

# An overview of ICHEP2024

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**Tai-Hua Lin**

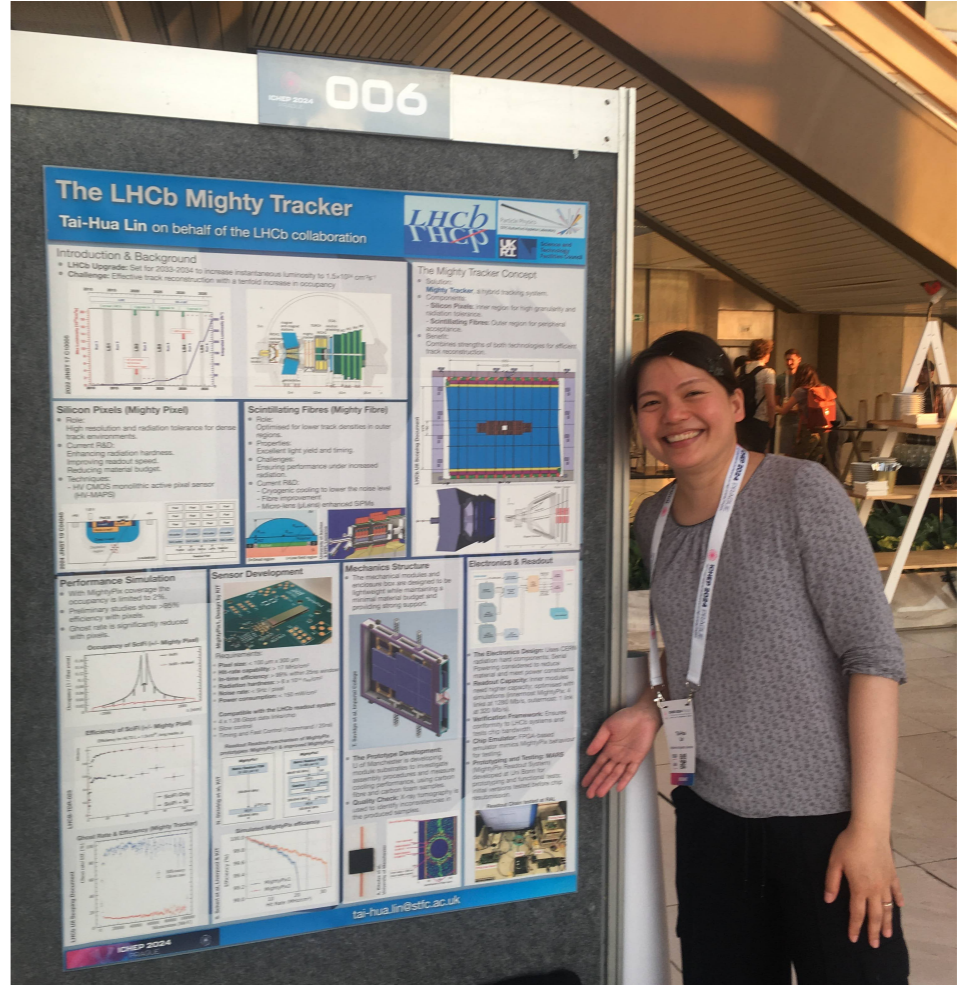
Rutherford Appleton Laboratory, STFC

09/10/2024

PPD Seminar

# ICHEP 2024 Introduction

- ✦ 42nd International Conference on High Energy Physics (ICHEP 2024)
  - ✦ 281 Posters
  - ✦ 918 parallel talks
  - ✦ 40 plenary talks
  - ✦ 1388 Participants (~2 from PPD)
- ✦ ICHEP
  - ✦ first held in 1950
  - ✦ Biannual conference since 1960
  - ✦ In Prague on 17th to 24th July 2024
  - ✦ ICHEP 2026 will be in Natal, Brazil



*Draft Beer 0,3l/0,5l*

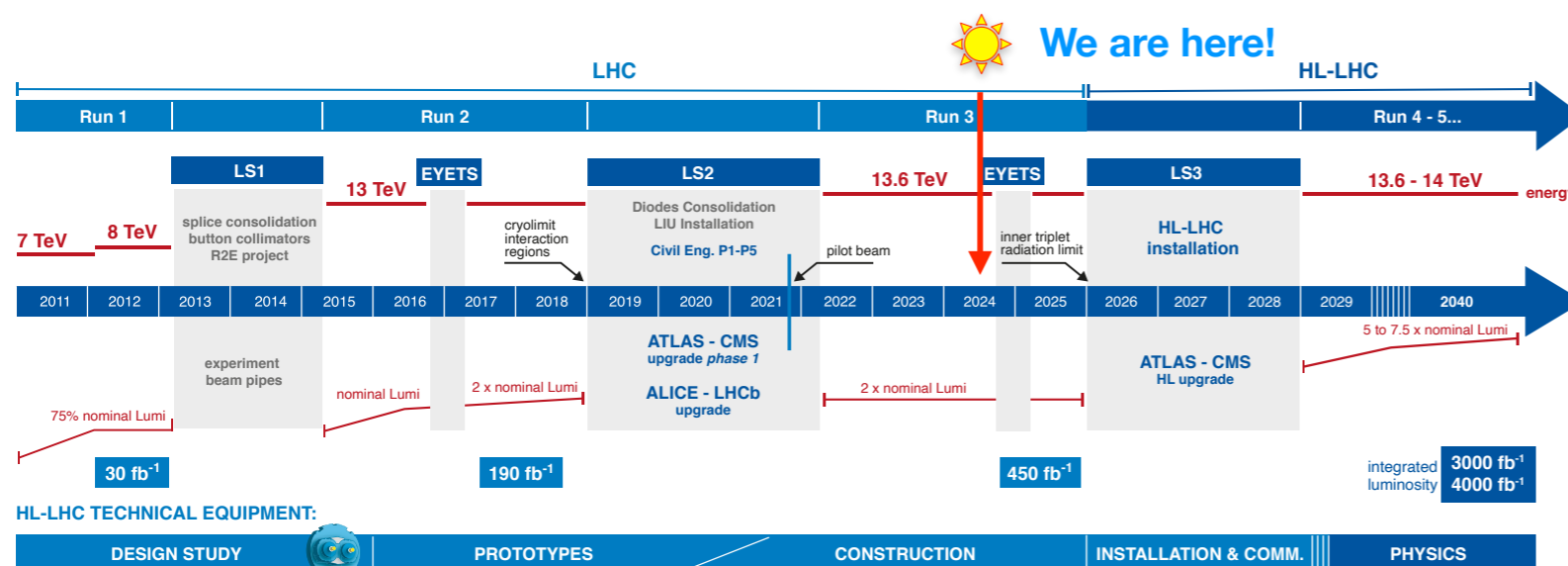
Budvar Redix	/ light 39 Kč / 59 Kč
Budvar 33 hořký ležák	/ light 49 Kč / 69 Kč
Budvar Tmavý ležák	/ dark 49 Kč / 69 Kč

Voda perlivá / Sparkling water 0,33l / 0,75l 59 Kč / 109 Kč  
Voda neperlivá / Still water 0,33l / 0,75l 59 Kč / 109 Kč

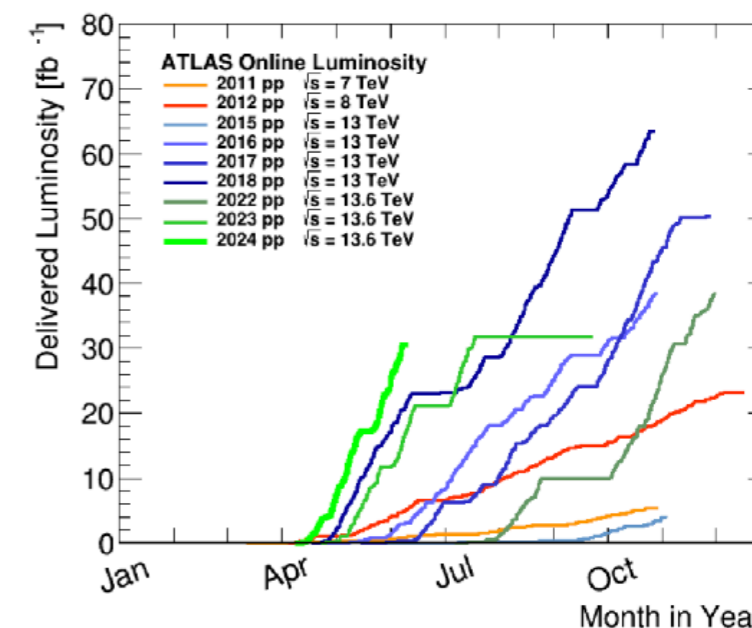
Source: menu from a restaurant near St Nicholas Bell Tower

- ✧ **Particle Physics Highlights:**  
Results from Higgs, neutrino physics, and dark matter
- ✧ **Cosmology & Astrophysics:**  
New insights from DESI, gravitational wave observations, and cosmology
- ✧ **Accelerator Physics:**  
LHC upgrades, upcoming collider designs (HL-LHC, FCC, ILC)
- ✧ **Experimental Innovations:**  
New detection techniques, AI, and ML applications in HEP
- ✧ **Future Directions:**  
Summary of discussions on future projects, including the physics potential of HL-LHC and FCC
- ✧ **Outreach:**  
Ideas to improve our PPD exhibition room
  
- ✧ <https://indico.cern.ch/event/1291157/>  
<https://www.youtube.com/@ICHEP>
- ✧ **This talk is heavily referencing [Marumi Kado's](https://indico.cern.ch/event/1291157/contributions/5958406/attachments/2903497/5123335/ICHEP-Summary.pdf) summary talk:**  
<https://indico.cern.ch/event/1291157/contributions/5958406/attachments/2903497/5123335/ICHEP-Summary.pdf>

- LHC is refining its results on a **clean and well calibrated dataset of  $\sim 140 \text{ fb}^{-1}$**  at 13 TeV the Run 3 nearly equaling the Run 2 dataset  $\sim 120 \text{ fb}^{-1}$  at 13.6 TeV.

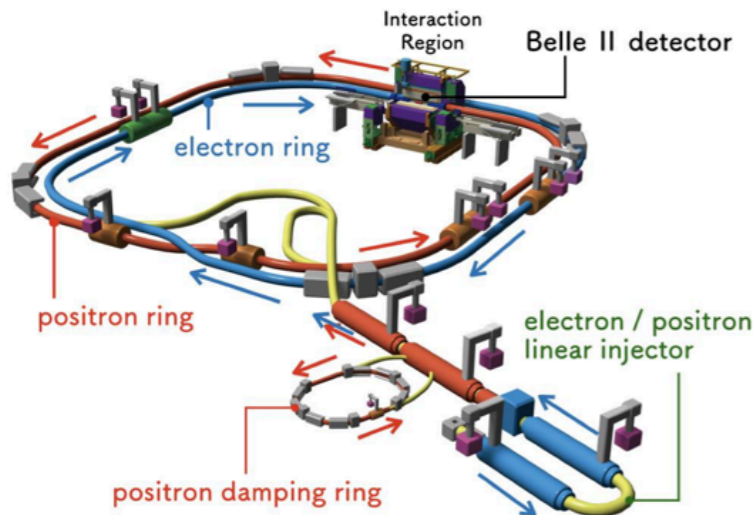


**2024 - High availability operation, Full mastery of considerable inherent operational risks**



- Currently about half way into entire LHC operations schedule, **Major upgrades leading** to the High Luminosity during the third long shutdown now on the horizon! (See backup for more details)

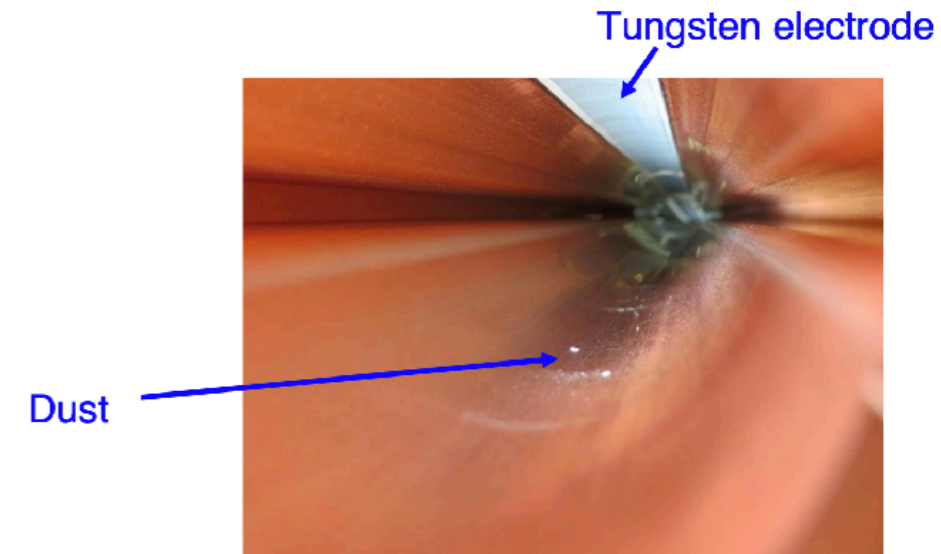
Approximately x10 Luminosity delivered (in terms of results x20)



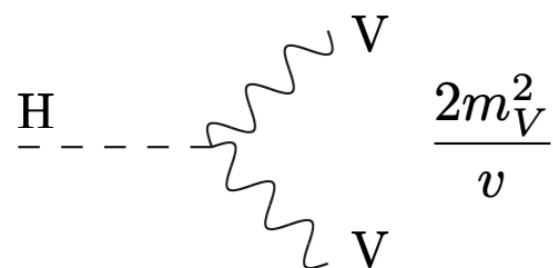
- Super KEK-B and Belle II world's highest instantaneous luminosity ( $4.7 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ ), difficult year with a focus on understanding **Sudden Beam Losses** (Belle II is currently running with VTX off).

## Sergei also discussed other facilities:

- RHIC AA (and future EIC, NICA)
- High intensity facilities Fermilab, JPARC, PSI, TRIUMF, CERN PS and SPS
- VEPP and BEPC continuing to bring very useful data (discussed in this talk).
- DAΦNE facility still brings physics potential (PADME) and progress in accelerator R&D!



Countermeasures to the SBL will be implemented during the 2024 summer maintenance period.



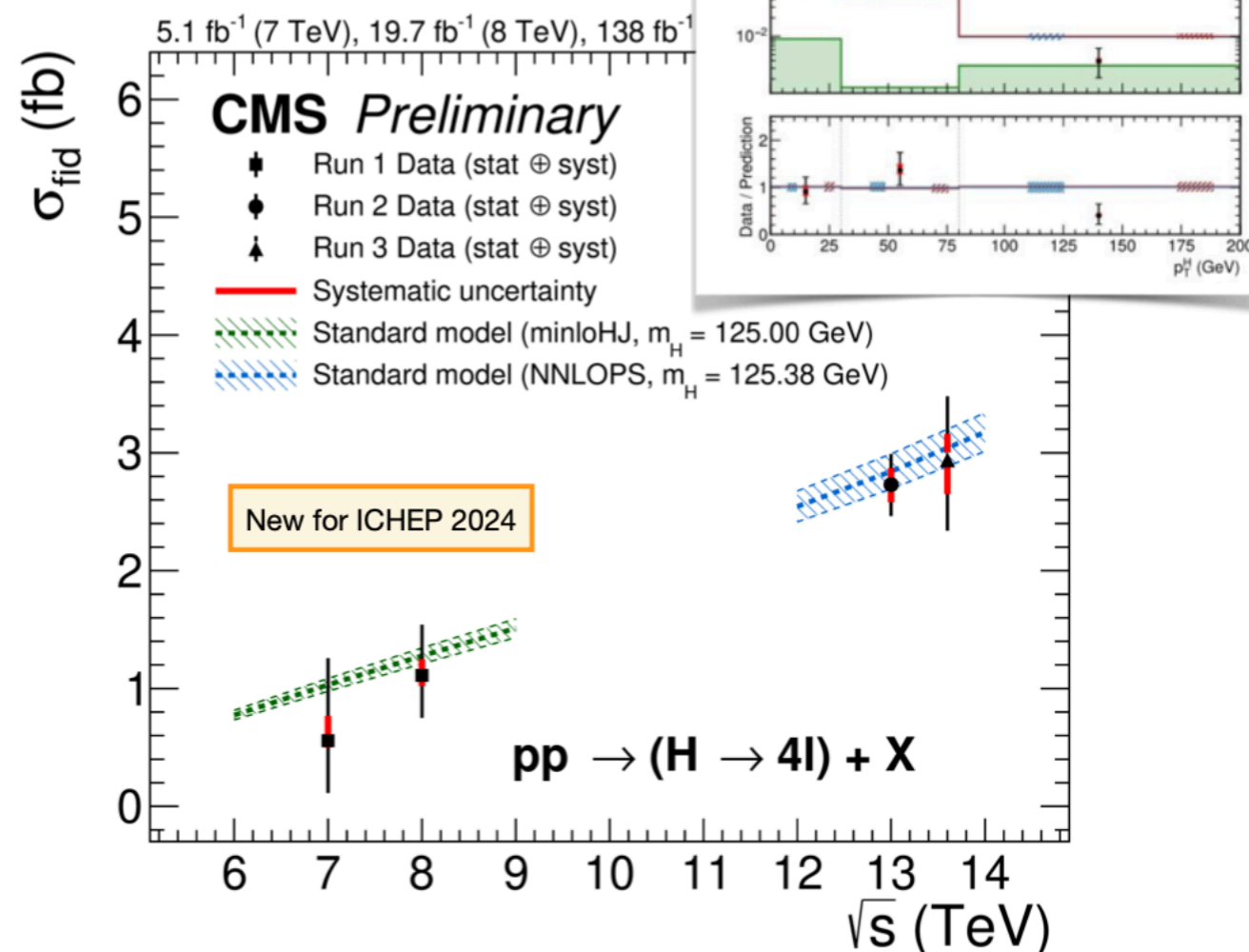
$$|\partial_\mu \phi|^2$$

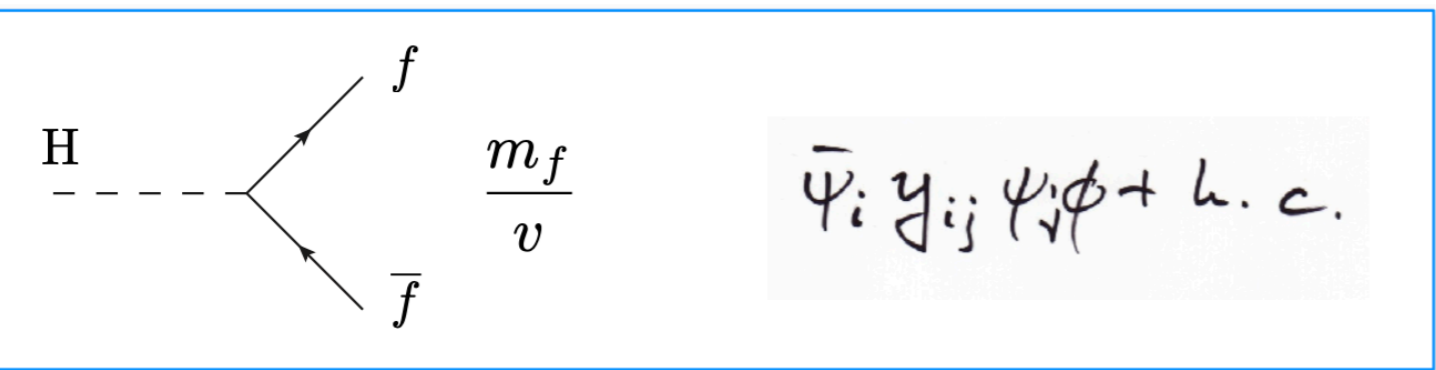
This term could not exist without a vev

$K_{W,Z}$	Current	HL-LHC	FCC (ee)
	6%	1.5%, 1.7 %	0.4%, 0.2 %

- Unambiguous proof of the existence of the Higgs condensate!
- Most precisely known Higgs coupling tells us how elementary the Higgs boson is!
- (Matthew McCullough):
  - It is of utmost importance to measure the most precisely measured coupling (hZZ) to probe the Higgs compositeness.
  - Precision in Higgs physics is key.
- Precision at HL-LHC is limited by TH (Higher Order, PDFs and  $\alpha_s$ )

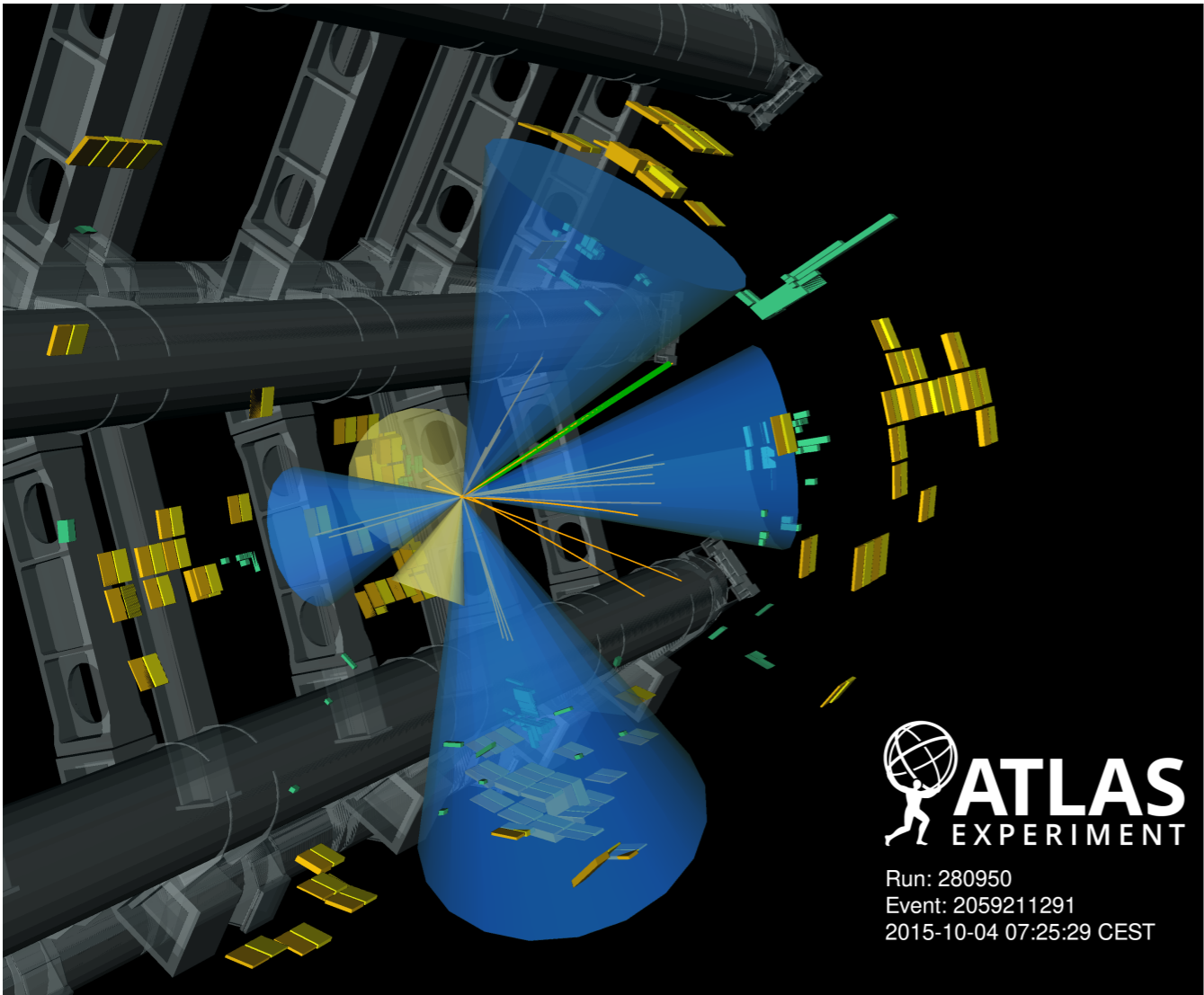
## New measurements at 13.6 TeV (Run 3)





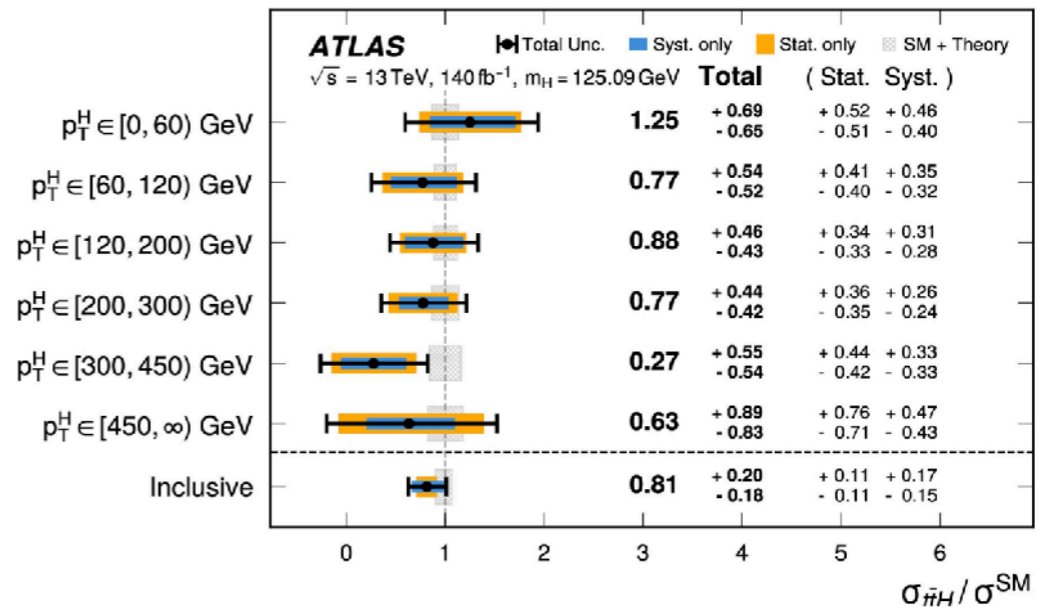
	Current	HL-LHC	FCC (ee)	FCC (hh)
$\kappa_t$	11%	3.4%	-	1%
$\kappa_b$	11%	3.7%	0.7%	-
$\kappa_\tau$	8%	1.9%	0.7%	-
$\kappa_\mu$	20%	4.3%	8.9%*	-

## Top Yukawa Coupling at the LHC



### ttH news from ATLAS!

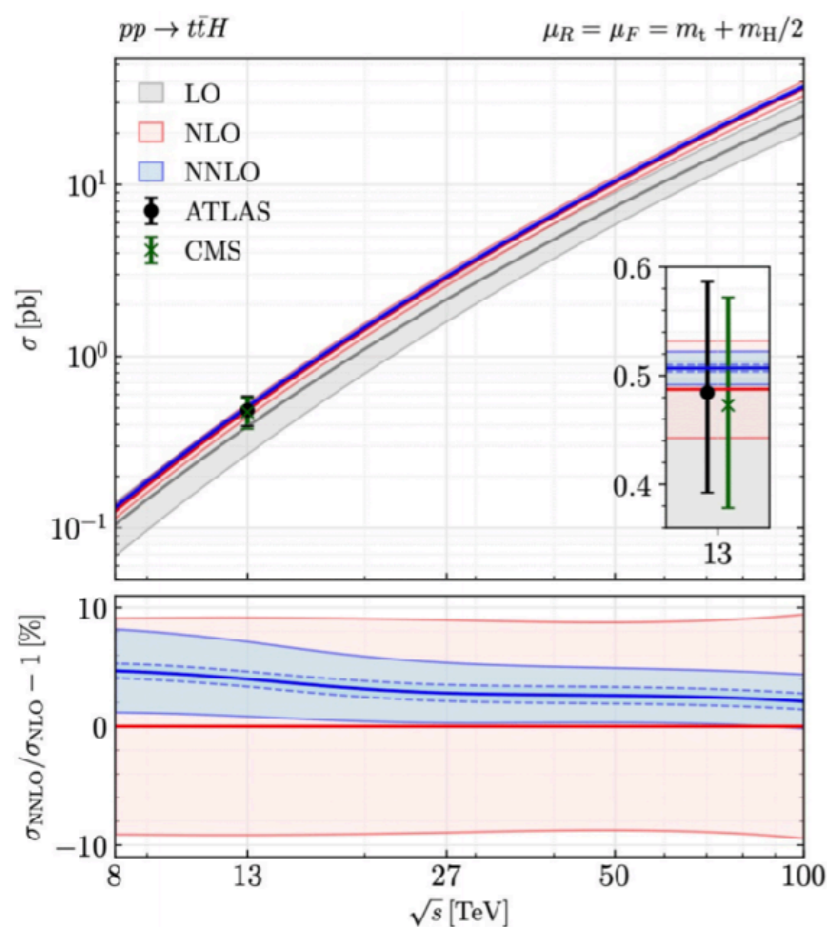
Very complex final state Main challenge: tt+bb background



**Best single ttH measurement!**  
Overall uncertainty improved by factor 1.8, 4.6 $\sigma$

## Top Yukawa Coupling at the LHC

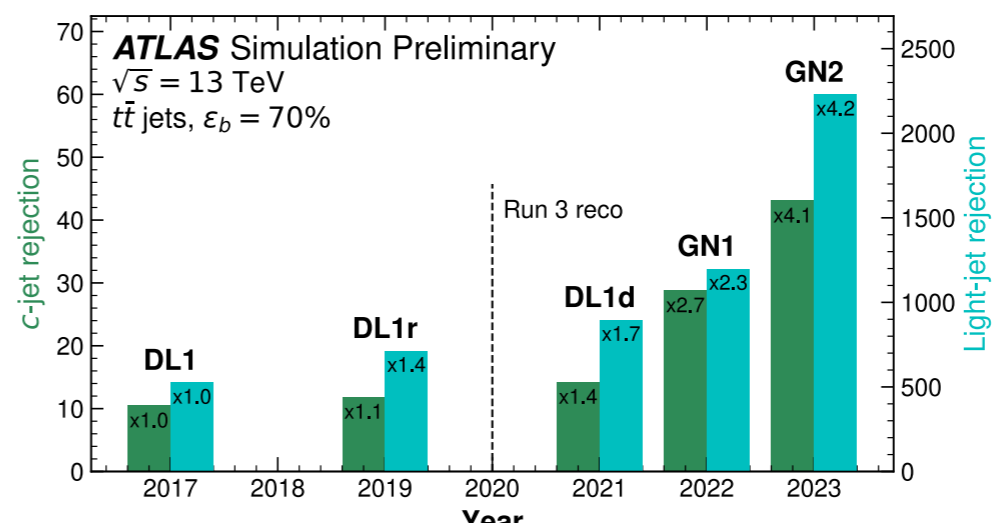
### Recent example from Mathew (ttH)



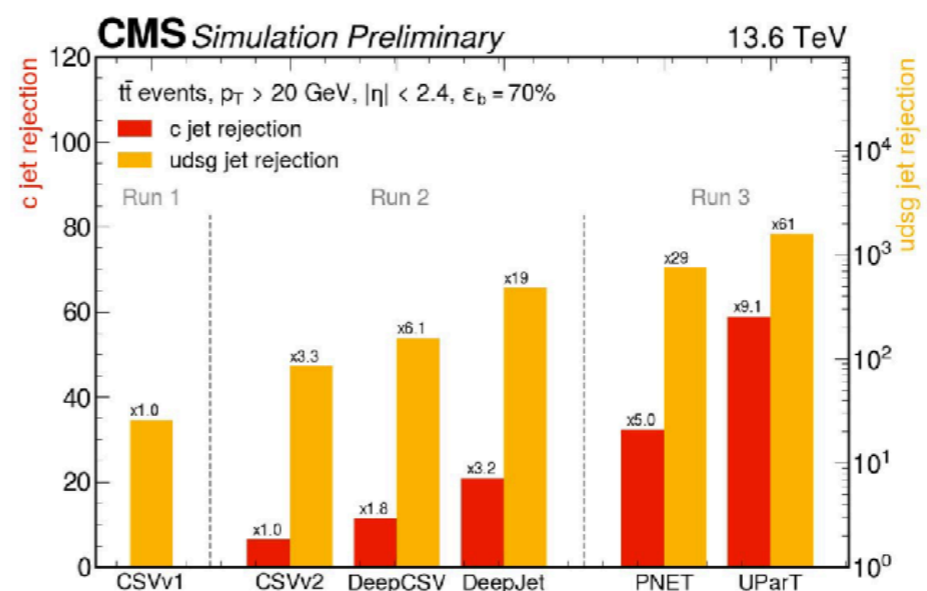
**Message 2 (again)** (Matthew McCullough): Precision in Higgs physics is key.

### ttH from ATLAS

AI in HEP reconstruction has a significant impact!



The most shown plot at this conference!

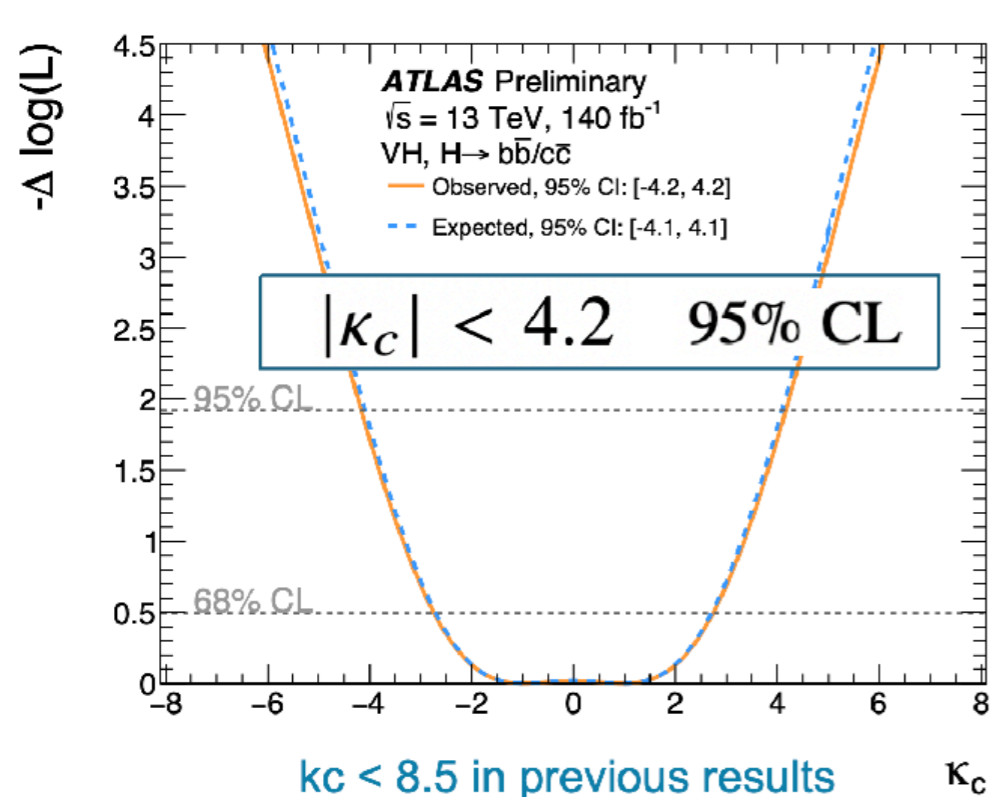


There are 4 b-quark jets in the  $t\bar{t}H(bb)$  event topology!



