## Particle Identification

**PPD Lectures 2025** 

Tai-Hua Lin

RAL - STFC

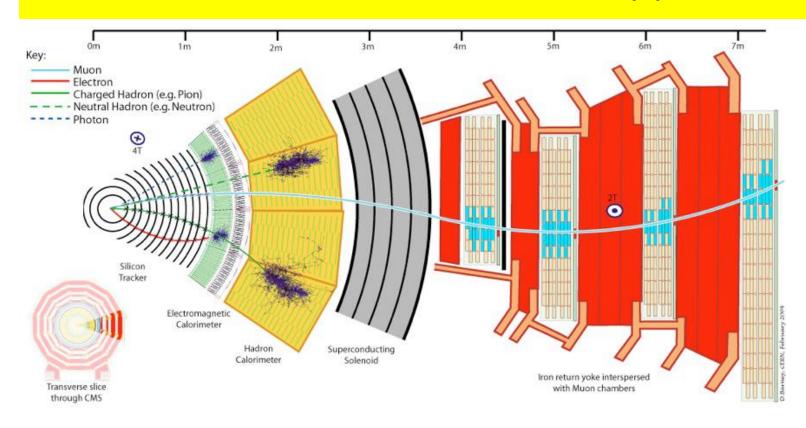


#### **Outline**

- Introduction
- Particle ID techniques:
  - > Time of flight
  - → dE/dx
  - Transition radiation
  - Cherenkov radiation
    - RICH detectors
    - DIRC detectors
- Give examples of experiments and detectors



## Introduction (i)



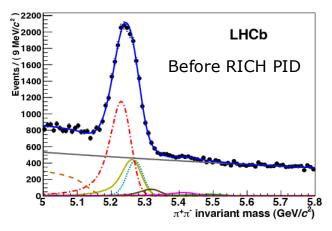
A "generic" particle physics experiment has "built-in" a lot of particle ID for:

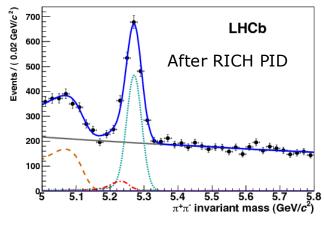
Electrons, photons, neutrons and muons

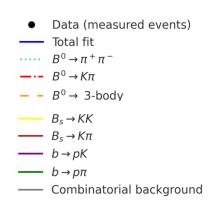


# Introduction (ii)

- Particle ID for hadrons
  - Distinguish between kaons, pions, protons, deuterons
    - Required for flavour physics
- Specific experiment requirements
  - Distinguish between electrons and charged pions
  - Distinguish between electrons and muons
    - For neutrino experiments
    - For rare decay searches







$$B \to \pi^+\pi^- (h^+h^-)$$



## Time of flight

Time difference for two particles with masses  $m_1$  and  $m_2$  for length L and momentum P (which is given by the tracking system)

$$\Delta t = \frac{L}{\beta_1 c} - \frac{L}{\beta_2 c} = \frac{L}{c} \left[ \sqrt{\left(1 + \frac{m_1^2 c^2}{P^2}\right)} - \sqrt{\left(1 + \frac{m_1^2 c^2}{P^2}\right)} \right]$$

$$P^2 \gg m^2 c^2$$
 
$$\Delta t \sim \frac{Lc(m_1^2 - m_2^2)}{2P^2}$$

Mass resolution depends on time resolution of detectors and distance L Typical values: L=3.5 m,  $\Delta t$ =100 ps, for 3  $\sigma$  separation,  $P_{max}$ =2.1 GeV/c

Requires fast detectors: scintillation, Cherenkov, fast collection of ionisation

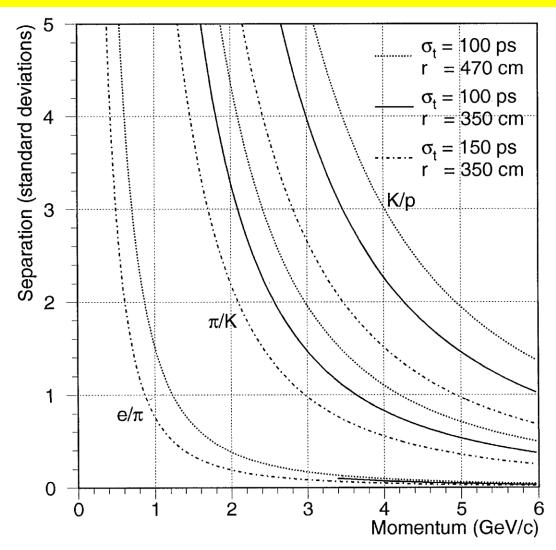


#### PID with TOF

Expected PID performance for various path lengths and time resolutions

In 100 ps light travels 3 cm

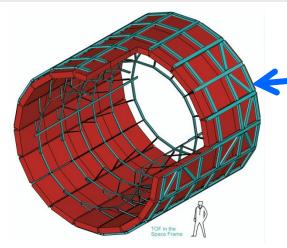
TORCH, a proposed future detector, can push  $\pi/K$  separation to 9 GeV/c with L~10 m and time resolution per track ~ 20 ps



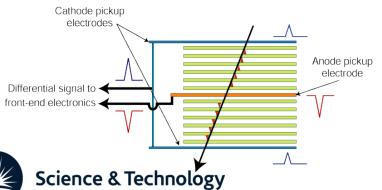


# **TOF: ALICE experiment**

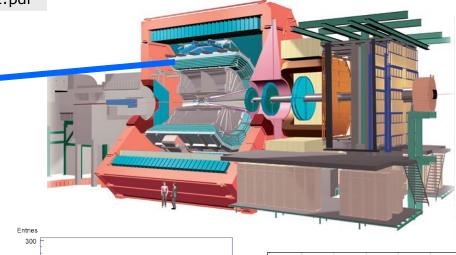
https://cds.cern.ch/record/545834/files/alice-tdr-008-add-1.pdf

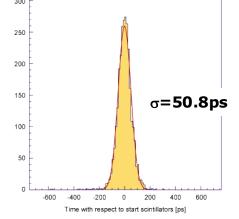


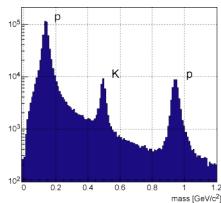
Multigap Resistive Plate Chambers (MRPCs) Resistive plates made of glass. 2 X 5 gaps : 250 mm Readout by HPTDC chip (High Performance Time to Digital Converter)



**Facilities Council** 





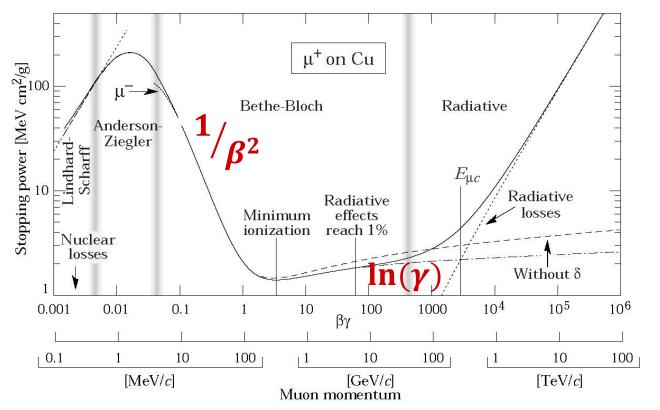


 $3\sigma$  n/K separation up to 2.2 GeV/c and K/p separation up to 4 GeV/c.

