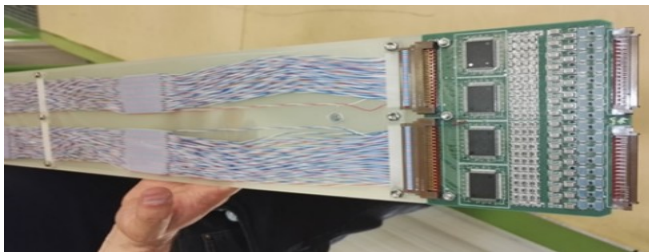
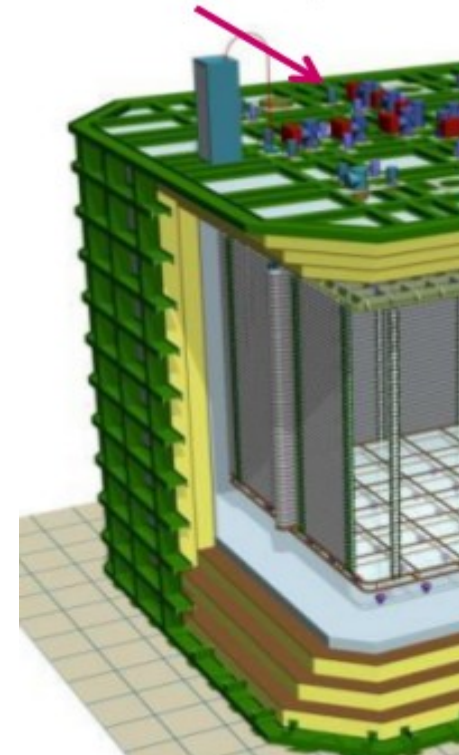
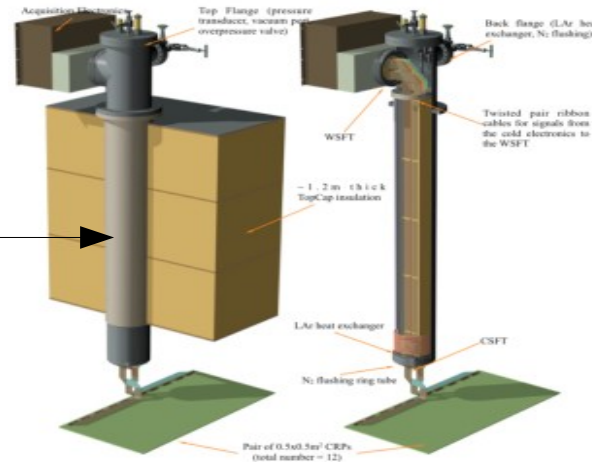
Accessible Cold Electronics

Specially developed signal chimneys

- 16-channel Cryogenic ASIC amplifiers close to the anodes but are also externally accessible!

On the tank deck (warm): digital electronics

- Based on uTCA standard
- 1 uTCA crate/ signal chimney
- AMC card – 64 channels, 12 bit ADC, 2.5 MHz
- 10 AMC cards (64 channels/card)
- Total of 12 uTCA crates (7,680 channels)

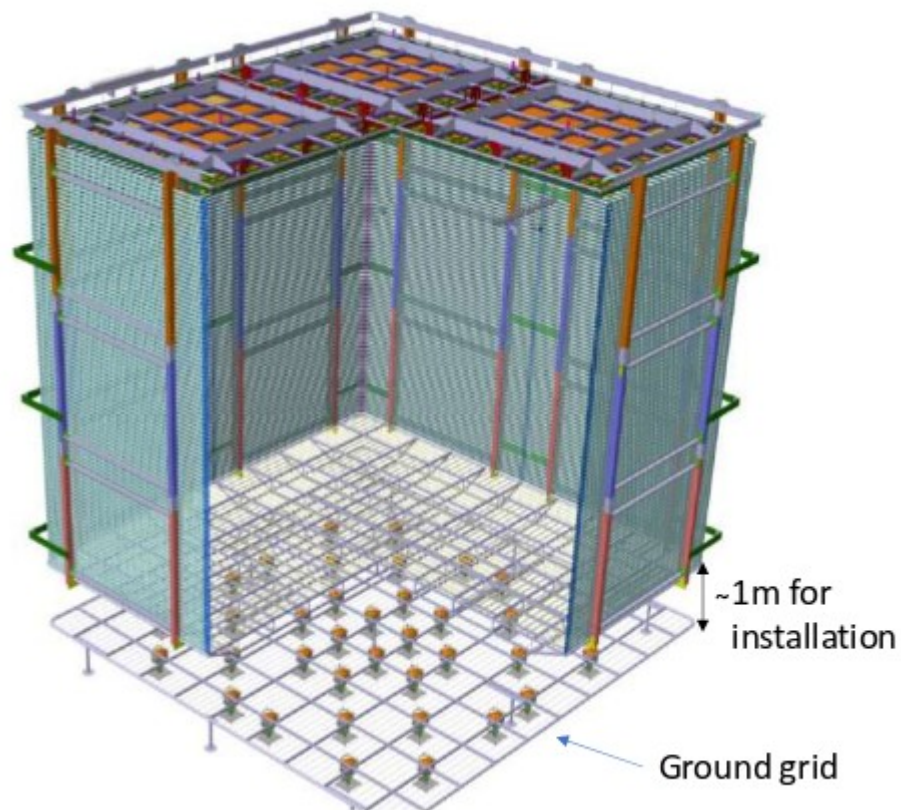
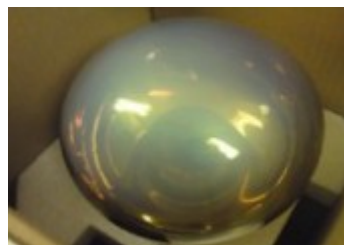


Successfully tested on the 4t demonstrator

Light Read Out

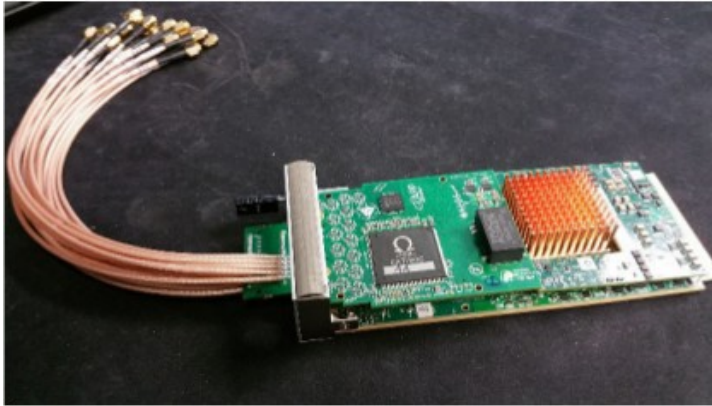
Light signal gives the t_0 of an event
→ determines the z co-ordinate

36 PMTs (Hamamatsu R5912-02
MOD 8 inch) coated with TPB – as
used in 4t demonstrator



Light Read Out

1uTCA crate for Light Read-Out electronics
Designed to integrate into Global DAQ



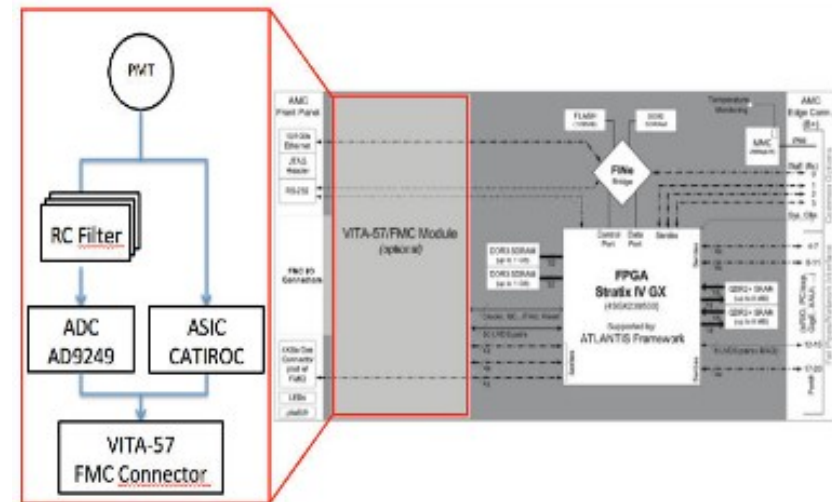
Under-development
16 channels
Anti-aliasing low pass filter

ADC: AD9249 65 MHz (max), 14 bits
provides waveform with a window of $\pm 4\text{ms}$
around the external trigger down-sampled
to 400 ns

ASIC: CATIROC
Provides auto-triggered channel-wise Q, t

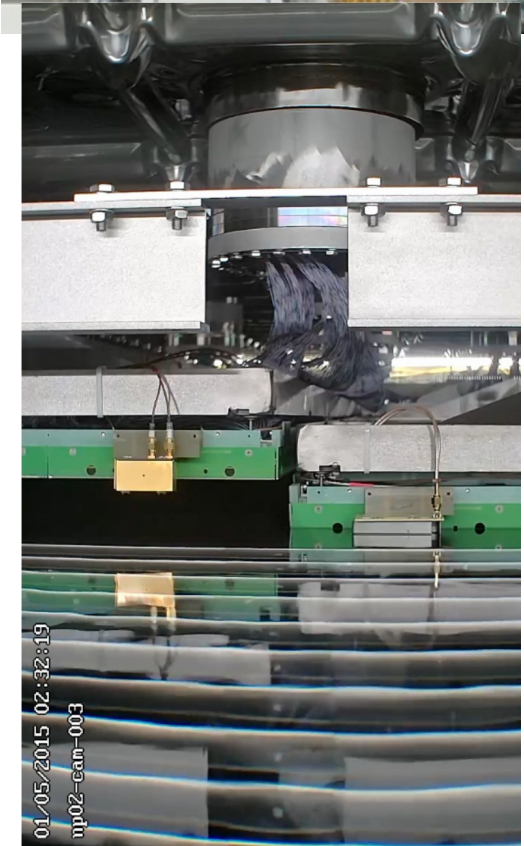
Readout time of Dual Phase is $\leq 4\text{ms}$
Light Read Out needs to timestamp
cosmic ray muons occurring before and
after trigger