## Yukawa Coupling to Charm at the LHC

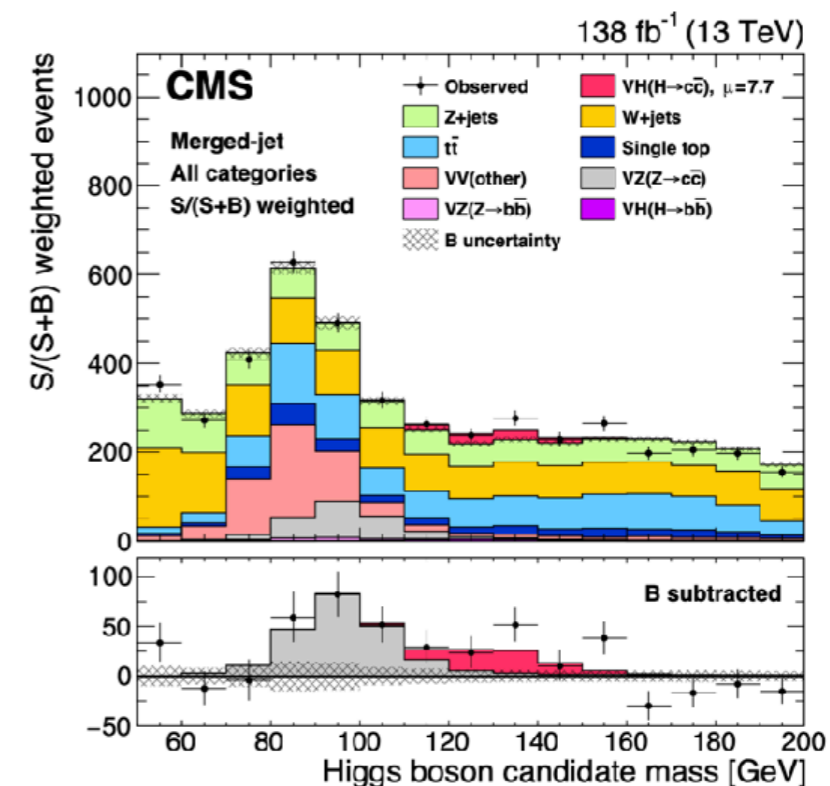
Refined analysis of Run 2 data with now Graph NN charm tagging!



$$\mu_{VH}^{cc} = 1.0^{+5.4}_{-5.2} = 1.0^{+4.0}_{-3.9} \text{ (stat.)}^{+3.6}_{-3.5} \text{ (syst.)}$$

Improvement by a factor of 2 w.r.t. previous result

Use of state-of-the-art ML techniques [Particle Net](#) uses Dynamic Graph CNN

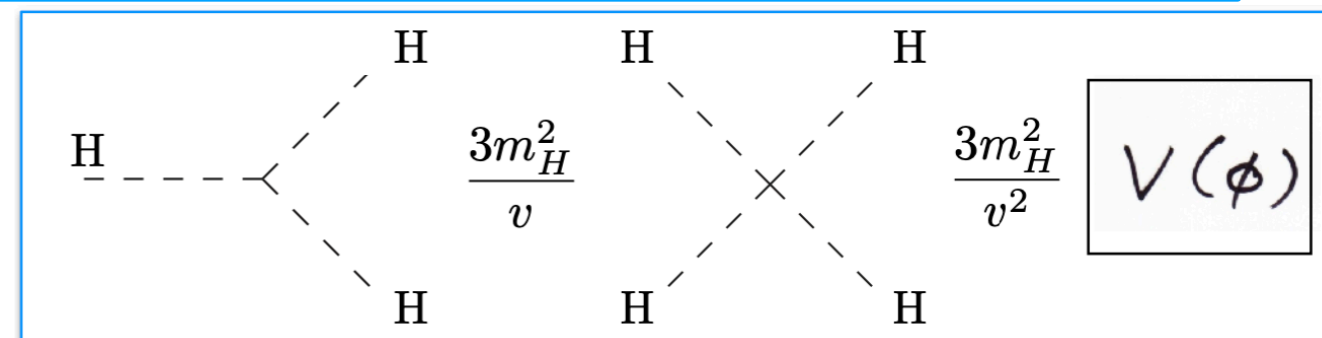


Constraints on charm Yukawa  $1.1 < \kappa_c < 5.5$

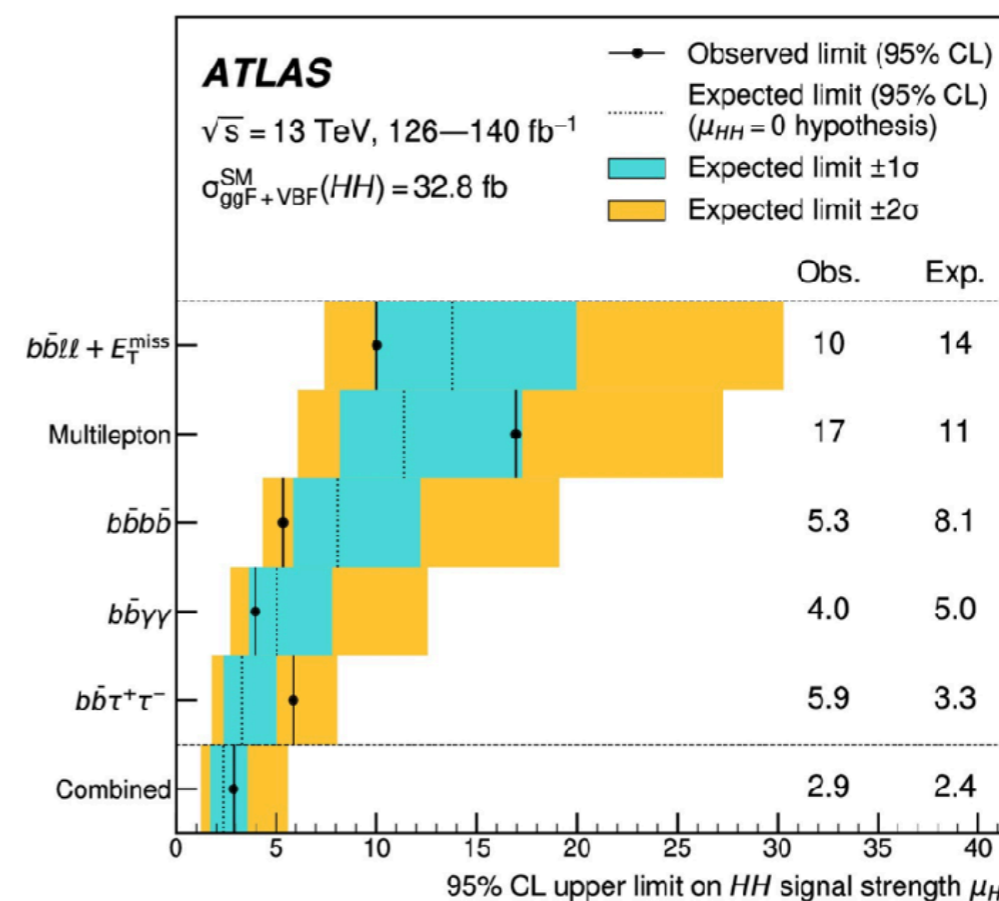
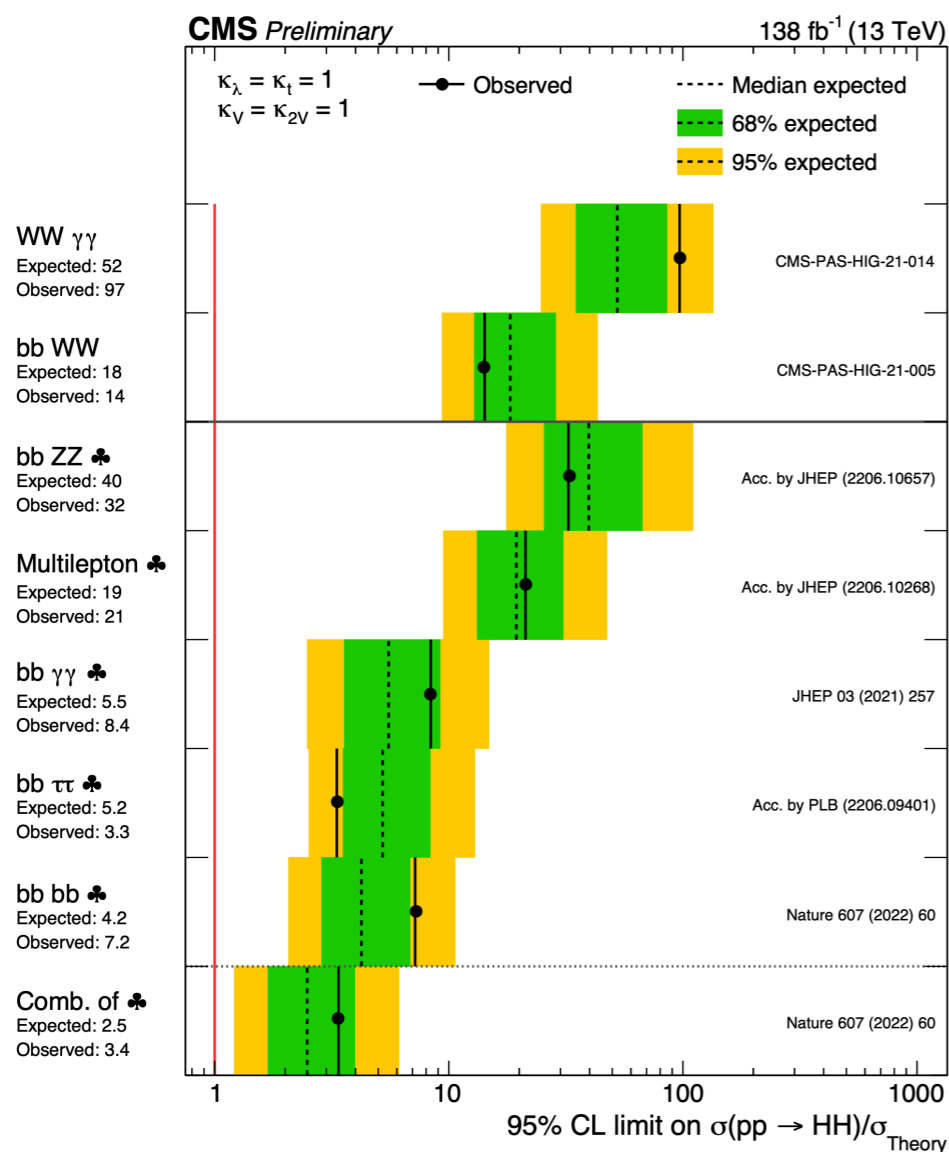
Yields a precision on  $\kappa_c$  of ~40% per experiment at HL-LHC

New perspective at the LHC!

- ✦ **Higgs Self Coupling and HH Production**
- ✦ **Despite the fact that in “Vanilla SUSY and vanilla composite models it is difficult to have large deviations in trilinear w.r.t. vector boson coupling”**
- ✦ **Large trilinear deviations are possible while deviations of the Higgs to Z coupling remain small.**



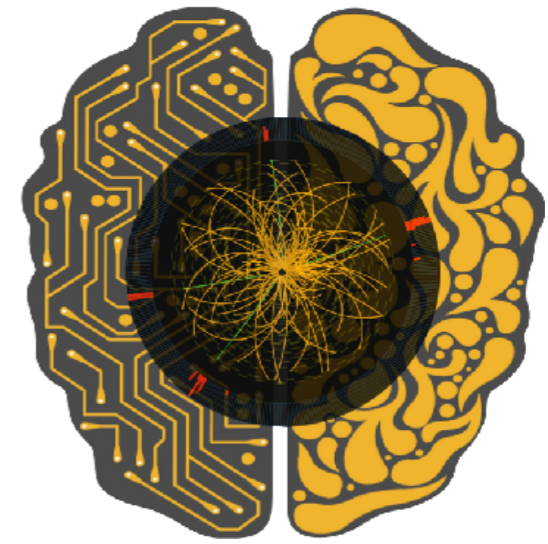
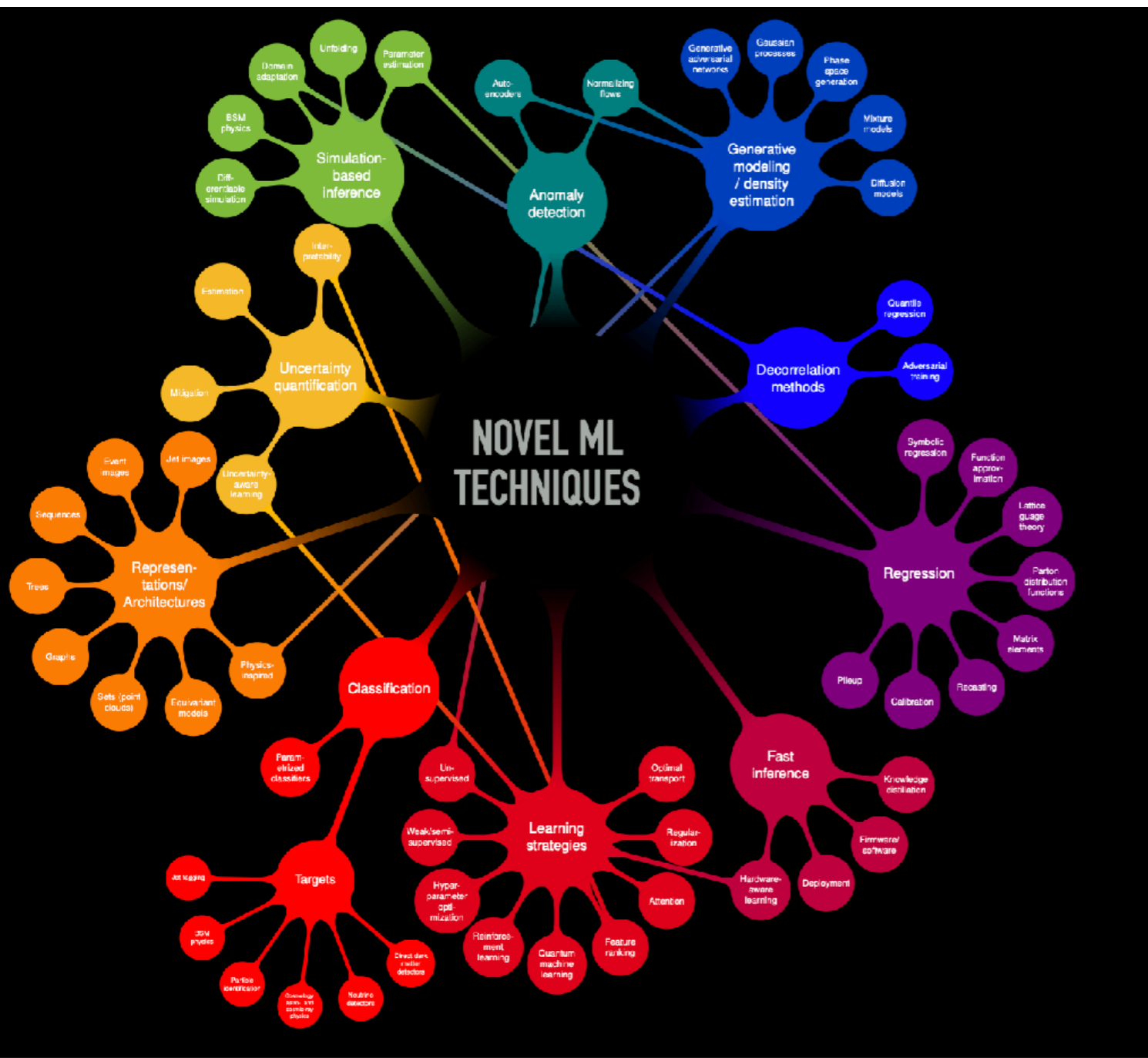
“Arguably the most important of them all!”



Observed limits start deviating from expectation!!

Both experiments have  $\sim 1\sigma$  sensitivity to a signal (Obs. ATLAS  $0.4\sigma$  and CMS  $\sim 1\sigma$ ) with Run 2!!

Naive comb. ATLAS-CMS sensitivity with Run 3 close  $2.5\sigma$  with improvements (and as much data as possible) **aim at  $3\sigma$**

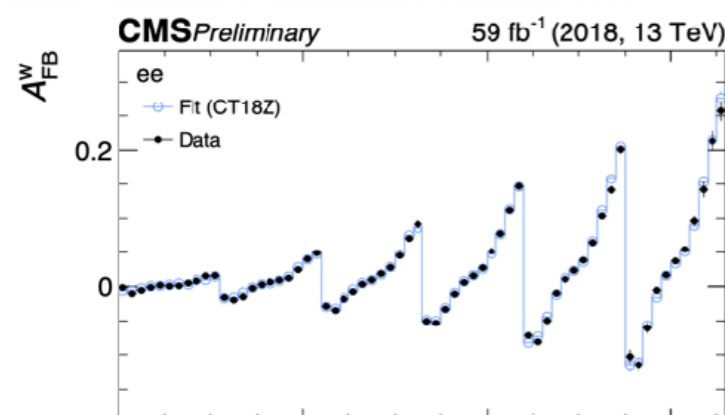


Array of ML opportunities beyond classification and regression, in simulation, unfolding, anomaly detection, etc.

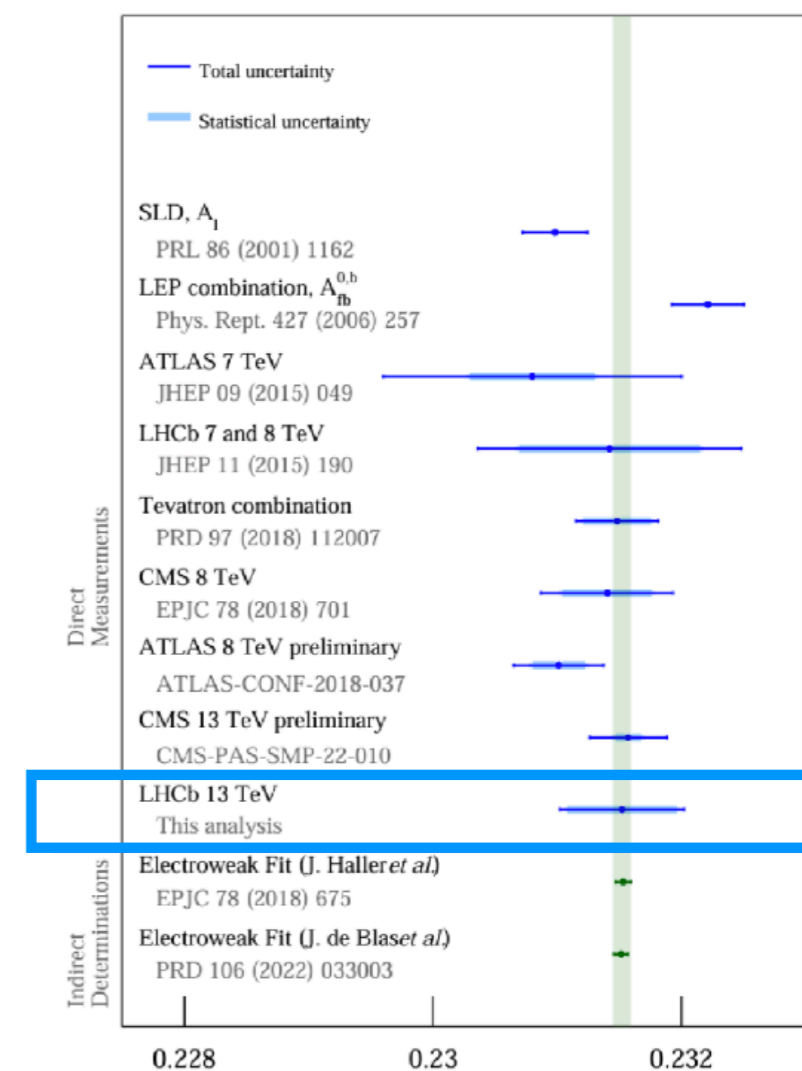
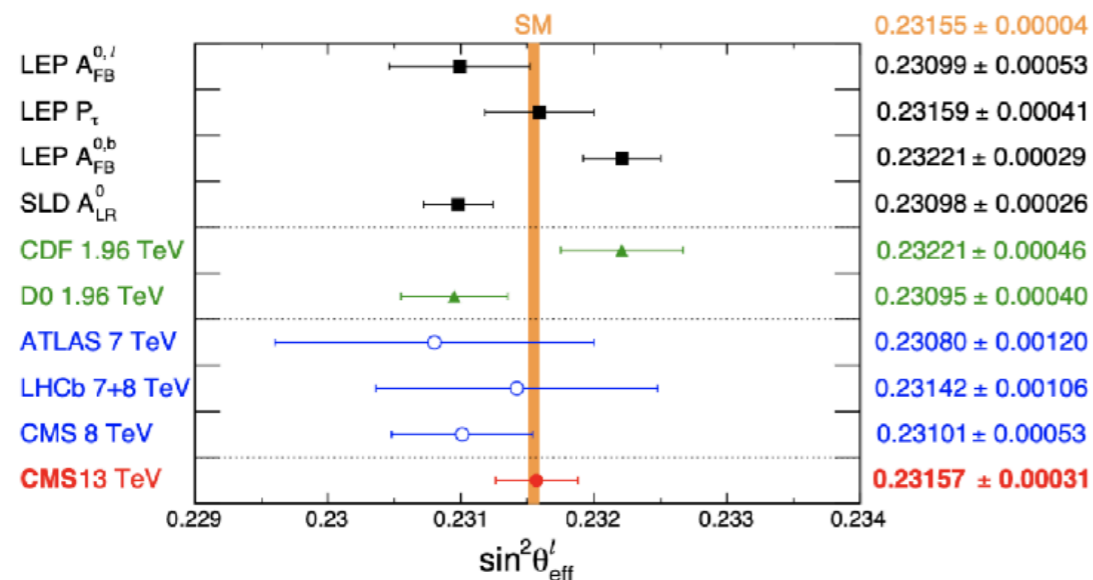
New ideas also have to be concerned with **robustness and interpretability**

## New measurements of $\sin^2 \theta_W$ by CMS and LHCb through $A_{FB}$

$A_{FB}$  in pp collisions is a tricky question, the forward region direction is given by the valence quark i.e. the system boost direction.



Precision comparable to the most precise single measurements at LEP  $A_{FB}^b$  and SLD  $A_{LR}$  determination



$$\sin^2 \theta_{eff}^l = 0.23152 \pm 0.00044 \pm 0.00005 \pm 0.00022$$

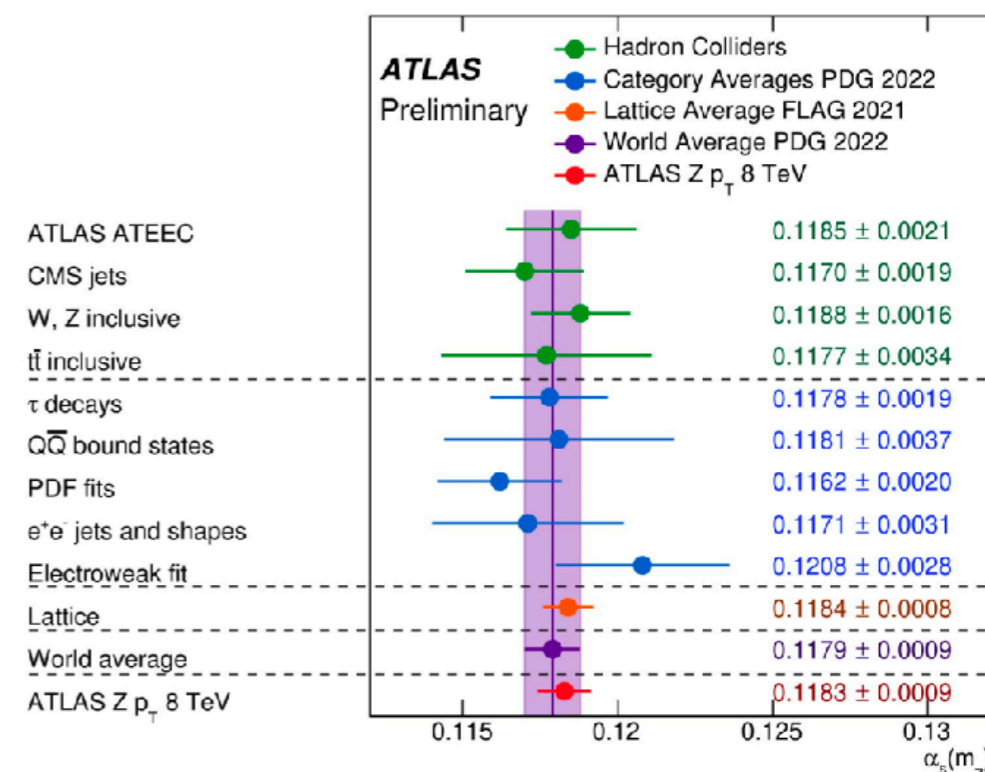
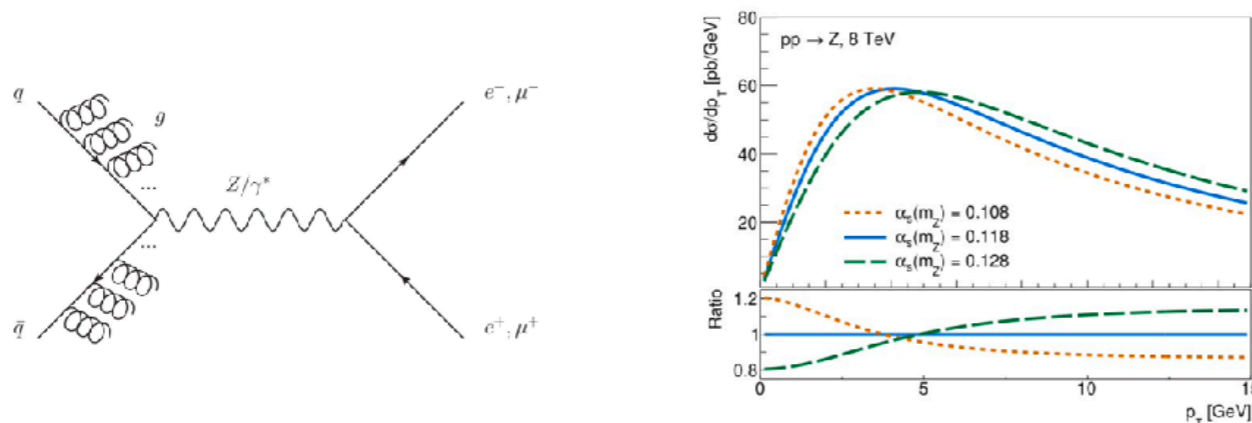
## Conditions for a competitive measurement of $\alpha_s$

- 1) is proportional to the strong coupling constant;
- 2) can be predicted theoretically with a percent precision (NNLO and higher);
- 3) is independent (nearly independent) of poorly-known parton distribution functions;
- 4) refers to low(er) region of hard momentum region;
- 5) does not suffer from unknown non-perturbative effects.

**“Inclusive Z transverse momentum seems to fit the bill”!**

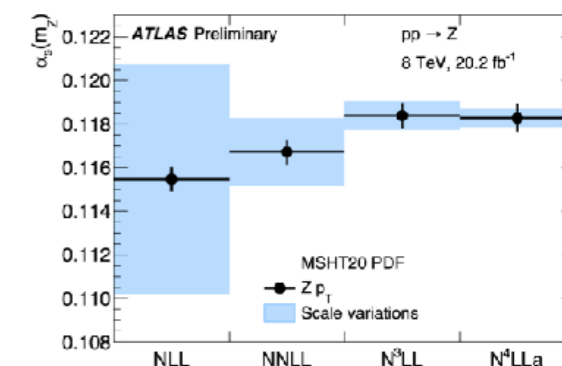
$$\frac{d\sigma_Z}{\sigma_Z dp_\perp} \sim \frac{\alpha_s(p_\perp)}{2\pi p_\perp} \ln \frac{M_Z}{p_\perp}$$

Using **Sudakov peak** in  $p_T$ , based on **resummed calculations**

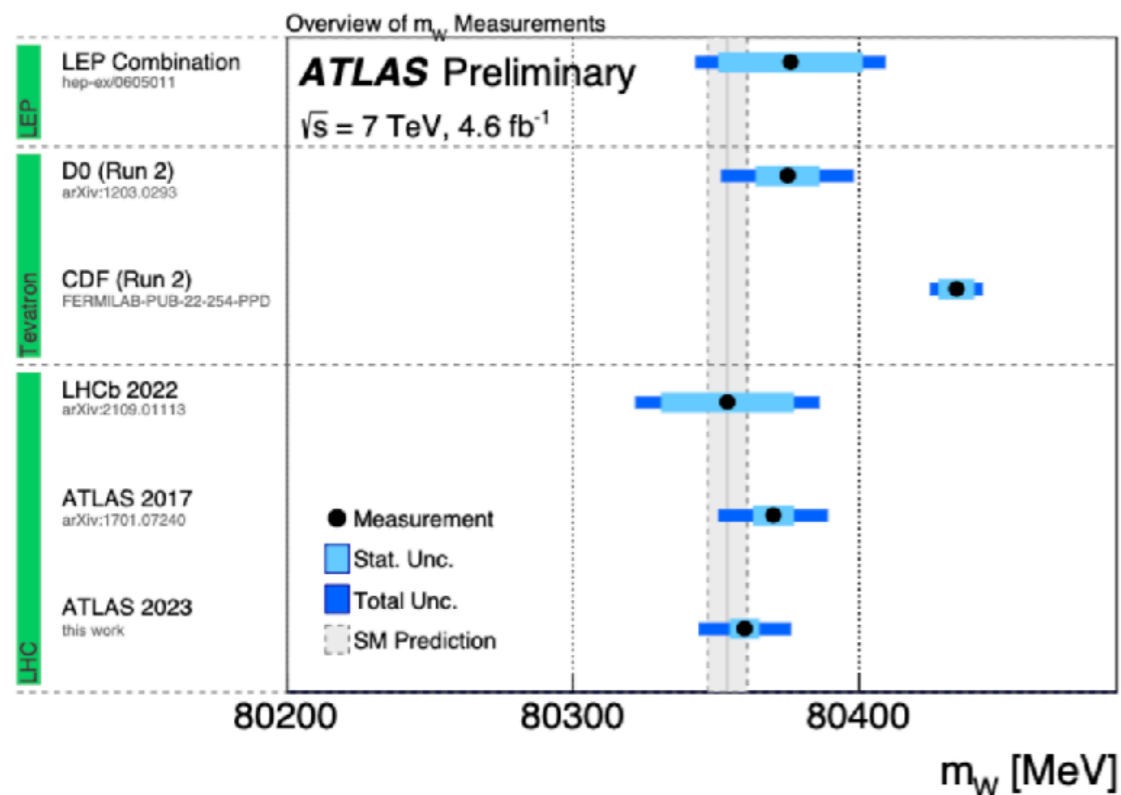


**Precision on par with lattice QCD and world average!**

**Such precision would not be possible without precise TH predictions!**



## Measurements at LHC from ATLAS and LHCb



The measurement relies on the ratio of W/Z pT (as noted by Kirill) non trivial QCD and EW corrections can modify this spectrum and bring correction of up to ~20 MeV

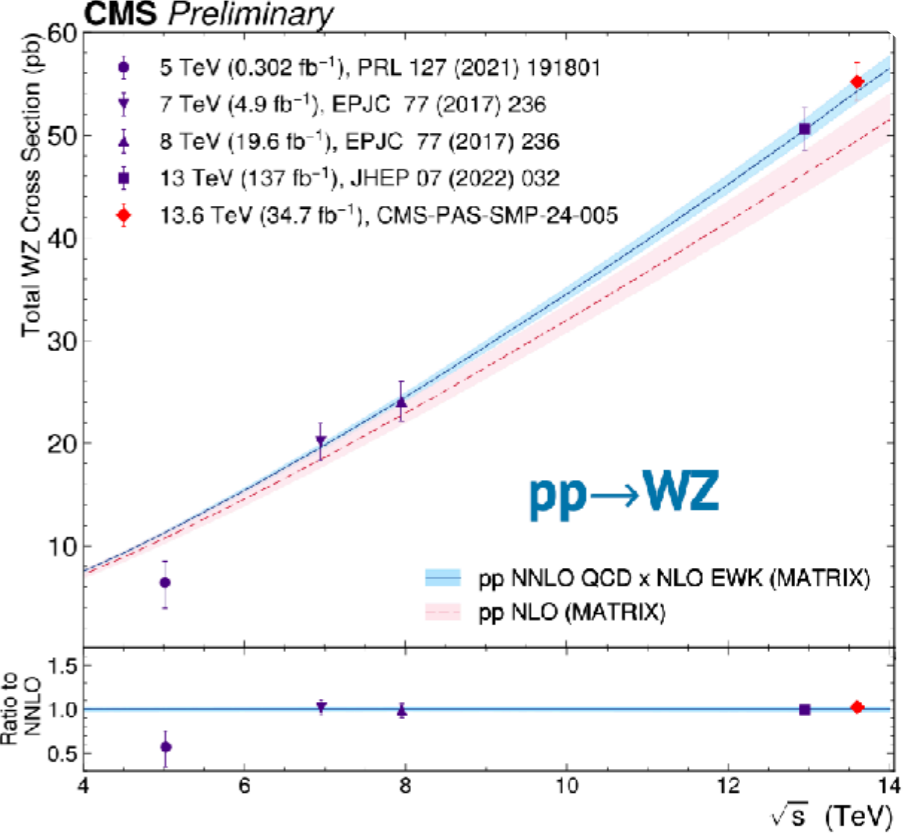
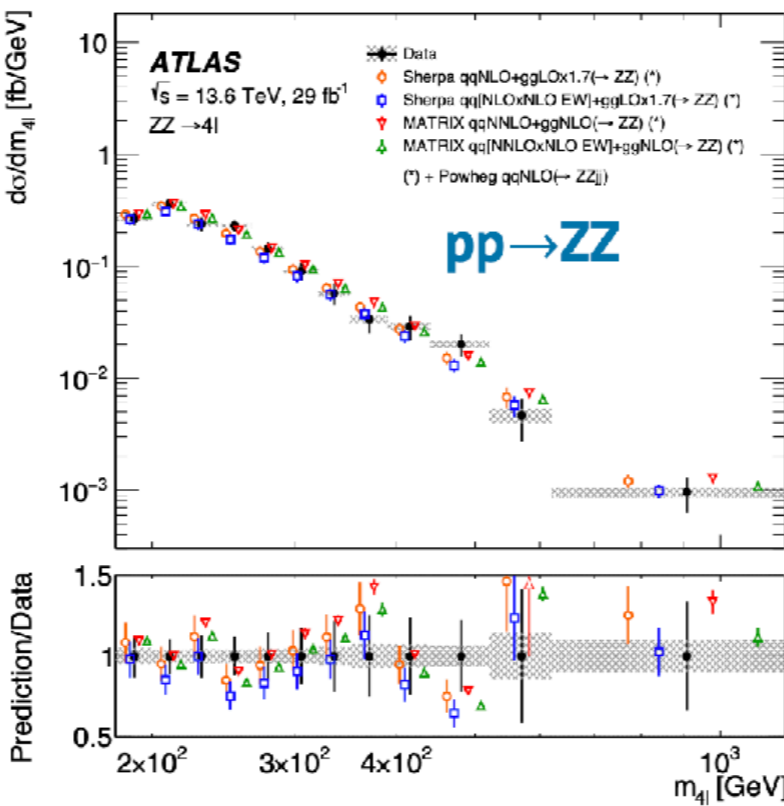
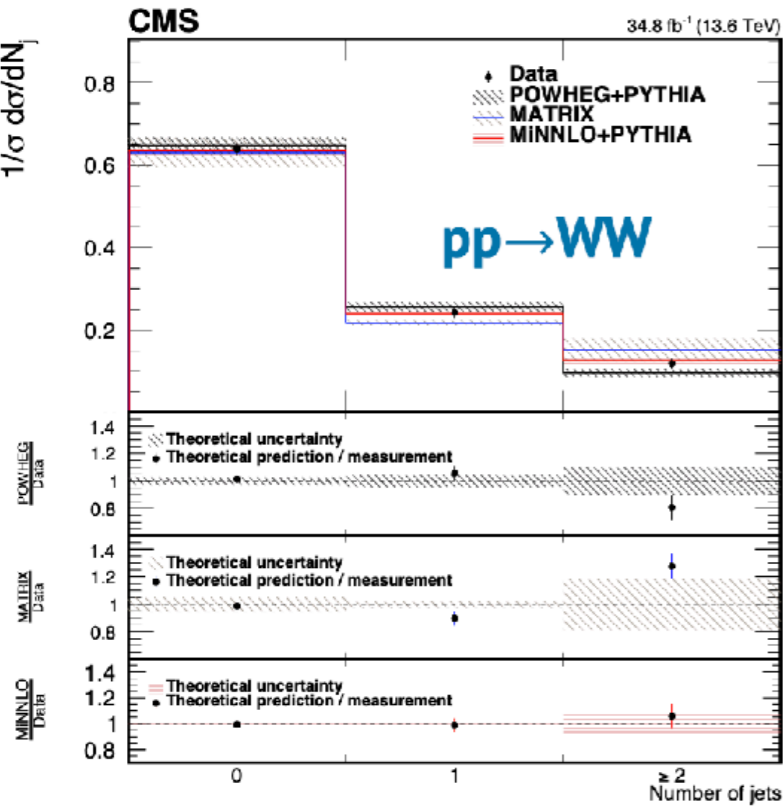
Before discussing the tension of the CDF measurement with the SM, need to address the tension between measurements!

The tension between ATLAS and CDF W mass is of  $4\sigma$

Significant evidence of measurement systematic bias!

$$m_W = 80360 \pm 5_{(\text{stat.})} \pm 15_{(\text{syst.})} = 80360 \pm 16 \text{ MeV}$$

CERN [press release](#)



First measurements of the diboson processes at Run 3 (with 2022 data)

In top pair production at the LHC, top quarks are **not produced polarised**, however a **spin correlation** exists.