# Ionisation Measurement (dE/dx)

$$-\frac{dE}{dx}(eVcm^{2}g^{-1}) = Kq^{2}\frac{Z}{A\beta^{2}}\left[\frac{1}{2}\ln\frac{2m_{e}c^{2}\beta^{2}\gamma^{2}T_{max}}{I^{2}} - \beta^{2} - \frac{\delta(\beta\gamma)}{I^{2}}\right]$$
Ionisation Constant for material

Density correction



$$T_{\rm max} \approx 2m_e c^2 \beta^2 \gamma^2$$

Max energy in single collision

$$K = 4\pi N r_e^2 m_e c^2$$

### dE/dx in silicon

Bethe-Block formula (for  $2 \gamma$  m/M << 1)

$$\frac{\mathrm{d}E}{\mathrm{d}x} = \frac{4\pi Ne^4}{mc^2\beta^2}z^2 \left(\ln\frac{2mc^2\beta^2\gamma^2}{I} - \beta^2\right)$$

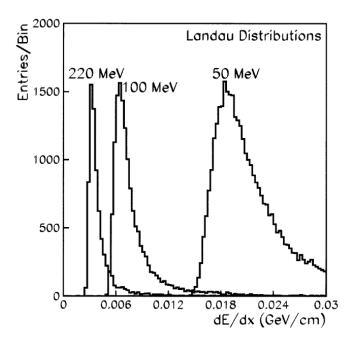


Fig. 1. Landau  $\langle dE/dx \rangle$  distributions for pions of momenta 220, 100 and 50 MeV going through 300  $\mu$ m of silicon.

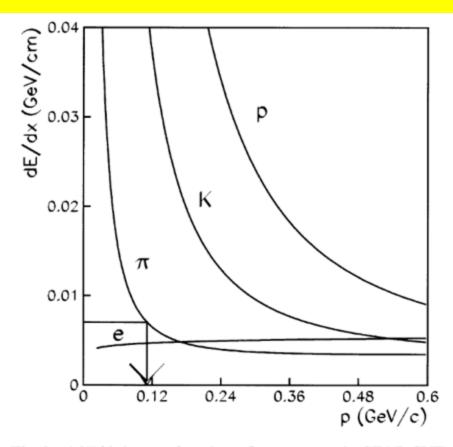


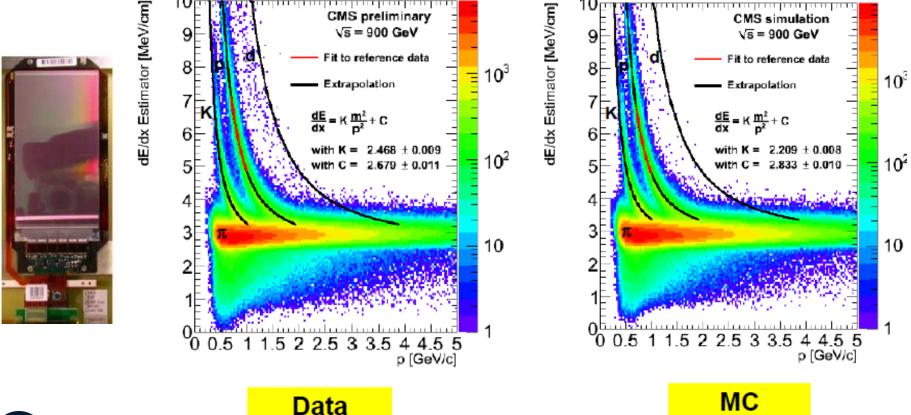
Fig. 3.  $\langle dE/dx \rangle$  as a function of momentum in STAR-SVT. For low momentum pions one can use the pion curve to extract the momentum from the measured dE/dx as shown.

Ref: NIMA 469(2001) 311-315

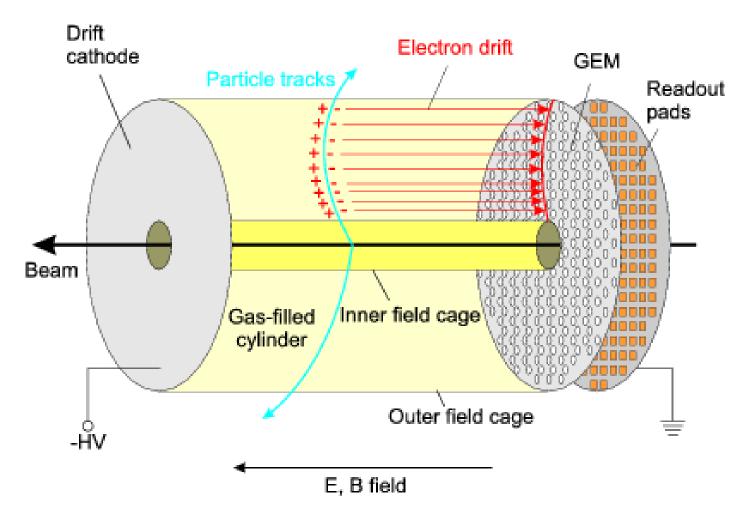


### dE/dx: CMS Si tracker

- Each Silicon sensor gives a dE/dX measurement.
- Estimate the Most Probable Value from several (10-25) measurements (Truncated Mean: Ignore upper 40%)



### **Drift chambers**





## dE/dx: Drift Chambers

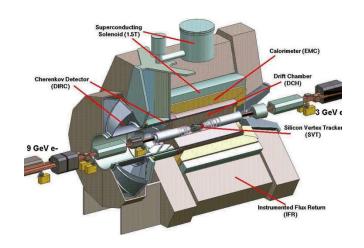
Larger Landau fluctuations compared to those from Silicon detectors.

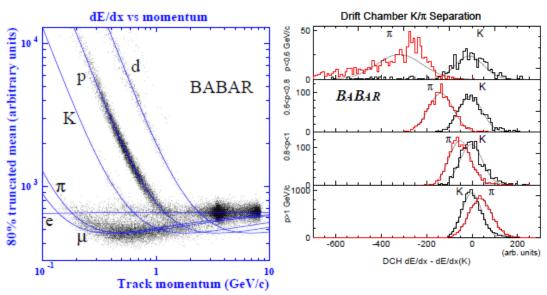
So many measurements needed to get the average.

Ref: IEEE-TNS VOL:47, NO:6, Dec2000.

#### **BABAR Drift Chamber:**

Gas mixture 80% helium, 20% isobutane, 3500–4000 ppm water vapor,  $\sim$  80 ppm  $O_2$ 





Good  $\pi/K$  separation up to  $\sim$  700 MeV/c.

dE/dx resolution  $\sim$  7.5%



## Combined TOF and dE/dx

NA 49: Heavy ion experiment,

TOF: Scintillator thickness=2.3 cm, time resolutions= 59 ps, 95 ps

TPC for dE/dx.

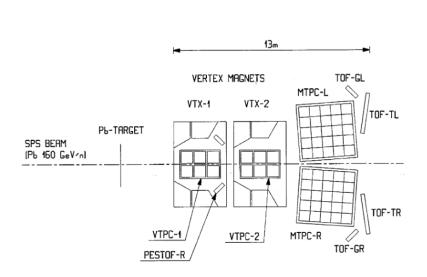


Fig. 2. Central part of the NA49 experimental setup.

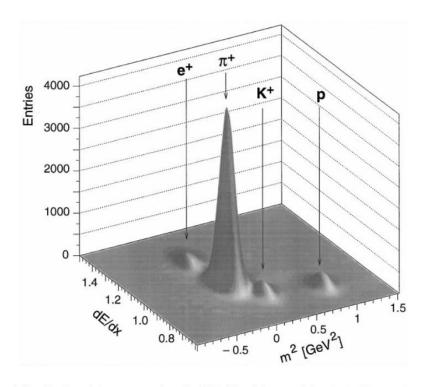


Fig. 5. Particle separation in NA49 with combined dE/dx and TOF information.

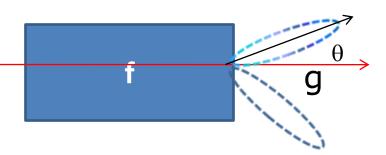


## **Transition Radiation (TR)**

- Transition Radiation: Radiation in the x-ray region when ultra relativistic particles cross the boundary between 2 media with different dielectric constants.
- Mainly for  $e-\pi$  separation in 0.5 GeV/c  $\rightarrow$  200 GeV/c.