ASIC – measurements at single ph.e. level
(calibration, low energy signals)



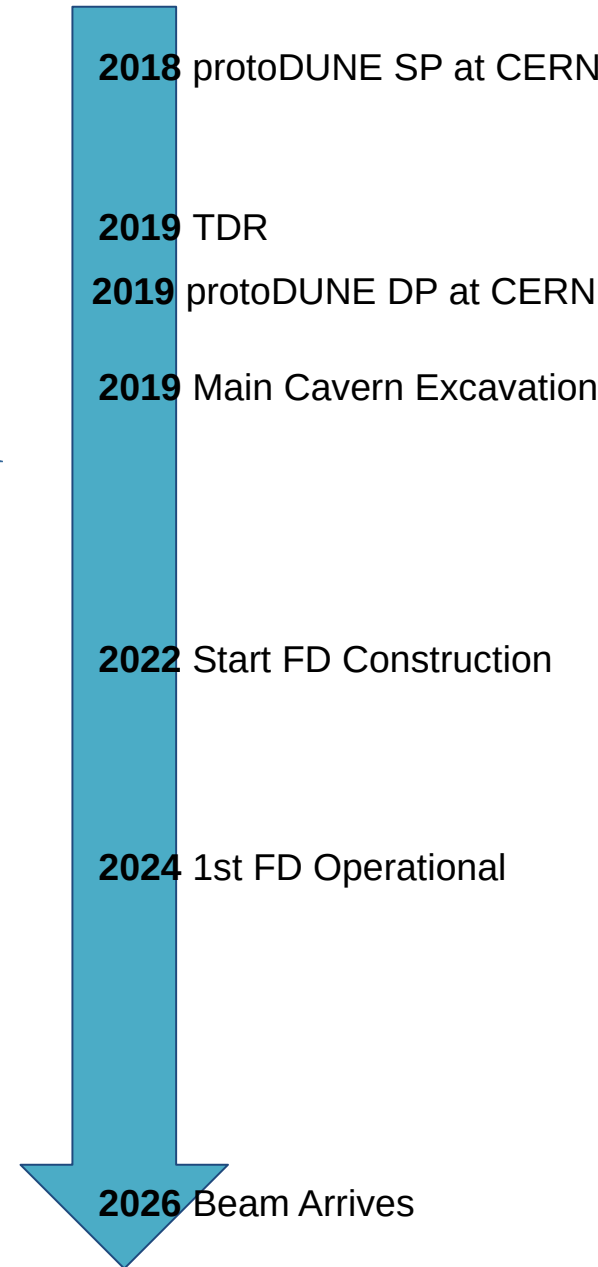
Dual Phase Status

- ProtoDUNE Dual Phase has been operated from September 2019
 - Experiencing some technical issues
- Runs of Cosmic Ray muon data (~ 7 kHz muon rate)
- With CR muons
 - We can explore 3D response of detector
 - Determine operational parameters such as gain, purity etc
 - Demonstrate track reconstruction
 - Search for Michel electrons (low energy signal)
 - And more..



Summary

- DUNE will measure neutrino and anti-neutrino oscillations over a 1300 km baseline with gigantic (4x10kt fiducial) LAr TPCs
- Large underground detectors will provide other interesting physics (Nucleon decay, neutron-antineutron oscillations, SN etc)
- Large-scale detector prototypes, ProtoDUNE, are being tested at CERN
- DUNE TDR is on the archive
arXiv:2002.03010 arXiv:2002.03008
arXiv:2002.02967 arXiv:2002.03005
volume 5 (Dual Phase) to come



Backups

Dual Phase

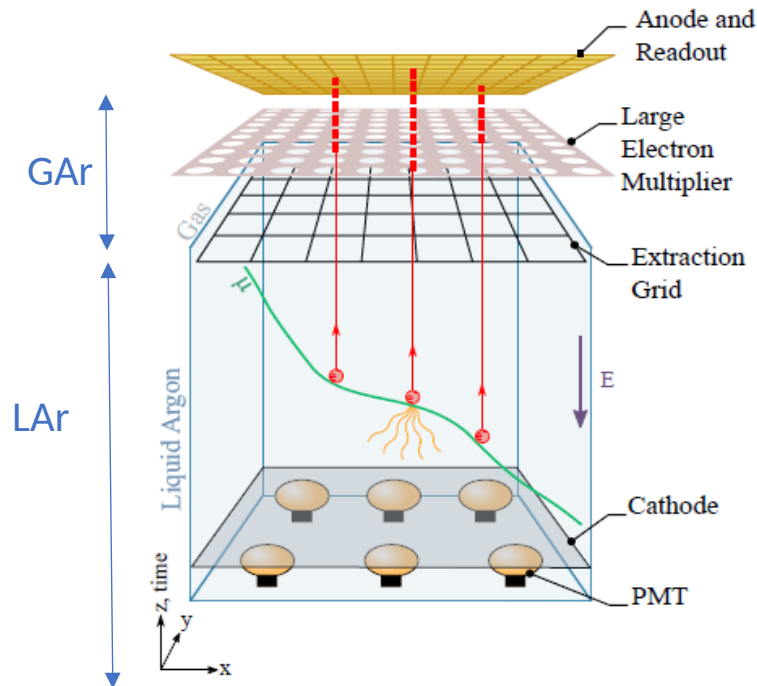
In Liquid

Ionization: ~8k electrons/mm from a mip

Recombination: electrons captured by their parent ions

Lifetime: electrons captured by impurities during drift

Single Phase measures >4ms lifetime!



In Gas

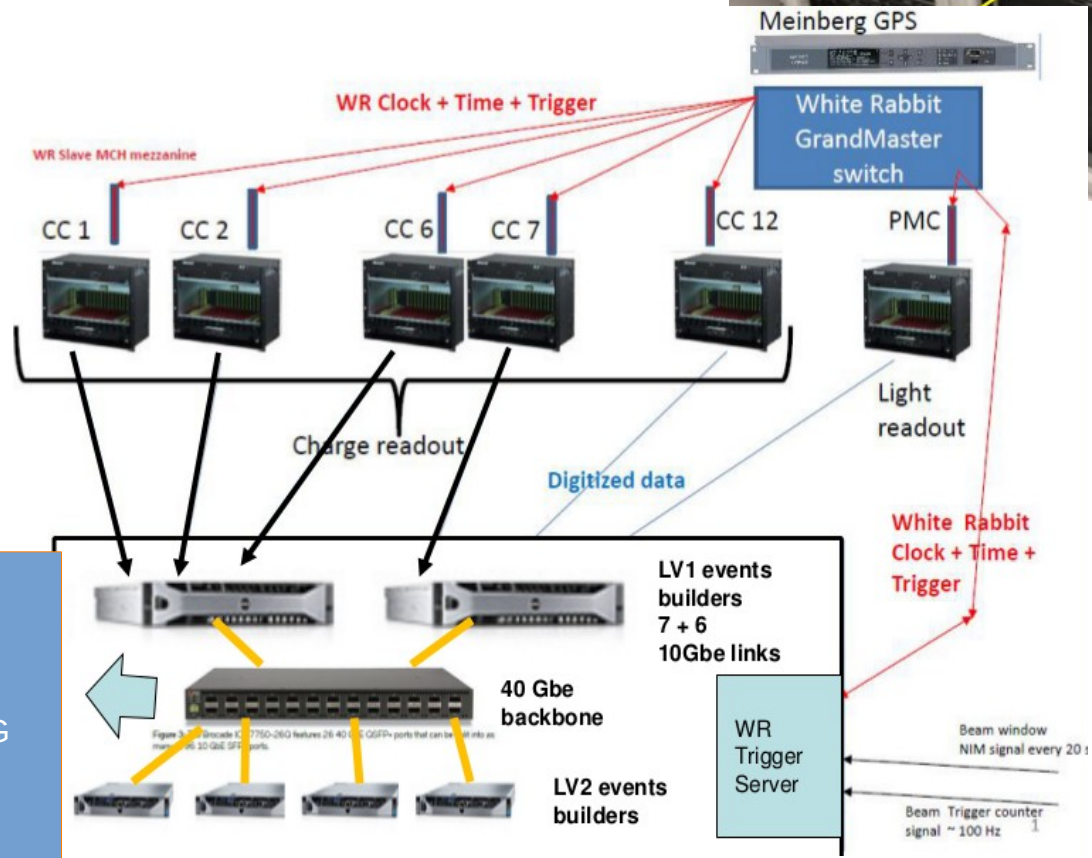
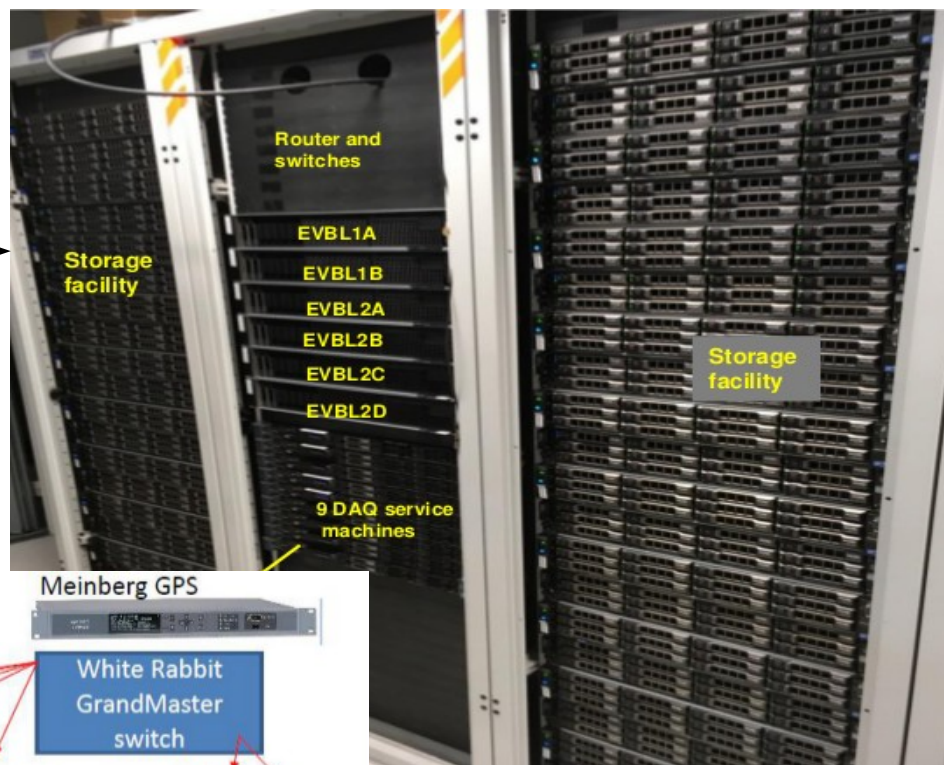
Extraction: probability of extraction > 90%