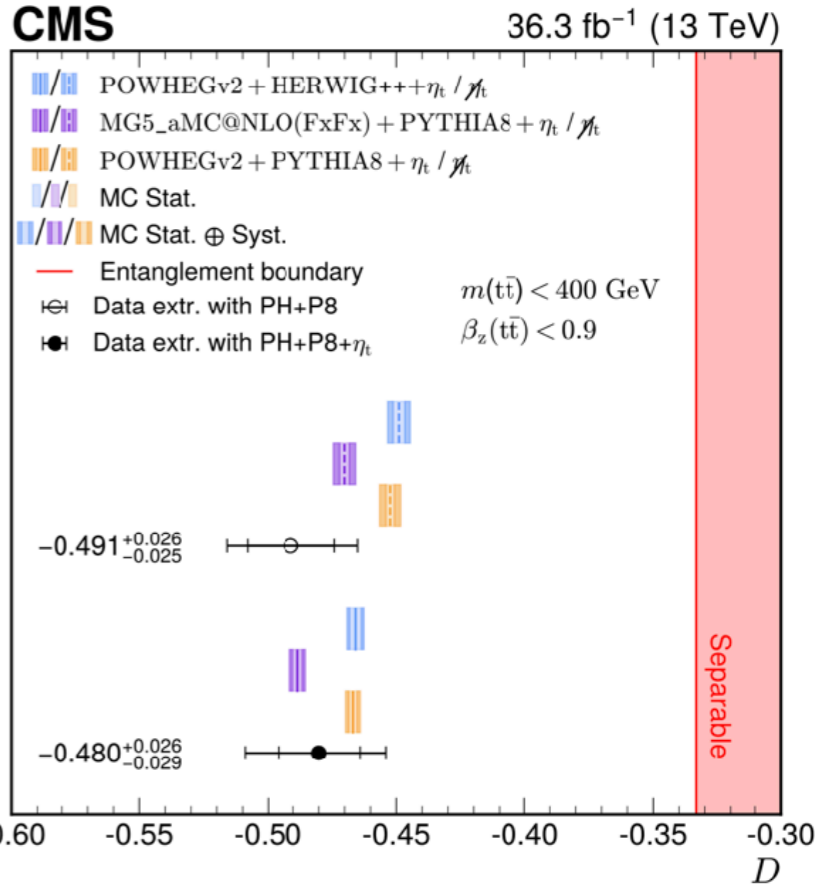
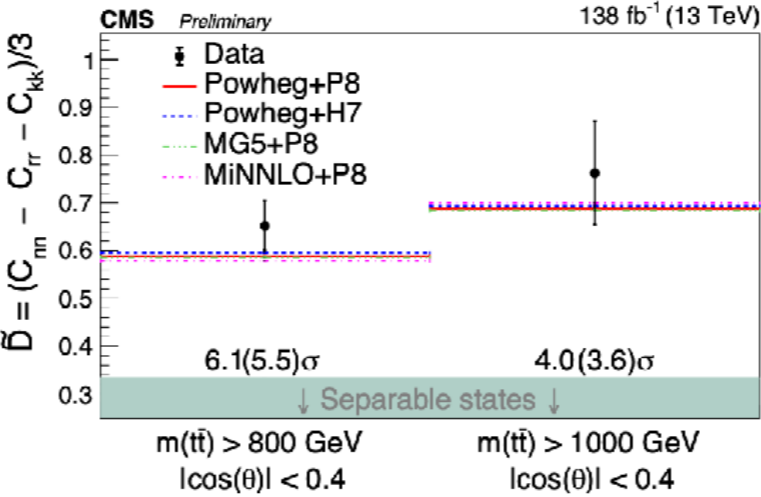
At threshold the  $gg \rightarrow t\bar{t}$  production is dominated by the “singlet” spin configuration, **which is a pure, superposed and maximally entangled Bell state:**

$$\frac{1}{\sqrt{2}} (|\uparrow\downarrow\rangle - |\downarrow\uparrow\rangle)$$

From the measurement of the spin density matrix we can probe whether this correlation is of quantum nature or not!

Initially measured near threshold where it is easier! CMS went beyond with:

- At production threshold in  $t\bar{t} \rightarrow b\ell\nu b\ell\nu$  events
- At high  $m_{t\bar{t}}$  with  $t\bar{t} \rightarrow b\ell\nu bq\bar{q}$  events, (phase space dominated 90% by space-like events)

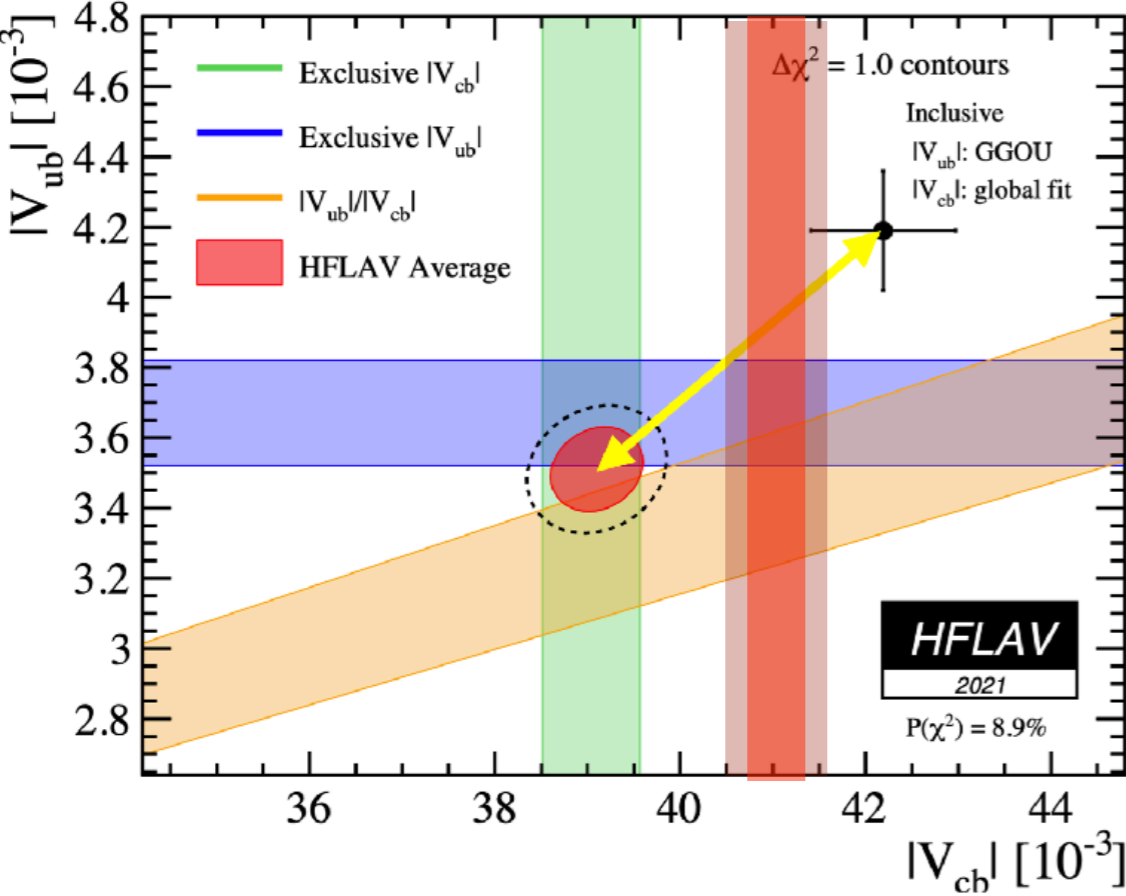


Very important elements (space-like) to go beyond entanglement towards the **violation of Bell Inequalities!** (With higher sensitivity)



## New measurements from Belle II and Babar

$|V_{cb}|$  and  $|V_{ub}|$  discrepancy  $\sim 3\sigma$  between exclusive and inclusive (have different TH uncertainties)!



Limiting factor in precision flavour physics!

Inclusive extraction from  $B \rightarrow X_c \ell \nu$ , where  $X_c$  goes to anything ([Bernlochner et al.](#) and [Bordone et al.](#)).

New exclusive  $|V_{cb}|$  results from BaBar and Belle II using fully differential information for the first time!

$$B \rightarrow D \ell \nu$$

$$|V_{cb}| = (41.1 \pm 1.2) \times 10^{-3}$$

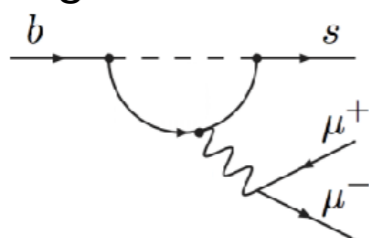
$$B \rightarrow D^* \ell \nu$$

$$|V_{cb}| = (41.0 \pm 0.7) \times 10^{-3}$$

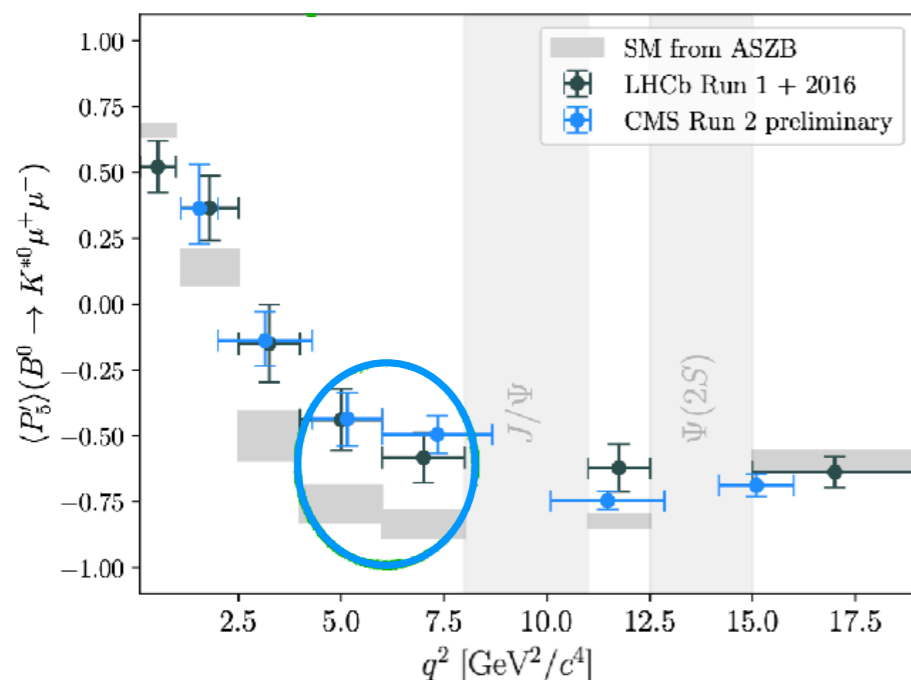
“Perhaps on the right path to resolve this puzzle”

## RK\* from LHCb cancelled end of 2022

Still a 2-3 Standard Deviation in the angular distribution and absolute branching fractions of  $B^0 \rightarrow K^{0*} \mu^+ \mu^-$  ([paper](#))



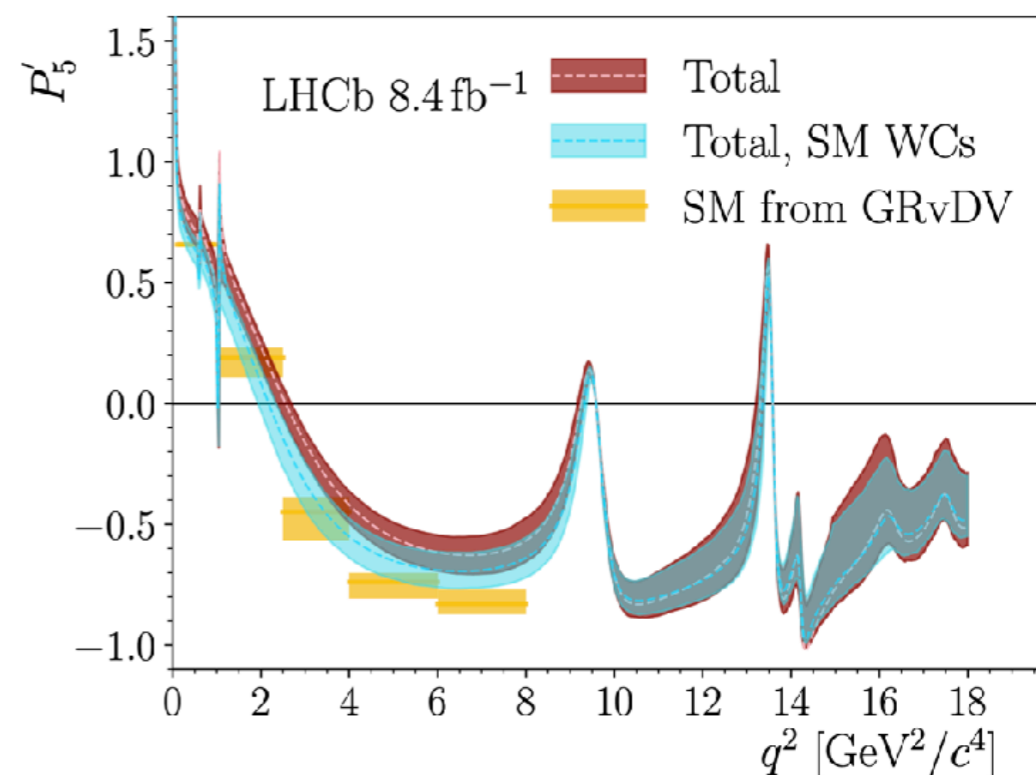
## $B^0 \rightarrow K^{0*} \mu^+ \mu^-$ also from CMS



$P_5'$  angular observable essentially free from form factor uncertainties.

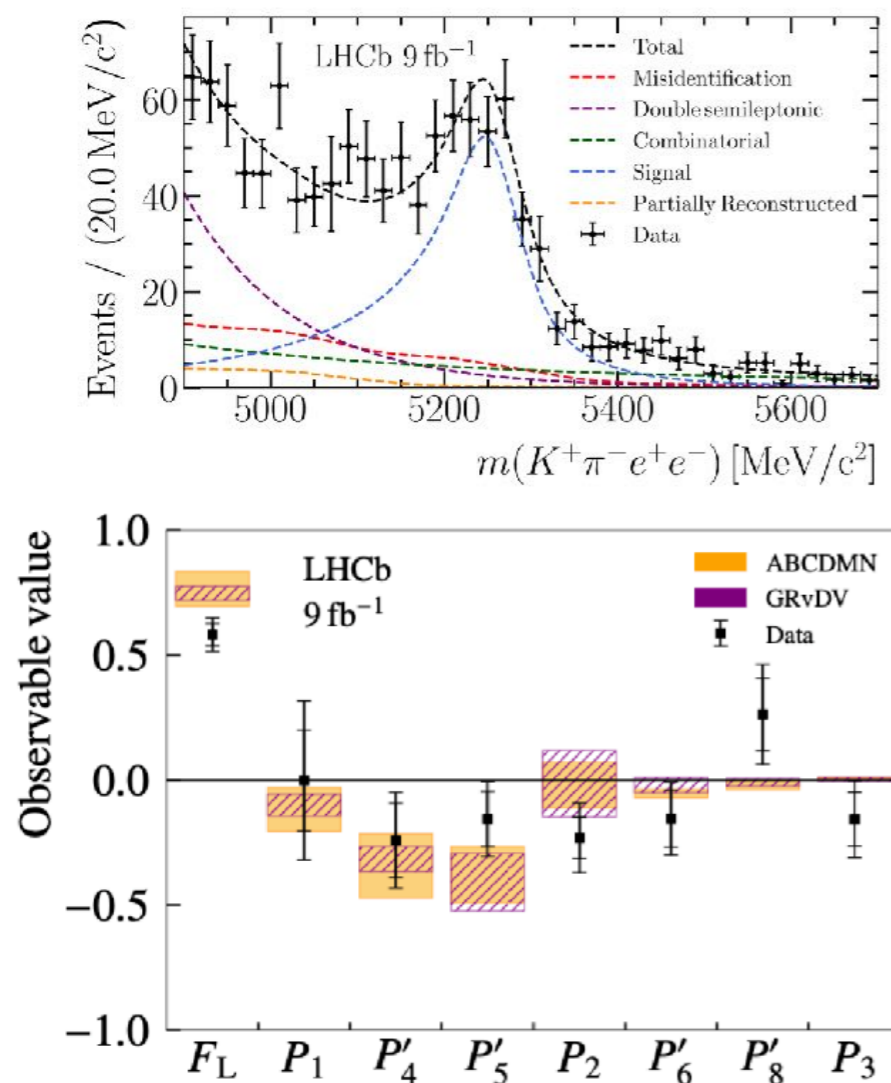
## Comprehensive analysis of local and nonlocal amplitudes $B^0 \rightarrow K^{0*} \mu^+ \mu^-$ ([paper](#))

Careful long-distance contributions weaken these tensions are not considered



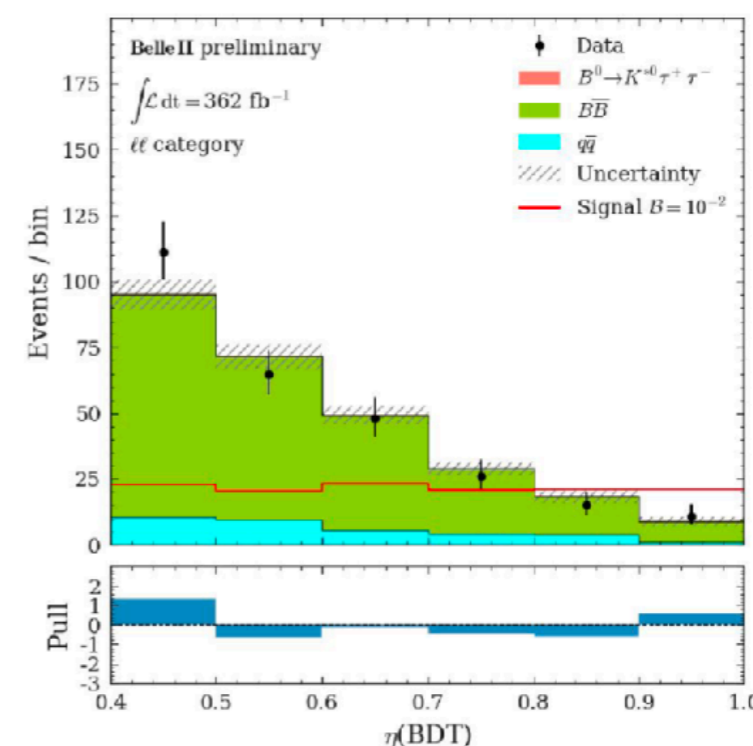
Tension now down to  $2.1\sigma$

## New $B^0 \rightarrow K^{0*} e^+ e^-$ measurement from LHCb



## New $B^0 \rightarrow K^{0*} \tau^+ \tau^-$ measurement from Belle II

Analysis particularly sensitive to new physics affecting the  $B^0 \rightarrow K^{0*} \mu^+ \mu^-$  decay!



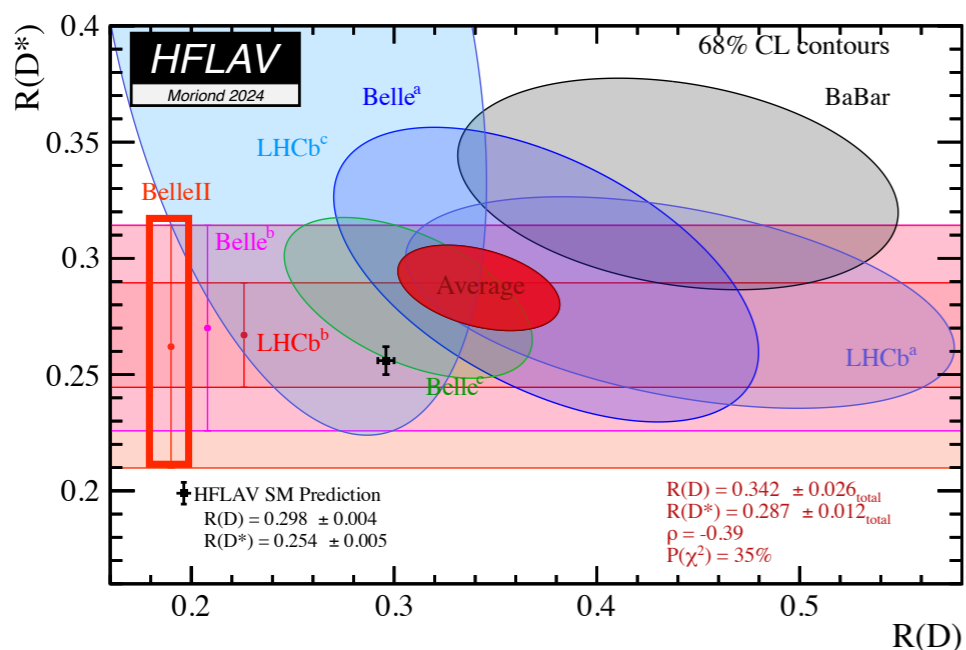
$$\mathcal{B}(B^0 \rightarrow K^{*0} \tau^+ \tau^-) < 1.73 \times 10^{-3}$$

Limit twice improved over the Belle result!

## First Belle II $R_{D^*}$ measurement!

Both TH and EXP clean!

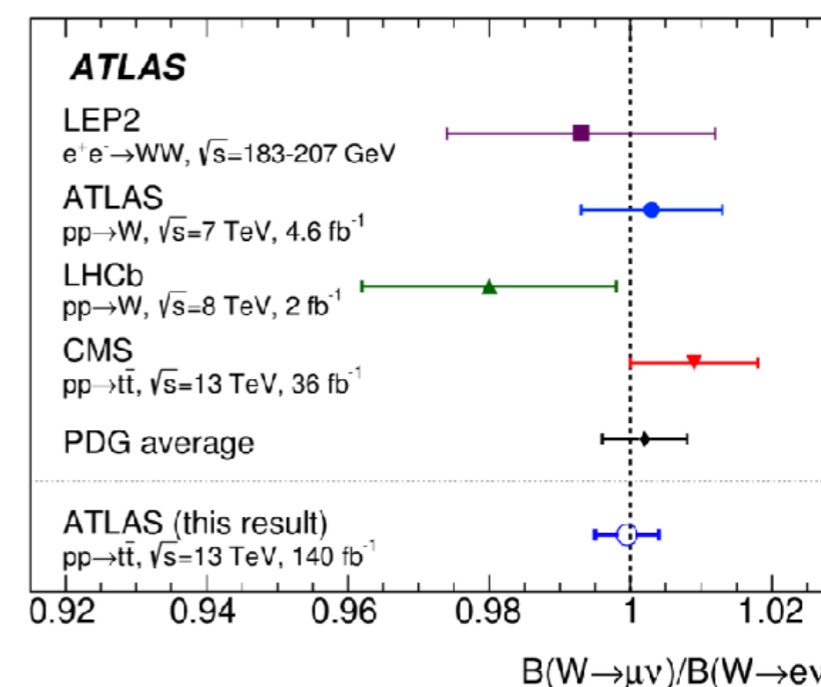
$$\mathcal{R}(D^{(*)}) = \frac{\mathcal{B}(B^0 \rightarrow D^{(*)-} \tau^+ \nu_\tau)}{\mathcal{B}(B^0 \rightarrow D^{(*)-} \mu^+ \nu_\mu)}$$



$$R_D^* = 0.26 \pm 0.04^{+0.04}_{-0.03}$$

- Systematic uncertainty related mainly to size of control samples
- Comparable precision to equivalent Belle result with 1/4 the sample

## Lepton universality measurements for “on-shell” W bosons in top decays



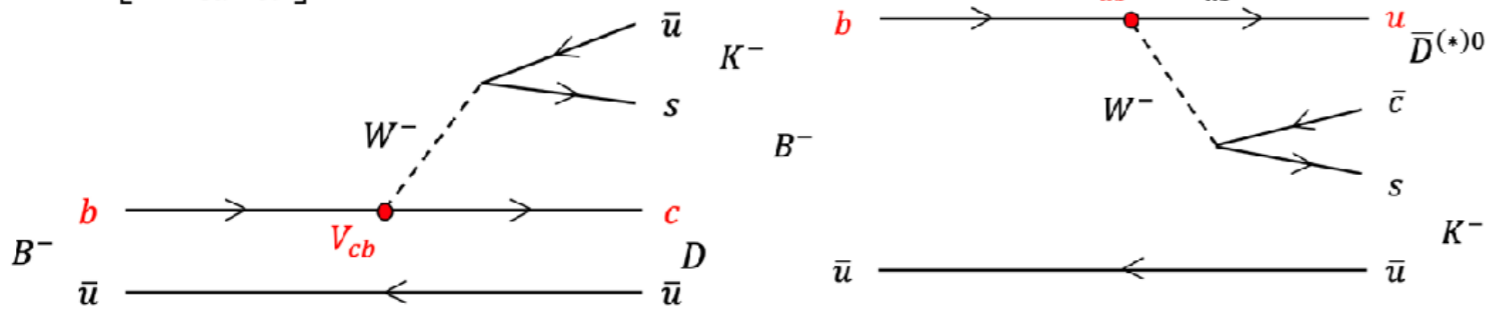
ATLAS result more precise than current world average

Also with W decays from tops CMS measures  $|V_{cs}|$  !

$$R_W^c = \frac{\mathcal{B}(W \rightarrow cq)}{\mathcal{B}(W \rightarrow cd)\mathcal{B}(W \rightarrow ud)} \quad |V_{cs}| = 0.959 \pm 0.021$$

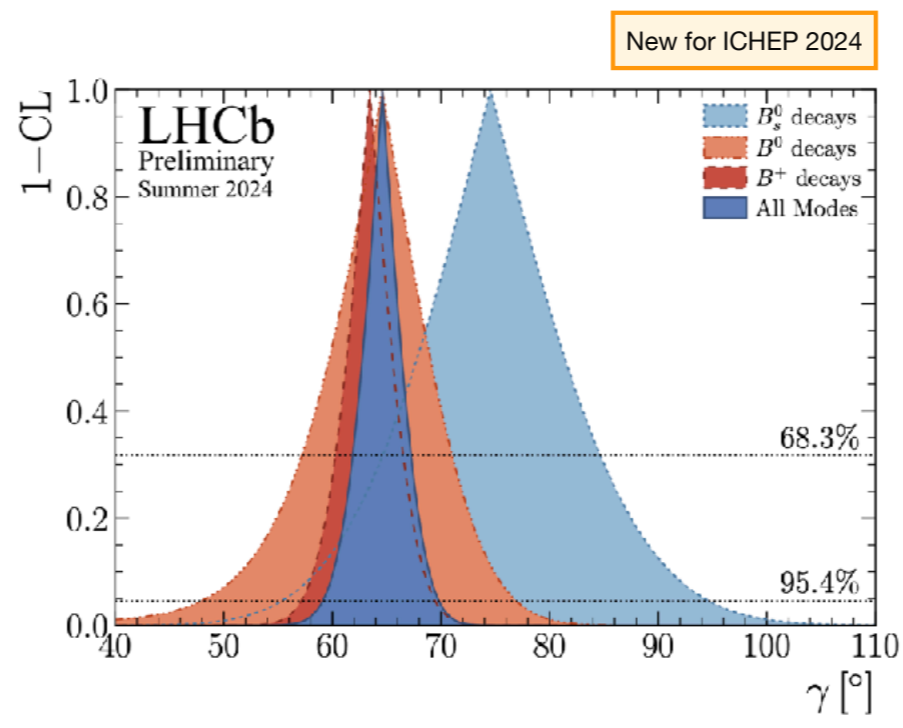
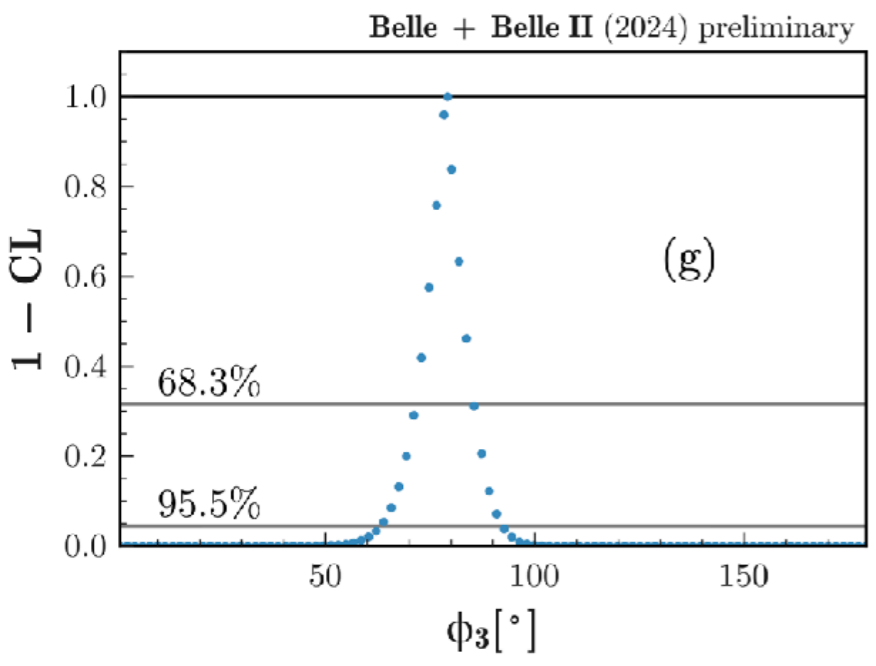
Recent Measurements of  $\gamma$  in the golden channel  $B^\pm \rightarrow DK^\pm$

$$\gamma = \arg \left[ -\frac{V_{ud}V_{ub}^*}{V_{cd}V_{cb}^*} \right]$$



Lack of Lattice QCD needs makes it a **“pristine observable”** in flavour physics!

Charm input from BESIII/CLEO is critical



Combination from Belle II

$$\gamma = (78.6^{+7.2}_{-7.3})^\circ$$

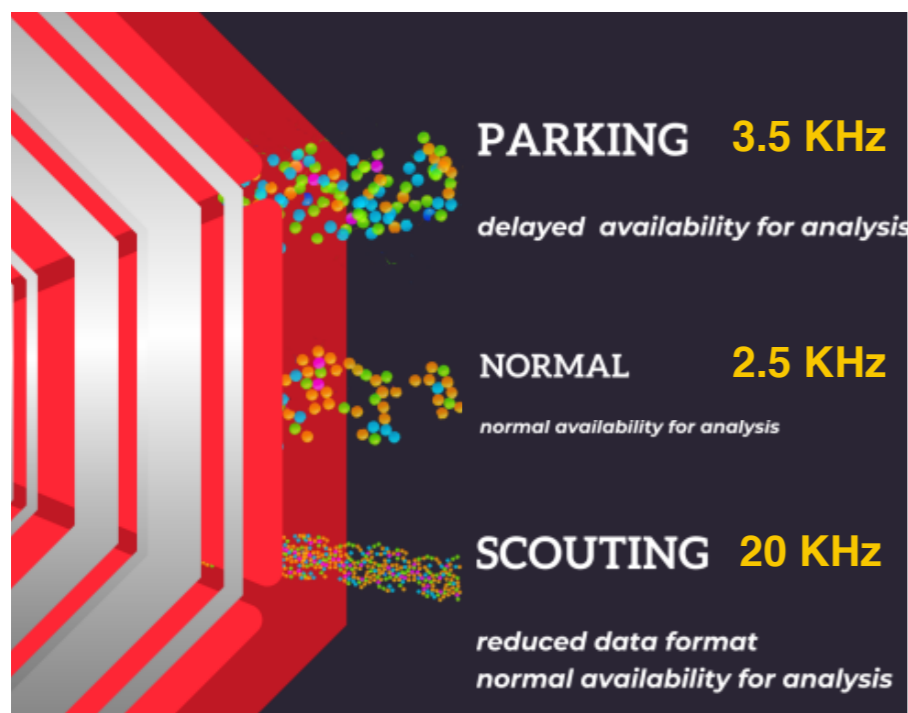
Combination from LHCb!

$$\gamma = (64.7 \pm 2.8)^\circ$$

**Measurement from LHCb has surpassed the target goal for Run 2!!**

From CKM fitter  $\gamma = (66.3^{+0.7}_{-1.9})^\circ$

CMS-EXO-23-007

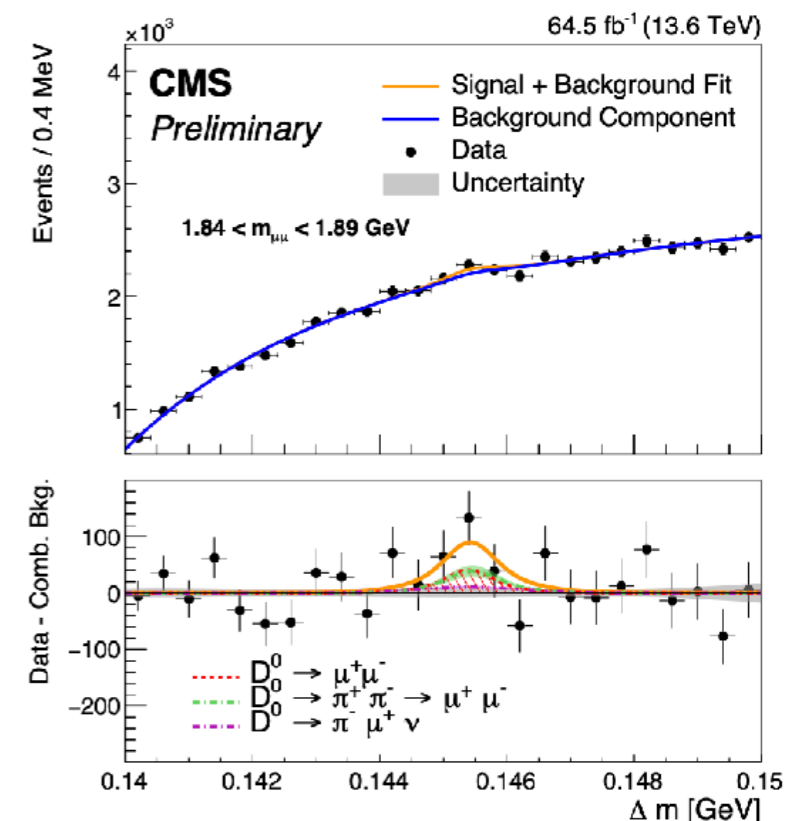


## First Results using Run3 Parking

Search for  $D^0 \rightarrow \mu^+ \mu^-$  - improved by 35% over previous best limit

Further improvements foreseen!

CMS-BPH-23-008



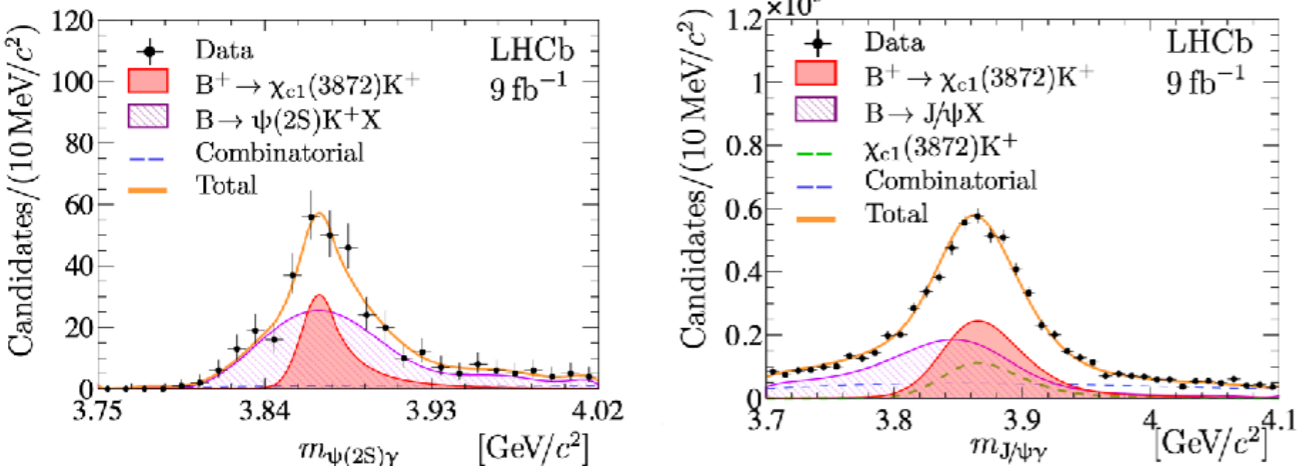
Large scale project to support the developments (with powerful hardware and software) to bring experiments to the next level with efficient data flows and structures, with ambitious and large ML models e.g. GNN tracking!

See [talk](#) by D. Rohr

Observation of  $\chi_c(3872)$  radiative decays at LHCb!

$\chi_c(3872)$  discovered 20 years ago, could it be D meson molecular state?

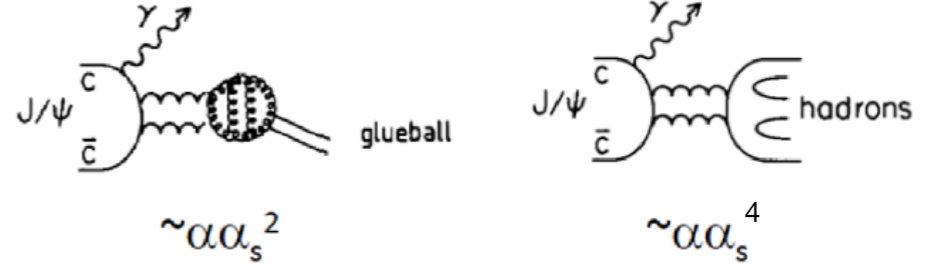
Now, the LHCb collaboration is closer to finding out what it is made up of!