Full explanations and the derivations from Maxwell's equations:

NIMA 326 (1993) 434-469 and in references



$$\frac{dW}{d\omega d\theta} = \frac{2\alpha}{\pi} f_0(\theta)$$

The radiation is peaked at a small angle:

$$\theta = 1/\gamma$$

$$f_0(\theta) = \theta^3 \left( \frac{1}{\gamma^{-2} + \theta^2 + \xi_g^2} - \frac{1}{\gamma^{-2} + \theta^2 + \xi_f^2} \right)$$

$$\xi_i = \omega_i / \omega_i = \text{plasma freq of medium i}$$



### **TR Detection**

Integrating the previous equation, one can get , for  $\xi_g=0$ ,

$$W_{TR} = 2.43 \times 10^{-3} \omega_f \gamma$$

- W: energy spectrum radiated by a charged particle
- γ = E/m , Lorentz factor of the particle.
   This makes PID possible by measuring W.
- Lighter particle give larger signal
- $\omega_f$  = plasma frequency = 28.8 ( $\rho$  Z/A)<sup>0.5</sup> eV
- ρ=density, Z=atomic weight, A=atomic number

For example, for  $\,\omega_{\rm f}$  = 0.02 keV and  $\gamma$  = 5000, most of the photon energy is in the range 10 keV<  $\omega$  <100 keV (ie. 0.1  $\omega_{\rm c}$ <  $\omega$  <  $\omega_{\rm c}$  where  $\omega_{\rm c}$  = cut-off frequency ).

Number of photons produced = 
$$N(>\omega) = \frac{\alpha}{\pi} \left\{ \ln \frac{\omega_c}{\omega} \left( \ln \frac{\omega_c}{\omega} - 2 \right) + \frac{\pi^2}{12} + 1 \right\}$$

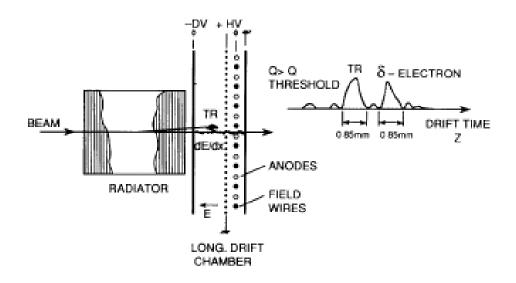
For  $\omega_c$ =100 keV and  $\omega$ =1 keV, N= 0.03 for a single surface.

Hence to get sufficient number of photons, large number of interfaces are used: a stack of many foils with gaps in between.



#### TRD Example 1: HELIOS experiment (NA34)

- The minimum thickness of the foils and air gaps are determined by the size of the 'formation zone' and the interference effects. (typically, foils can be 10-20  $\mu$ m thick and are made of polypropelene)
- Behind a TRD foil stack there is a MWPC or drift chamber where the TRD signal is detected along with the signal from the charged track.

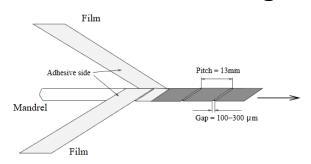


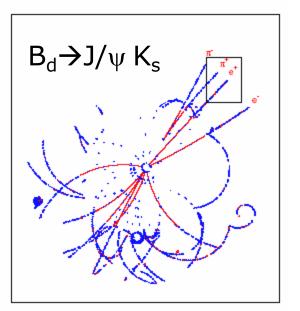
Drift space = 10 mm, Anode space = 6 mm Drift time= $0.5 \rightarrow 1 \mu s$ May use FADC or discriminators

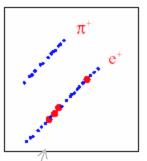


#### **TRD Example 2: ATLAS TRT**

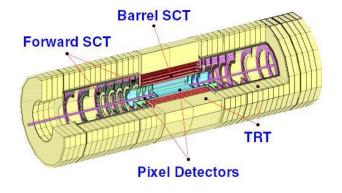
#### TRT straw wall design







Blue dots: ionizing hits Red dots: TR hits



#### Barrel and endcap TRT



In an average event, energy deposit from: Ionization loss of charged particles  $\sim 2.5 \text{ keV}$ TR photon > 5 keV.

(Photon emission spectrum peaks at 10-30keV)

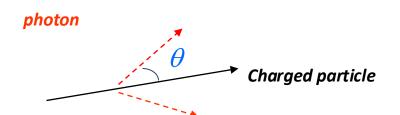


## **Cherenkov radiation**





#### **Basics of Cherenkov radiation**



$$\cos( heta) = rac{1}{n(\lambda)eta} \quad nig(E_{ph}ig) = c/c_M \quad \text{Refractive index}$$
 
$$\beta = v/c = P/E = P/\sqrt{P^2 + m^2} = rac{1}{\sqrt{1 + (m/p)^2}}$$

 $\beta$  = velocity of the charged particle in units of speed of light (c) vacuum

P, E, m = momentum, energy, mass of the charged particle

C<sub>M</sub> = speed of light in the medium (phase velocity)

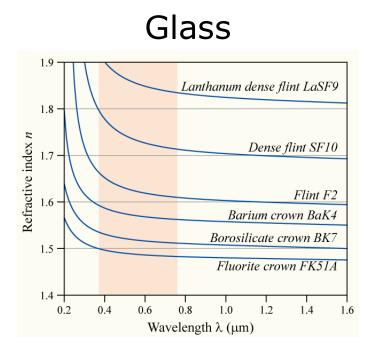
 $E_{ph}$  = photon energy

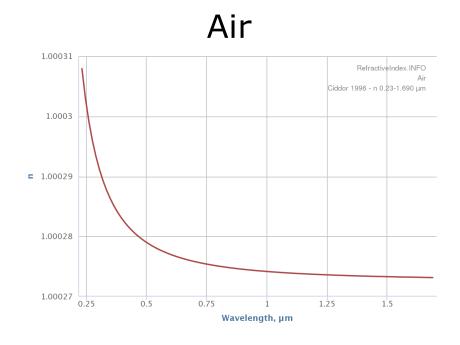
 $\lambda$  = Photon wavelength

Theory of Cherenkov Radiation: Classical Electrodynamics by J. D. Jackson (Section 13.5)



#### Refractive index





- Dependence of the refractive index on photon frequency means the Cherenkov angle is not constant
- The spread of angles is referred to as Chromatic Error



## **Number of photons**

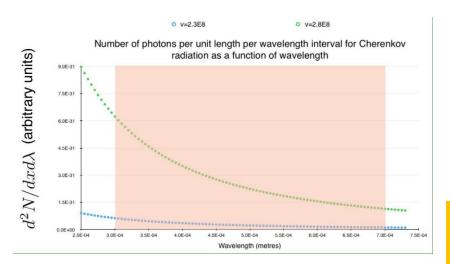
Number of emitted photons:

$$\frac{\partial^2 N}{\partial x \partial \lambda} = 2\pi a \left( 1 - \frac{1}{\beta^2 n^2} \right) \frac{1}{\lambda^2}$$

Between two wavelengths:

$$\frac{\partial N}{\partial x} = 2\pi a \left( 1 - \frac{1}{\beta^2 n^2} \right) \left( \frac{1}{\lambda_L} - \frac{1}{\lambda_H} \right)$$

The number of photons increases with energy



Using frequency instead of wavelength

$$\frac{dE}{dx} = \frac{q^2}{4\pi} \int_{v > \frac{c}{n(\omega)}} \mu(\omega) \omega \left( 1 - \frac{c^2}{v^2 n^2(\omega)} \right) d\omega$$

Typical example: Charged particle with momentum of few GeV/c or more emitting Cherenkov photons with few eV of energy