Amplification: Townsend avalanches of electrons

Nominal Eamp is 33 kV/cm and ENC is 1500 electrons

S/N > 10 for a mip on both views

Inside the NP-02 DAQ room



ONLINE STORAGE AND PROCESSING FARM

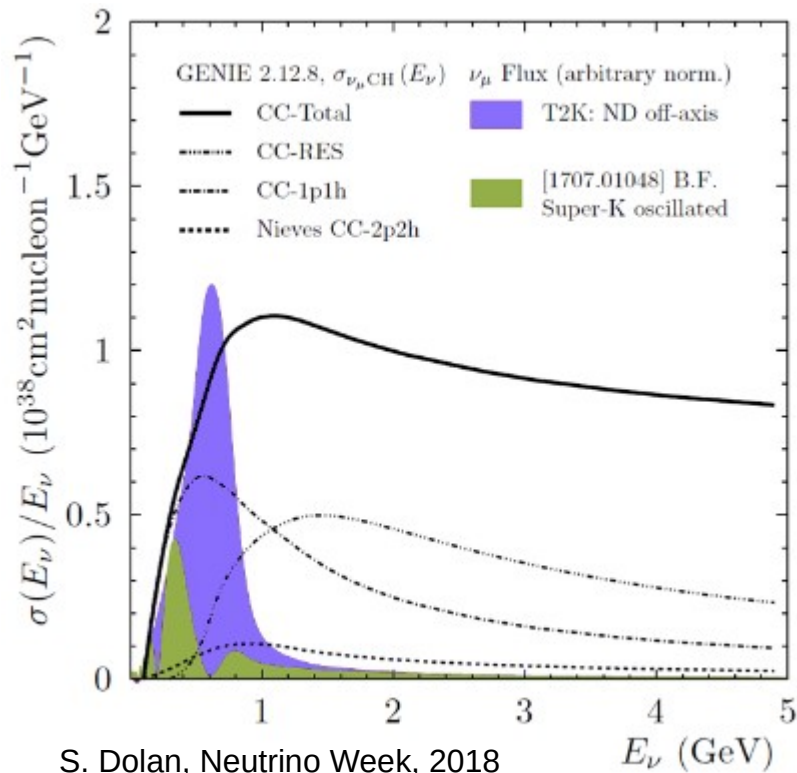
Long Baseline Experiments

T2K – State of the Art

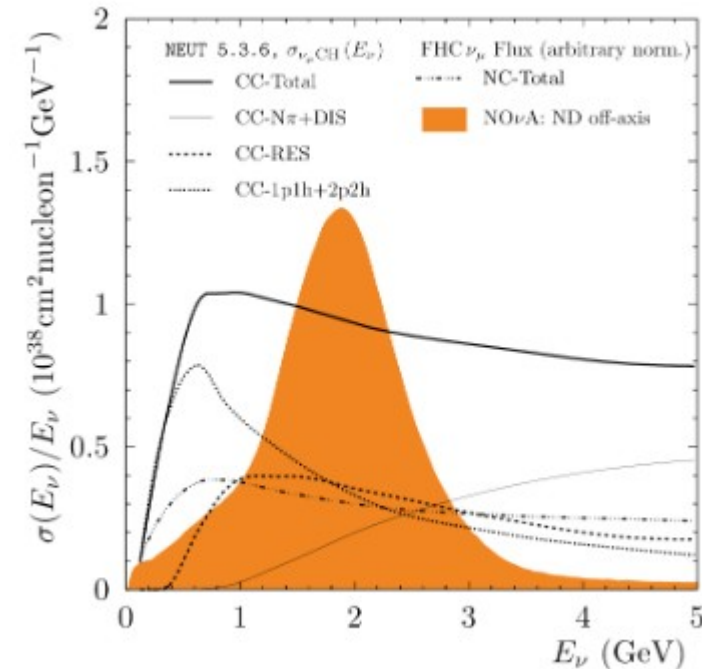
NOVA

- Interactions are CCQE or CCRES

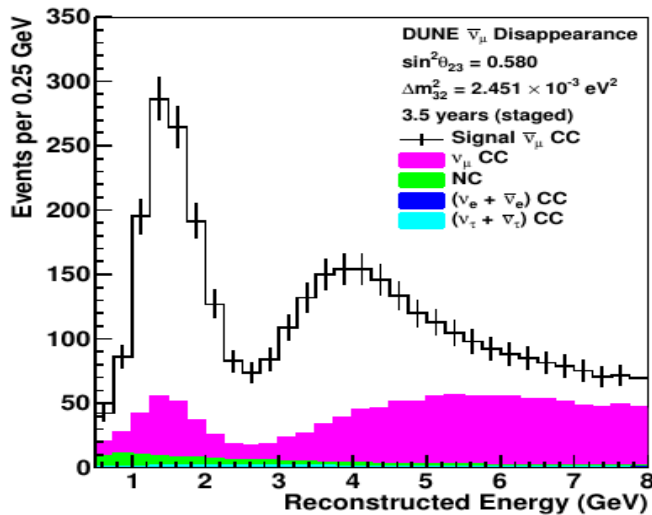
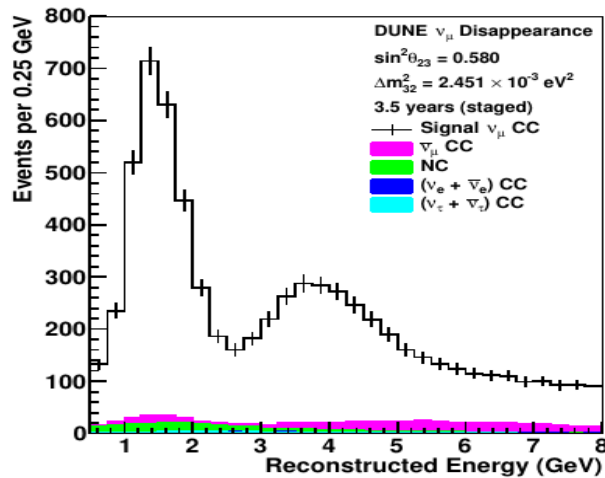
- Higher energy neutrinos means most interactions are not CCQE



- Deep Inelastic Scattering occurs



ν_μ Appearance



Expected Events (3.5 years staged)

ν mode

ν_μ Signal	6200
$\bar{\nu}_\mu$ CC background	389
NC background	200
$\nu_\tau + \bar{\nu}_\tau$ CC background	46
$\nu_e + \bar{\nu}_e$ CC background	8

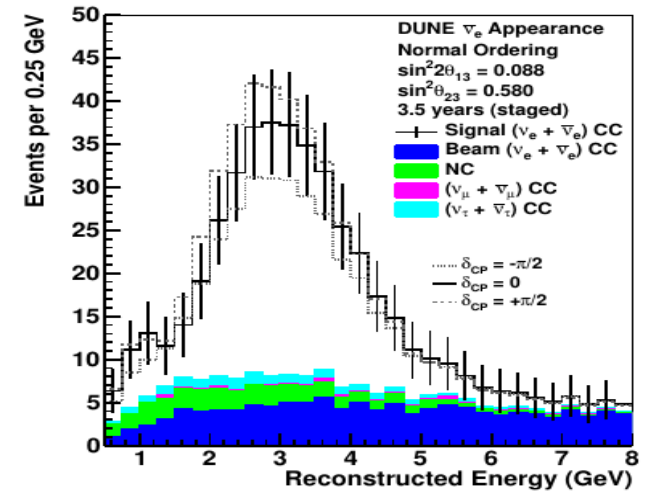
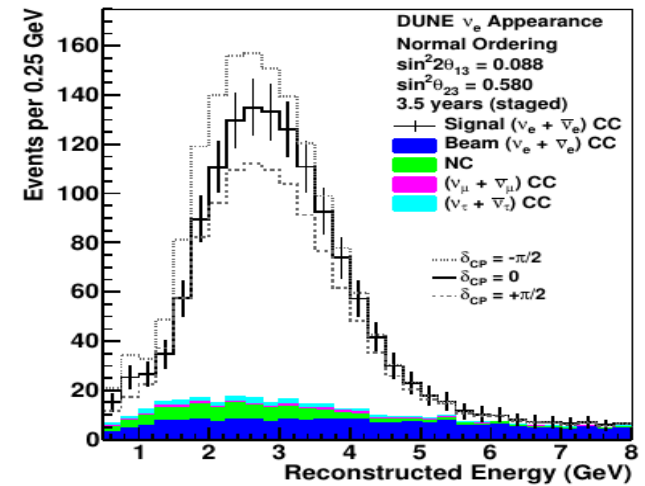
$\bar{\nu}$ mode