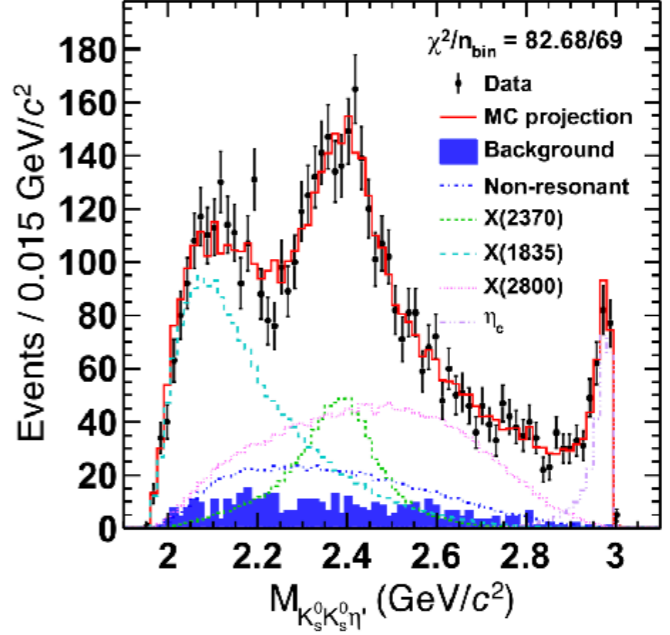
$$R_{\psi\gamma} = \frac{\mathcal{B}_{B^+ \rightarrow (\chi_{c1}(3872) \rightarrow \psi(2S)\gamma)K^+}}{\mathcal{B}_{B^+ \rightarrow (\chi_{c1}(3872) \rightarrow J/\psi\gamma)K^+}} = 1.67 \pm 0.21 \pm 0.12 \pm 0.04.$$

Strong indication of a sizeable charmonium or tetraquark compact component of the X(3872)!

Discovery of a Glueball-like particle X(2370) at BESIII

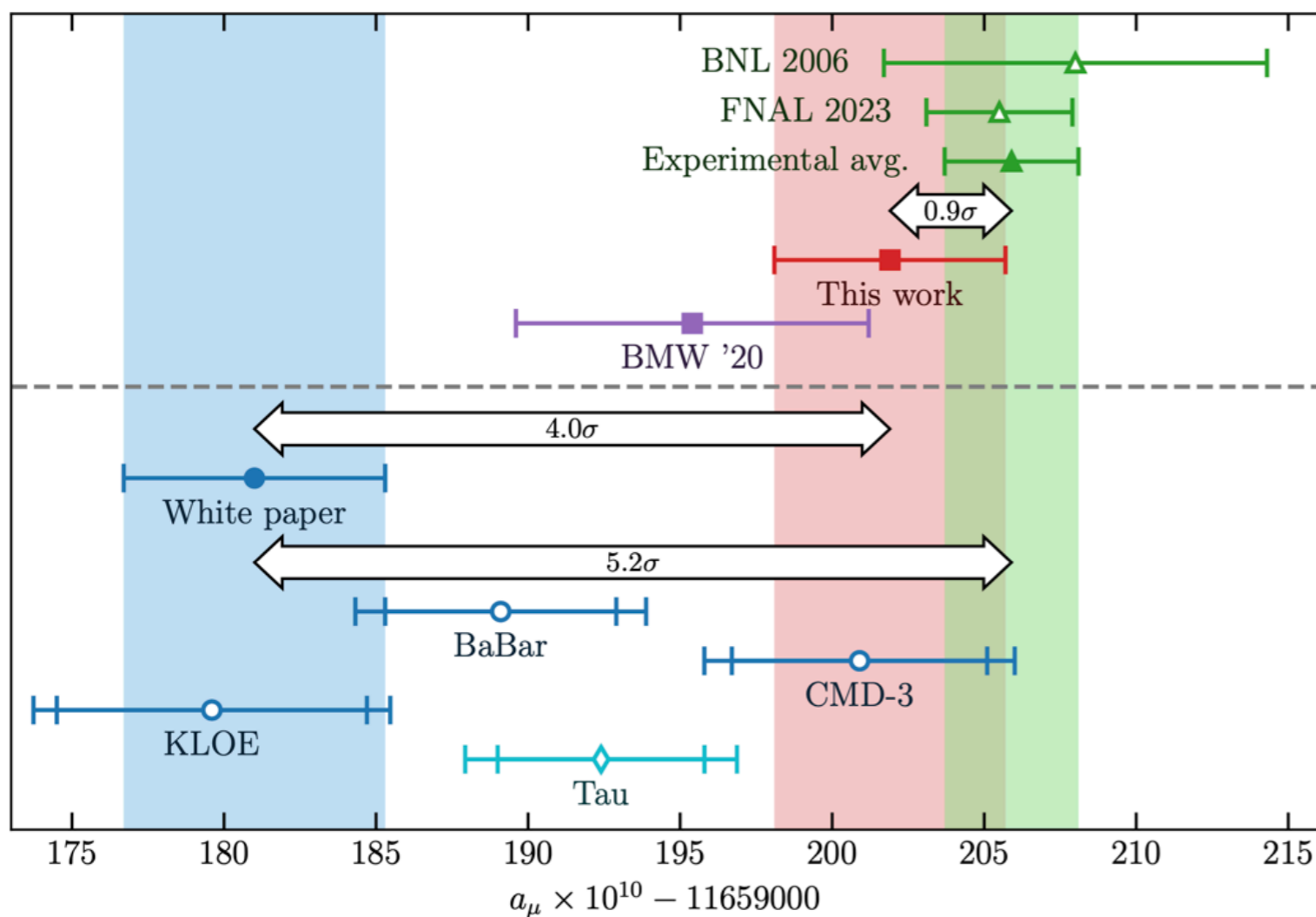


Radiative  $J/\psi$  decays are gluon rich!



Candidate for elusive lightest pseudoscalar glueball predicted by LQCD

Several important news from the front of  $(g_\mu - 2)$  predictions from BaBar and Lattice in conjunction with data!



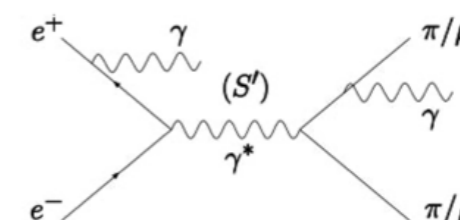
Scanning  $e^+e^- \rightarrow \pi^+ \pi^-$  ECM = 0.32-2 GeV  
**CMD-3** at VEPP-2000  $e^+e^-$  collider

Better detector performance Larger statistics (x30 CMD-2)

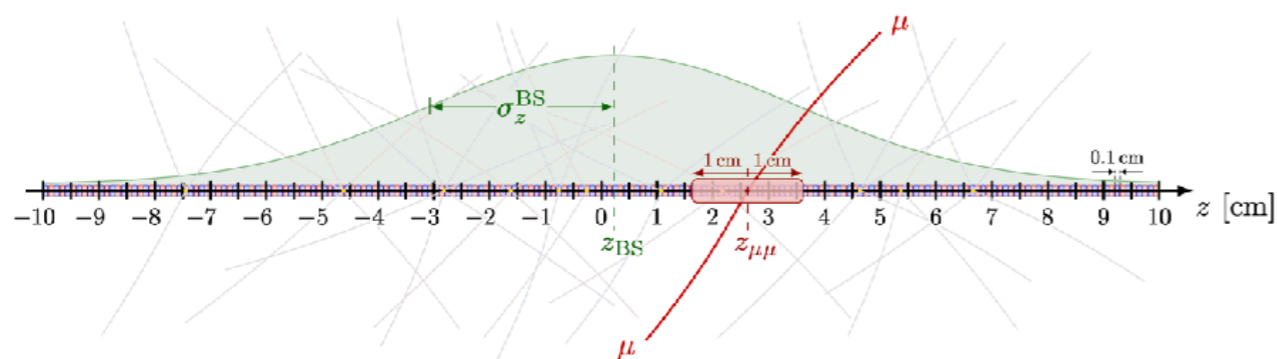
New BMW result including finer lattice and long distance effects from  $e^+e^-$  data!

New BaBar studies of higher order radiation and impact on the vacuum polarisation predictions of  $(g-2)$ !

**NNLO Radiative corrections need to be better understood and accounted in analyses!**







Beautiful analysis selecting isolated low multiplicity vertices, sensitive to photo-production of tau pairs!

Large gain in sensitivity! Only ~3 times larger than the Schwinger term (QED part)

However still almost 3-4 orders of magnitude above sensitive corrections e.g. EW!

$$\begin{aligned}
 a_\tau^{\text{QED}} &= 1.1732 \times 10^{-3}, \\
 a_\tau^{\text{had}} &= 3.2(4) \times 10^{-6}, \\
 a_\tau^{\text{EW}} &= 4.7 \times 10^{-7}.
 \end{aligned}$$

## CMS

138 fb<sup>-1</sup> (13 TeV)

• Observed — 68% CL — 95% CL

**OPAL**  
 $ee \rightarrow Z \rightarrow \tau\tau\gamma$   
 PLB 434 (1998) 188

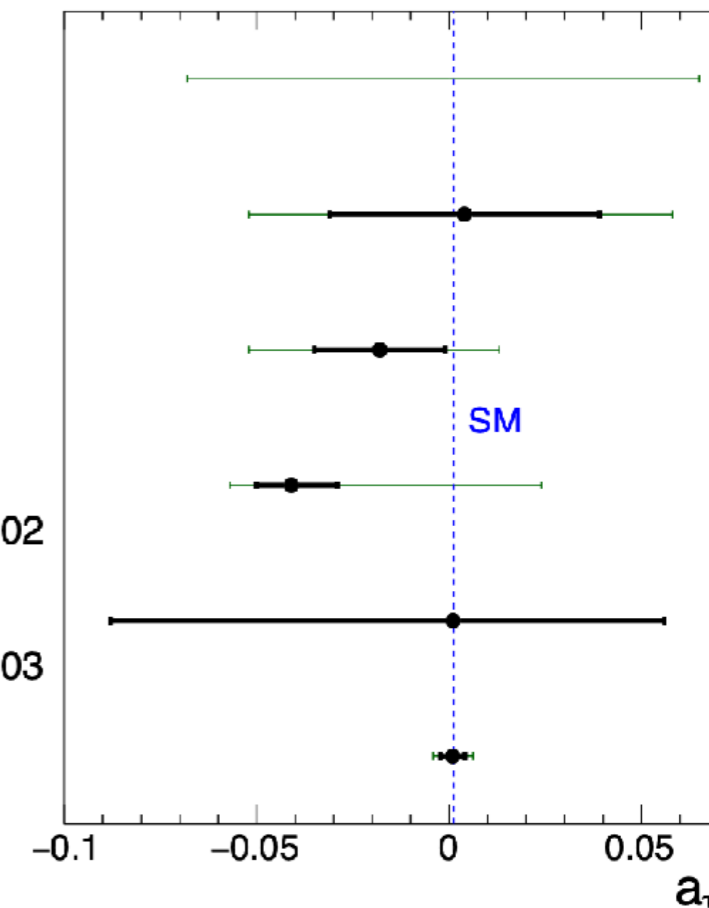
**L3**  
 $ee \rightarrow Z \rightarrow \tau\tau\gamma$   
 PLB 434 (1998) 169

**DELPHI**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from e)  
 EPJC 35 (2004) 159

**ATLAS**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from Pb)  
 PRL 131 (2023) 151802

**CMS**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from Pb)  
 PRL 131 (2023) 151803

**CMS**  
 $\gamma\gamma \rightarrow \tau\tau$  ( $\gamma$  from p)  
 This result



$$a_\tau = \frac{g_\tau}{2} - 1 = 0.0009^{+0.0032}_{-0.0031}$$

## Accelerator Neutrino Oscillations

The current two main players  $\nu_\mu$ -beam experiments!

### NOvA

#### Off Axis

Fermilab to Ash River

Improved sensitivity to mass ordering!

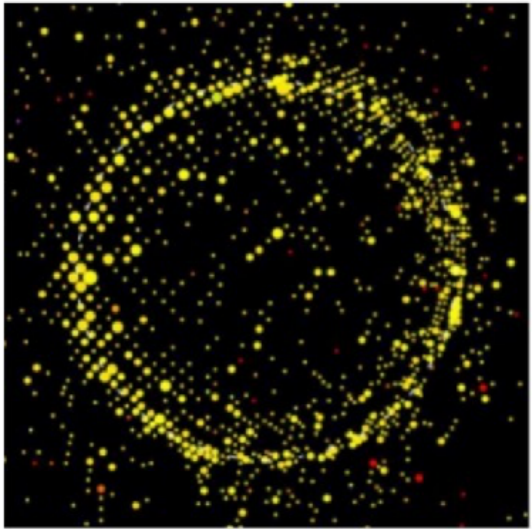
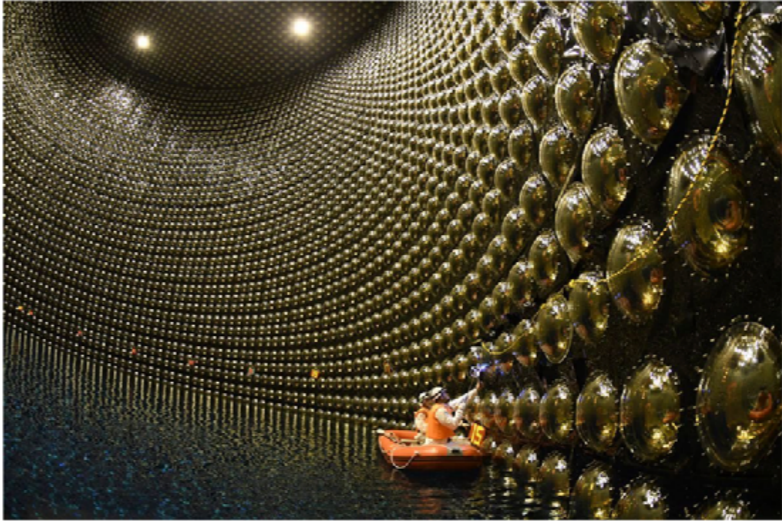
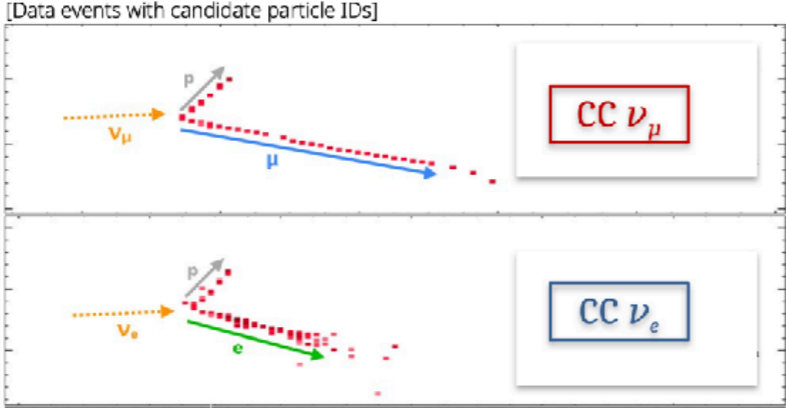
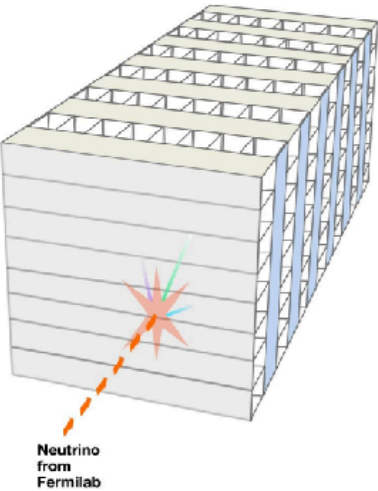
810 km/GeV - E 2 GeV - 0.8° off-axis

### T2K

Slightly off axis (J-PARC to Super K)

490 km/GeV - E 0.6 GeV - 2.5° off-axis

Sensitive to  $\nu_\mu$  disappearance and  $\nu_e$  appearance!



$\nu_e$ -like

## Accelerator Neutrino Oscillations

The current two main players  $\nu_\mu$ -beam experiments!

### NOvA

#### Off Axis

Fermilab to Ash River

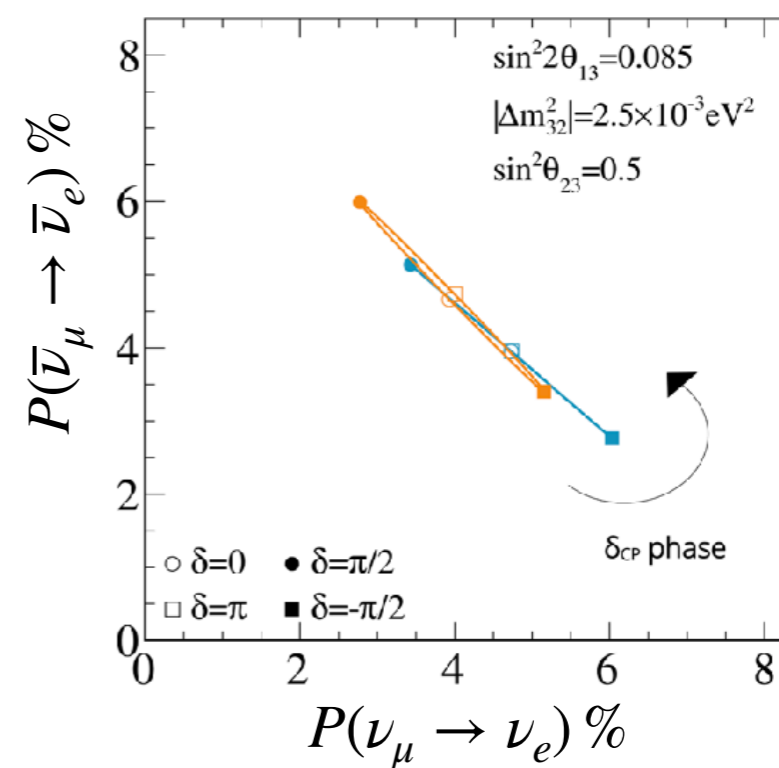
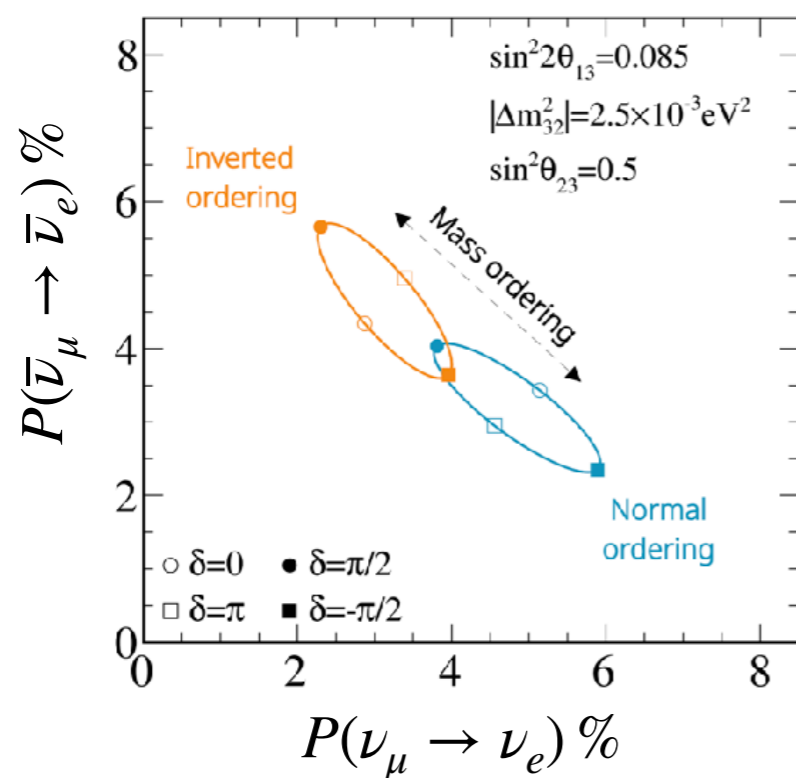
Improved sensitivity to mass ordering!

810 km/GeV - E 2 GeV -  $0.8^\circ$  off-axis

### T2K

Slightly off axis (J-PARC to Super K)

490 km/GeV - E 0.6 GeV -  $2.5^\circ$  off-axis



## Accelerator Neutrino Oscillations

The current two main players  $\nu_\mu$ -beam experiments!

New combination of the two experiments for this conference!

### NOvA

#### Off Axis

Fermilab to Ash River

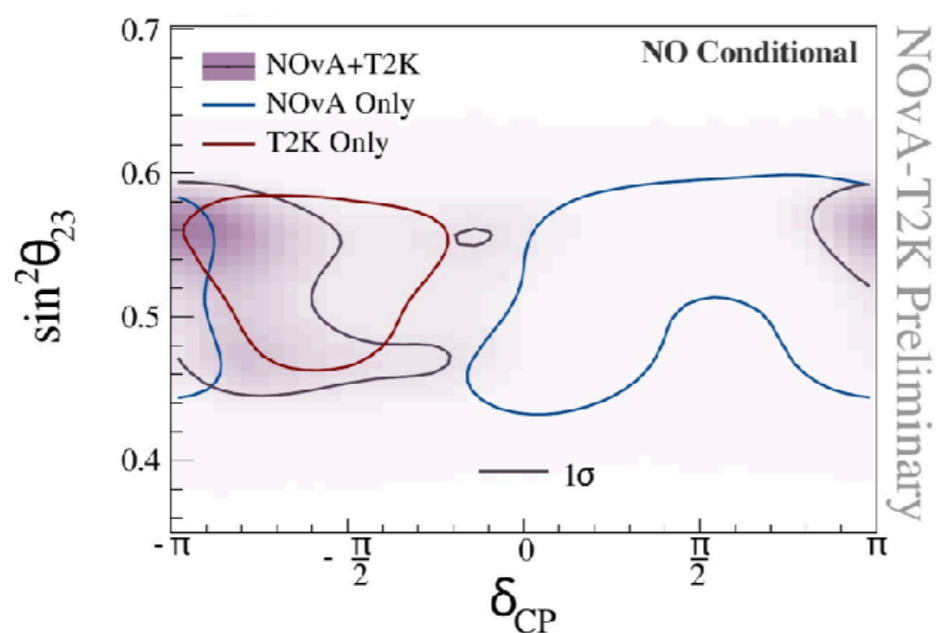
Improved sensitivity to mass ordering!

810 km/GeV - E 2 GeV -  $0.8^\circ$  off-axis

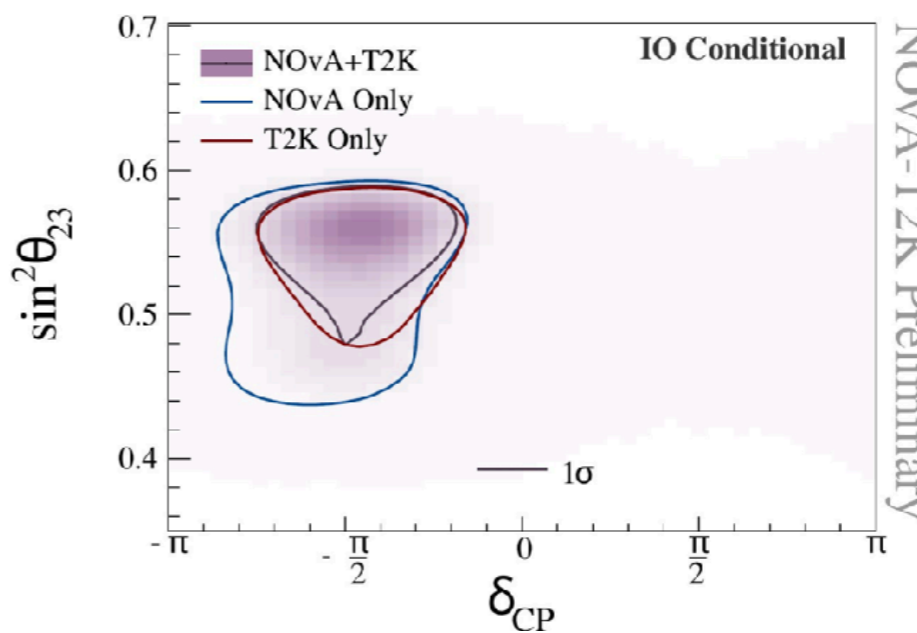
### T2K

Slightly off axis (J-PARC to Super K)

490 km/GeV - E 0.6 GeV -  $2.5^\circ$  off-axis



Mild preference for Inverted Ordering but influenced by  $\theta_{13}$  constraint

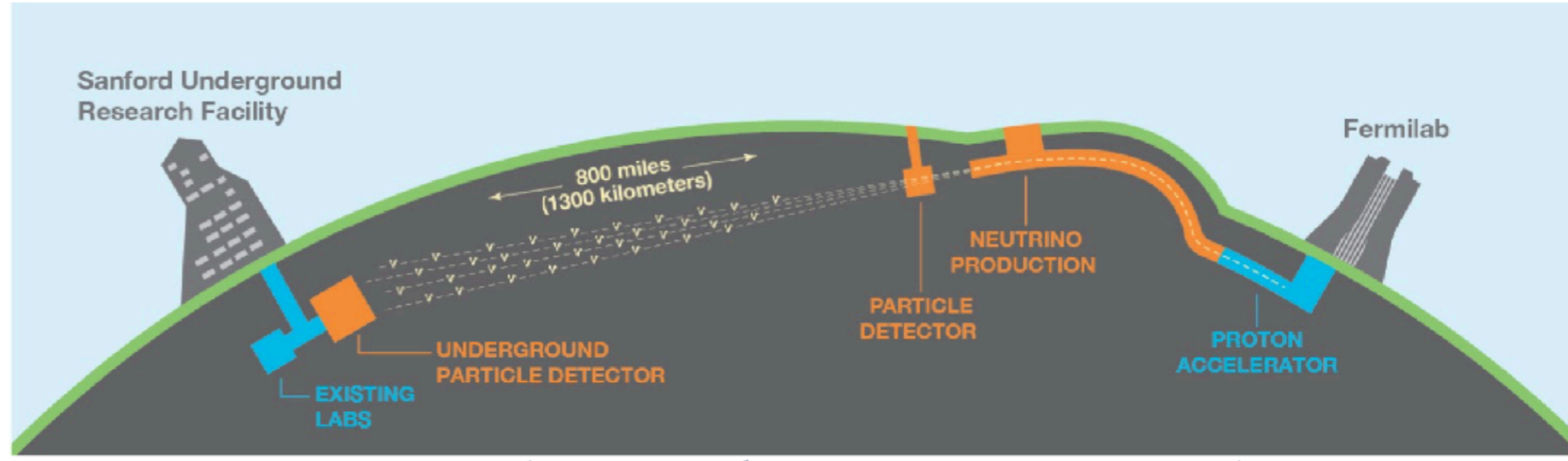


CP-conserving points are *outside*  $3\sigma$  intervals in IO  
Expect CPV if ordering is inverted

- Need definitive measurements! Two large next-generation projects are under preparation:

**DUNE:**

- > 2 MW beam
- Liquid-Argon TimeProjection Chamber (LArTPC) technology
- $\geq 40$  kton far detector fiducial mass
- First physics in ~2029

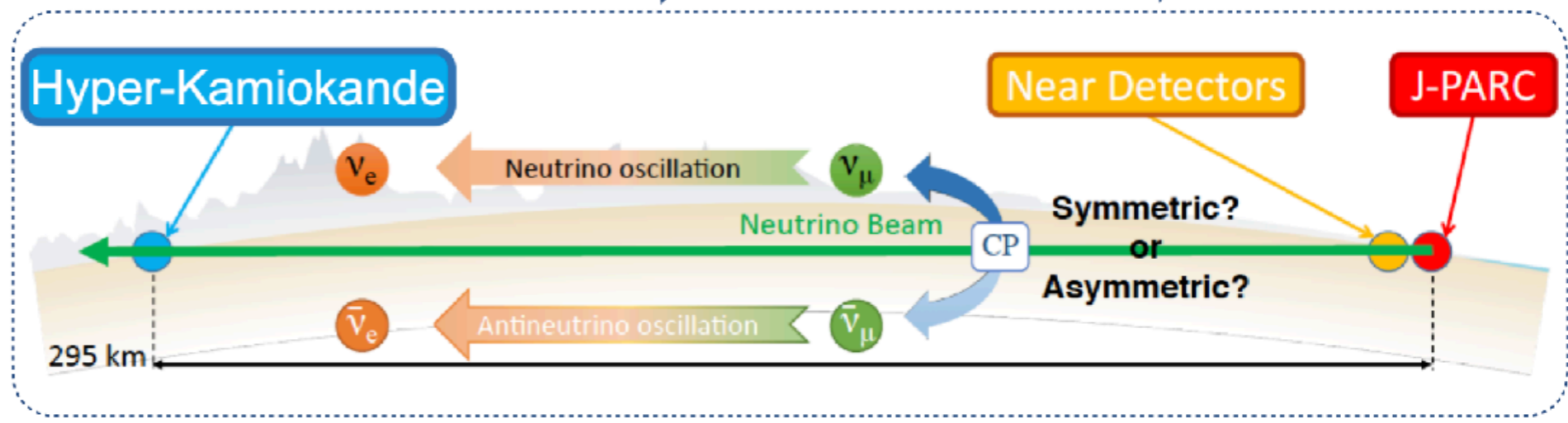


Large degree of complementarity:

- large matter effects (small)
- wide band, higher energy energy spectrum (narrow band, lower energy)
- LArTPC detection systematics (Water Cherenkov)

**Hyper-Kamiokande:**

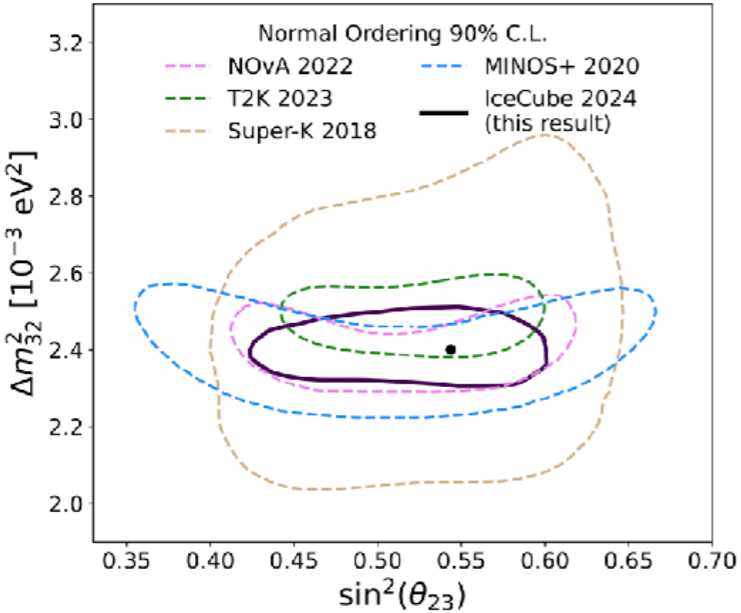
- 1.3 MW beam
- Water Cherenkov far detector
- 190 kton far detector fiducial mass
- First physics in ~2027



From T. Nakadaira's talk at ICHEP 2024

## ICECUBE and Deep Core

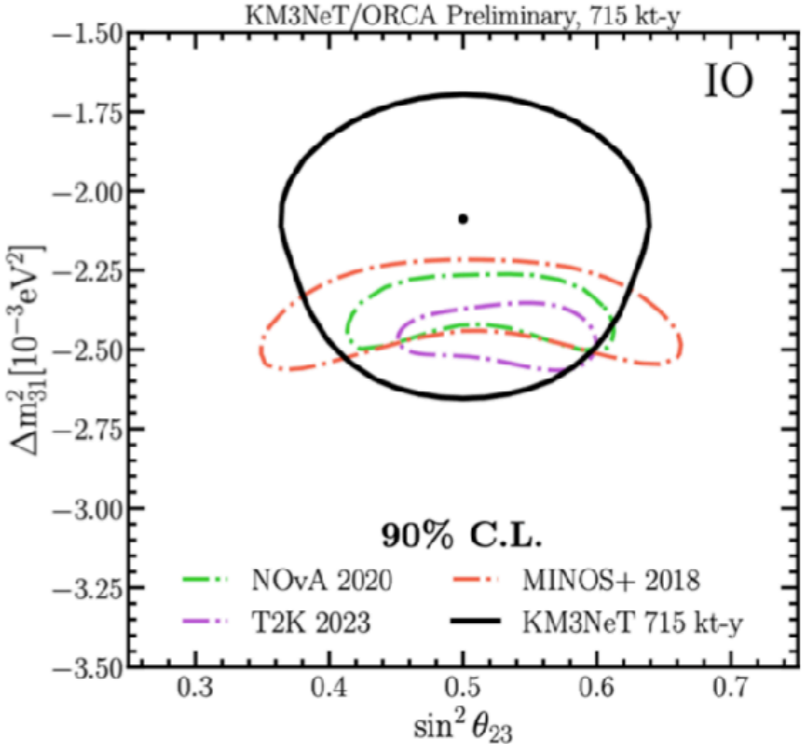
- 1km<sup>3</sup> of ice instrumented with strings of Digital Optical Modules (DOMs), each with a PMT
- DeepCore: 8 densely region at the center (threshold ~8 GeV)



- Comparable and compatible results between accelerator and atmospheric neutrinos.

## Km3net and ORCA

- 115 strings optimized for neutrino oscillation measurements
- Each DOM has 31 3-inch PMTs
- About 20% of DOMs already installed



**“Atmospheric experiments will make leading contributions to the global neutrino oscillation landscape before 2030.**

**In particular, atmospheric experiments will improve our knowledge of the  $\theta_{23}$  mixing angle and should be able to make a definitive measurement of the neutrino mass ordering by the end of the decade.**

**Together with JUNO, they will also improve our knowledge of the  $\Delta m_{32}^2$  mass splitting.”**

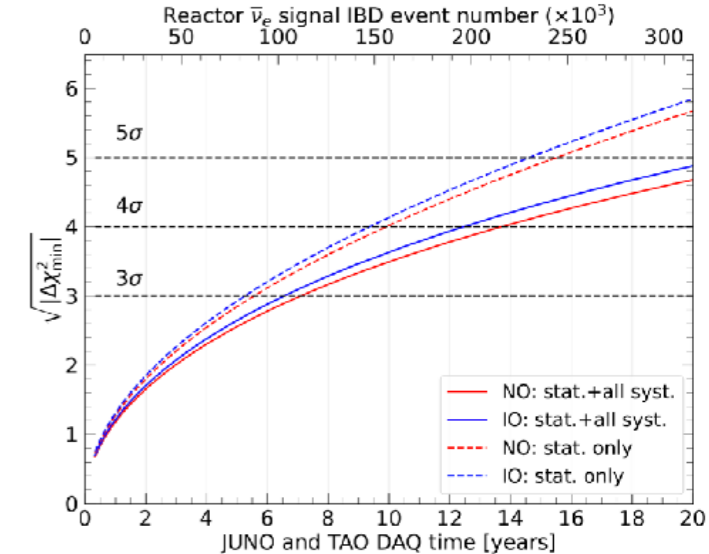
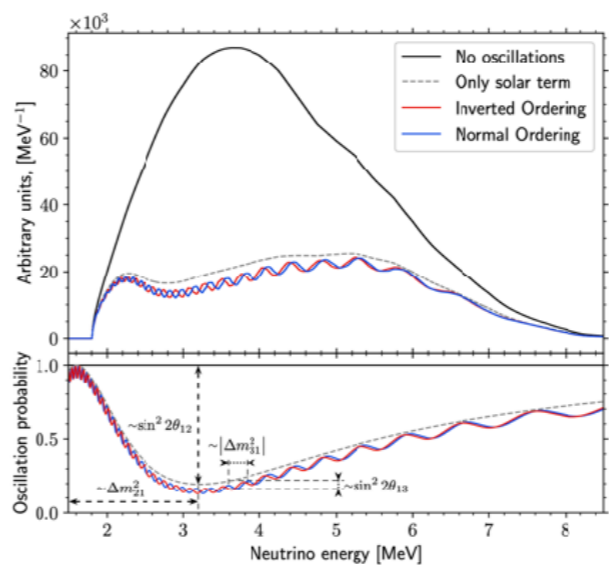


## JUNO Medium baseline experiment

20-kton Liquid Scintillator neutrino observatory located in Southern China

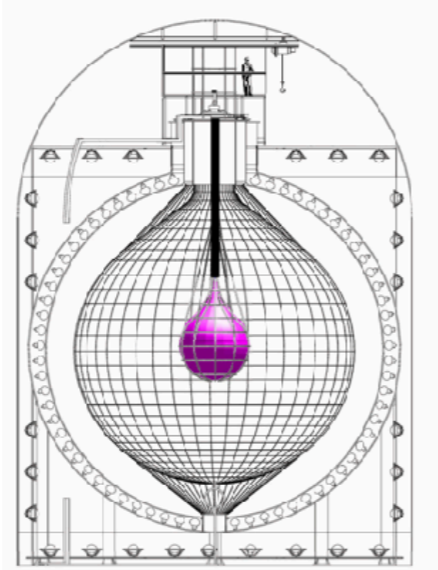
Mass hierarchy from the electron anti-neutrino disappearance pattern through the interference effect of quasi-vacuum oscillation of reactor antineutrinos

JUNO reactor neutrino oscillation analysis alone provides a median  $3\sigma$  sensitivity to NMO in 6.5 years!