# **Detected photons**

$$\frac{\partial N}{\partial x} = 2\pi a \left( 1 - \frac{1}{\beta^2 n^2} \right) \left( \frac{1}{\lambda_L} - \frac{1}{\lambda_H} \right) \quad \text{For } \lambda_L = 400 \text{ nm and } \lambda_H = 700 \text{ nm}$$

$$\frac{N}{L} = 490 \sin^2 \theta$$

If the photons are reflected by a mirror with reflectivity  $R(\lambda)$ , and detected by a photon detector with quantum efficiency  $QE(\lambda)$  through a window with transmission  $T(\lambda)$ 

$$N=490\,L\,R\,Q\,T\sin^2\theta$$
 =  $N_0L\sin^2\theta$  And if we assume the mean angle  $\theta_{\rm c}$ 

$$N = N_0 L \sin^2 \theta_c$$

 $N_0$  is a figure of merit for a Cherenkov detector; e.g.  $N_0=200$ /cm is a good value

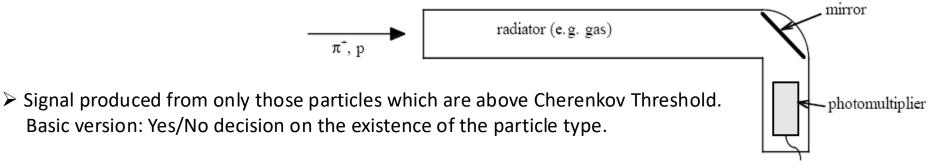


## A variety of Cherenkov detectors

- Detector design:
  - > Threshold Counters
  - Imaging detectors
    - Differential Cherenkov detectors
    - Ring Imaging Cherenkov detectors (RICH)
    - Detector for Internally Reflected light (DIRC)
- Types of photon detectors:
  - Gas based
  - Vacuum
  - Solid state
- Applications:
  - Accelerator HEP detectors
  - Astroparticle Physics detectors
  - Neutrino detectors



## **Threshold Cherenkov Counters**



- ➤ One counts the number of photoelectrons detected.
- Improved version: Use the number of observed photoelectrons or a calibrated pulse height to discriminate between particle types.
- $\triangleright$  For typical detectors: N<sub>o</sub> = 90 cm<sup>-1</sup>,

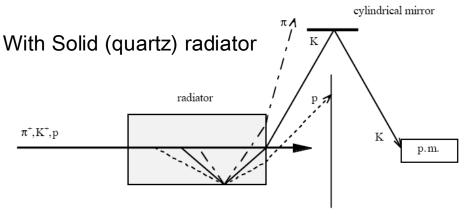
$$N_{ph}$$
 per unit length of the radiator =  $N_o * (m_1^2 - m_2^2)/(p^2 + m_1^2)$ 

At p= 1 GeV/c,  $N_{ph}$  per unit length = 16 /cm for Pions and 0 for Kaons. At p= 5 GeV/c,  $N_{ph}$  per unit length = 0.8 /cm for Pions and 0 for Kaons.

$$\rightarrow \Delta \beta / \beta = \tan^2 \theta / (2 * sqrt (N_{ph}))$$



## **Differential Cherenkov detectors**

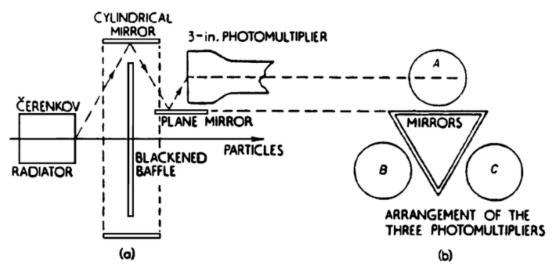


- $\triangleright$  Very small acceptance in  $\beta$  and direction of the charged particle. (Narrow range in velocity and direction intervals ).
- ➤ Mostly used for identifying particles in the beam lines.

$$\rightarrow \Delta \beta / \beta = (m_1^2 - m_2^2) / 2 p^2 = \tan \theta \Delta \theta$$

m<sub>1</sub>,m<sub>2</sub> (particle masses )<< p ( momentum)

 $\triangleright \Delta \beta / \beta$  from 0.011 to 4\* 10 <sup>-6</sup> achieved.

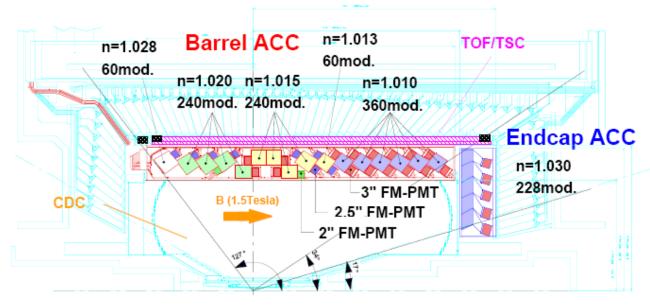


- Discovery of anti-proton in 1955 by Chamberlain, Segre et. al. at Berkeley.
- ➤ Nobel Prize in 1959

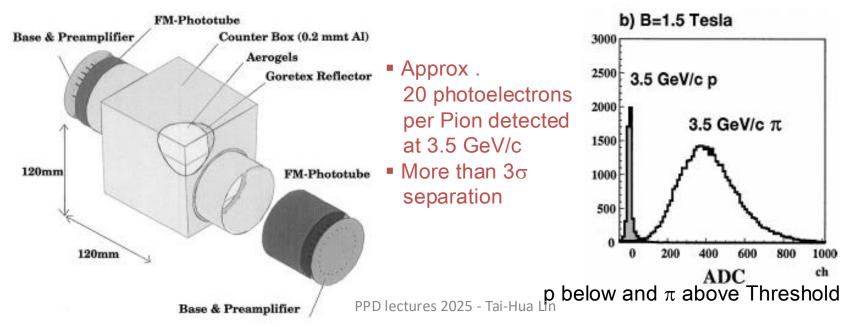


#### **Threshold Counters**

#### **BELLE: Threshold Cherenkov Detector**

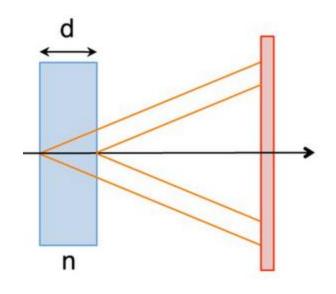


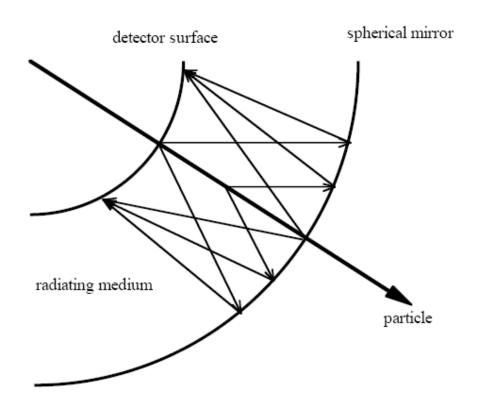
- Five aerogel tiles inside an aluminum box lined with a white reflector(Goretex reflector)
- Performance from testbeam



## **RICH Detectors**

#### Proximity focusing geometry

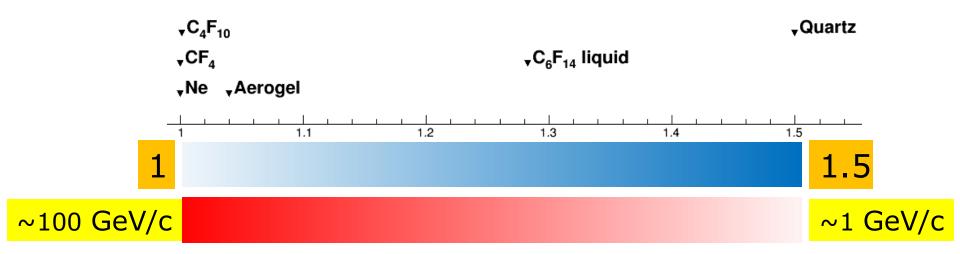




- ➤ Measures both the Cherenkov angle and the number of photoelectrons detected.
- > Can be used over particle identification over large surfaces.
- > Requires photodetectors with single photon identification capability.