$\bar{\nu}_\mu$ Signal	2303
ν_μ CC background	1129
NC background	101
$\nu_\tau + \bar{\nu}_\tau$ CC background	27
$\nu_e + \bar{\nu}_e$ CC background	2

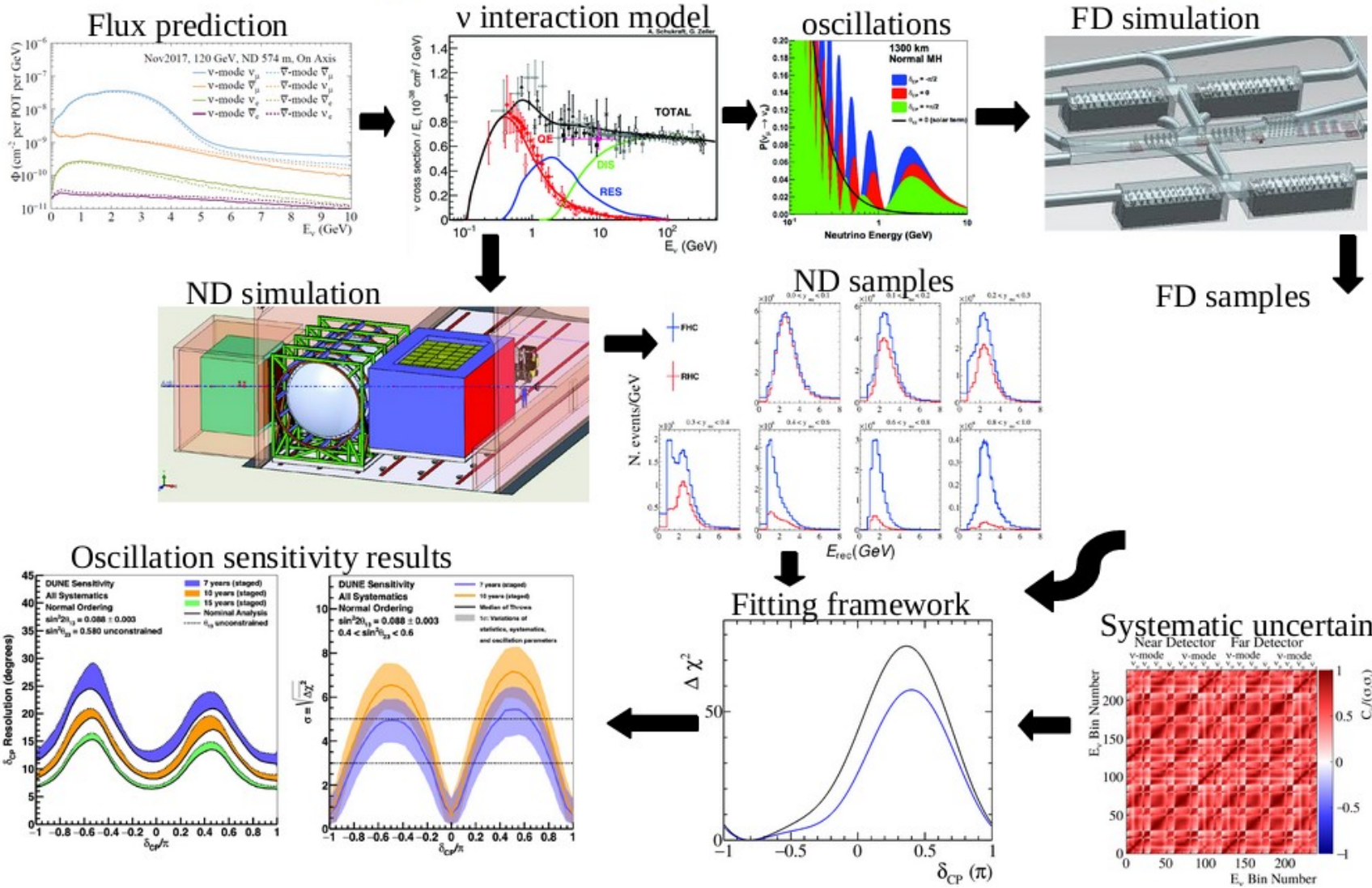
ν_e Appearance

Expected Events (3.5 years staged)

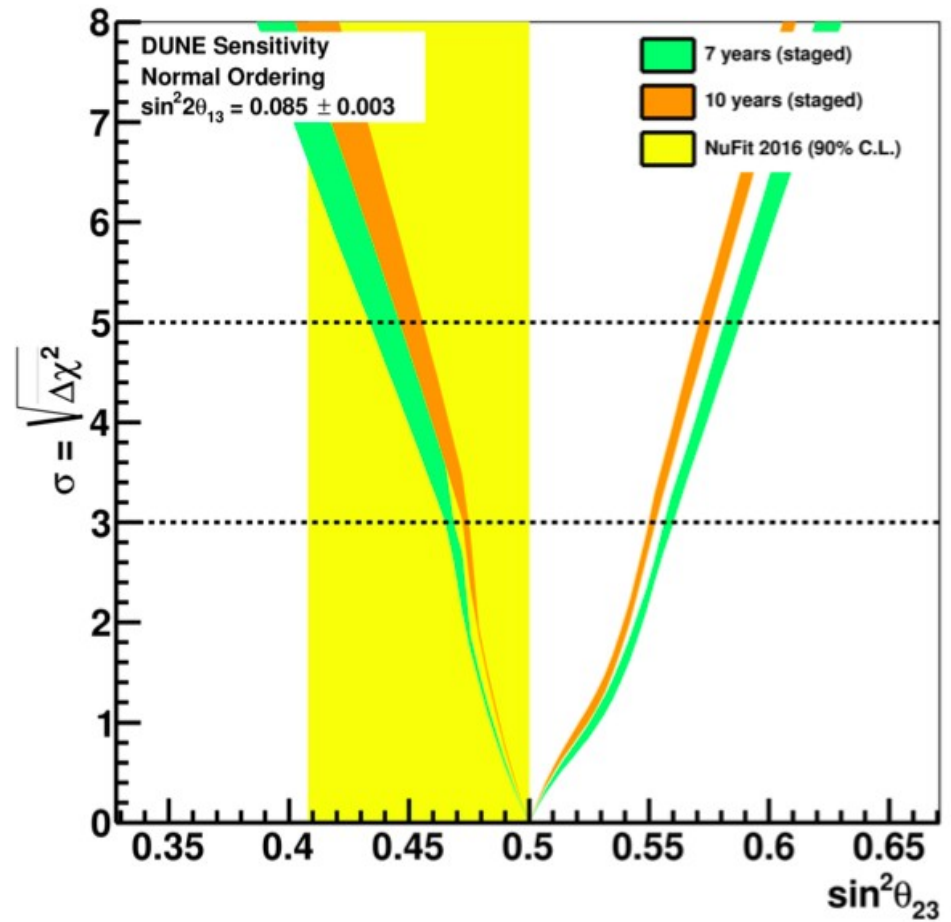
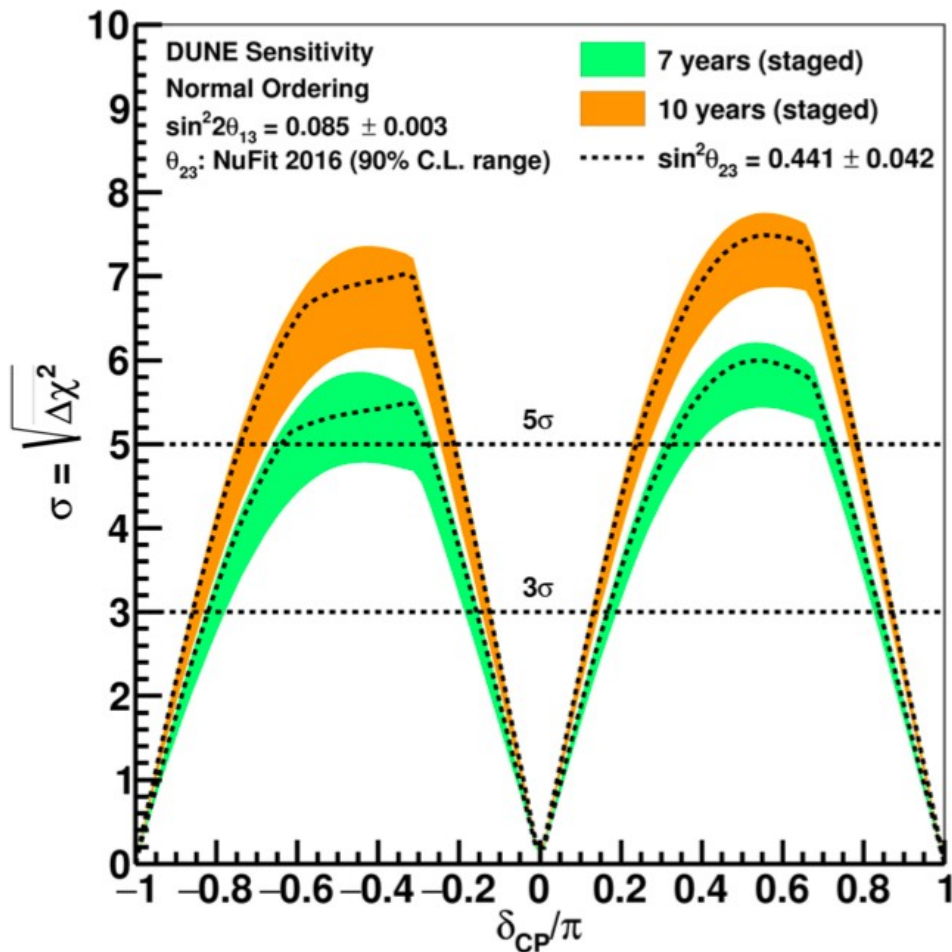
ν mode	
ν_e Signal NO (IO)	1092 (497)
$\bar{\nu}_e$ Signal NO (IO)	18 (31)
Total Signal NO (IO)	1110 (528)
Beam $\nu_e + \bar{\nu}_e$ CC background	190
NC background	81
$\nu_\tau + \bar{\nu}_\tau$ CC background	32
$\nu_\mu + \bar{\nu}_\mu$ CC background	14
Total background	317
$\bar{\nu}$ mode	
ν_e Signal NO (IO)	76 (36)
$\bar{\nu}_e$ Signal NO (IO)	224 (470)
Total Signal NO (IO)	300 (506)
Beam $\nu_e + \bar{\nu}_e$ CC background	117
NC background	38
$\nu_\tau + \bar{\nu}_\tau$ CC background	20
$\nu_\mu + \bar{\nu}_\mu$ CC background	5
Total background	180