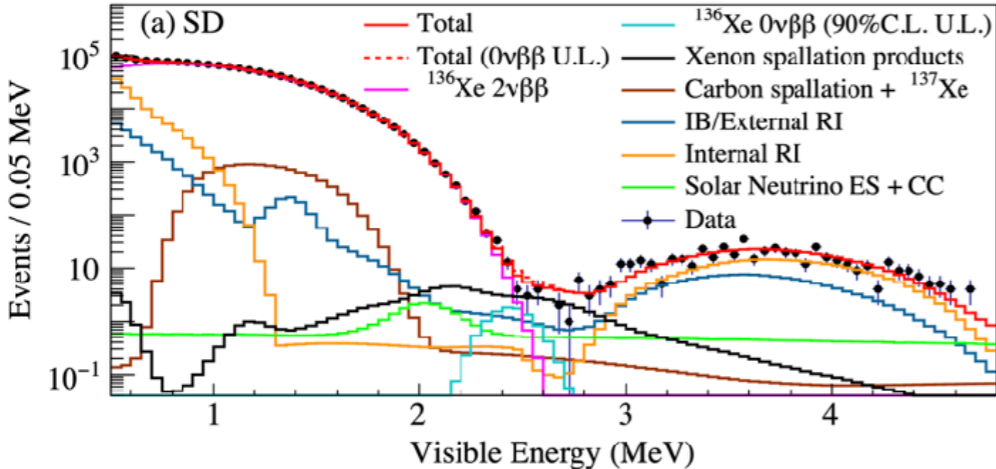


## KamLandZen

Mini-balloon Radius = 1.90 m  
 Xenon mass = 750 kg  
 Data taking started in 2019



The largest number of  $\beta\beta$  nuclei. Low BG by distillation and filtration of both Liquid Scintillator and Xenon.



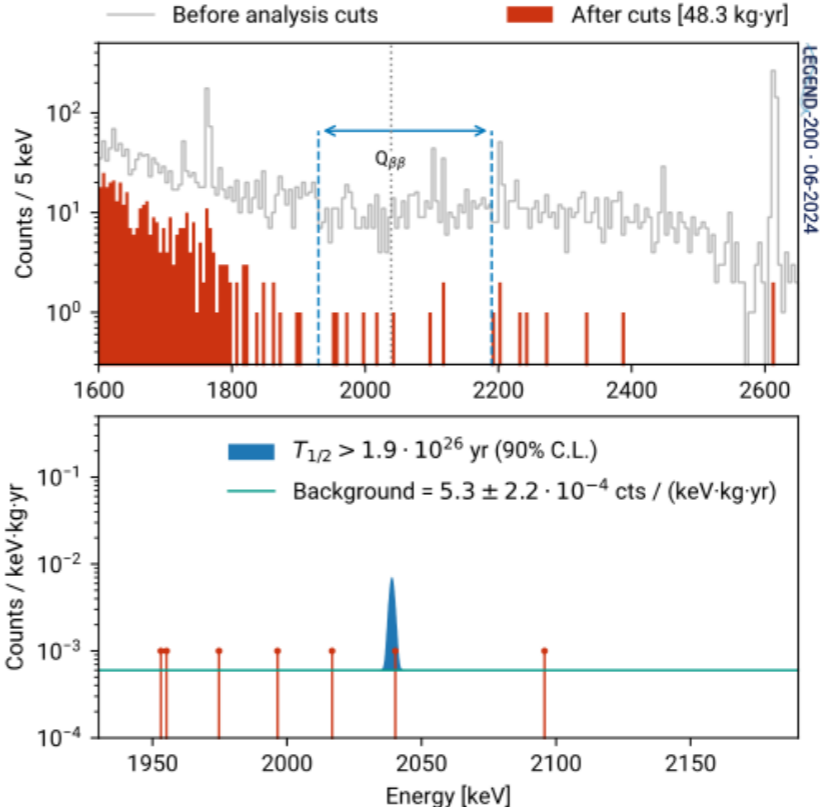
Limiting factor is the muon spallation of Xenon background.

$$\langle m_{\beta\beta} \rangle < 28 - 122 \text{ meV}$$

$$\text{KamLAND-ZEN-1T: } \langle m_{\beta\beta} \rangle < \sim 20 \text{ meV}$$

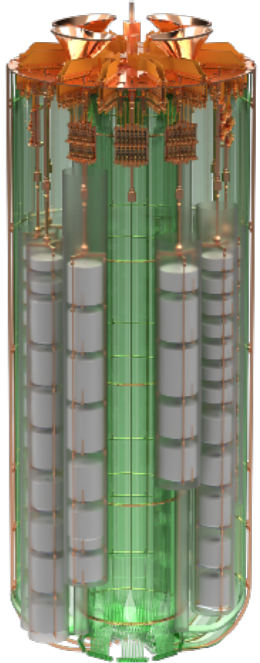
## Legend-200

Germanium Semiconductor, with enrichment to > 90% in  $^{76}\text{Ge}$  ( $Q_{\beta\beta}=2039 \text{ keV}$  Excellent energy resolution (0.1 % FWHM @  $Q_{\beta\beta}$ )



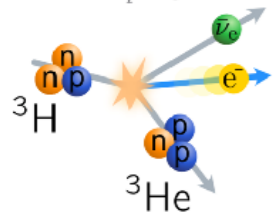
$$\langle m_{\beta\beta} \rangle < 75 - 178 \text{ meV}$$

$$\text{Legend 1000 } \langle m_{\beta\beta} \rangle < \sim 20 \text{ meV}$$





## KATRIN Experiment



### High-activity tritium source

- 30 μg of gaseous T<sub>2</sub>
- 10<sup>11</sup> T<sub>2</sub> decays/s

- Best fit:

$$m_\nu^2 = (-0.14^{+0.13}_{-0.15}) \text{ eV}^2 \text{ (stat. dom.)}$$

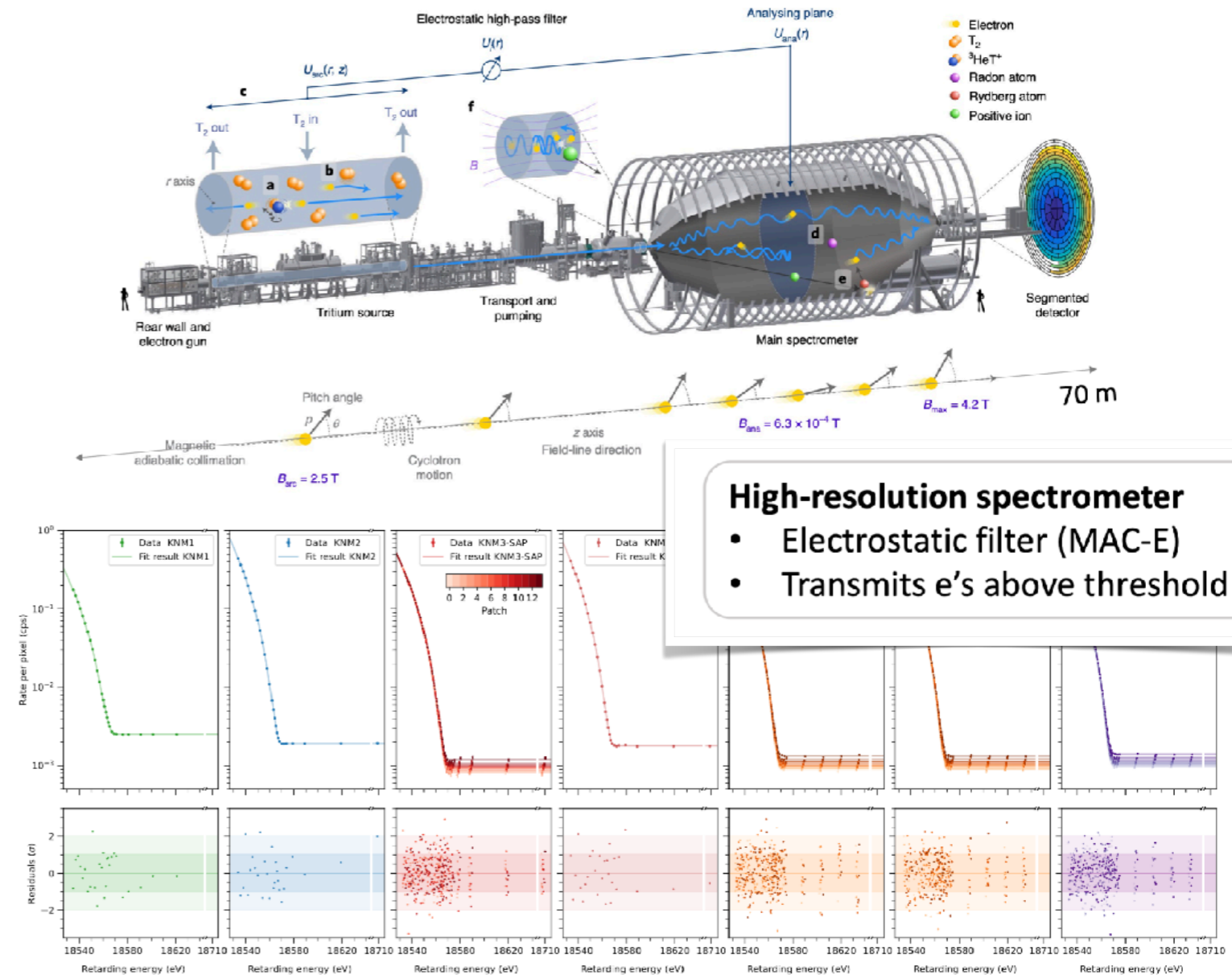
- New limit:

$$m_\nu < 0.45 \text{ eV (90\% CL)}$$

Neutrino-24 (2024), arXiv:2406.13516 (2024)

### Final goal (in 2026):

- < 0.3 eV sensitivity

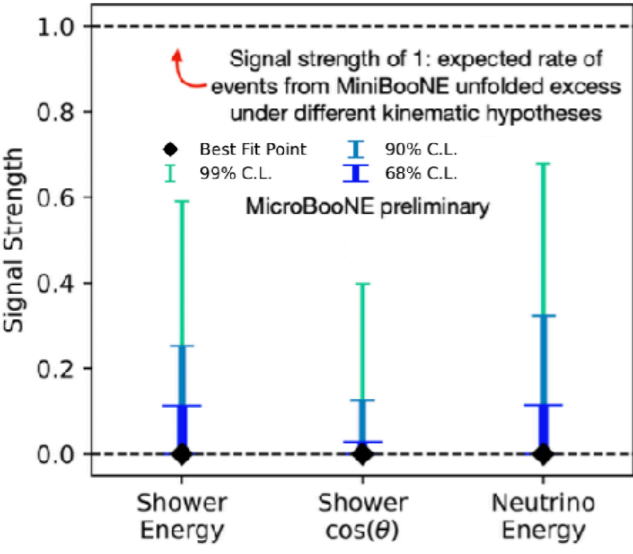
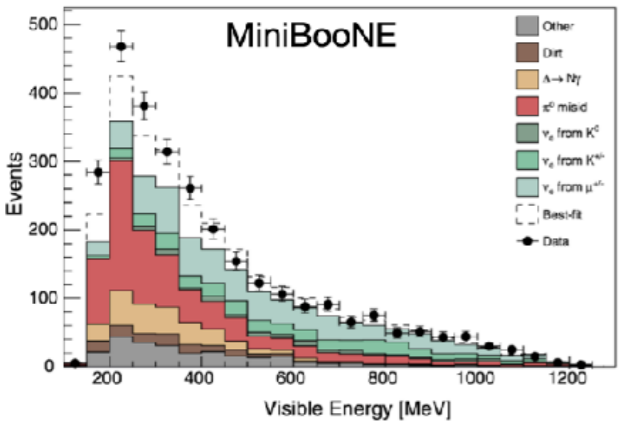


### High-resolution spectrometer

- Electrostatic filter (MAC-E)
- Transmits e's above threshold

## LSND/MiniBooNE

LSND/MiniBooNE observed  $6\sigma$  excess of electron (anti)neutrinos in muon (anti)neutrino beam!

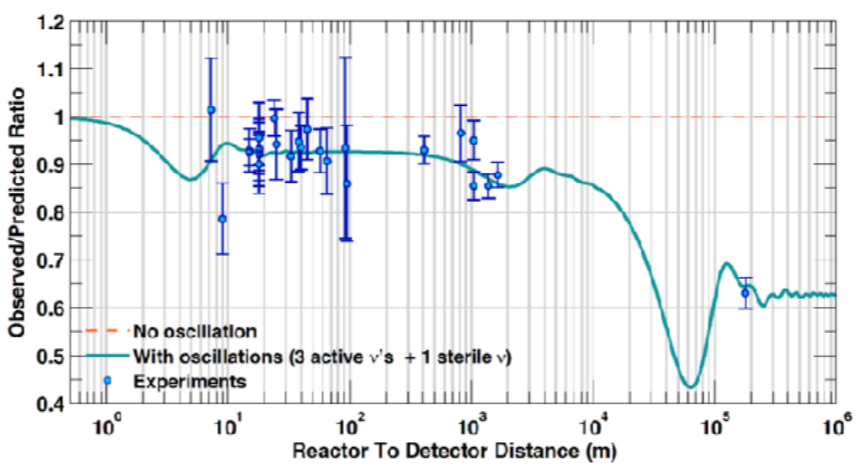


Excess not seen by MicroBooNE experiment

**SBN program** at Fermilab to completely settle the question!

## Reactor anti-neutrino deficit

6% deficit of reactor anti-neutrinos  $\sim 3\sigma$  with respect to flux prediction models.

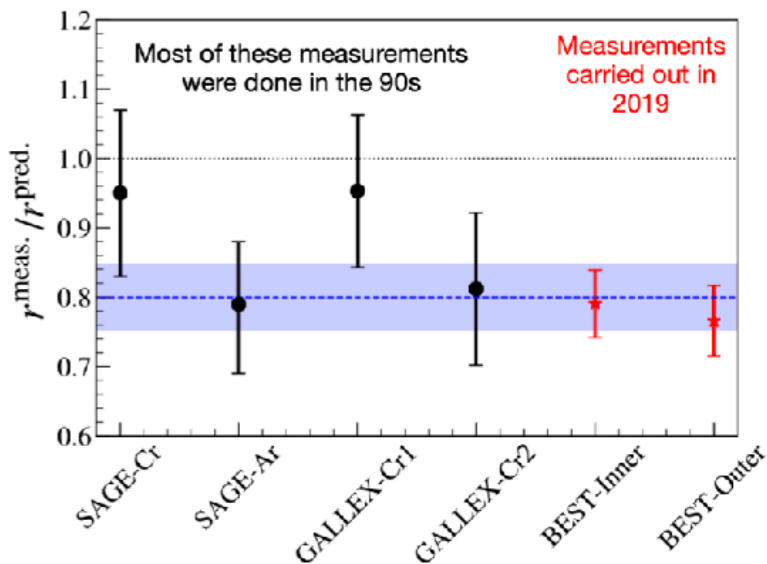


Recent fission data suggests that the flux were overestimated... **by about the right amount!**

Also essentially excluded by short baseline experiments searching for sterile neutrinos: DANSS, NEOS, PROSPECT, SoLid, STEREO.

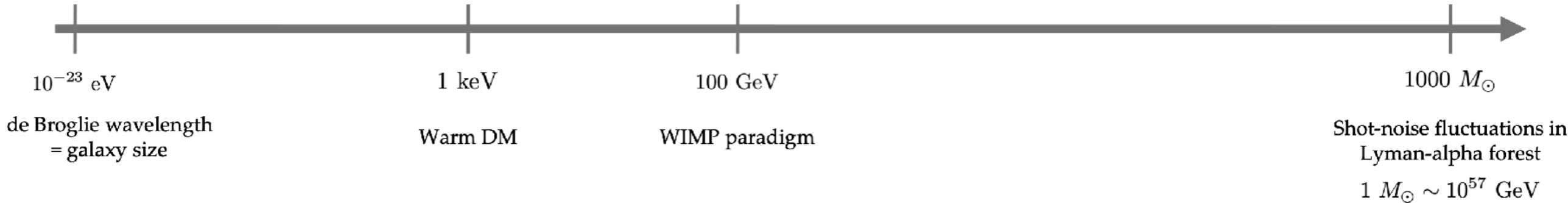
## Gallium anomaly

Seen by several experiments (including recent)



Also excluded by short baseline experiments.

The Range of Possibilities is **Stunning!**



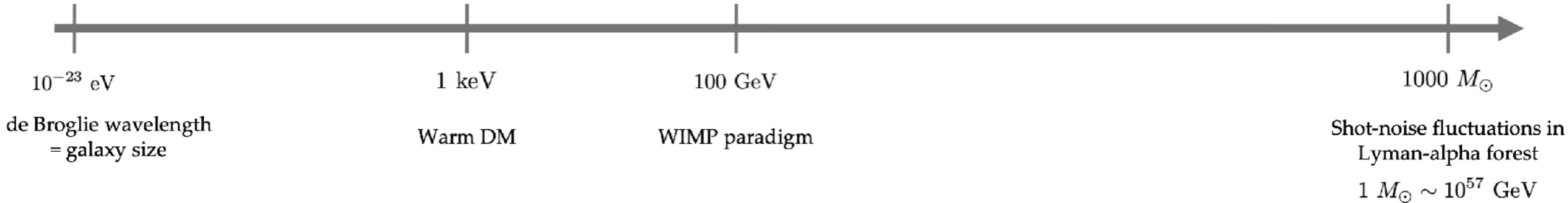
### A clear roadmap!

- Challenge #1: Fully Cover Electroweakino DM, e.g. with Cherenkov Telescopes
- Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection
- Challenge #3: Build out the suite of axion searches
- Challenge #4: Build out the suite of accelerator searches—**high energy and intensity**—for hidden sectors
- Challenge #5: Cover the abundance-driven light DM models in direct detection
- Challenge #6: Observe the Dark Matter Power on Small Scales

...

**Jocelyn and Stefania have covered (#2 - #5) !**

The Range of Possibilities is **Stunning!**



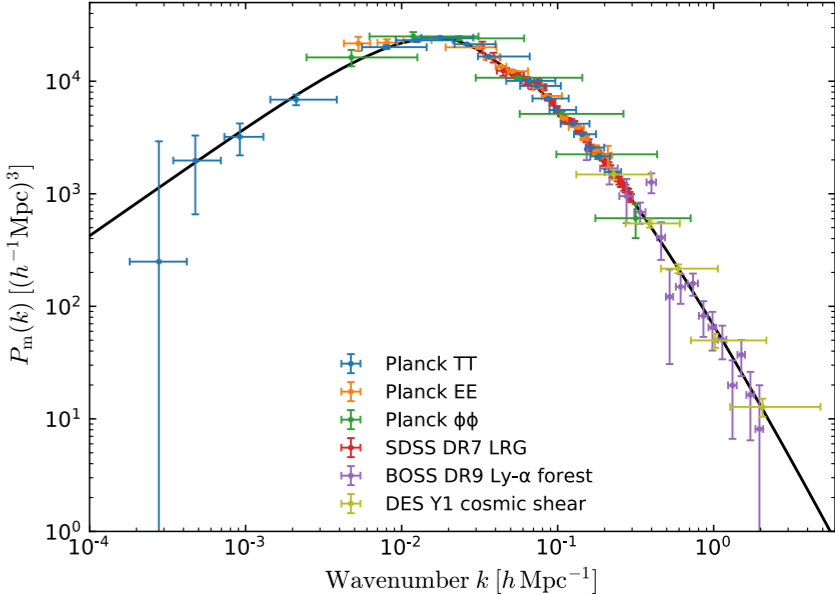
Challenge #1: Fully Cover Electroweakino DM, e.g. with Cherenkov Telescopes

- LHC is very sensitive to strongly produced SUSY particles, less so for pure electroweakinos
- Cherenkov telescopes can cover to larger masses!

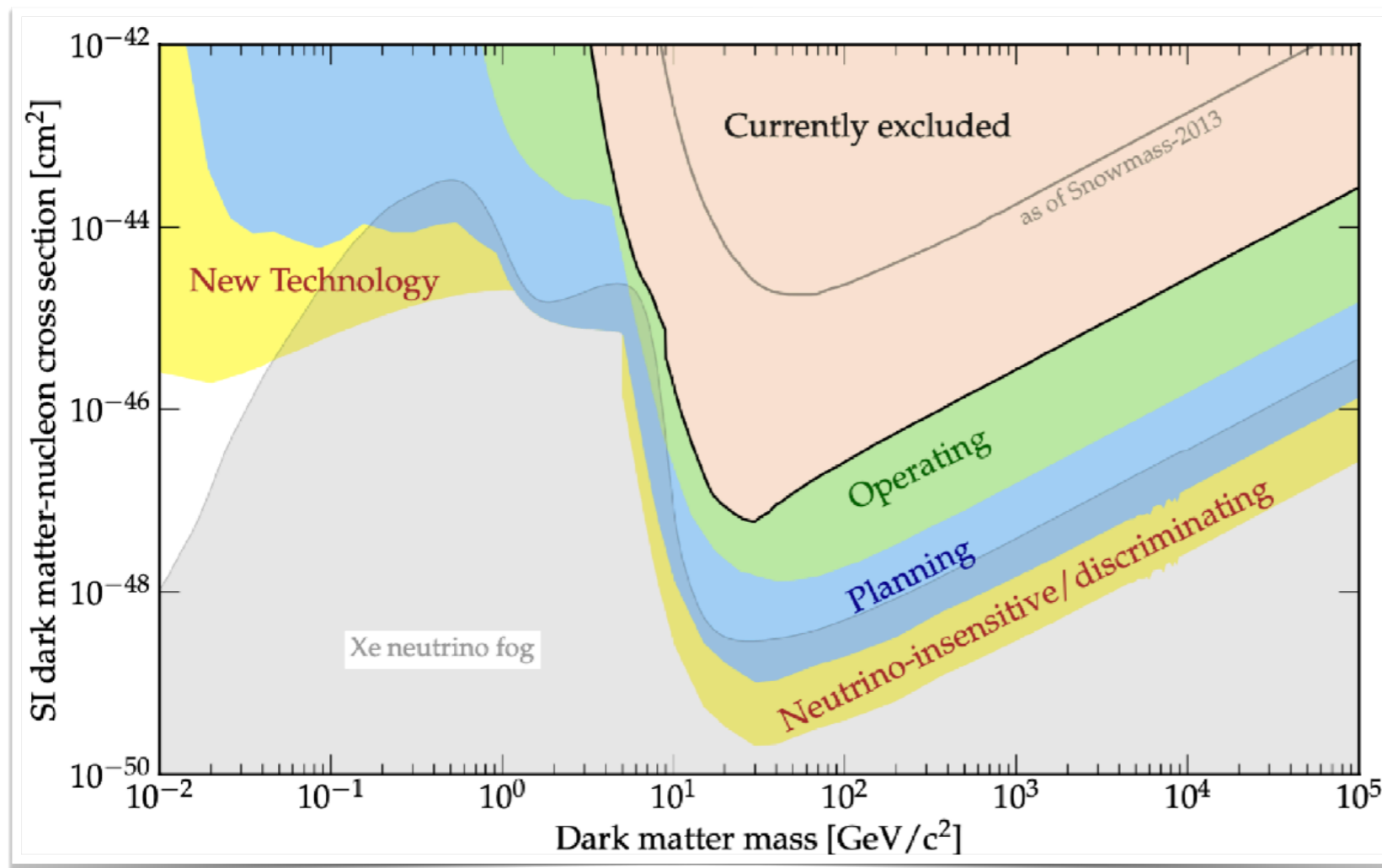
Challenge #6: Observe the Dark Matter Power on Small Scales

Theories as **PBH, axions**, predict different behavior than  $\Lambda_{\text{CDM}}$  at sub-halo scales, requiring new measurements of small scale!

(e.g. measuring the changes in the metric due to transiting DM substructure using Pulsar Timing Arrays)



## Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection



### - Currently operating detectors:

- Xenon nT, PandaX-4T, LZ-7T
- Argon DarkSide

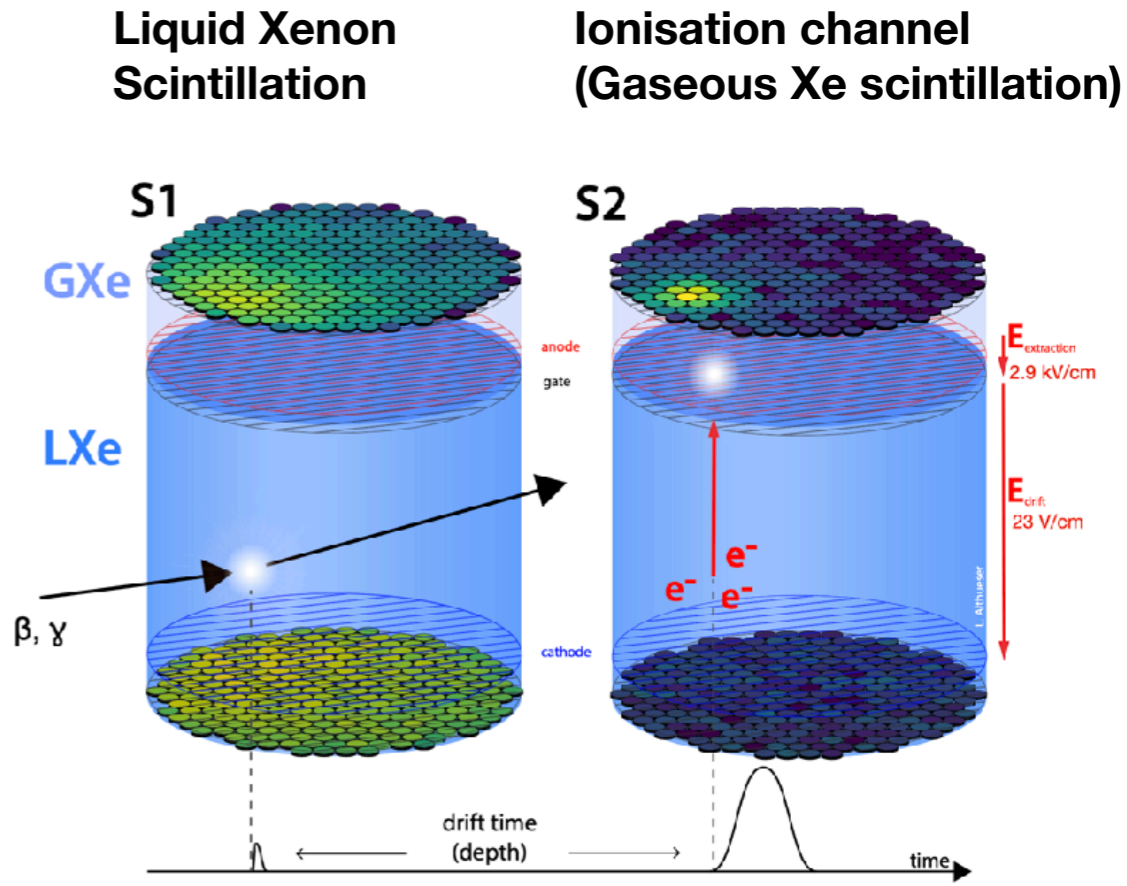
### - Next generation:

- PandaX-20T
- Darwin XLZD 40-60T

## Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection

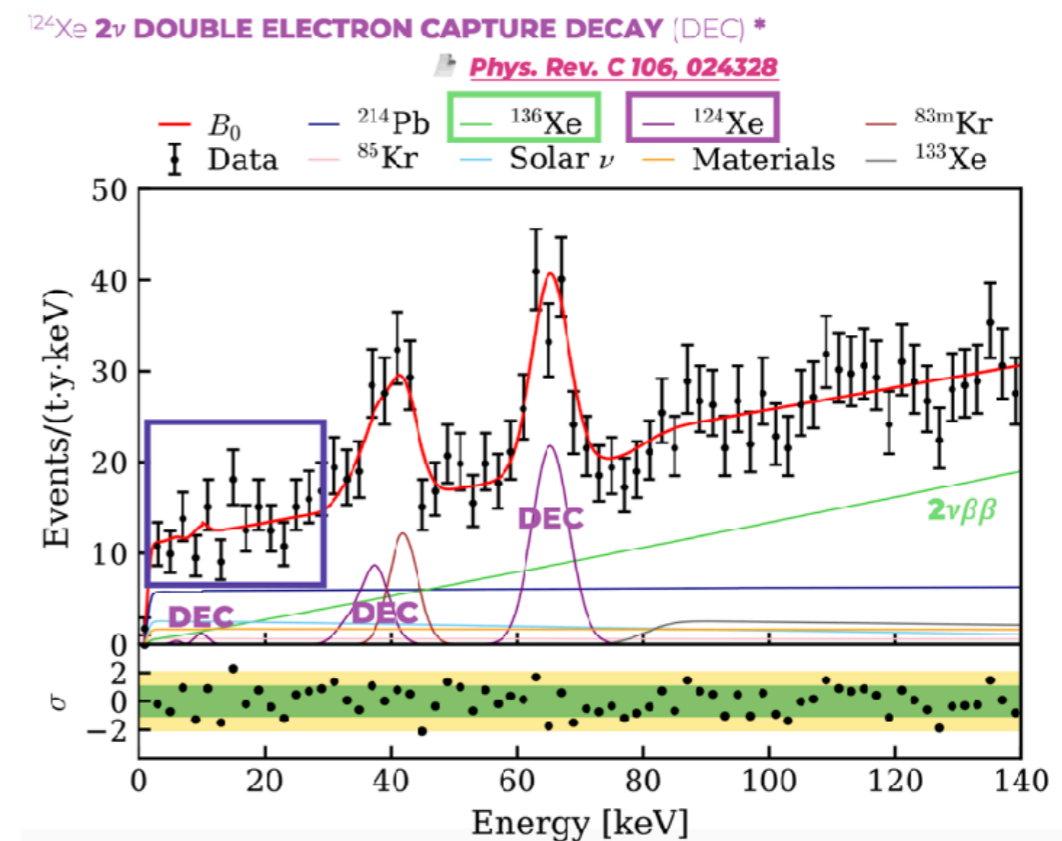
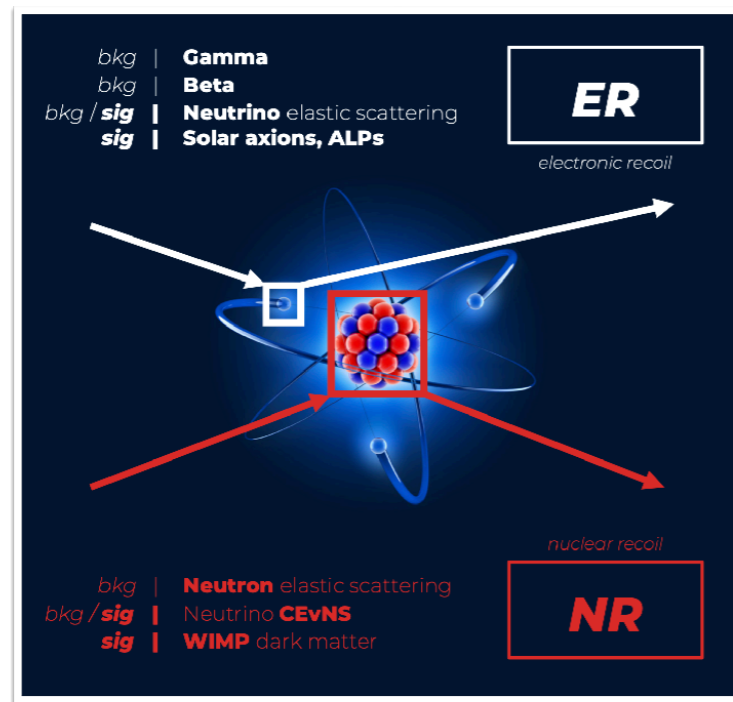


### Xenon nT

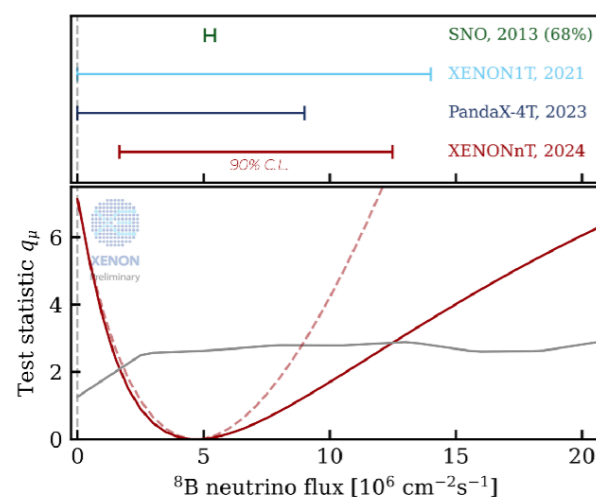


## Challenge #2: Search for WIMPs to the Neutrino Background in Direct Detection

## Electron Recoil in Xenon



### Nuclear Recoils from $^8\text{B}$ solar neutrinos CEvNS in Xenon!



**2.73  $\sigma$**   
SIGNIFICANCE

- Observation of DEC: Two of the orbital electrons are captured protons in the nucleus emission of 2 neutrinos
- Solar neutrinos background subdominant (backup PandaX-4T)

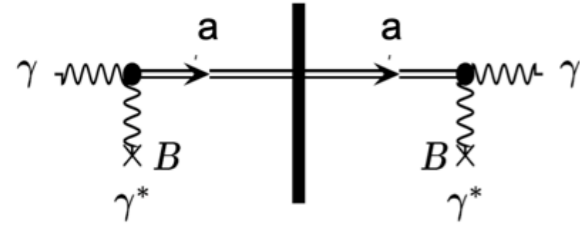
**37 Events observed** with 26 bkg and 11 signal expected!

Great achievement, but now need to learn how to fight it!