## Refractive index range



#### Momentum

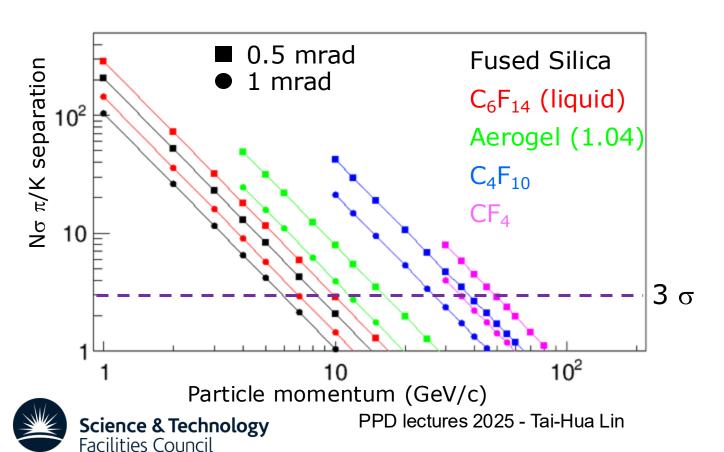


## RICH performance

$$N_{\sigma} \approx \frac{\left|m_1^2 - m_2^2\right|}{2P^2\sigma[\theta_c(tot)]\sqrt{n^2 - 1}}$$

For particles well above threshold

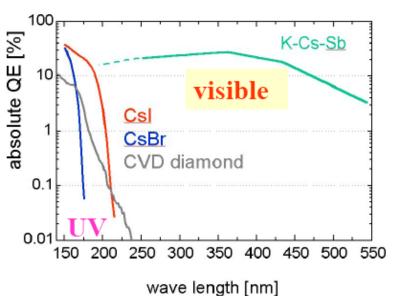
B. N. Ratcliff, NIMA 502 (2003) 211-221

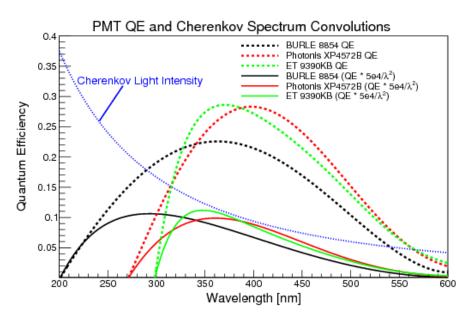


## **Detection of optical photons**

#### Photocathode:

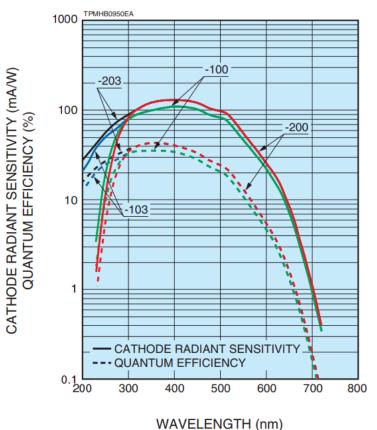
- A surface engineered to convert optical photons to electrons using the photo-electric effect
- Requires the energy of the photon to be higher than the "work function"
- Also possible in gas form
- Quantum Efficiency (QE): fraction of photons converted to electrons







# Latest photocathodes



	Spectral response		(A)	cteristics
Type No.	Range (nm)	Peak wavelength (nm)	Photo- cathode material	,
R11265U-100-M4	300 to 650	400	SBA	35
R11265U-200-M4	300 to 650	400	UBA	43
R11265U-103-M4	185 to 650	400	SBA	35
R11265U-203-M4	185 to 650	400	UBA	43
H13226A-100	300 to 650	400	SBA	35
H13226A-200	300 to 650	400	UBA	43
H13226A-103	185 to 650	400	SBA	35
H13226A-203	185 to 650	400	UBA	43

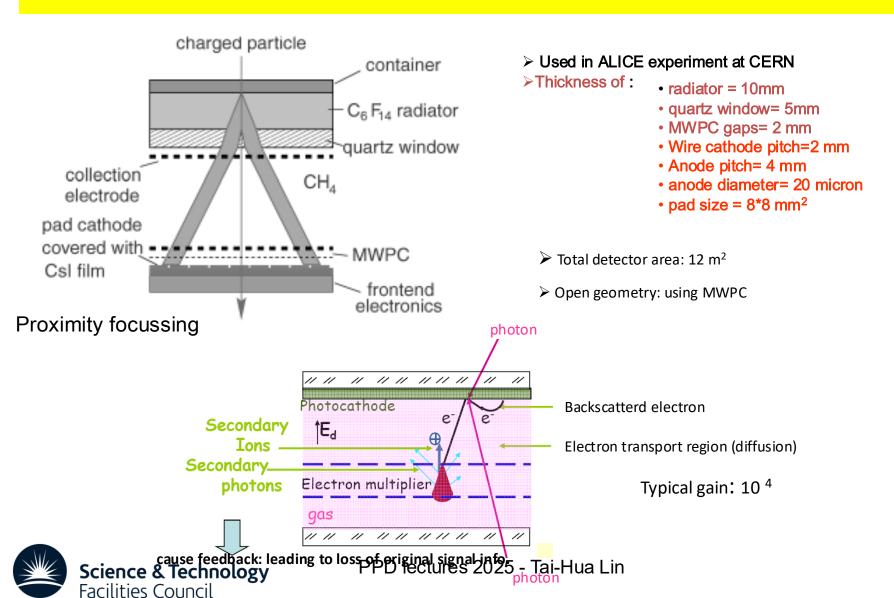


## **Detection of photo-electrons**

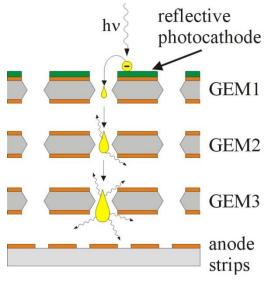
- Gas based detectors
  - Proportional counters
  - > MWPC
  - > GEM
  - Other kind of micro-pattern detector
- Vacuum:
  - Photomultipliers
  - Microchannel plates
  - Hybrid photon detectors (HPD)
- Solid state
  - Avalanche photo-diodes (APD)
  - > SiPMs



#### Gas based detectors



## **GEM** based gas detectors



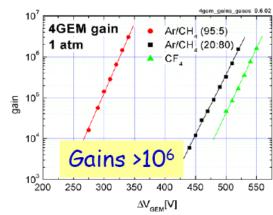
**GEM:** Gas Electron Multiplier



GEM with semi-transparent Photocathode (K-Cs-Sb)

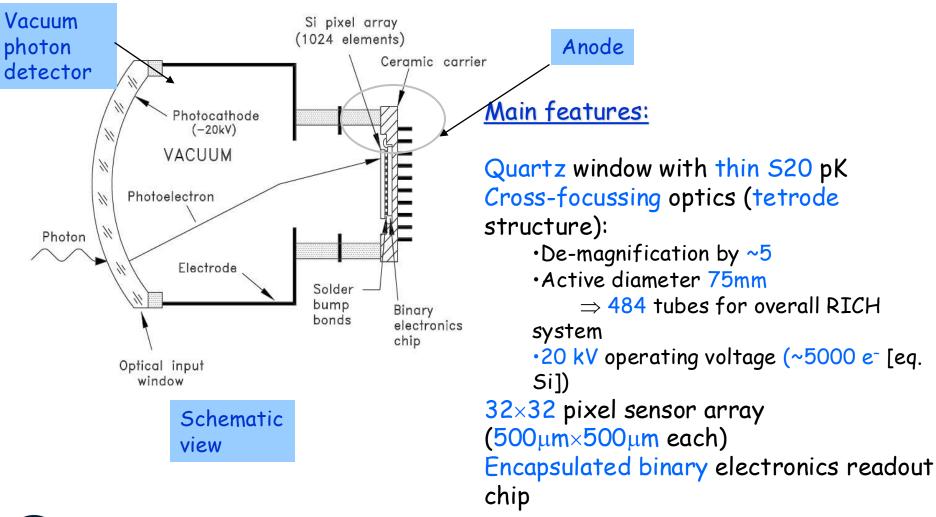
- > Photon and ion feed back reduced.
- ➤ Gated operation to reduce noise.