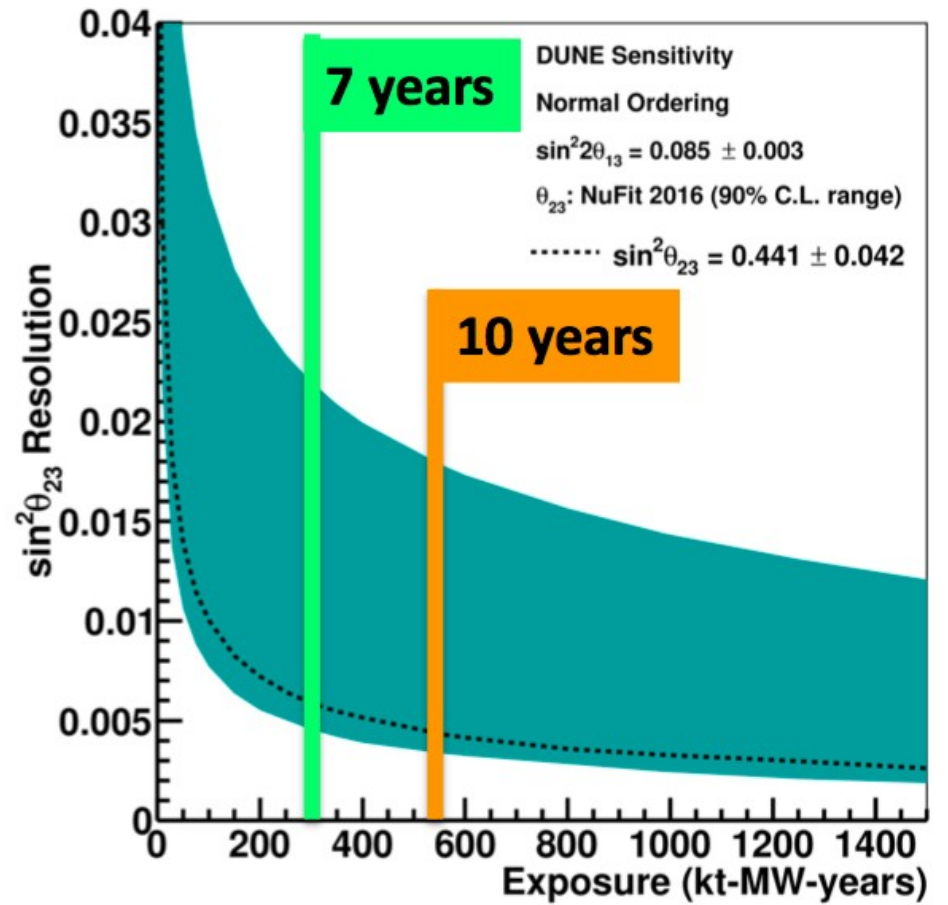
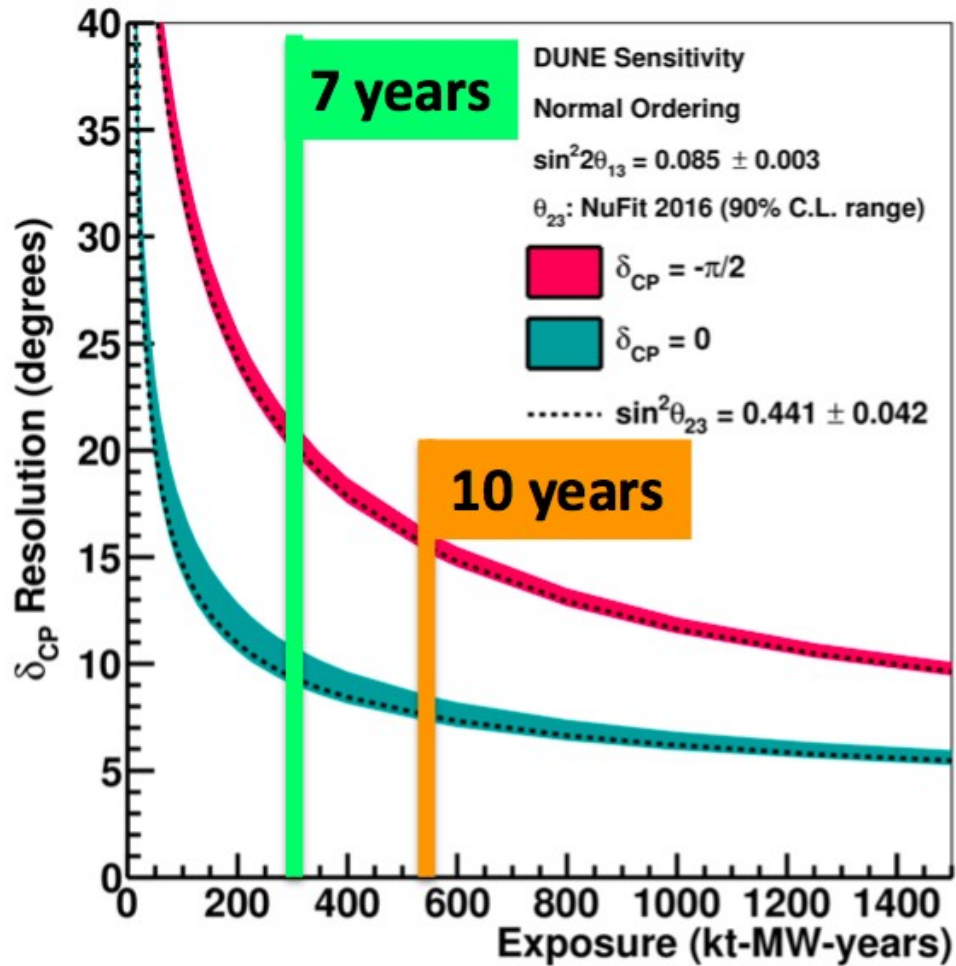