## Challenge #3: Build out the suite of axion searches

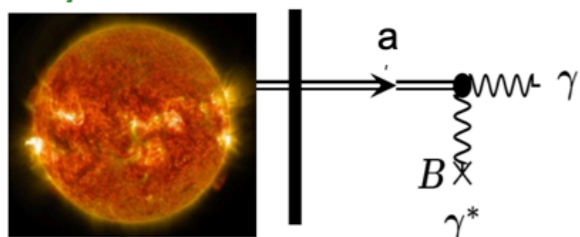
### Light shining through a wall (LSW)

[Anselm 85; van Bibber 87]



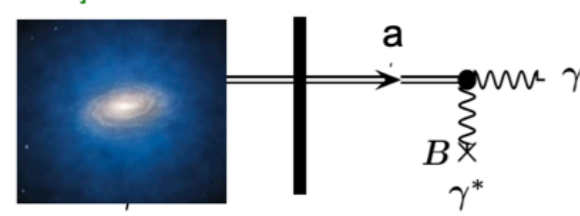
### Helioscope: Sun shining through a wall

[Sikivie 83]

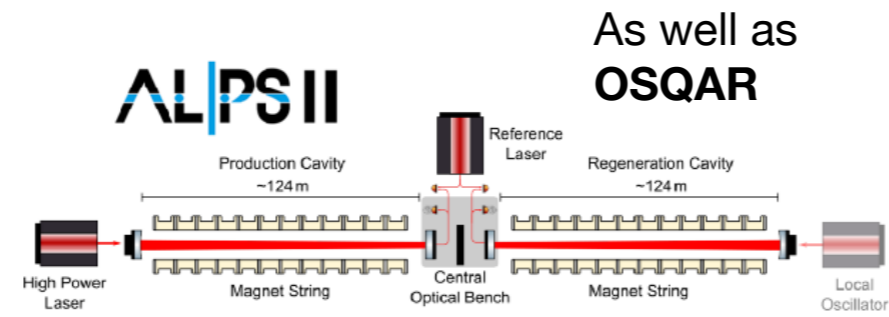


### Haloscope: DM shining through a wall

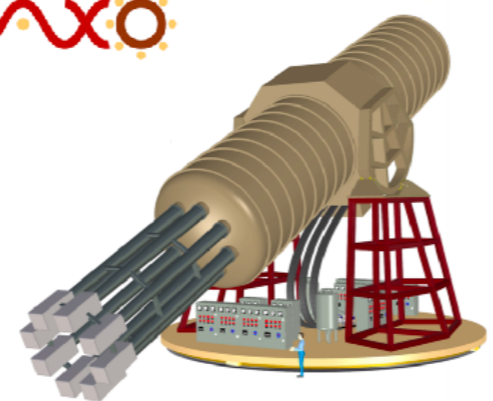
[Sikivie 83]



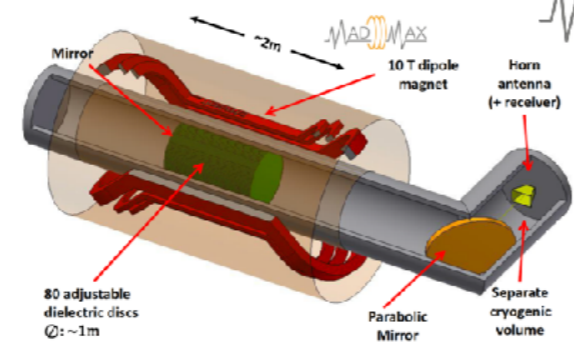
[Lindner]



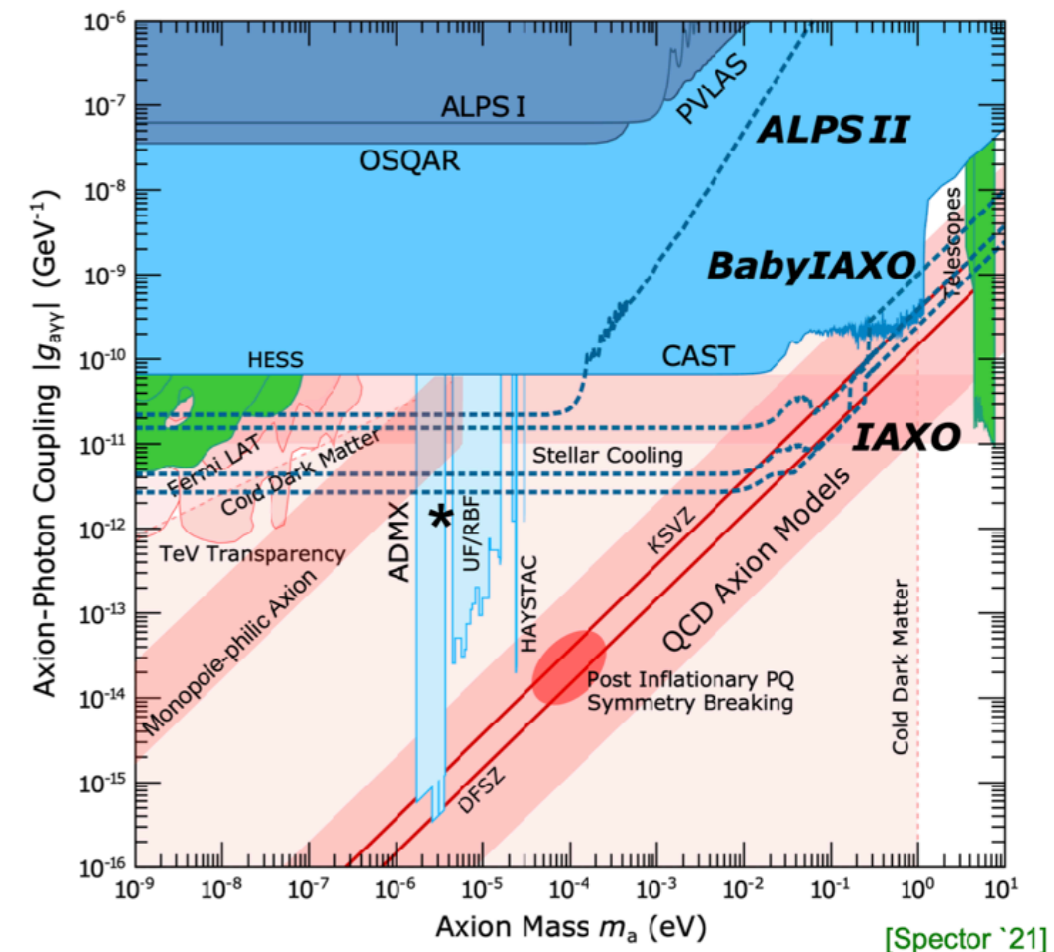
As well as OSQAR



As well as SUMICO, and CAST



As well as ADMX, HAYSTAC, QUAX, CAST-CAPP, RADES



[Spector '21]



Challenge #4: Build out the suite of accelerator searches—**high energy and intensity**—for hidden sectors

There are numerous gaps in the experimental exploration of dark sectors, both at the **electroweak-TeV** scale and at the **sub-GeV** scale.

NP particles can have a width that is suppressed by small mass splitting, multi-body final states, high NP scale small coupling, etc.

**Sub-GeV** new physics motivates searches at high intensity facilities and Meson factories!

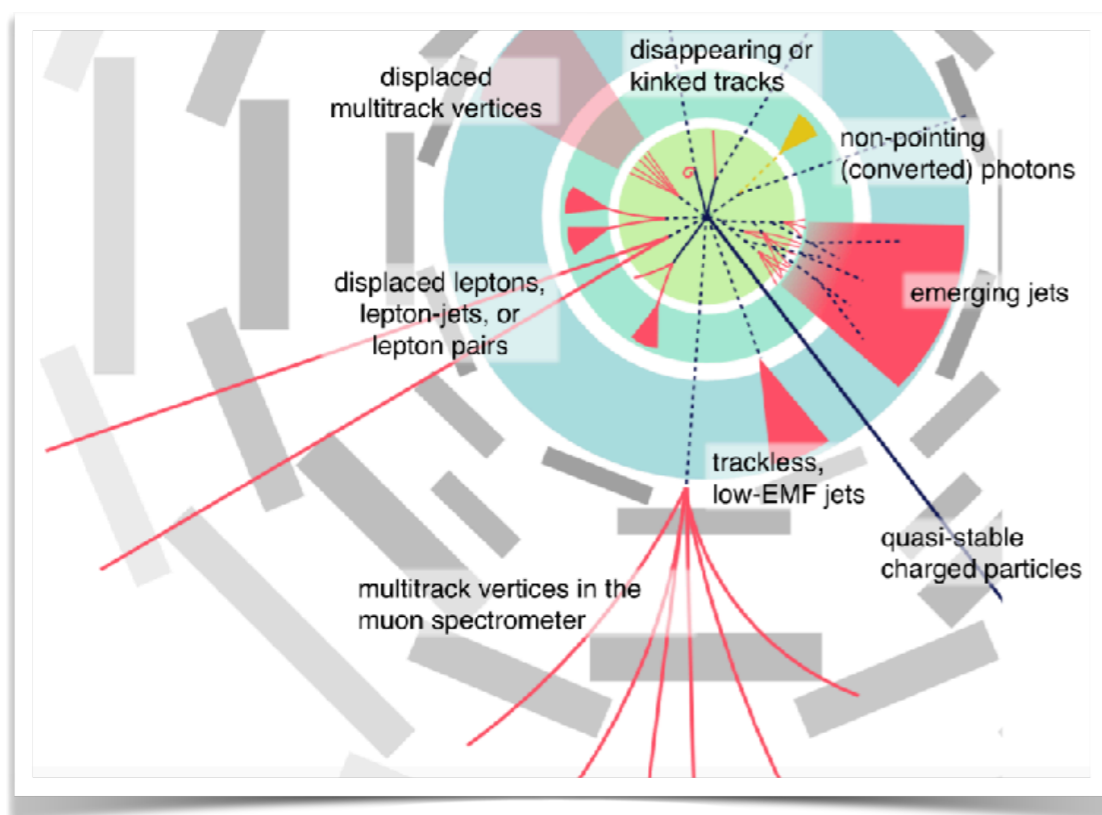
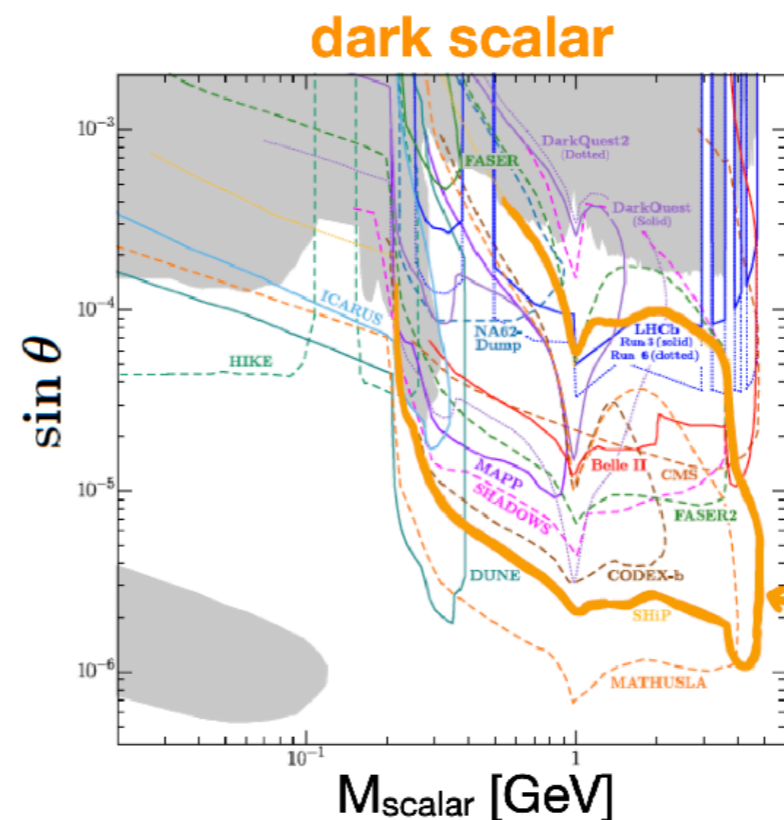


Image from H. Russel



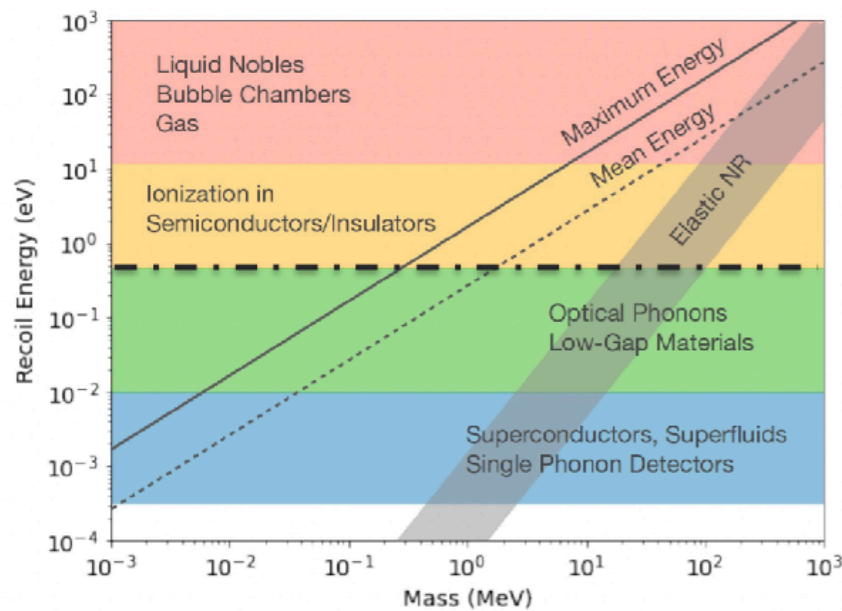
Complementarity with LHC auxiliary detectors + neutrino experiments

**SHiP**  
(scalar from meson decays)

Challenge #5: Cover the abundance-driven light DM models in direct detection

Hidden sector DM: huge range of techniques... **Models motivate from relic abundance** can be tested w/1 kg-yr exposures!

Quantum sensors can go to sub-eV recoil energies (a blossoming field!)

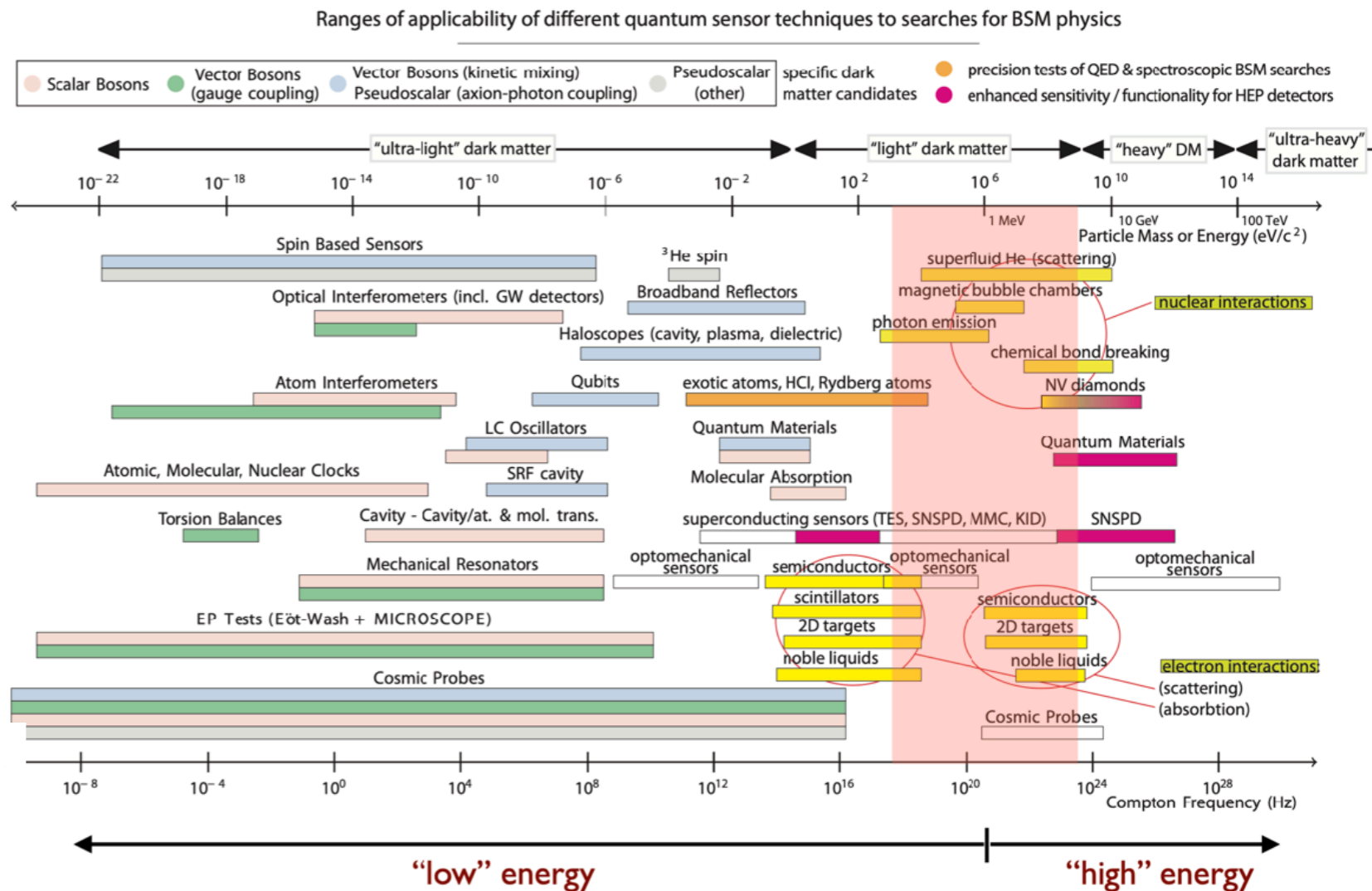


Daniel Baxter | IDM 2024    Essig et al, Snowmass CF1 WP2 (2022) [arXiv:2203.08297]

$\Delta E \sim 1$  eV  
e.g. Si, Ge, GaAs, diamond, Quantum Dots, organic scintillators...

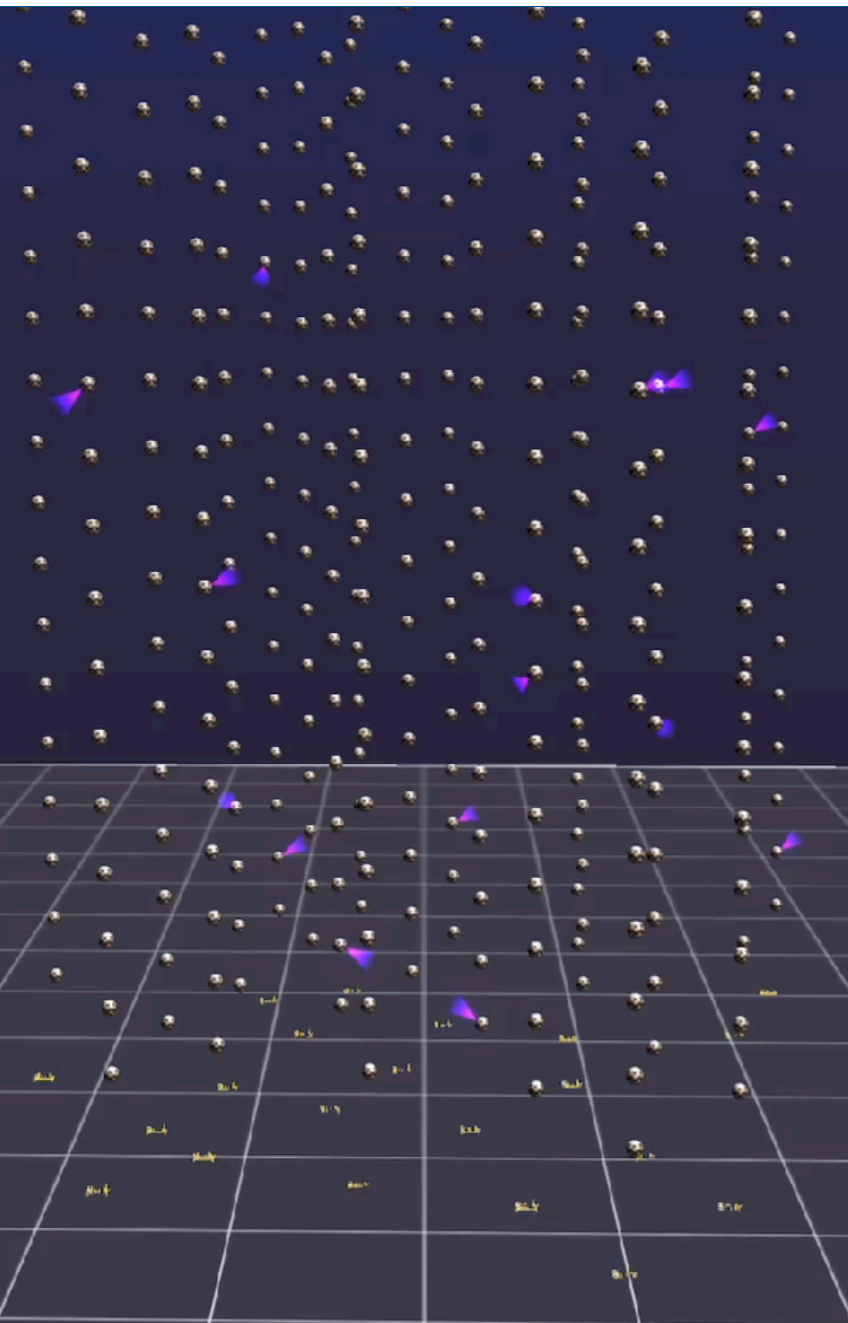
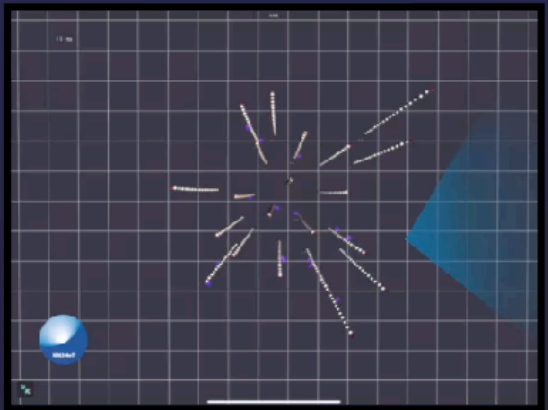
$\Delta E \sim 10 - 100$  meV  
e.g. GaAs, sapphire, Dirac materials, doped s/c, ...

$\Delta E \sim 1$  meV  
e.g. superfluids, superconductors



“low” energy

“high” energy

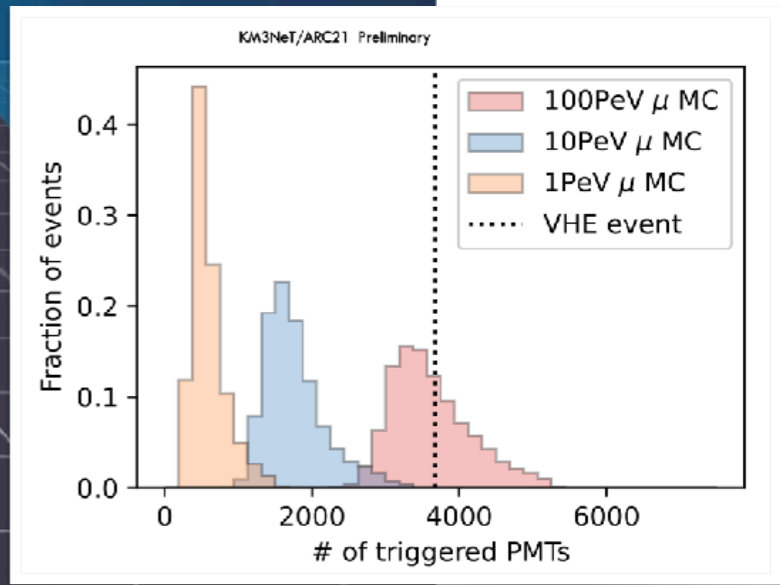


**NEWS IN FOCUS**

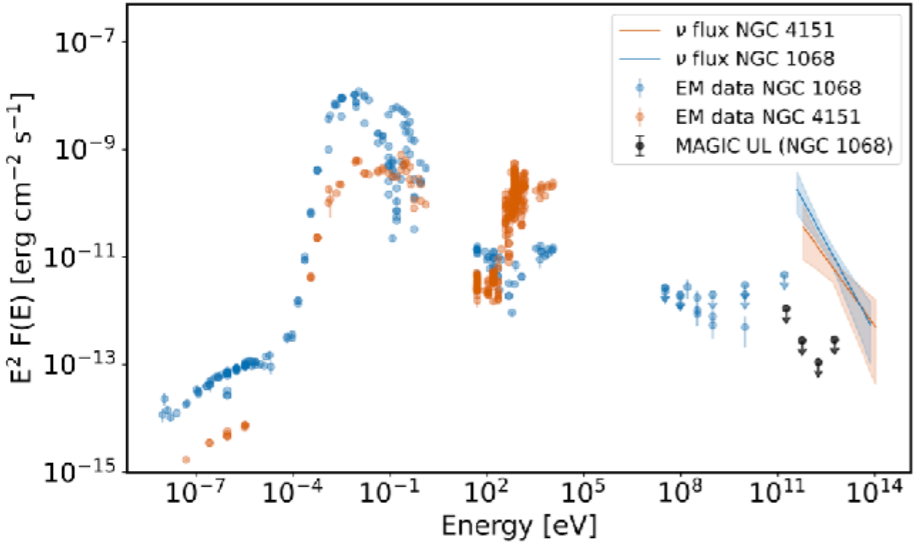
Two ARGONIE light detectors on board a ship, ready for deployment.

**'FANTASTIC' NEUTRINO  
COULD BE MOST  
ENERGETIC EVER FOUND**

Ultra-high-energy particle spotted by deep-sea detectors could point to a massive cosmic event

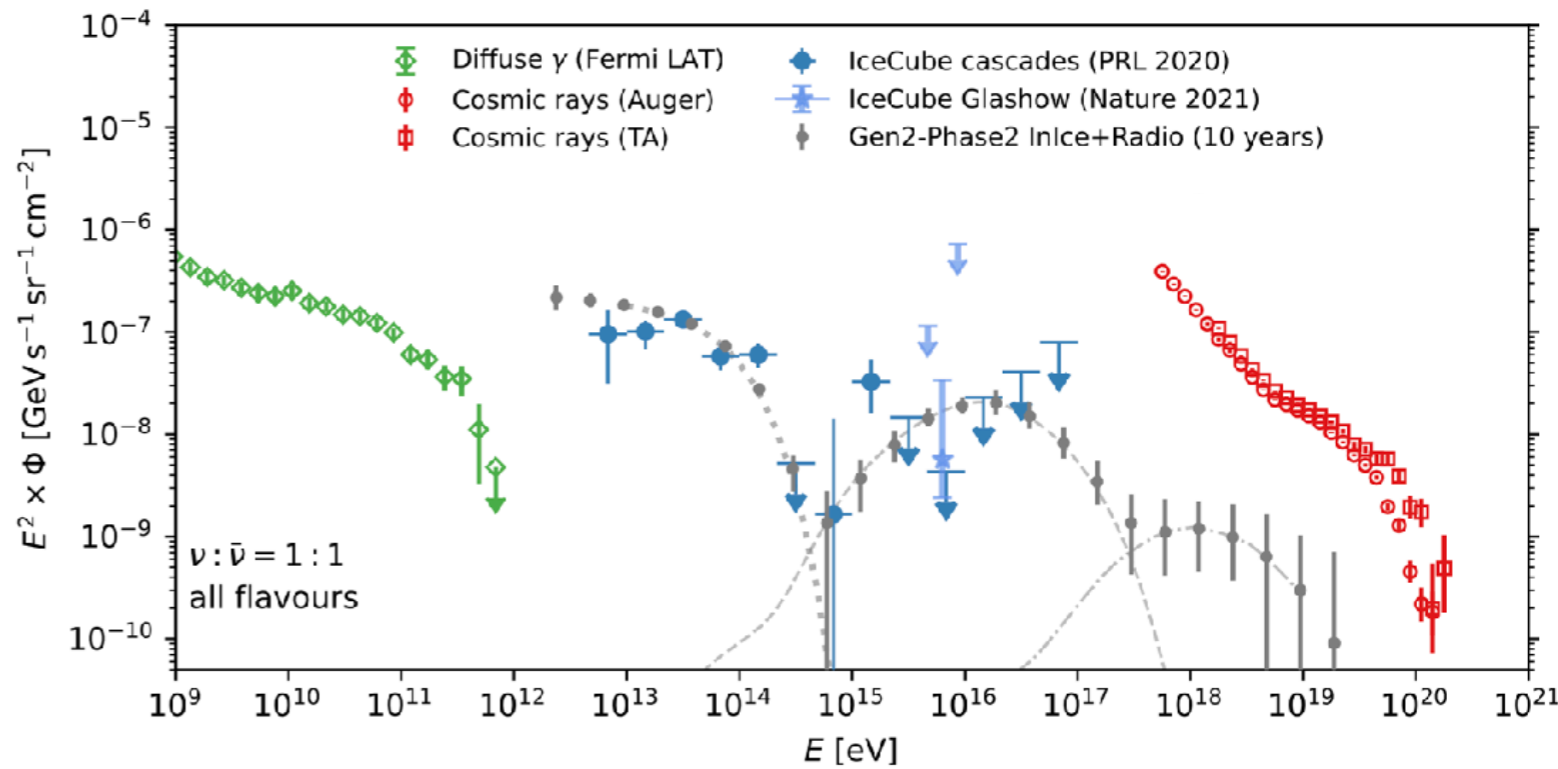


Observation by IceCube of NGC1068!



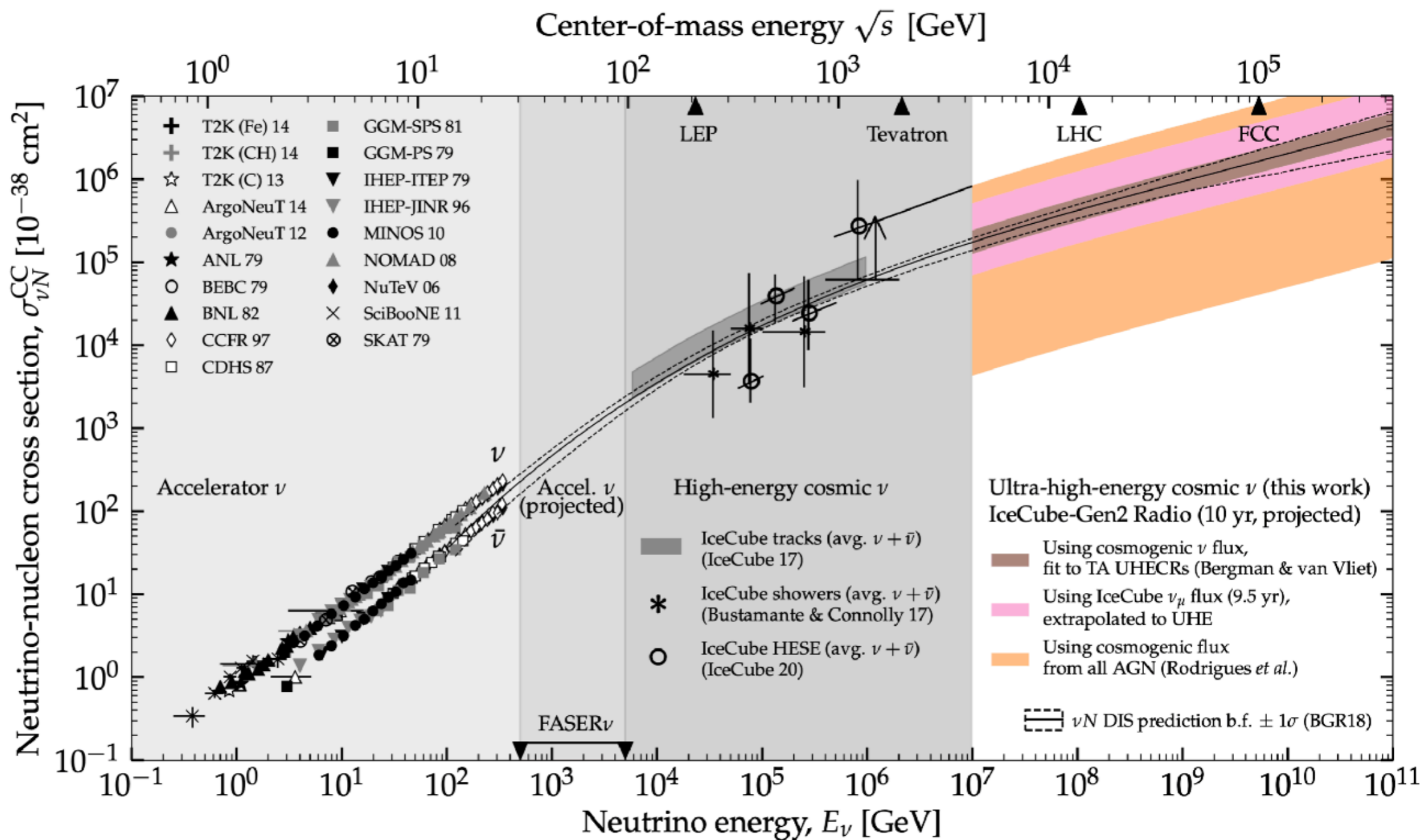
In 2022 ESO found a cloud of cosmic dust (obscuring/attenuating gamma rays) at the centre of NGC1068 hiding a supermassive black hole.

Evidence for diffuse astrophysical neutrinos: "10 PeV Neutrinos a gateway to 10^20 eV cosmic rays!"



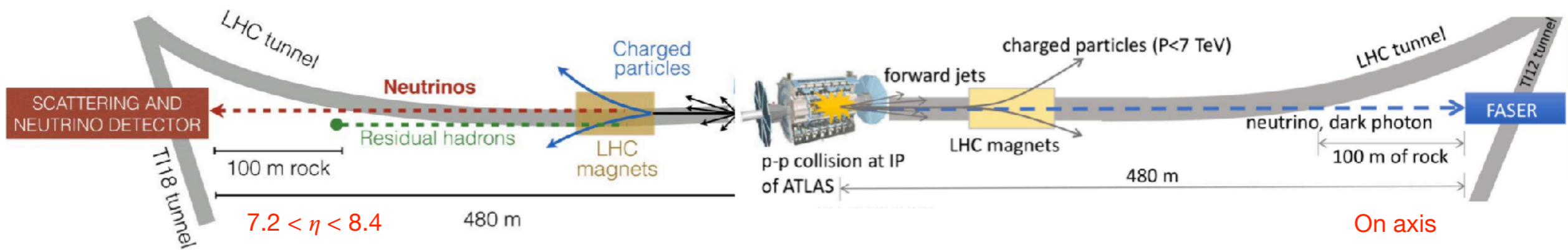
IceCube Gen2 in the race towards Ultra High Energy neutrinos with ARA, Trinity 18, GRAND, and POEMMA projects!