  (no readout outside a 'time window of signal')
- For now only closed geometry (in sealed tubes): Reduced fraction of useful area for photon detection (Active Area Fraction) compared to open geometry.



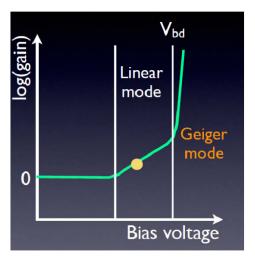


## **Pixel-HPD description**

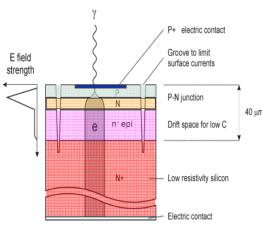


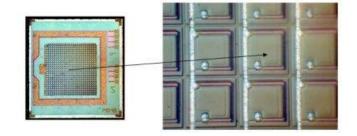


# Silicon Photo-Multiplier

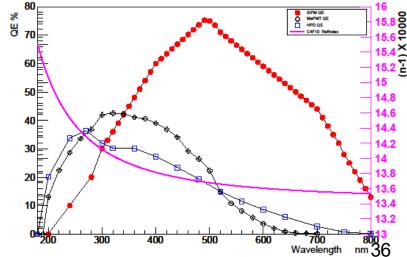


 Primary building block, GM-APD.

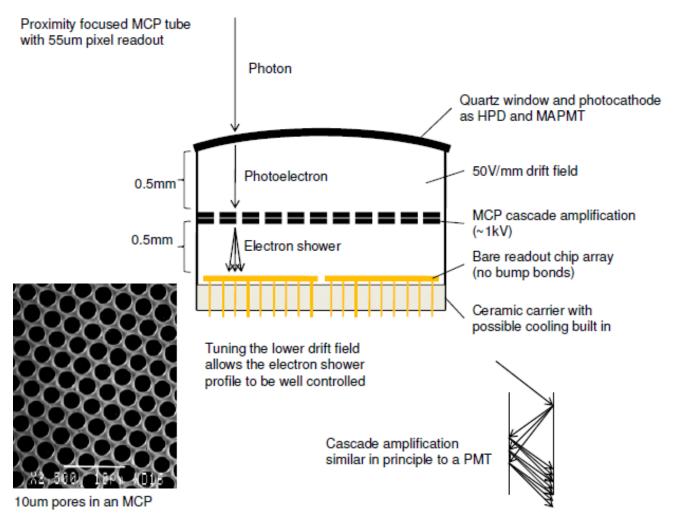




- ➤ Photon Detection Efficiency (PDE) for SiPM is better that of ordinary PMT.
- > Time resolution= ~ 100 ps.
- ➤ Works in magnetic field
- $\triangleright$  Gain =  $\sim 10^{6}$
- Reducing noise levels for single photon detection is still an issue especially after irradiation

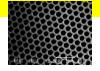


## Micro-Channel Plate (MCP) Detectors





## **MCP** detectors

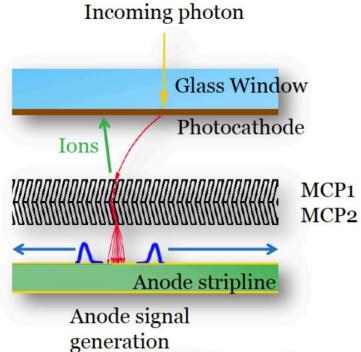


#### Typical Size:

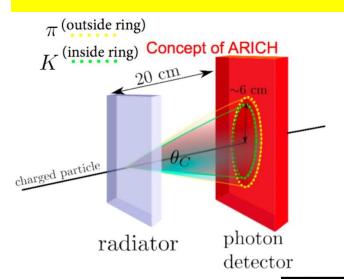
- 2 mm thickness,
   51 mm X 51 mm active area.
- 10 micron pores separated by 15 microns
- Chevron: 8 degree tilt: To increase th gain and reduce ion-feed back
- Gain: ~ 5 \* 10 5
- Typically ~ 1000 channels per MCP.

- Measure Space and time of the hits.
- Manufactured by industry (Photonics for example).
- Resolutions: Space: ~100 microns, Time: ~50 100 psec.
- Short flight path of photoelectrons: Resistant to magnetic fields up to 0.8 Tesla.
- Can work at 40 MHz readout rate.
- Can detect single photons (No noise from 'first dynode' as in MAPMT).
- Fast 'ageing' at large luminosity (eg: LHC) is an issue, but there are some solutions.

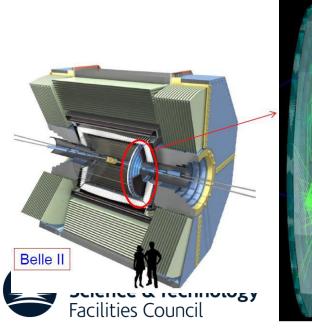


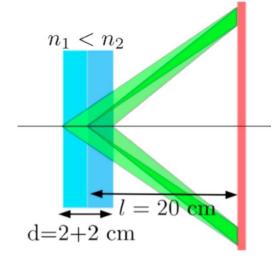


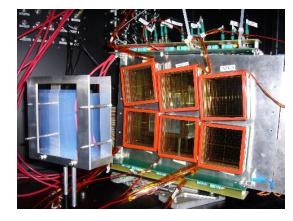
# Belle II ARICH (Endcap)



- ☐ Two tile focusing aerogel radiator
- Proximity focusing optics
- HAPD detectors
  - Bi-alkali photocathode
  - Work in magnetic field

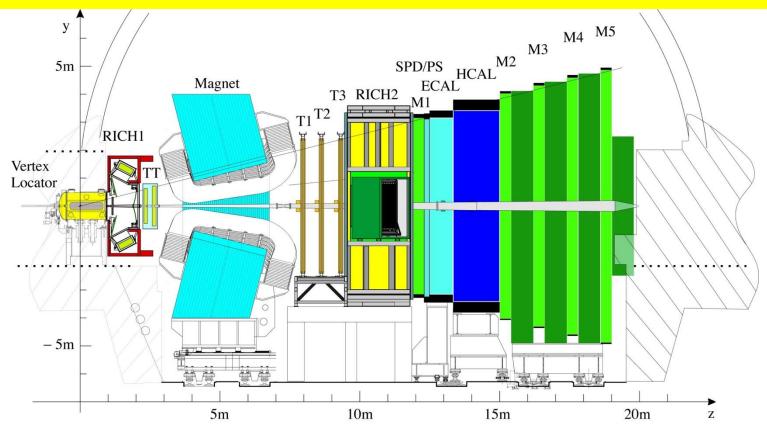






PD lectures 2025 - Tai-Hua Lin

# **LHCb Experiment**



- > Precision measurement of B-Decays and search for signals beyond standard model.
- ➤ Two RICH detectors covering the particle momentum range 1→100 GeV/c using aerogel, C<sub>4</sub>F<sub>10</sub> and CF<sub>4</sub> gas radiators.



# **LHCb RICH Specifications**

RICH1: Aerogel  $2 \rightarrow 10 \text{ GeV/c}$  $C_4 F_{10} < 70 \text{ GeV/c}$ 

RICH2:  $CF_4$  <100 GeV/c.