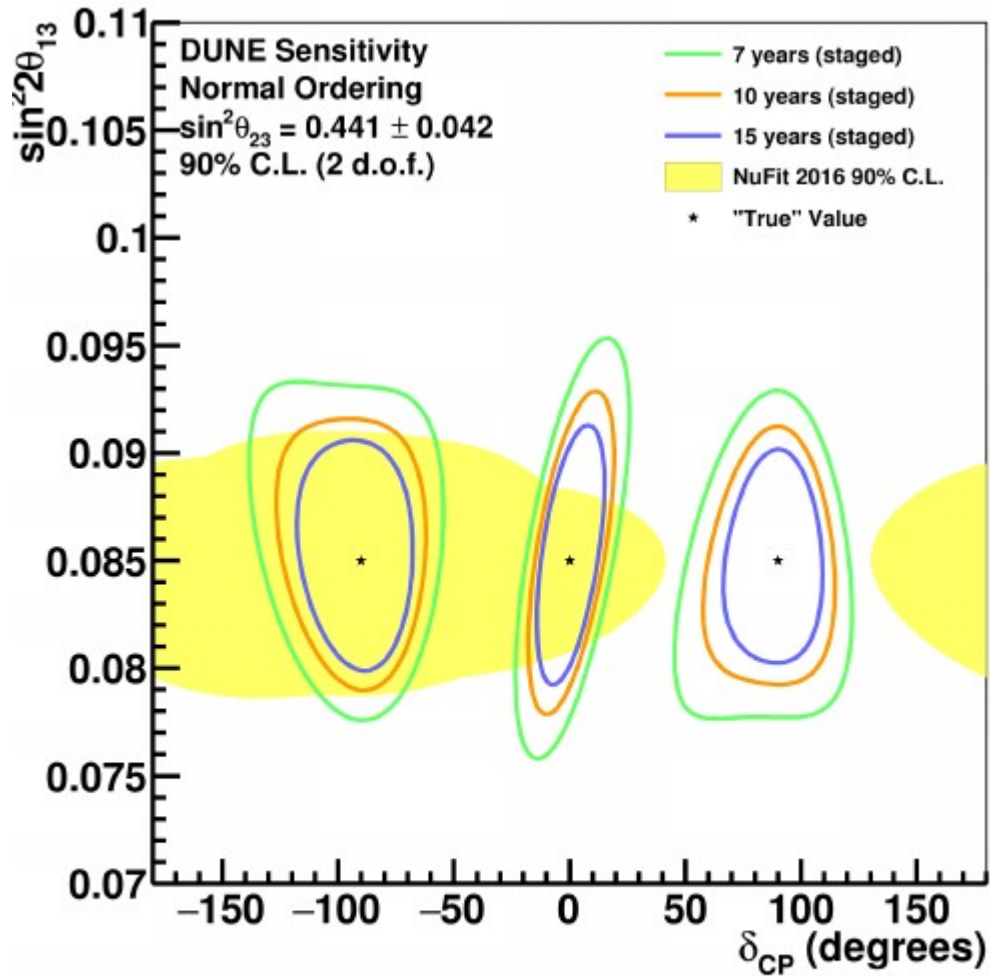
DUNE long-baseline oscillation analysis



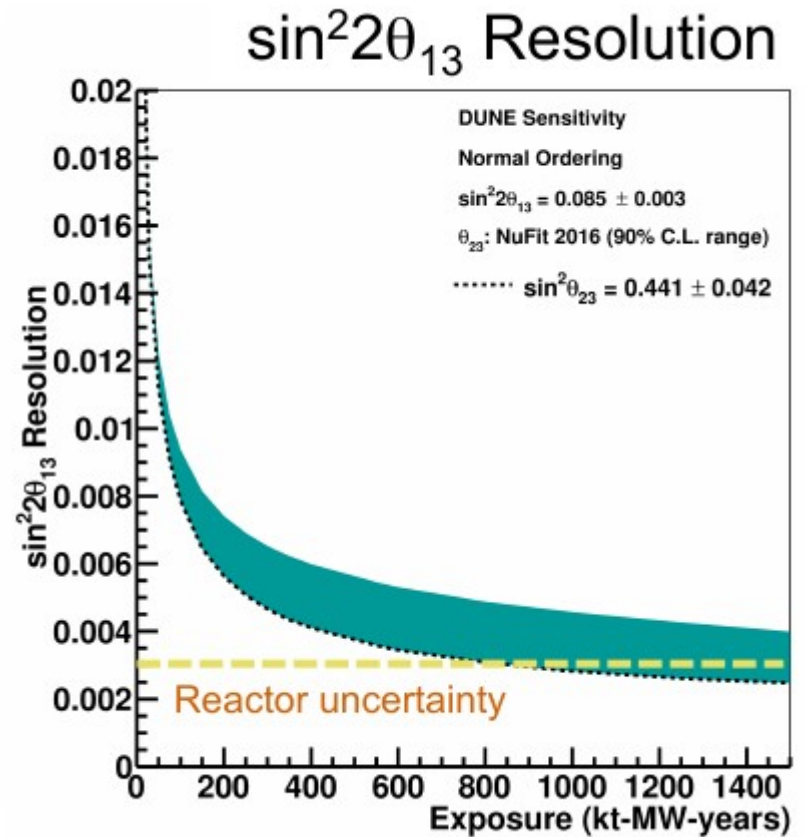
CP violation and theta_23







Simultaneous measurement of neutrino mixing angles and δ_{CP}



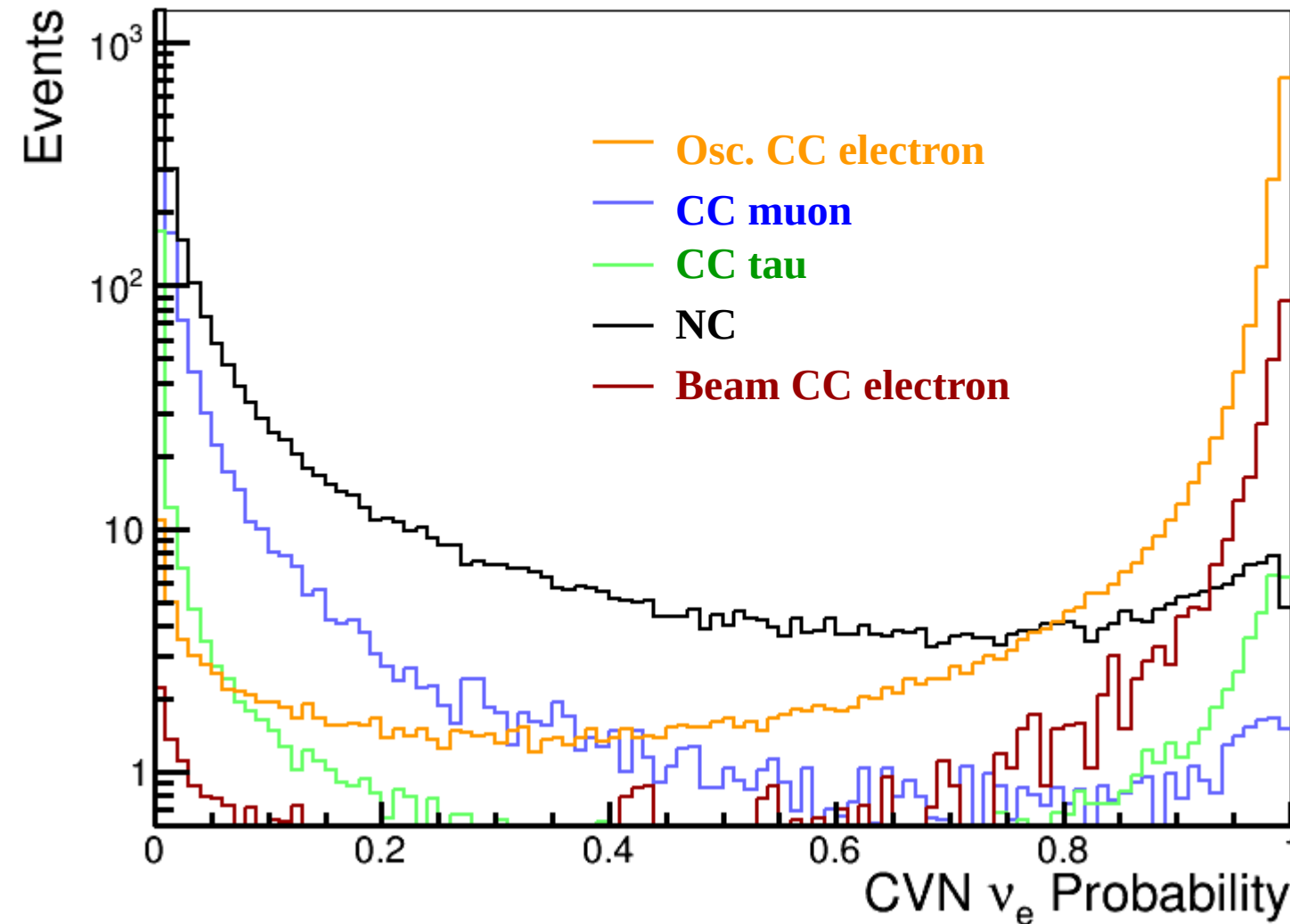
Systematic Uncertainties (CDR/IDR)

Source of Uncertainty	MINOS ν_e	T2K ν_e	DUNE ν_e
Beam Flux after N/F extrapolation	0.3%	3.2%	2%
Interaction Model	2.7%	5.3%	$\sim 2\%$
Energy scale (ν_μ)	3.5%	included above	(2%)
Energy scale (ν_e)	2.7%	includes all FD effects	2%
Fiducial volume	2.4%	1%	1%
Total	5.7%	6.8%	3.6 %
Used in DUNE Sensitivity Calculations			5% \oplus 2%

DUNE Conceptual Design Report (CDR) arXiv:1512.06148, the DUNE signal normalization uncertainty is taken to be 5% \oplus 2% in both neutrino and antineutrino mode, where 5% is the normalization uncertainty on the FD ν_μ sample and 2% is the effective uncorrelated uncertainty on the FD ν_e sample after fits to both near and far detector data and all external constraints

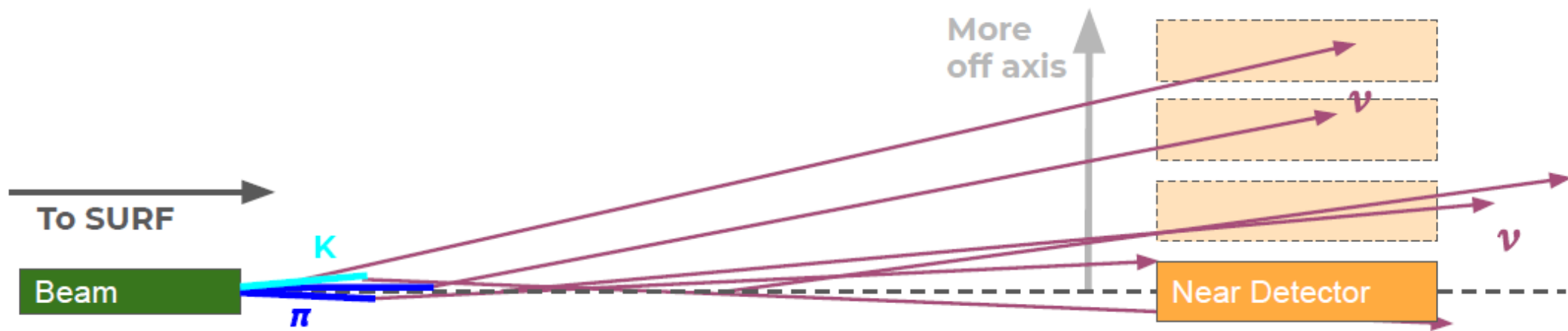
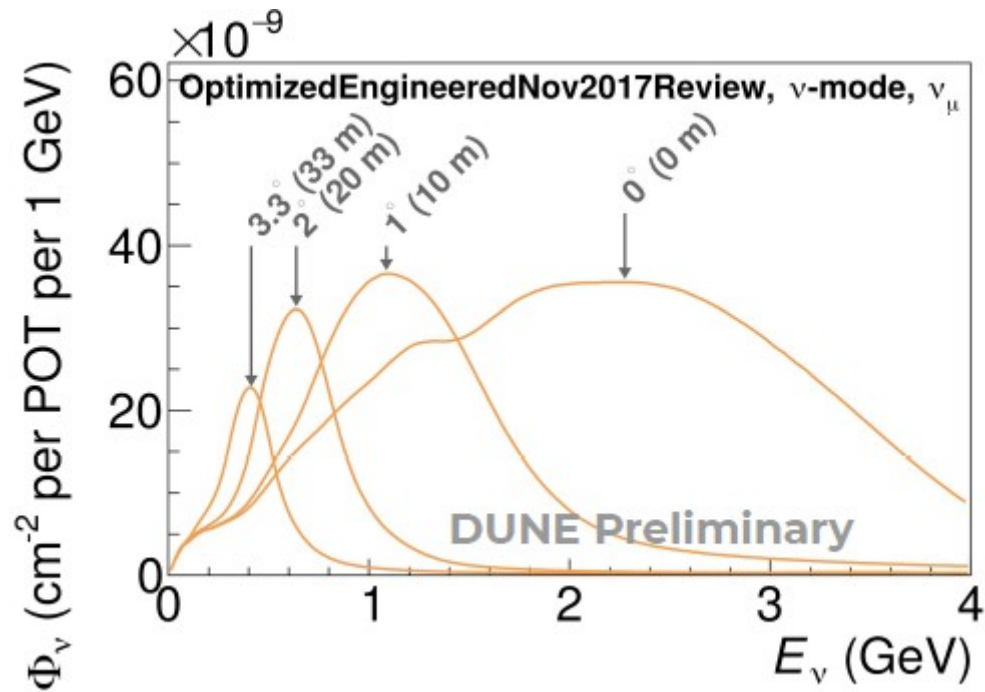
DUNE Conceptual Design Report (CDR)
arXiv:1512.06148