# The birth of Collider Neutrinos (at the LHC)



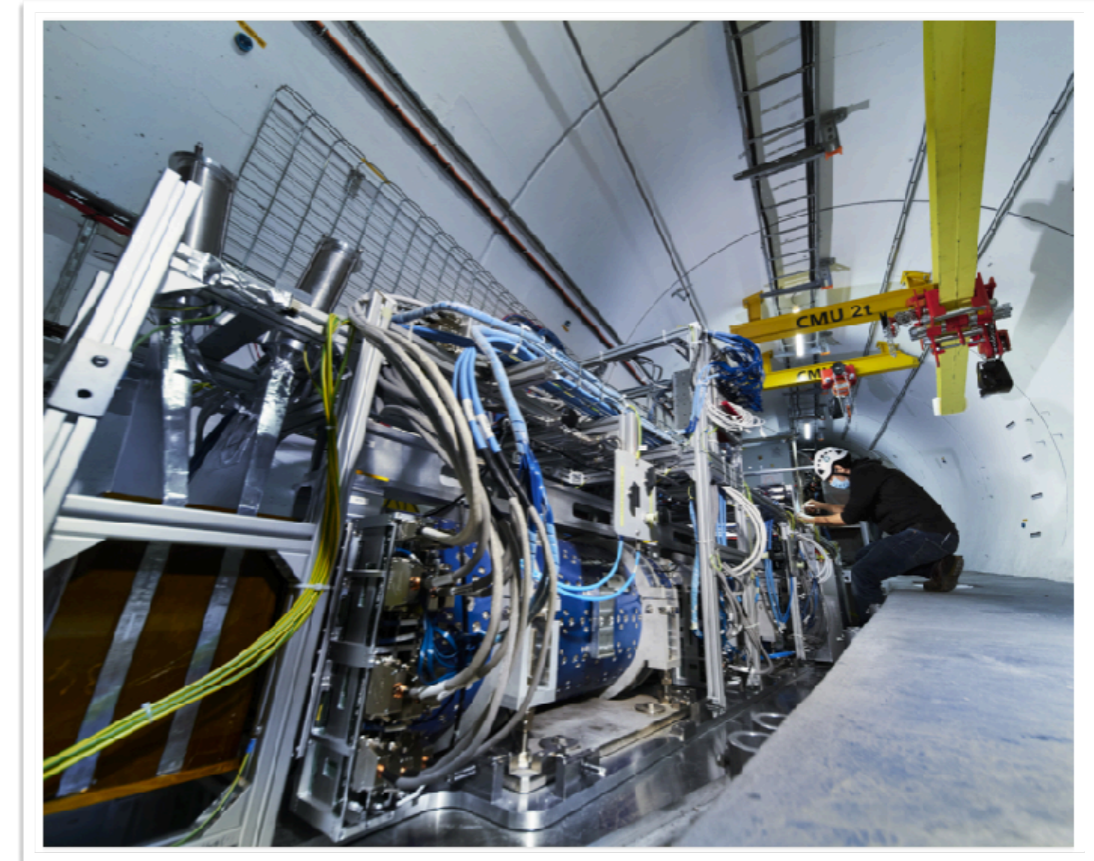
Accelerator neutrinos, up to O(100 GeV)

Cosmic neutrinos



**SND**

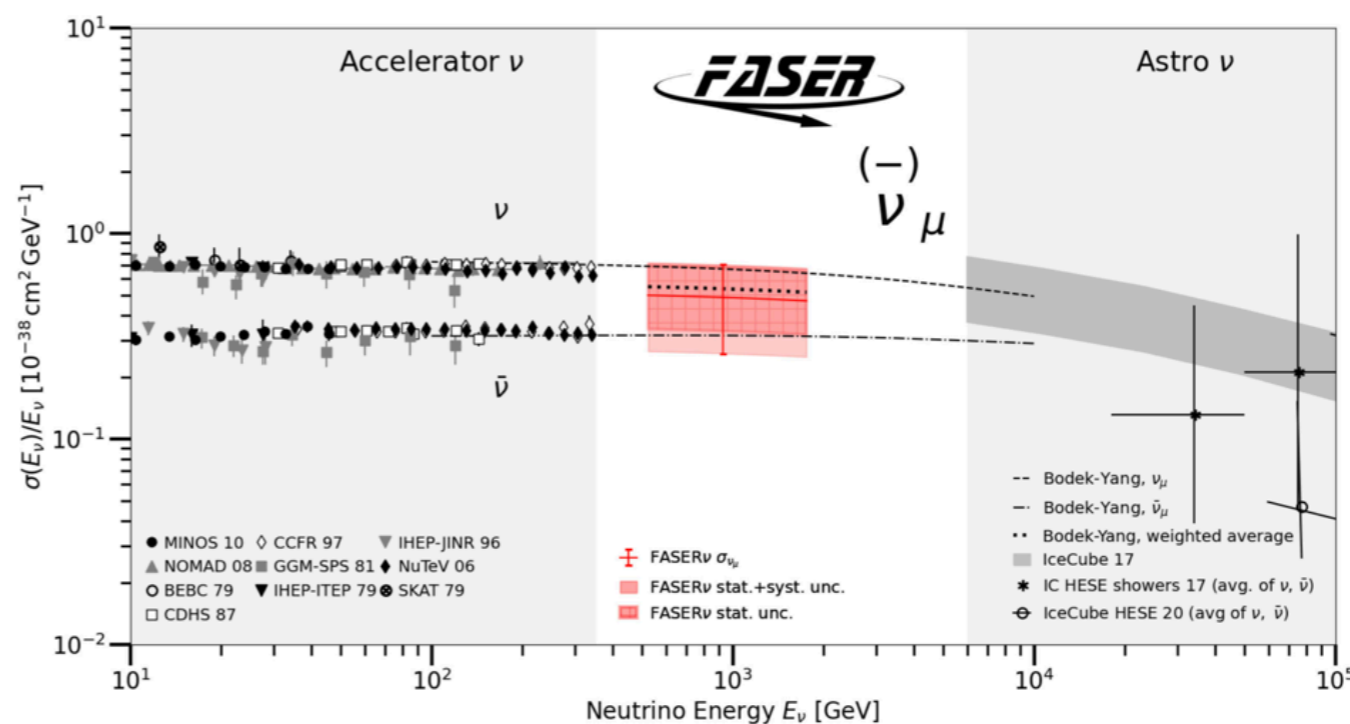
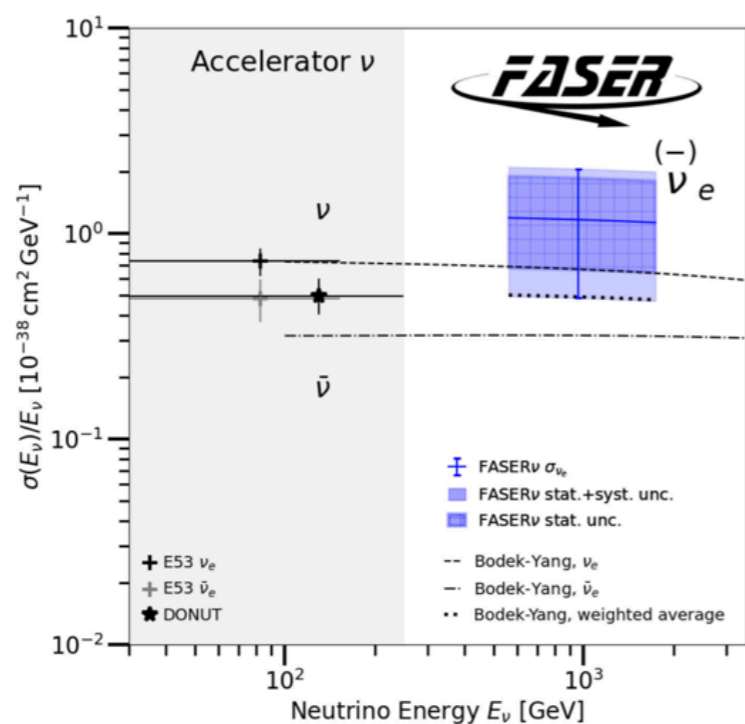
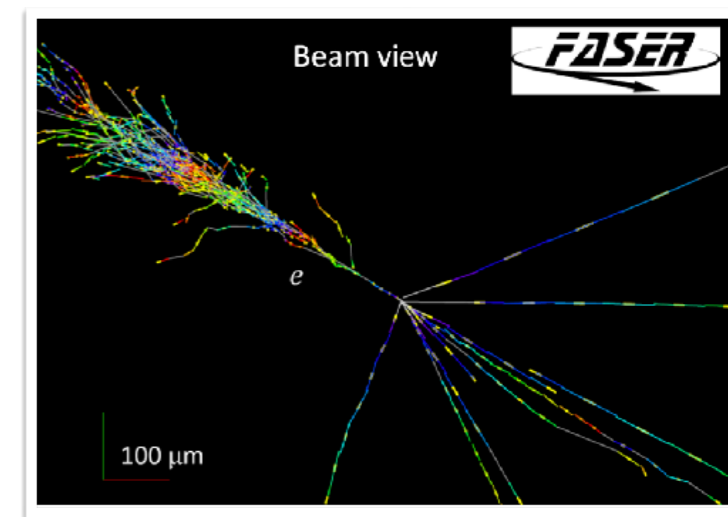
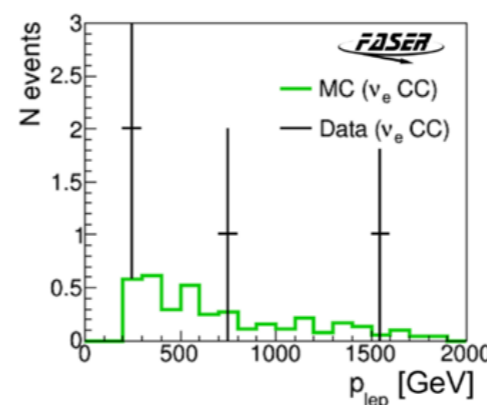
**Faser-v**



## Faser-v

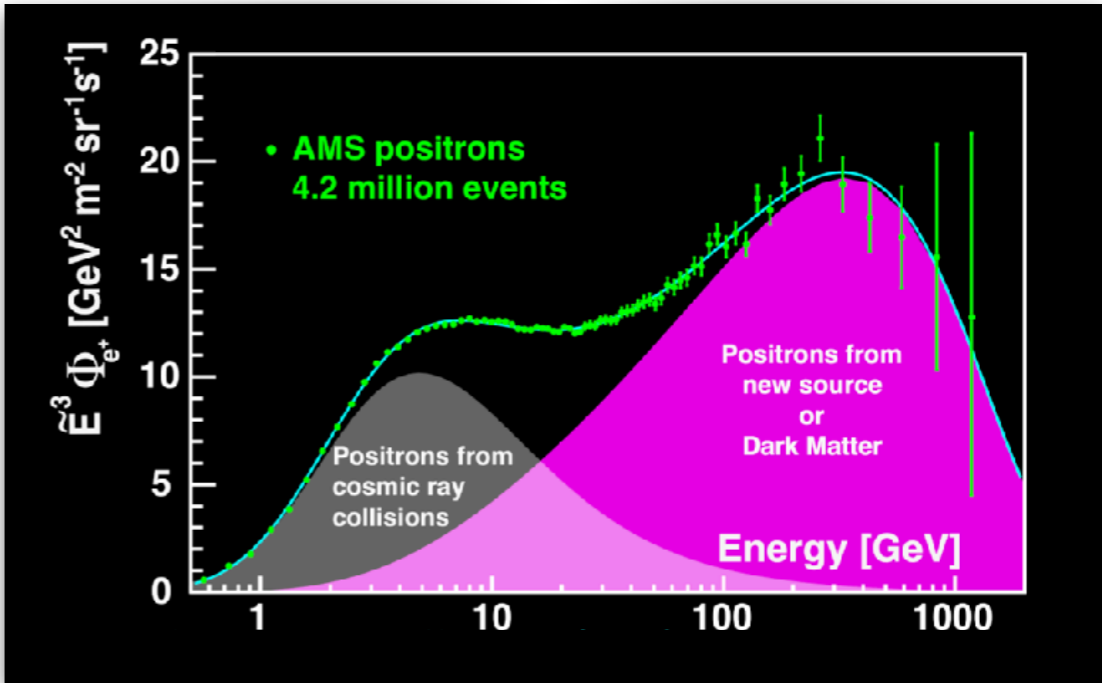
Results with emulsions (produced in Japan, sent to CERN and processed again in Japan)!

Results with emulsions (produced in Japan, sent to CERN and processed again in Japan)!



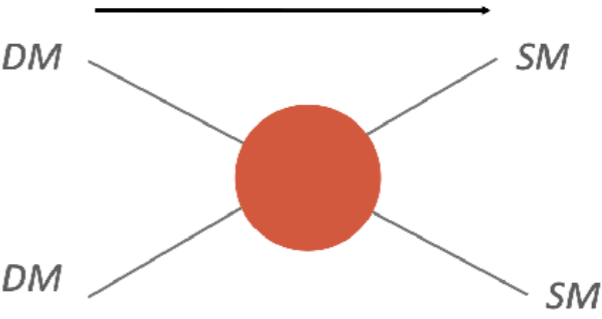


Latest data with complete dataset 2 times larger than previous release!



Interesting to note that there is a rather **sharp cut-off** indicative of mass range for DM candidate mass (in the DM annihilation hypothesis)

Relic abundance, indirect detection



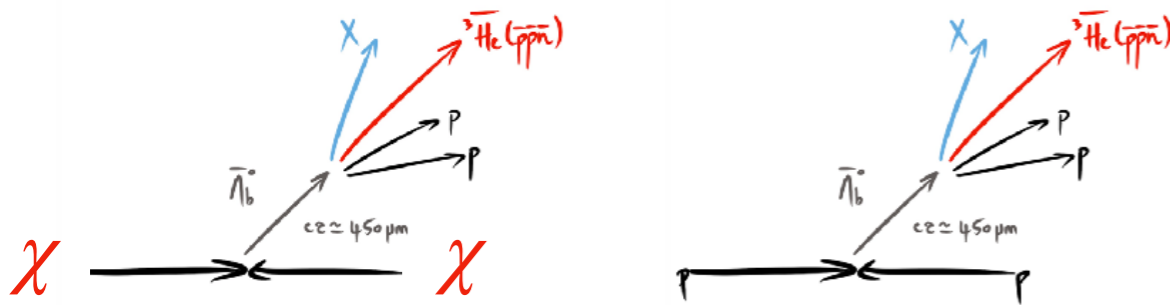
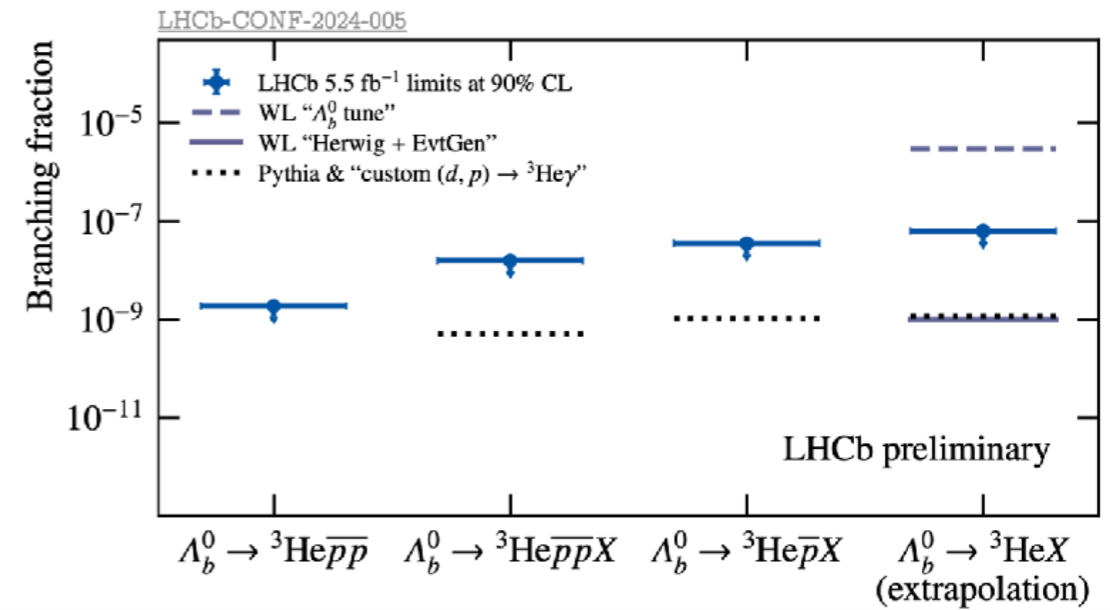




Anti-helium in Cosmic Rays could be a signature of physics beyond the standard model (e.g. **DM annihilation**)!

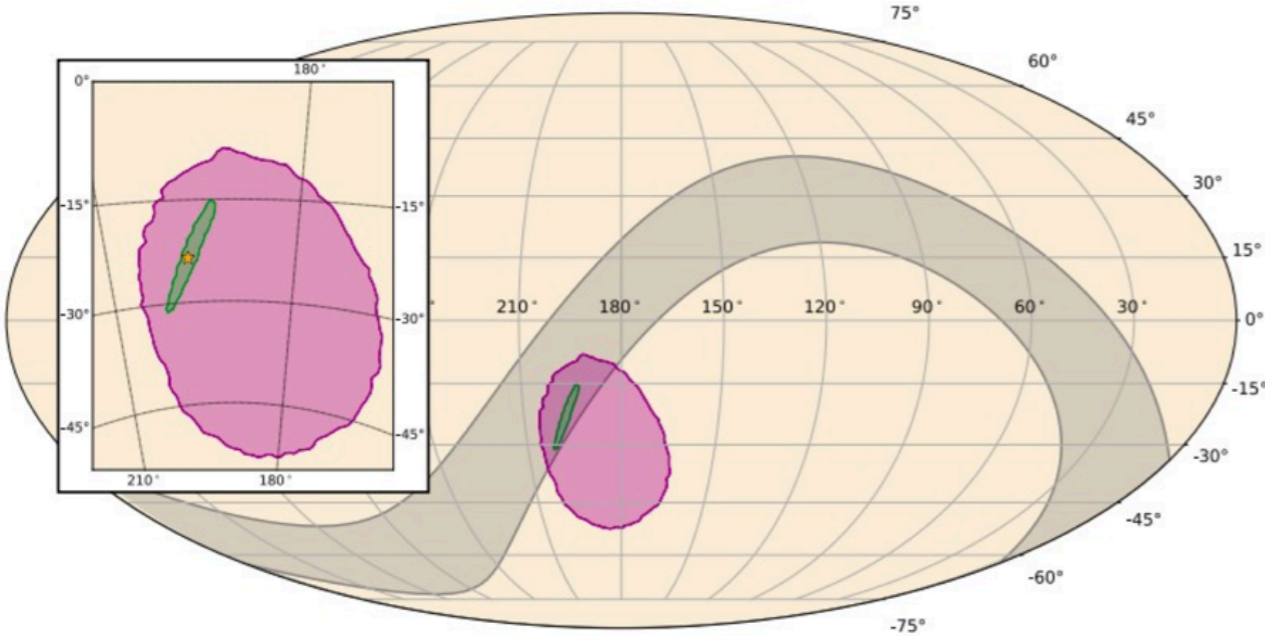
Interesting scenario where anti- $\Lambda_b$  are produced in Dark Matter scenario annihilation.

$\text{He}^3$  identified with correlated measurements of charge between VELO and Silicon Strips Tracker



Important to understand the production of  $\Lambda_b, ^3\overline{\text{He}}$  in cosmic rays

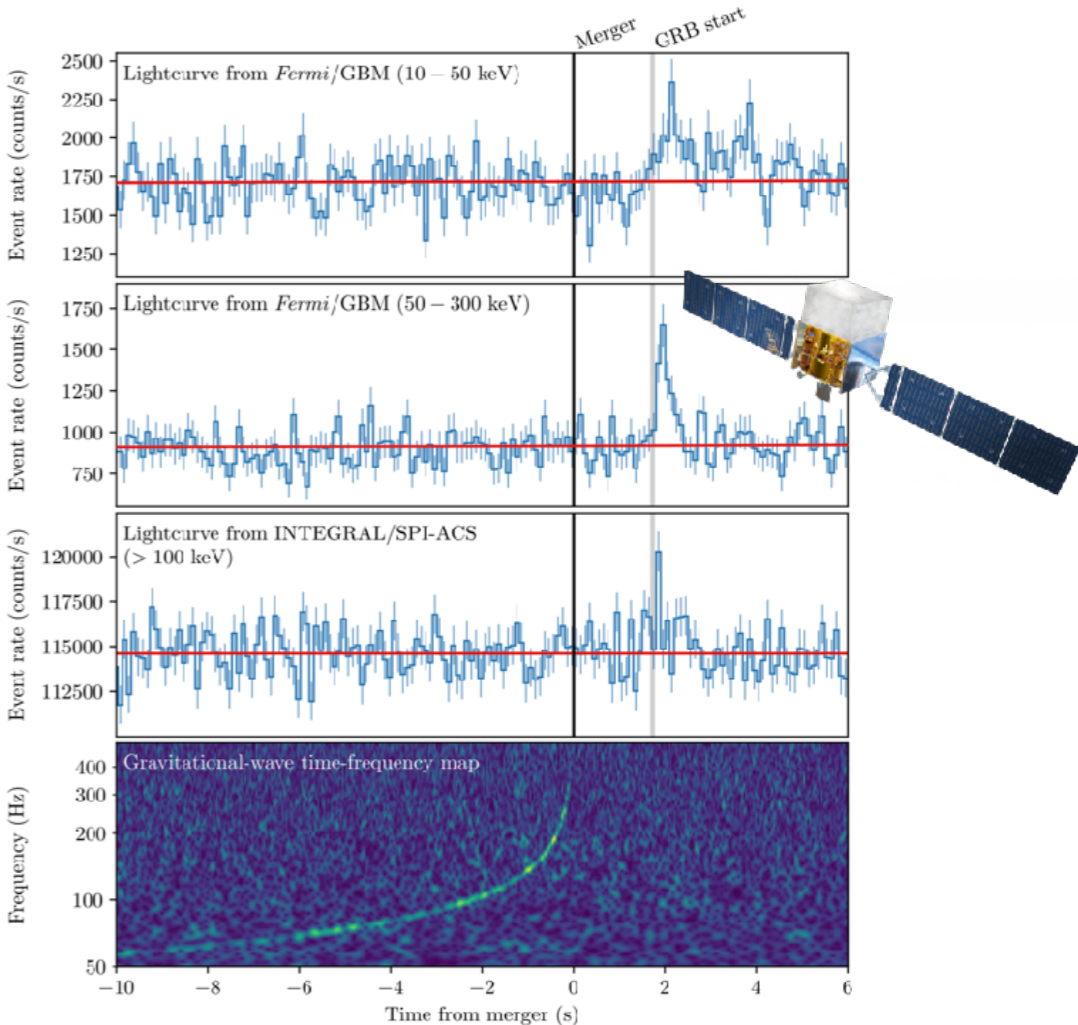
## Gravitational Waves and Gamma-rays from a Binary Neutron Star Merger: GW170817 and GRB 170817A



Signal from LIGO-Hanford, LIGO- Livingston data, Fermi Gam. Ray Burst monitor (NASA) and INTEGRAL (ESA)

Test of speed of gravity and light and equivalence principle

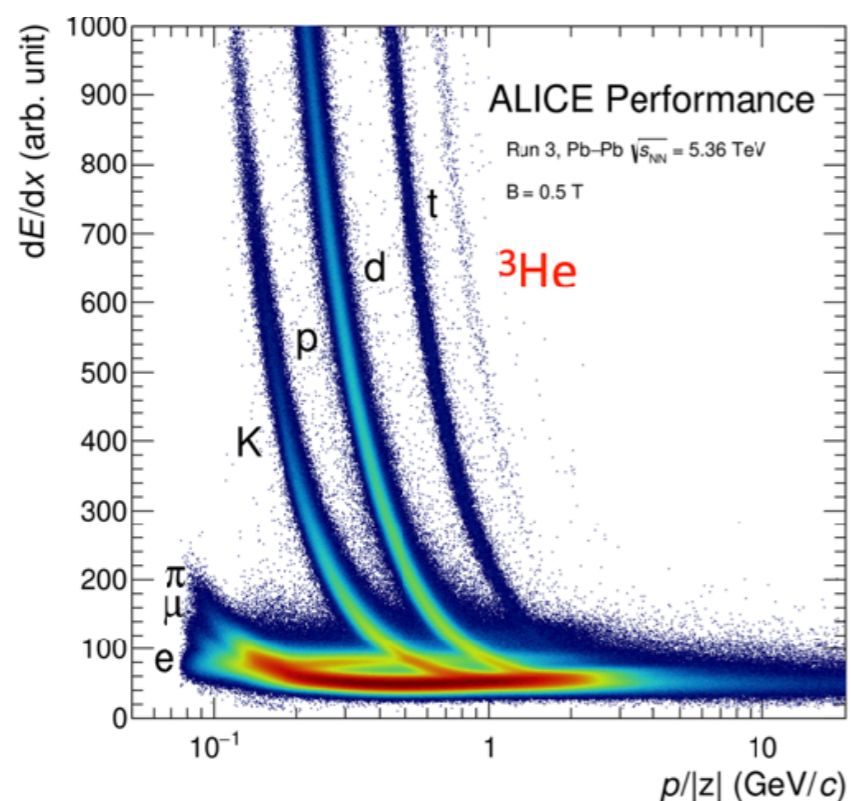
$$-3 \times 10^{-15} \leq \frac{v_{gw}}{v_{em}} - 1 \leq 7 \times 10^{-16}$$



Signal from LIGO-Hanford, LIGO- Livingston data, Fermi (NASA) and INTEGRAL (ESA)

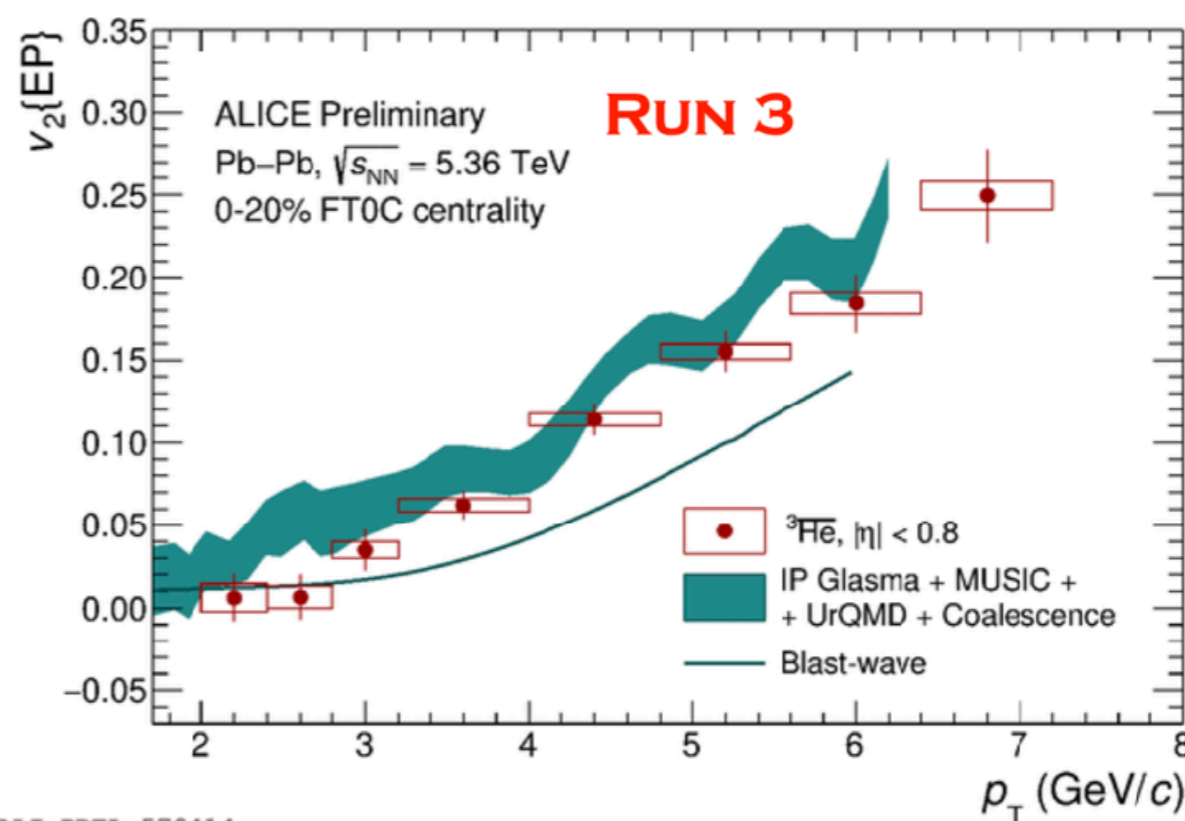
Also allows to constrain neutron star Equation of State!

Anti-Helium 3 production in heavy ion collisions, can give an interesting measure of the medium viscosity!



ALI-PERF-529714

Helium nucleus ( $Z = 2$ ), clean PID, rare probe

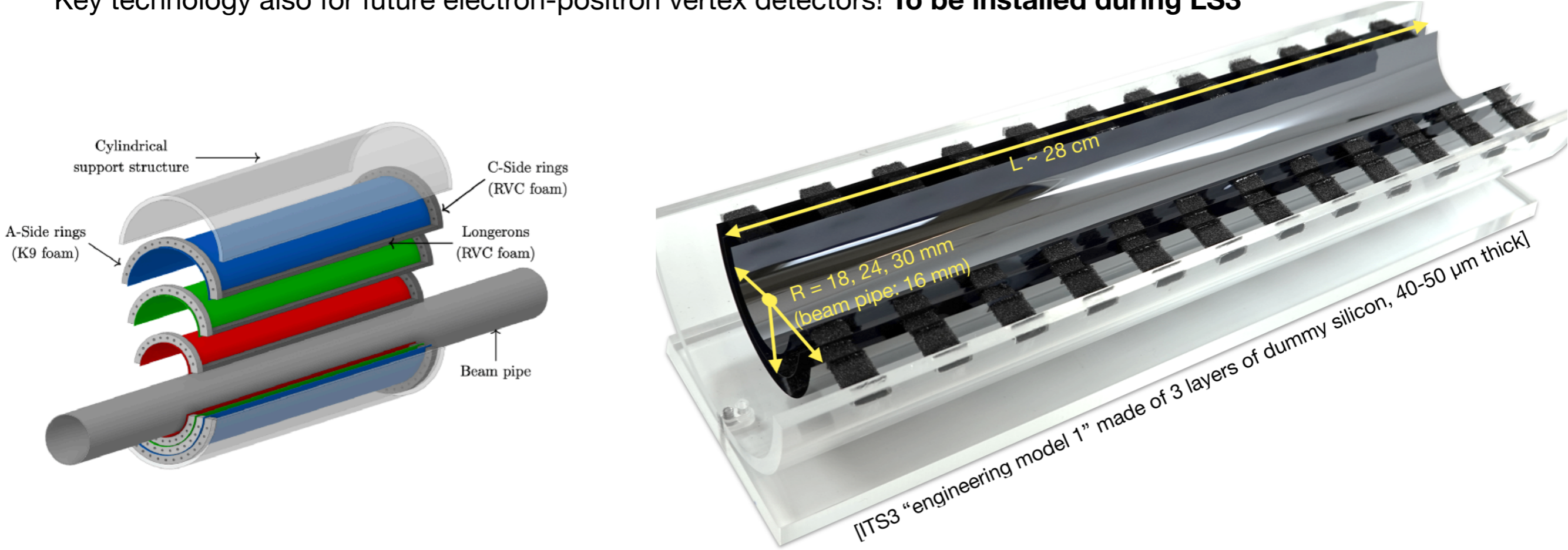


ALI-PREL-570414

Problem of **collective expansion**  
**microscopic models**



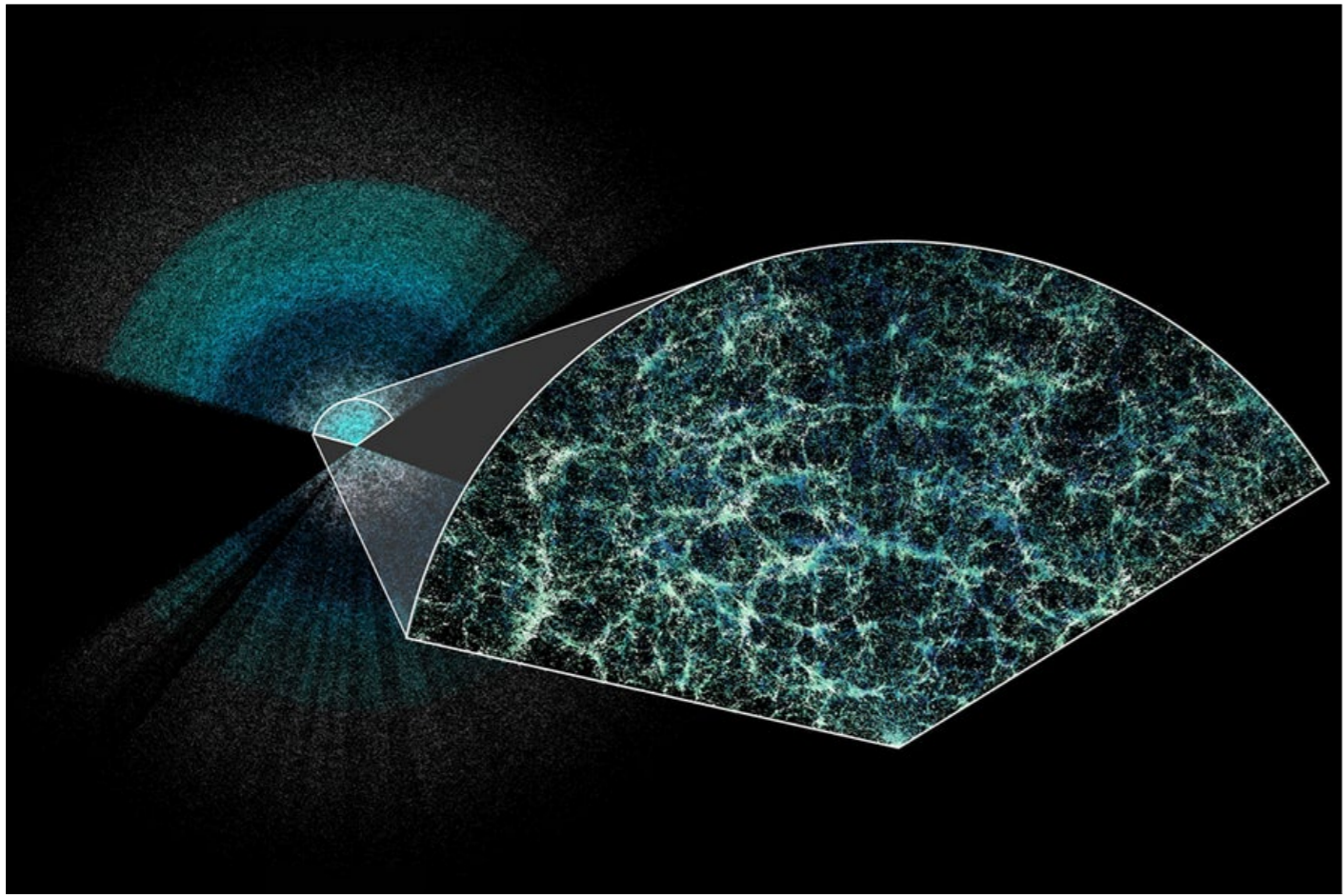
Monolithic CMOS (MAPS - Monolithic Active Pixel Sensors) technology for the ITS3 upgrade of ALICE - Key technology also for future electron-positron vertex detectors! **To be installed during LS3**



The design of the new vertex detector aims to reduce the material budget of the **first detection layer to an unprecedented minimum of 0.05% X<sub>0</sub>**, and to **get closer to the interaction point at a radial distance of 18 mm**.

Requires having in the active area only the thin MAPS silicon sensor (<50µm)

**DESI** - Dark Energy Spectroscopic Instrument in operations since 2021 and first publication of Year 1 Results!  
In one year more specific data than all previous experiments!



DESI 3D map of our universe to date.