Aerogel C<sub>4</sub>F<sub>10</sub> CF<sub>4</sub>

L 5 86 196 cm

 $\theta_{c}^{\text{max}}$  242 53 32 mrad

 $\pi_{Th}$  0.6 2.6 4.4 GeV/c

K<sub>Th</sub> 2.0 9.3 15.6 GeV/c

Aerogel: Rayleigh Scattering

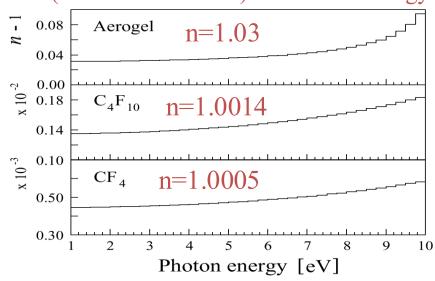
Not used from 2015 onwards

 $T = A e^{(-C t/\lambda^4)}$ 

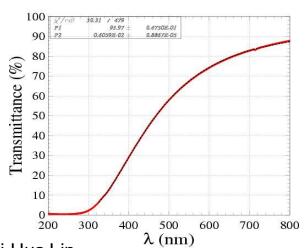
Typically: A = 0.94,  $C = 0.0059 \mu m^4 / cm$ 

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(Refractive Index-1) vs. Photon Energy

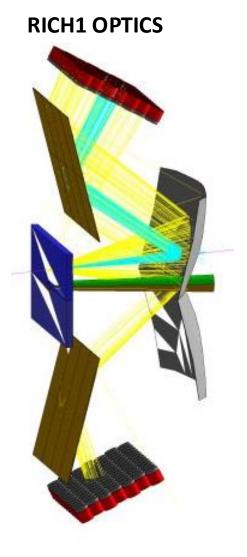


Aerogel Transmission

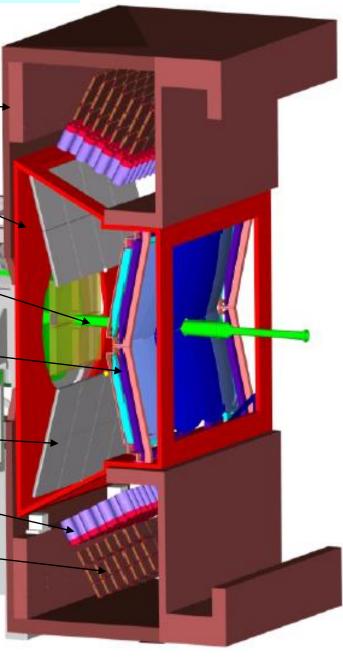


41

#### LHCb-RICH1 SCHEMATIC



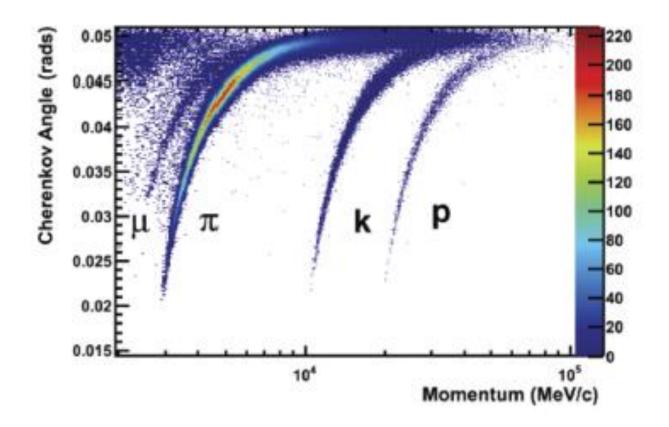
**Magnetic Shield Gas Enclosure Beam Pipe Spherical Mirror Flat Mirror Photodetectors Readout Electronics** PPD lectures 2025 - Tai-Hua Lin



■ Spherical Mirror tilted to keep photodetectors outside acceptance (tilt ~ 0.3 rad)

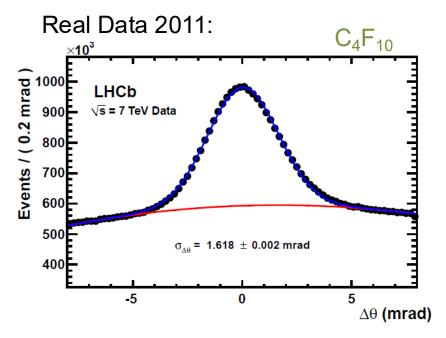
#### Performance of LHCb RICH

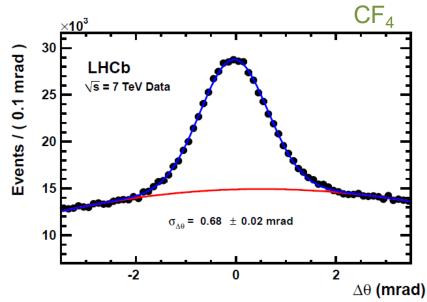
> From isolated Cherenkov rings from RICH1 in Real Data



#### LHCb-RICH resolution

#### Single photon resolutions





From simulation in 2011:

 $\sigma_{\Delta\theta}$  = 1.53 mrad

From simulation in 2011:  $\sigma_{\Delta\theta} = 0.68 \text{ mrad}$ 

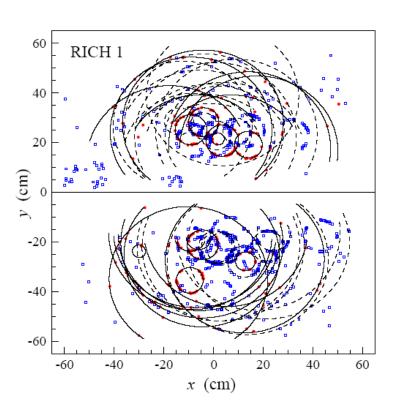
Resolution components: Chromatic: ref. index variation

Emission Point: tilt of the mirror

Pixel size: granularity of the pixel

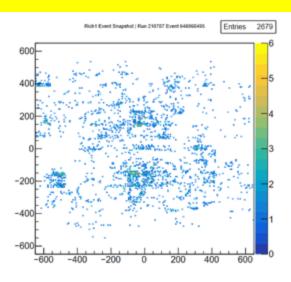
PSF: spread of photo electron direction inside HPD

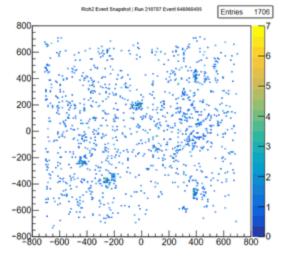
# Pattern recognition



Red: From particles from Primary and Secondary Vertex Blue: From secondaries and background processes

(sometimes with no reconstructed track)







#### Pattern Recognition in Accelerator based Cherenkov Detector

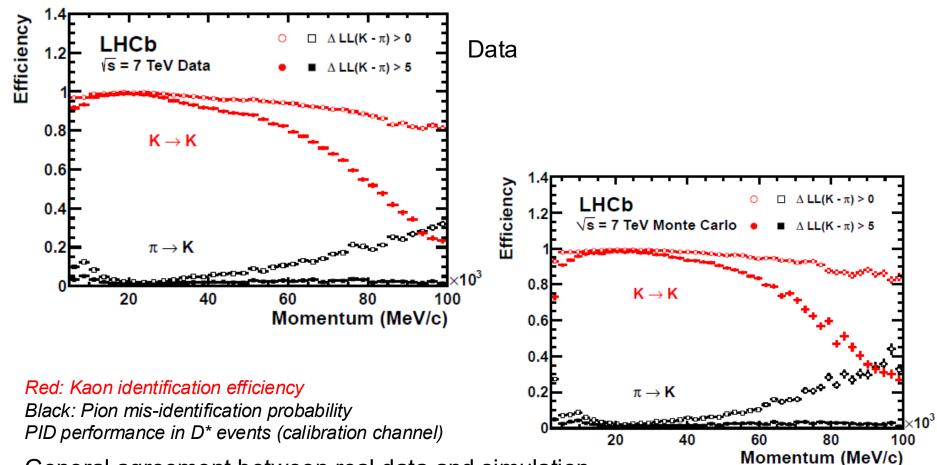
• Events with large number of charged tracks giving rise to several overlapping Cherenkov Rings on the Photo detector plane. Problem: To identify which tracks correspond to which hits and then identify the type (e,  $\pi$ , p etc.) of the particle which created the tracks.