ν_e charged-current selection



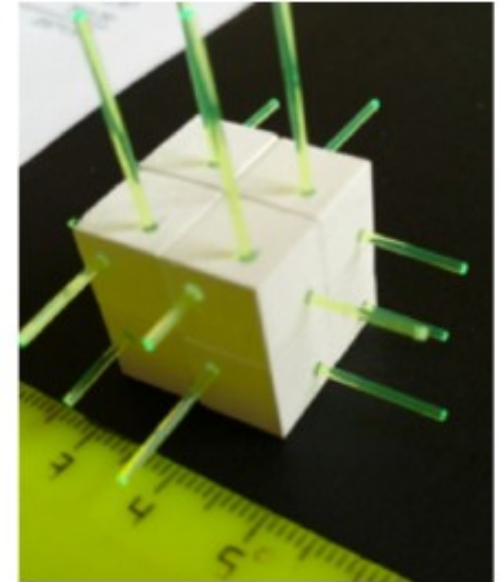
- Selects both oscillated and intrinsic ν_e
- Backgrounds include $\pi^0 \rightarrow \gamma\gamma$, $\tau \rightarrow e$

DUNE-PRISM: off-axis ND

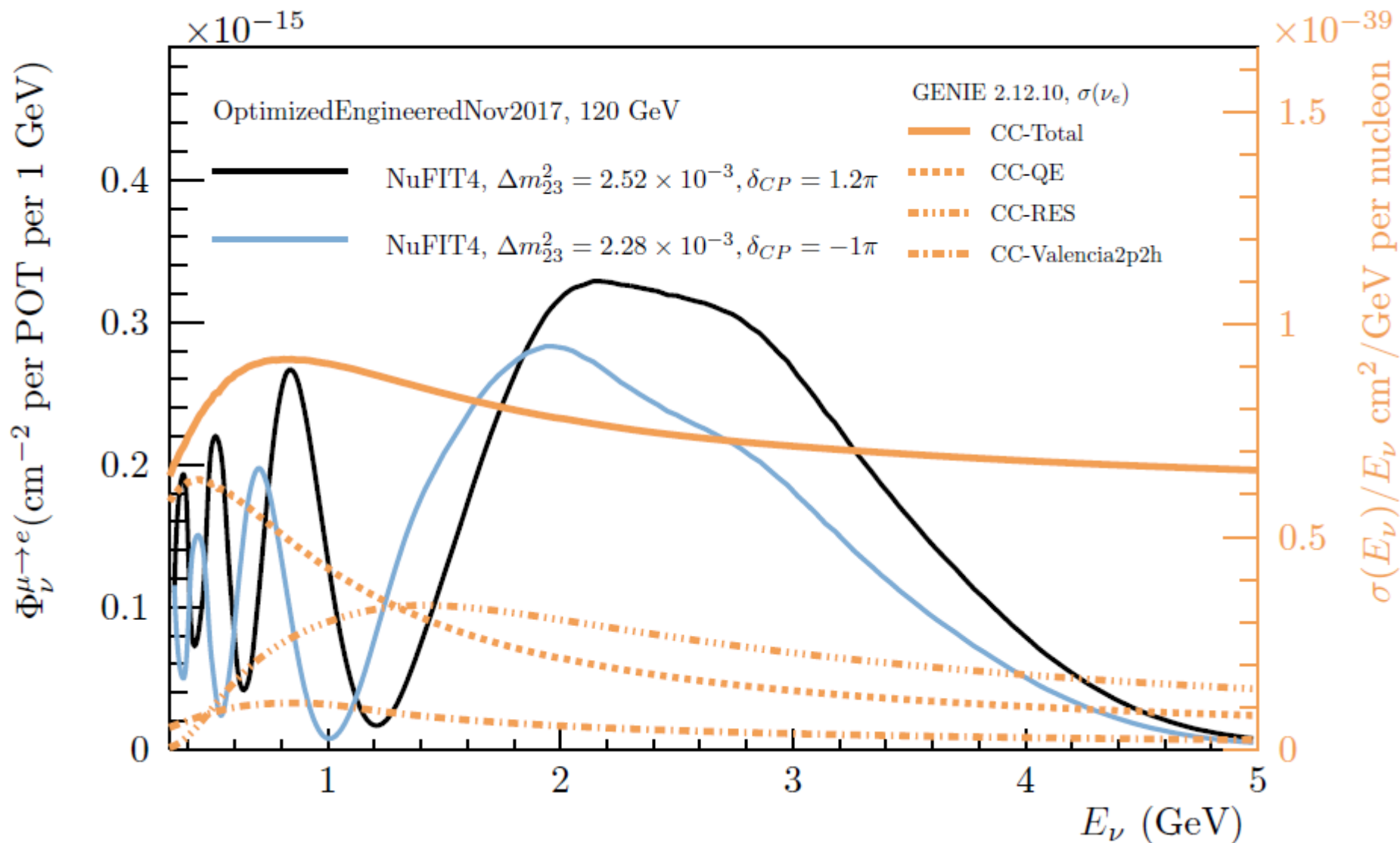


3D scintillator tracker (3DST)

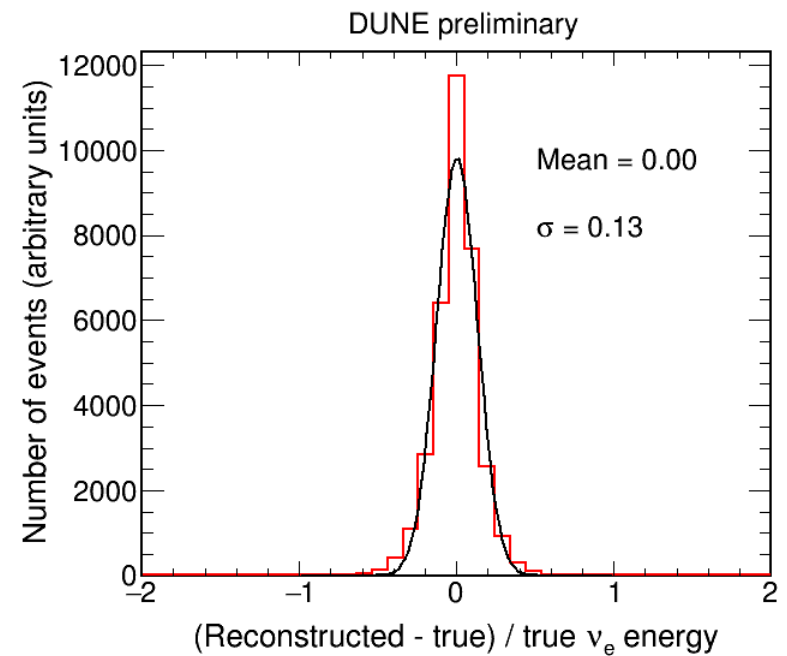
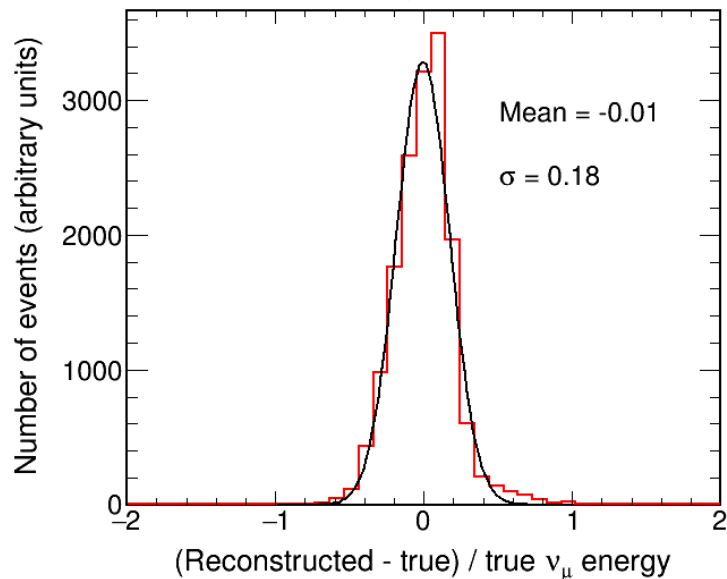
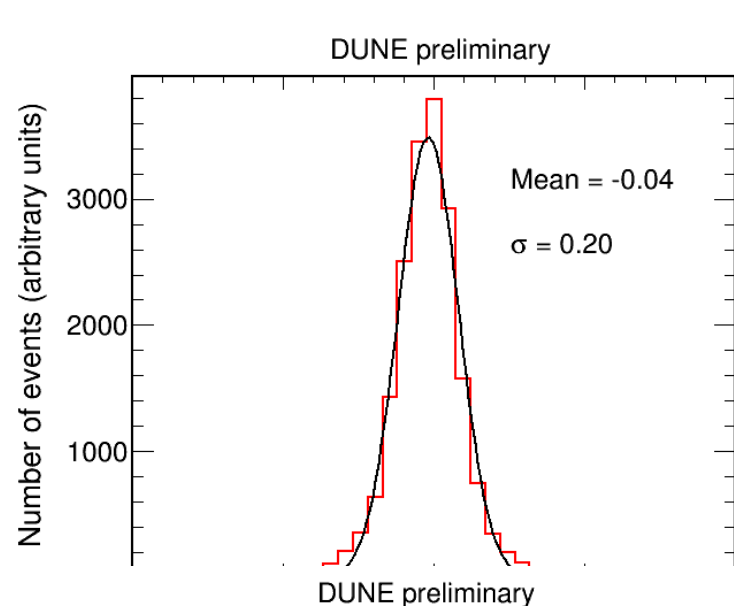
- 1 cm³ scintillator cubes in a large array, read out with orthogonal optical fibers in three dimensions
- Same concept being pursued by T2K ND280 upgrade, called “Super-FGD”
- Excellent 4π acceptance –no hole at 90°
- Very fast timing: capable of tagging neutrons from recoils, and measuring energy from time-of-flight
- Could be placed in front of (or inside?) gas TPC, or operated in its own magnet with muon spectrometer