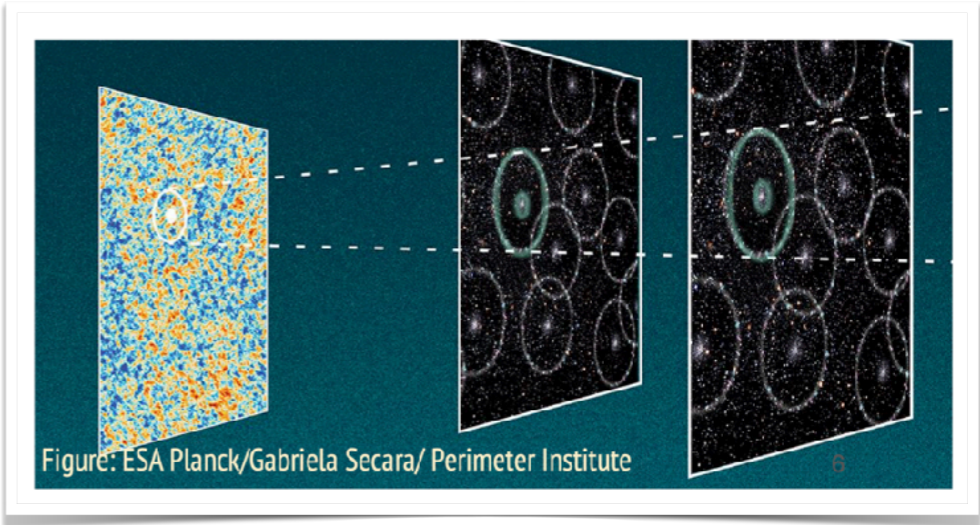
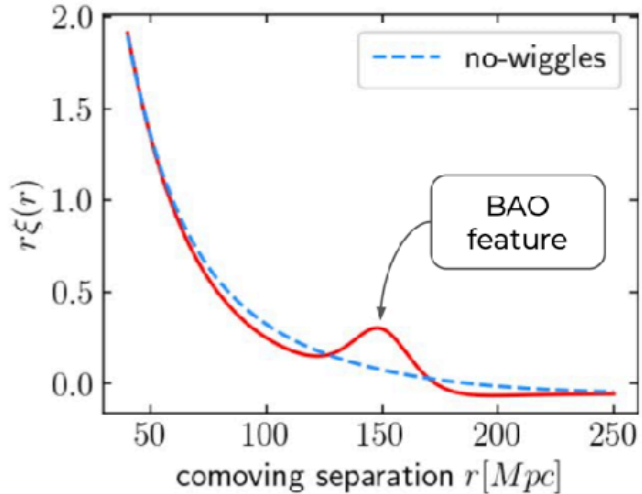


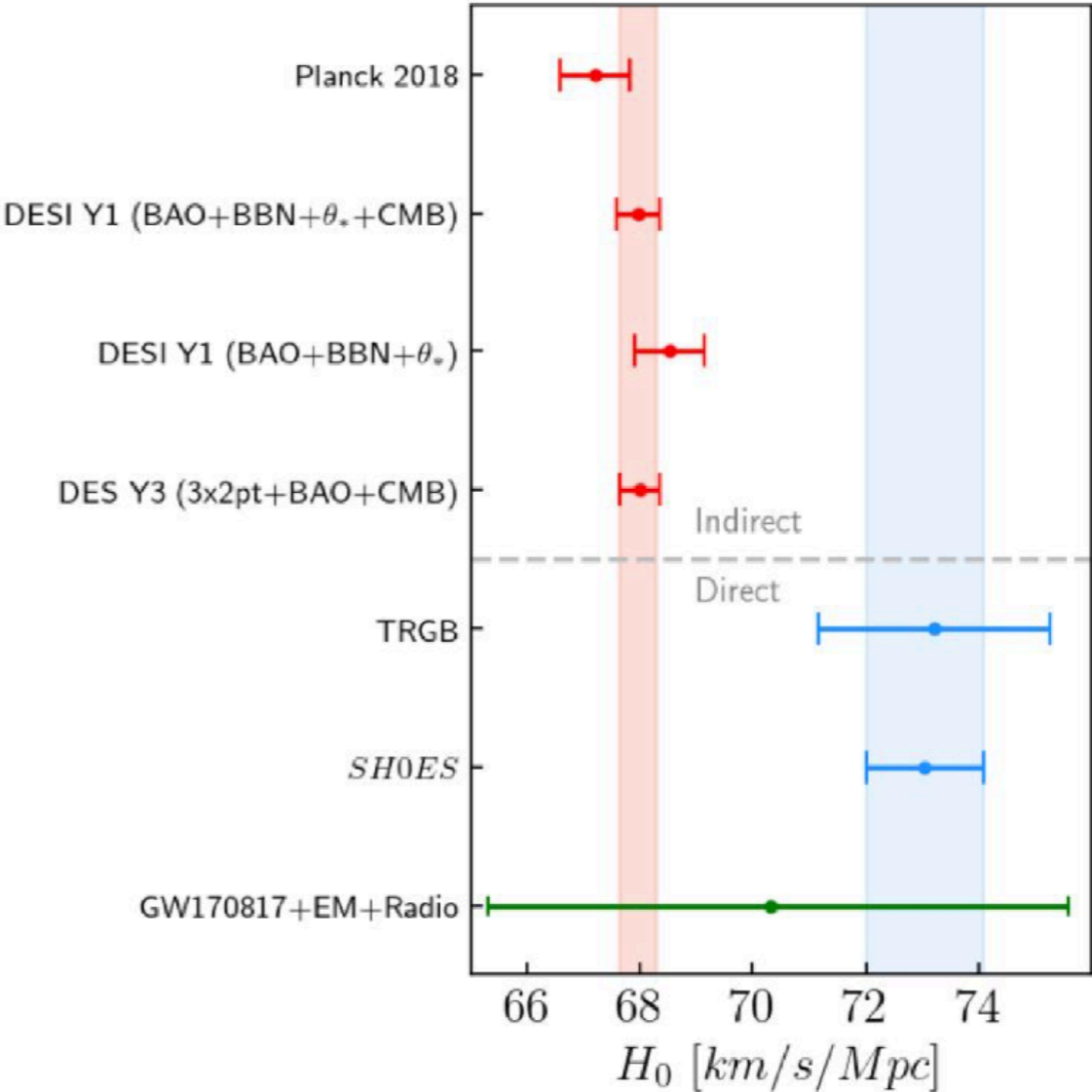
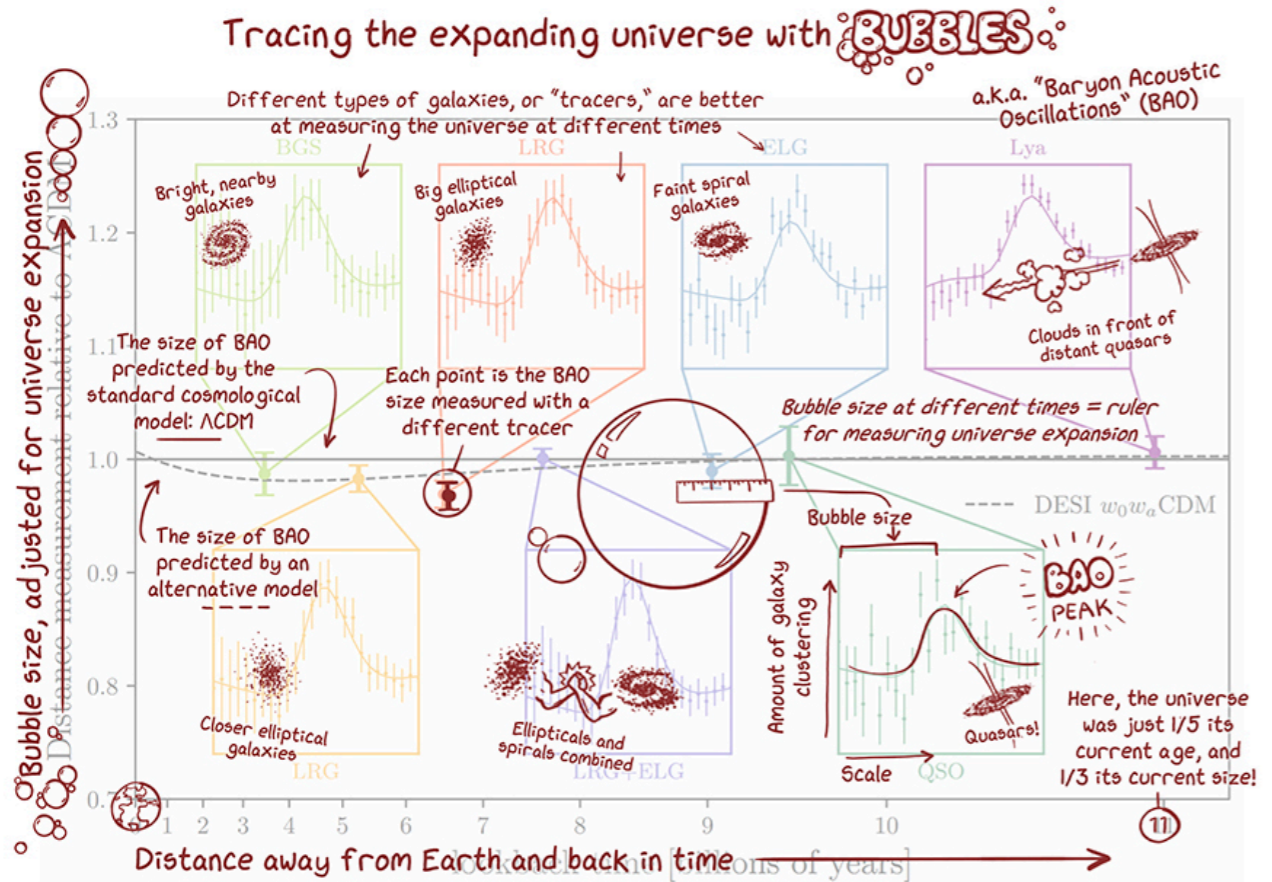
Figure: ESA Planck/Gabriela Secara/ Perimeter Institute

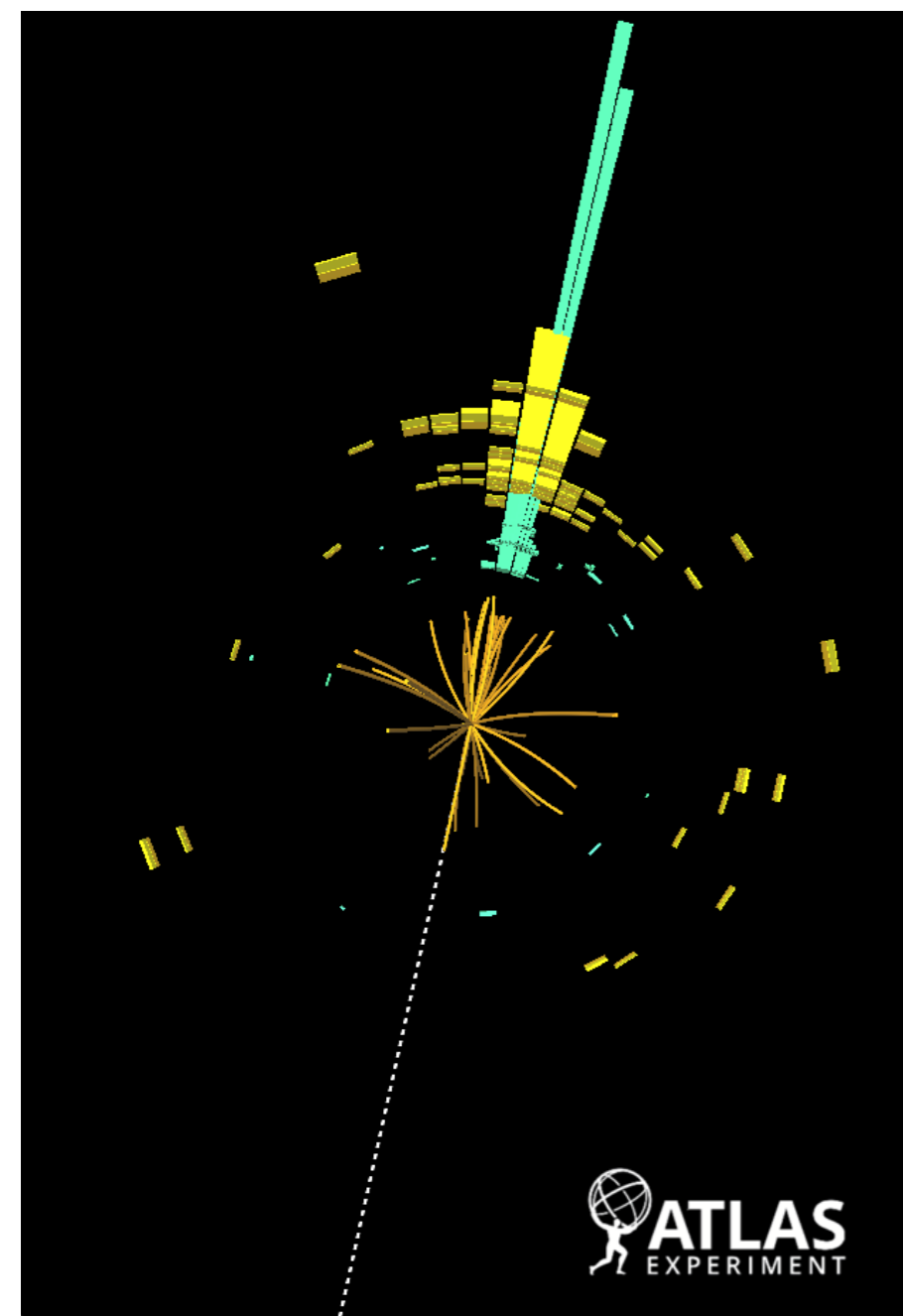
**Baryon acoustic oscillations**  
Early universe overdensities leave imprints (frozen when baryons and radiation decouple ~400,000y) in the distribution of galaxies.



- 6 million spectra grouped in 7 different redshifts [0.4<z<4.2] (LGR, BGS, QSO, Ly $\alpha$ )
- Fully exploits the capabilities of DESI and increases sensitivity!

The Hubble tension does not significantly change with much improved Cosmological BAO measurements!





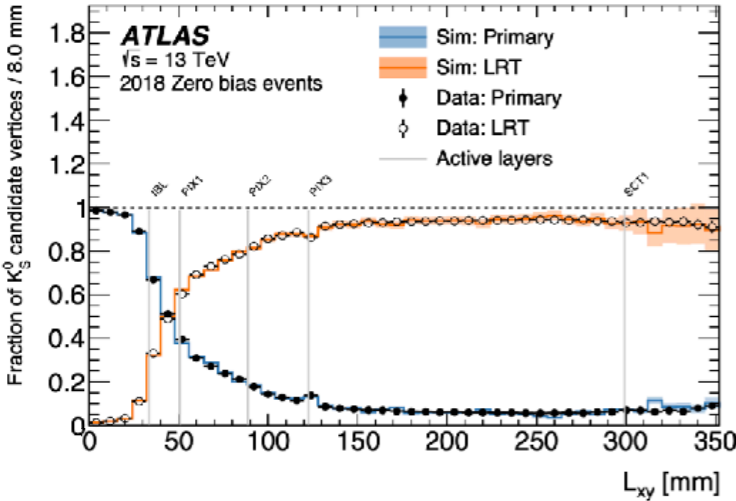
Livia presented more than 20 new searches for new physics in extended higgs sectors, Dark sector, SUSY, Heavy fermions, and EFT. Impressive harvest of searches for new physics mostly at Run 2!

Leaving no stones unturned!

Reference	Topic	Experiment	Model	Explored energy range (GeV)
<a href="#">HDBS-2021-07</a>	$H \rightarrow aa \rightarrow bb\tau\tau$	ATLAS	Extended Higgs Sector	0 - 300
<a href="#">HDBS-2020-11</a>	$H^\pm \rightarrow cs$	ATLAS		0 - 300
<a href="#">HDBS-2023-19</a> <a href="#">HIG-24-002</a>	Combination of charged H $H \rightarrow ZZ \rightarrow 4l$	ATLAS CMS		0 - 3000
<a href="#">HIG-22-004</a>	$A \rightarrow Zh(\tau\tau)$	CMS		0 - 1500
<a href="#">SUS-24-001</a>	$\phi \rightarrow b\bar{b}$	CMS		0 - 1800
<a href="#">EXOT-2018-55</a>	Prompt Leptonjets	ATLAS	Dark Sector +ALPs	0 - 300
<a href="#">EXOT-2022-04</a>	Neutral LLP into displaced jets	ATLAS		0 - 3000 - displaced
<a href="#">SUS-23-004</a>	mono- $t$	CMS		0 - 3000 dark matter
<a href="#">SUS-23-012</a>	mono- $h(\tau\tau)$	CMS		0 - 3000 dark matter
<a href="#">SUS-23-018</a>	$H \rightarrow Za \rightarrow ll\chi\chi$	CMS		0 - 1800
<a href="#">SUS-24-004</a>	pMSSM	CMS	Supersymmetry	0 - 3000
<a href="#">SUS-23-003</a>	Compressed SUSY w/ RJR	CMS		0 - 3000
<a href="#">ATLAS-2024-011</a>	Run3 displaced leptons	ATLAS		0 - 3000 - displaced
<a href="#">ATLAS-2024-008</a>	$VLL \rightarrow \tau b$	ATLAS	Heavy Fermions	0 - 1500
<a href="#">EXO-23-015</a>	$VLL \rightarrow \tau a(\gamma\gamma)$	CMS		0 - 3000 - displaced
<a href="#">B2G-22-005</a>	$t^* \rightarrow tg$	CMS		0 - 3000
<a href="#">EXO-23-010</a>	$ll + b - \text{jets, non - resonant}$	CMS	EFT	0 - 3000
<a href="#">EXO-24-007</a>	Low mass dijet+ISR	CMS	Z' Mediator	0 - 3000
<a href="#">EXO-22-006</a>	$Z' \rightarrow \mu\mu + b - \text{jets, resonant}$	CMS		0 - 3000

### Run 3 search!

Improving reconstruction techniques e.g. ATLAS Large Radius Tracking at Run 3 and reprocessed Run 2!

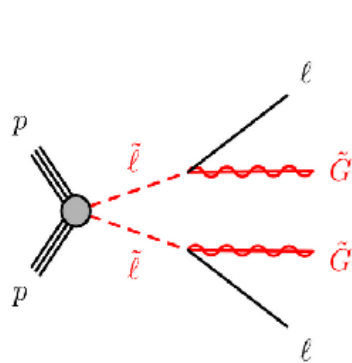


LRT performance tested with Ks reconstruction [\(Paper\)](#)

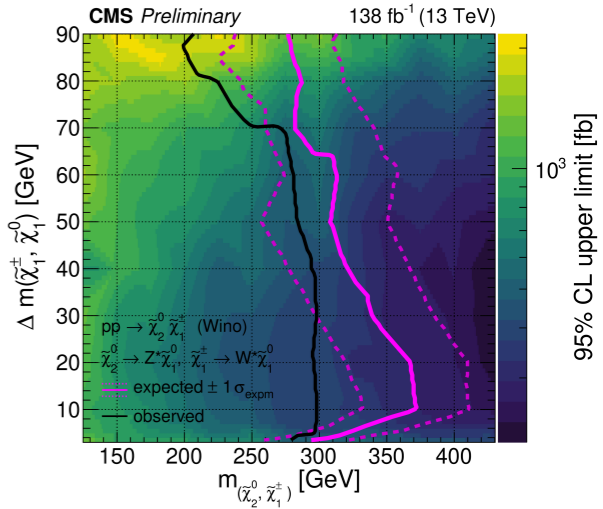
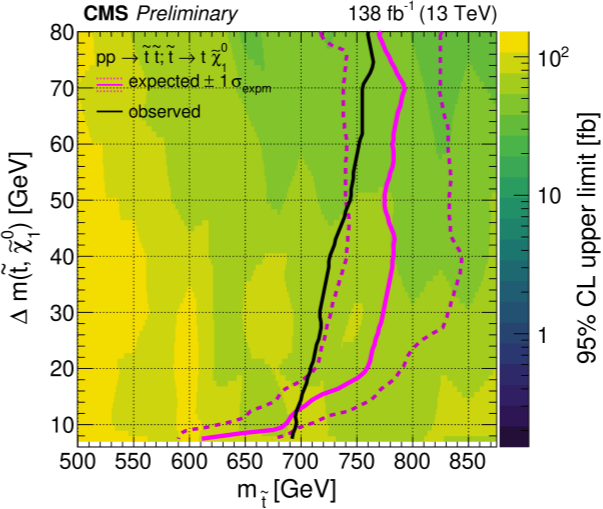
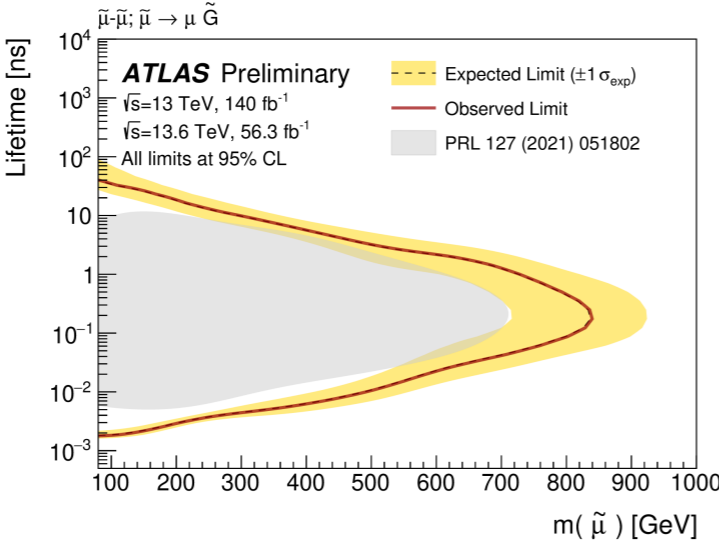
Search done for smuons, selectron and staus.

### Searches for EWK production SUSY compressed scenarios

Wide range of signatures targeting electroweakinos, sleptons, and top squarks with focus on events with a high transverse momentum system from initial-state-radiation jets and significant missing transverse momentum.



Long lifetime due to the small coupling to the low mass gravitino!

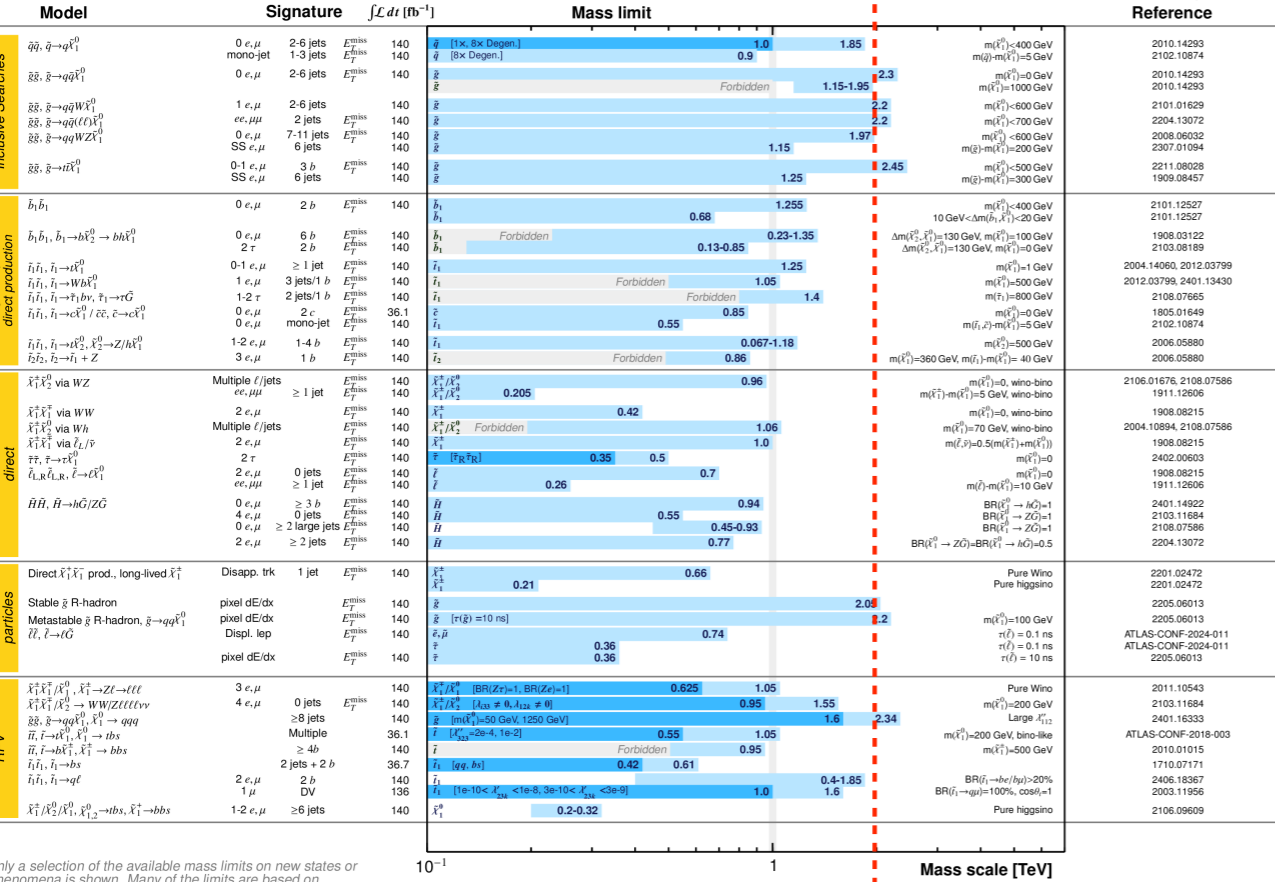




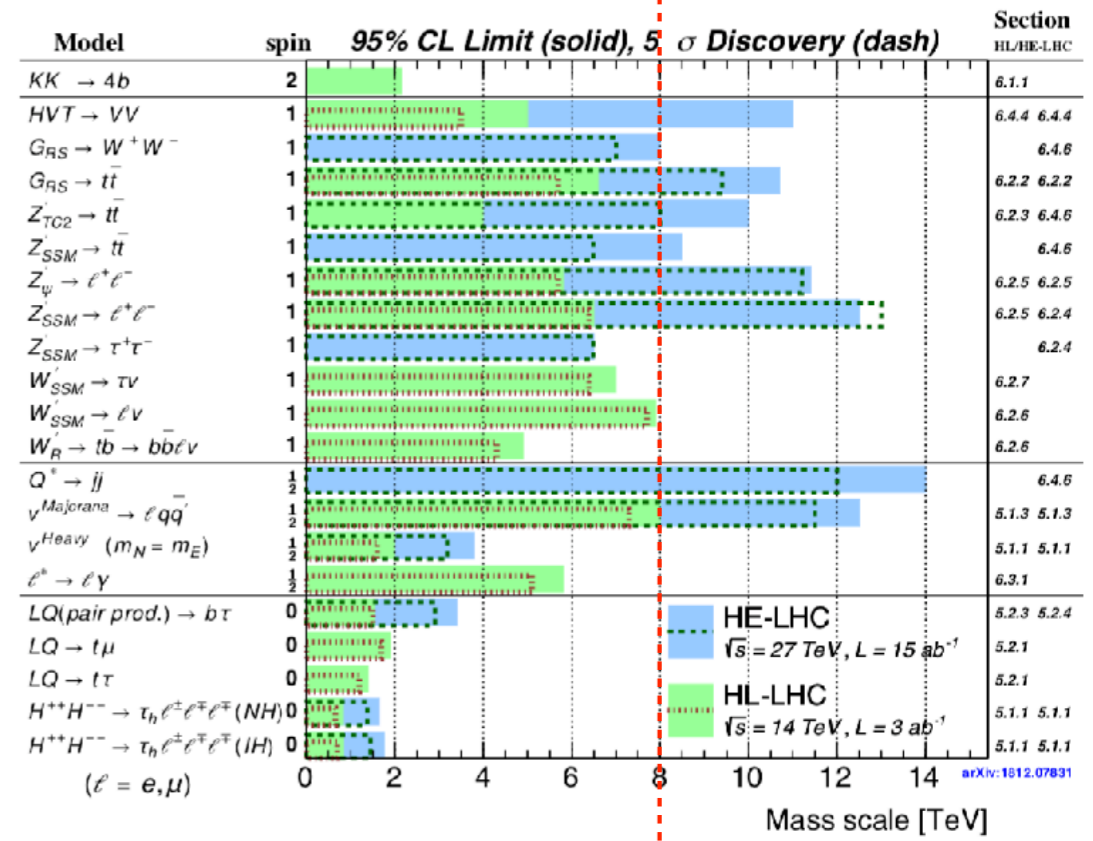
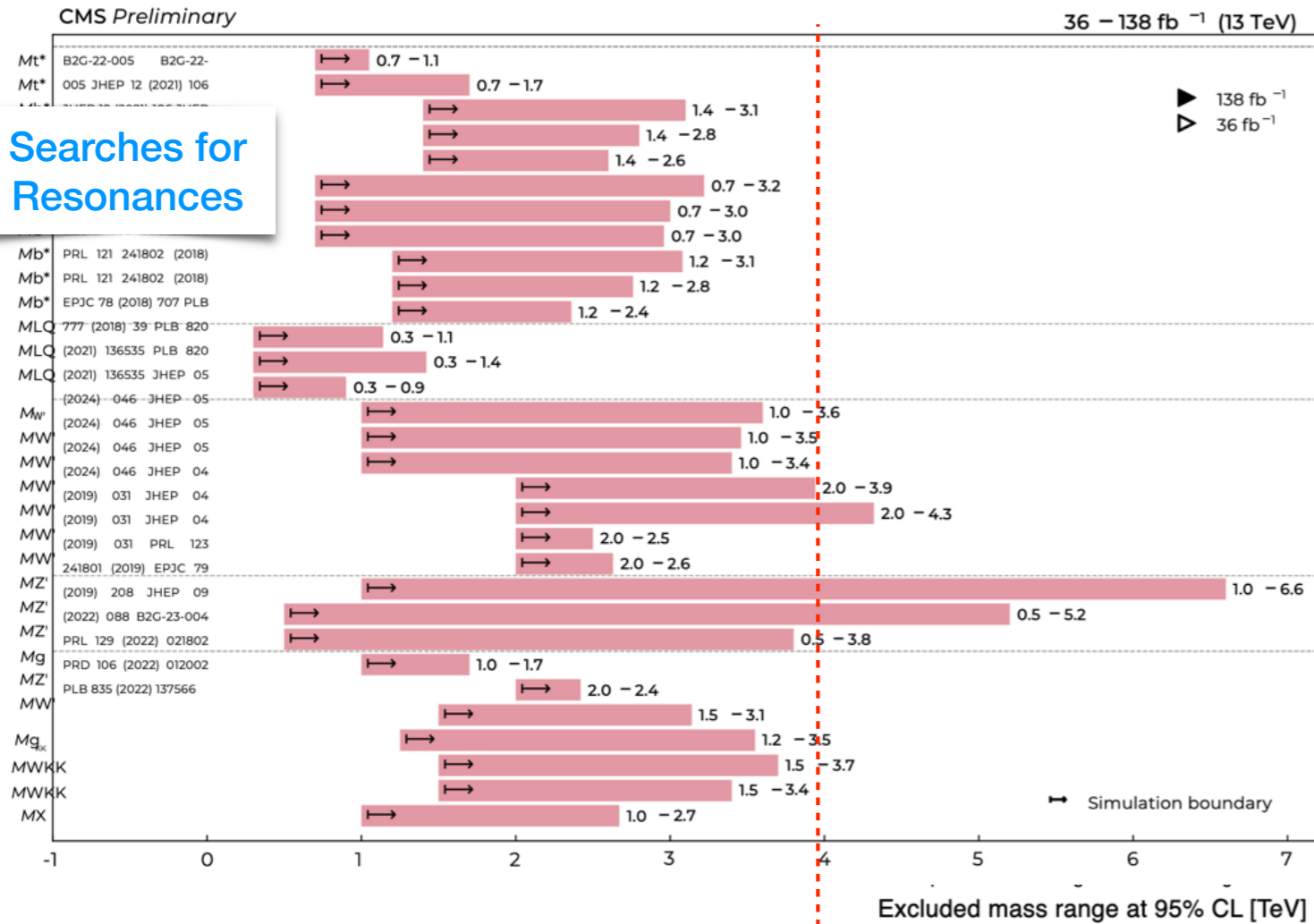
(in large variety of topologies and models)

ATLAS SUSY Searches\* - 95% CL Lower Limits  
July 2024

ATLAS Preliminary  
 $\sqrt{s} = 13$  TeV



(in large variety of topologies and models)

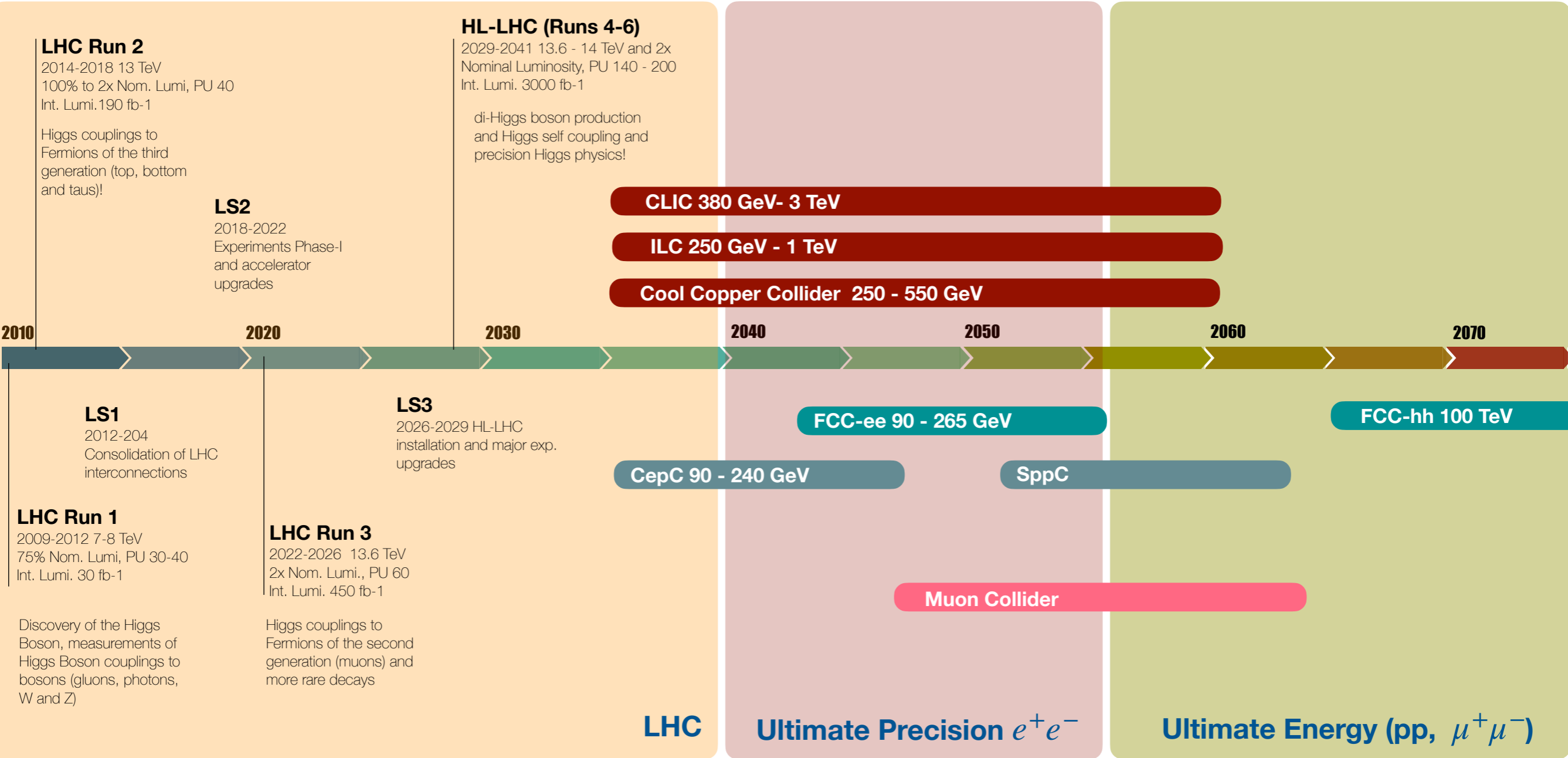


**Still Room for Discoveries!**

**Discovery potential of Z' and W' up to 6 TeV and 8 TeV**

**Discovery potential of Leptoquarks up to 1.5 - 2 TeV**

HL-LHC YR  
1812.07831



**The field is facing a defining moment in its history!**

**P5 - Reveal the secrets of the Higgs boson and much more... ultimate EW precision factory!**

“We would not consider the theory of electromagnetism established if we had only verified the strength of electromagnetic forces to within 10% accuracy.” (Salam, Wang, Zanderighi, [Nature](#))

**P5 - A realistic path to a 10 TeV parton center-of-momentum (pCM) collider.**

Closing quotes from the panel discussion ([video](#))

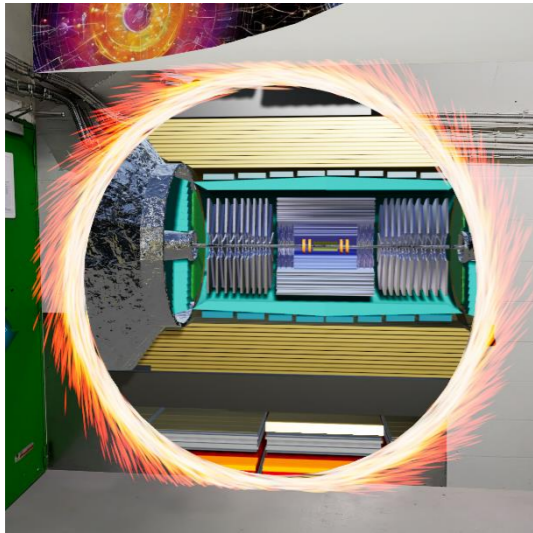
**Yifang Wang:** “In the Future, not very far from now there will be a Higgs factory!”

**Lia Meringa** ‘Need to continue to be bold and ambitious and dream big!’”

**Fabiola Gianotti:** “As we have seen at this conference the field is extremely vibrant and exciting!”

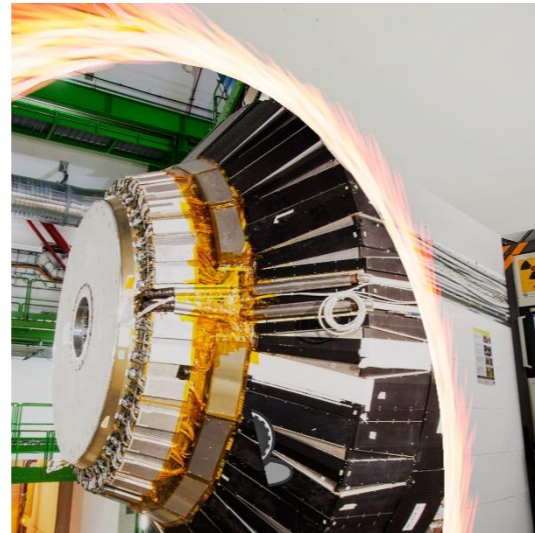


- ✦ Ian and I had several correspondence with the author before the conference
- ✦ Now we got the full instruction from Muhammad and a test play
- ✦ Introduce the CMS VR to the PPD exhibition room!



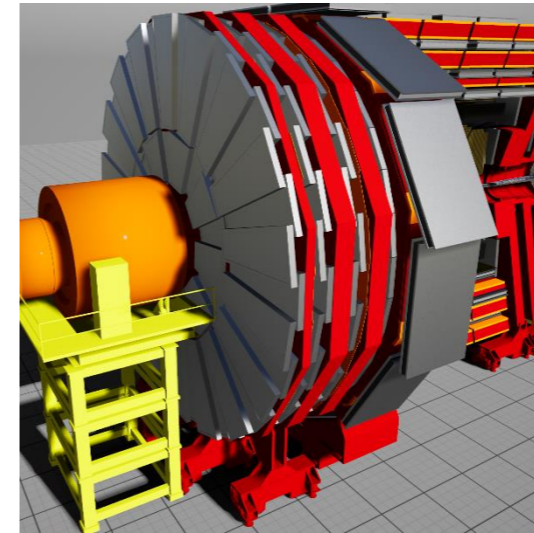
## CMS VR

- Visualise the CMS detector using **virtual reality**.
- Perfect for visits.



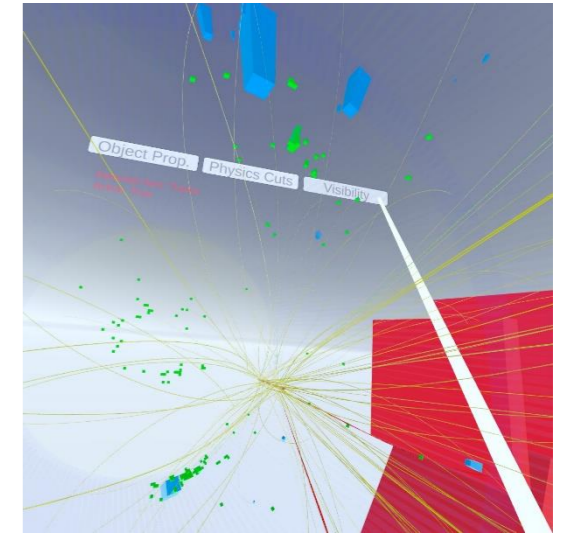
## Captivating visuals

- Engaging and **realistic visuals**.
- Coupled with interactive elements.



## Detailed detector

- **High-fidelity** CMS model.
- Explorable from all angles.



## Numerous applications

- Also useful for **event displays, engineering**.
- A wide range of audiences.

[https://indico.cern.ch/event/1291157/contributions/5887162/attachments/2899204/5083678/240718\\_CMSVR\\_ICHEP\\_AnsarIqbal.pdf](https://indico.cern.ch/event/1291157/contributions/5887162/attachments/2899204/5083678/240718_CMSVR_ICHEP_AnsarIqbal.pdf)

# Let's enjoy our physics! - Soshi Asai