Far detector spectra, with cross section predictions vs. energy



Neutrino energy resolution



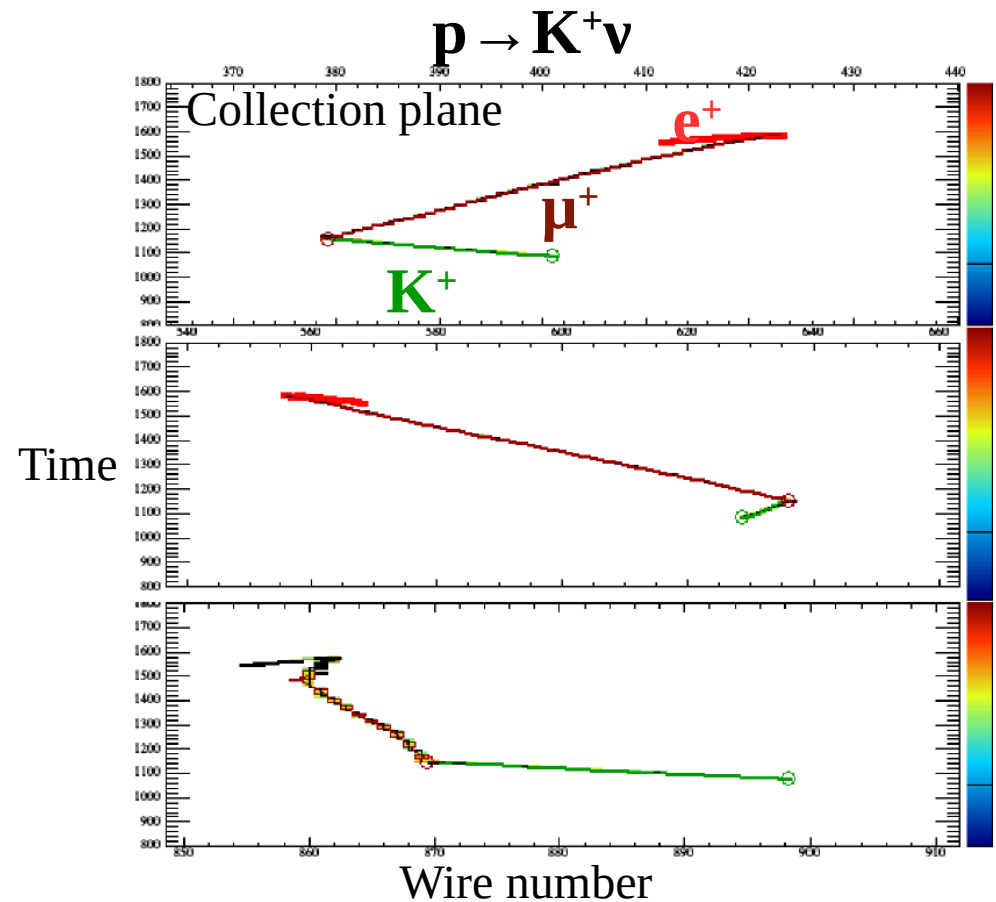
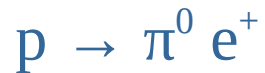
- Muons $\sim 18\%$ for contained tracks, 20% for exiting
- Electrons $\sim 13\%$

Nucleon decay

- DUNE sensitive to:



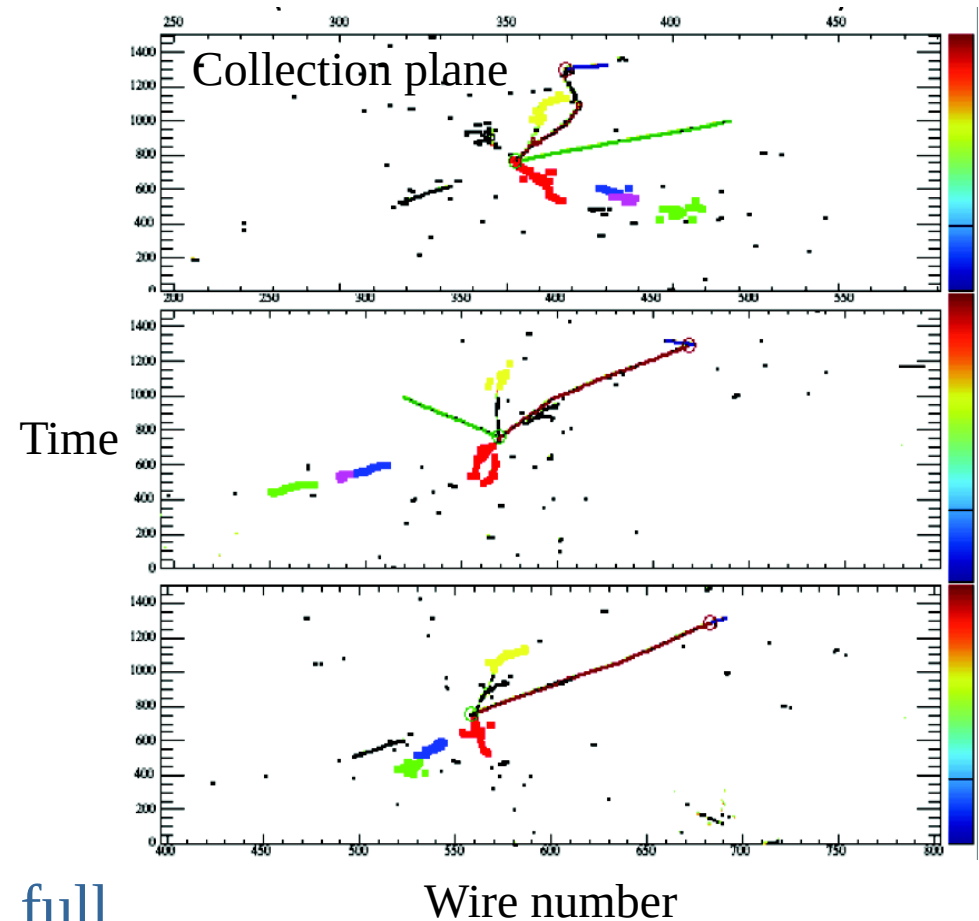
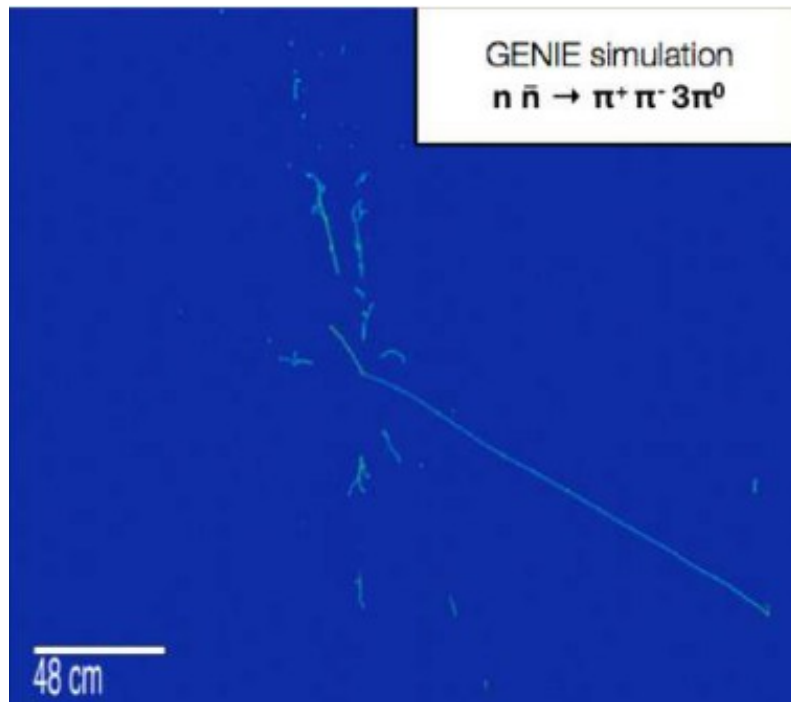
- and other channels:



upcoming TDR contains analyses with
full simulation & reconstruction

neutron-antineutron oscillations

Baryon number violating BSM process
Signature is a 'star' of charged and neutral pions



upcoming TDR contains analyses with full
simulation & reconstruction

and other BSM physics

New particles:

light dark matter, boosted dark matter,
heavy neutral leptons

Deviations from PNMS ν mixing paradigm:

non-standard ν interactions,
non-unitarity, CPT or Lorentz violation,
large extra dimensions,
 ν -trident production

Sterile neutrino searches
(ν_e appearance at ND)