- Hough Transform: (used by ALICE at CERN)
- Project the particle direction on to the detector plane
- > Accumulate the distance of each hit from these projection points in case of circular rings.
- Collect the peaks in the accumulated set and associate the corresponding hits to the tracks.
- (used by LHCb at CERN)
- Likelihood Method: > For each of the track in the event, for a given mass hypothesis, create photons and project them to the detector plane using the knowledge of the geometry of the detector and its optical properties. Repeat this for all the other tracks.
  - > From this calculate the probability that a signal would be seen in each pixel of the detector from all tracks.
  - > Compare this with the observed set of photoelectron signal on the pixels. by creating a likelihood.
  - > Repeat all the above after changing the set of mass hypothesis of the tracks. Find the set of mass hypothesis, which maximize the likelihood.

(Ref: R.Forty: Nucl. Inst. Mech. A 433 (1999) 257-261)



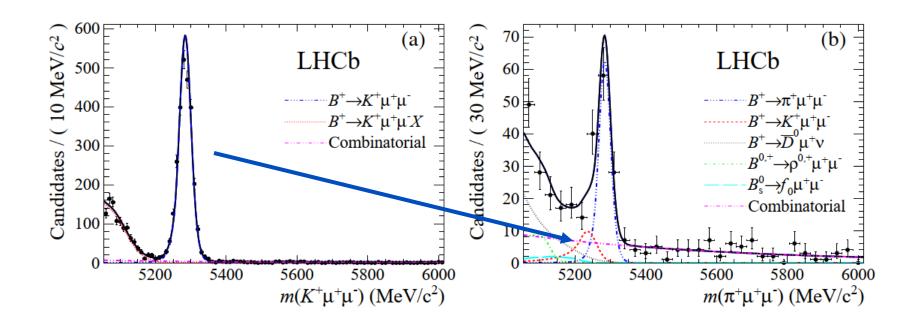
# LHCb RICH: PID performance



General agreement between real data and simulation



# Physics impact of hadron PID



First measurement of the differential branching fraction and CP asymmetry of the  $B\pm\to\pi\pm~\mu+~\mu-$  decay, JHEP 10 (2015) 034 [arXiv:1509.00414]



#### How to make a better RICH detector

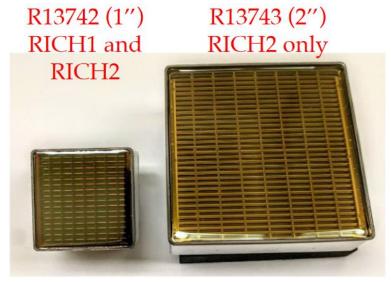
- Increase the number of photons
  - Better photon detectors (photo-cathodes)
  - Better area coverage by photon detectors
  - Better optical coupling of components
- Improve Cherenkov angle resolution
  - Smaller pixels
  - Lower chromatic error
    - Detectors sensitive to green rather than blue
  - > Improve aberrations
    - Light weight mirrors in the acceptance

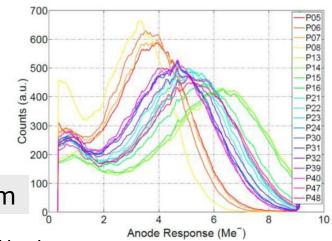


## The current LHCb RICH

# **Continuous 40 MHz** readout

- MaPMTs from Hamamatsu
- New radiation hard asic (CLARO) for MaPMT readout
- FPGA based digital board
- GBT for data transmission





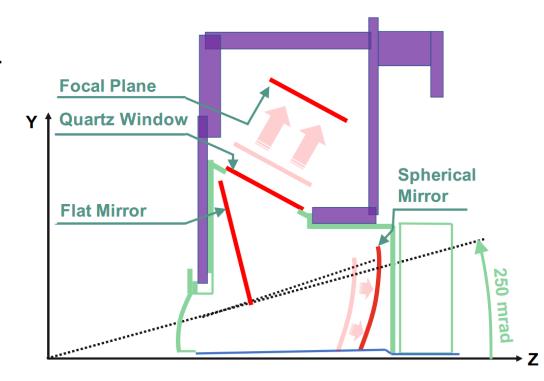
Single photon spectrum



## The current LHCb RICH

#### **More complex events**

- Completely new RICH1
- New mirrors with longer focus
- New position for the photon detectors
- New gas enclosure, exit window, cooling and support structures





# Cherenkov angle resolution

Contributions to resolution (in mrad)	RICH1-2015 HPD	RICH1- Upgrade MaPMT	RICH2-2015 HPD	RICH2- Upgrade MaPMT
Chromatic	0.84	0.58	0.48	0.31
Pixel	0.60, PSF=0.86	0.44	0.19, PSF=0.29	0.19
Emission Point	0.76	0.37	0.27	0.27
Overall	<b>1.60</b> →	0.78	<b>0.65</b> →	0.45

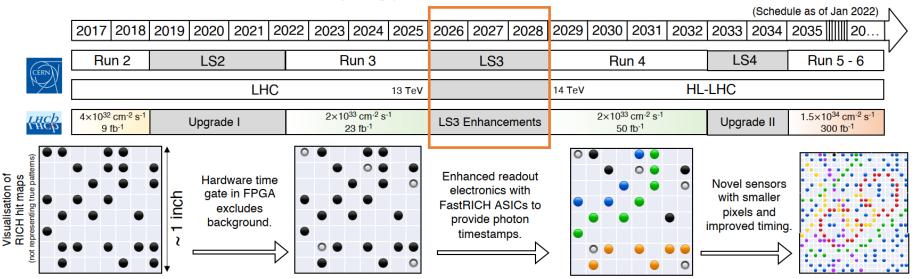
#### Improvements due to:

- ☐ Chromatic error; MaPMTs QE peaks at higher wavelength
- No Point Spread Function for MaPMTs
- ☐ Better emission point error with new RICH1 geometry



## **Evolution of the RICH photon detection**





- LS3 / Run 4 : focus on FastRICH readout electronics with fast timing and wide input dynamic range.
- LS4 / Run 5 : focus on sensor technology.
  Fast-timing is essential for the luminosity challenge after Upgrade II.

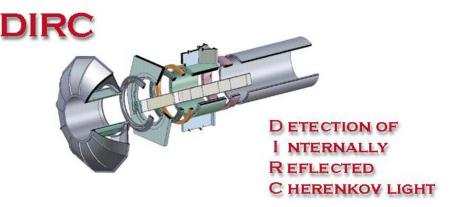


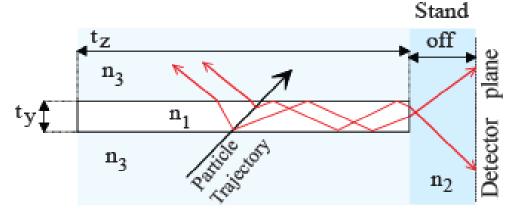
#### DIRC

- Direct Internally Reflected Cherenkov light
- Detectors outside the "acceptance"
- Can fit in a narrow space

The standoff region is designed to maximize the transfer efficiency between the radiator and the detector.

If this region has the same index of refraction as the radiator,  $n_1 \cong n_2$ , the transfer efficiency is maximized and the image will emerge without reflection or refraction at the end surface.

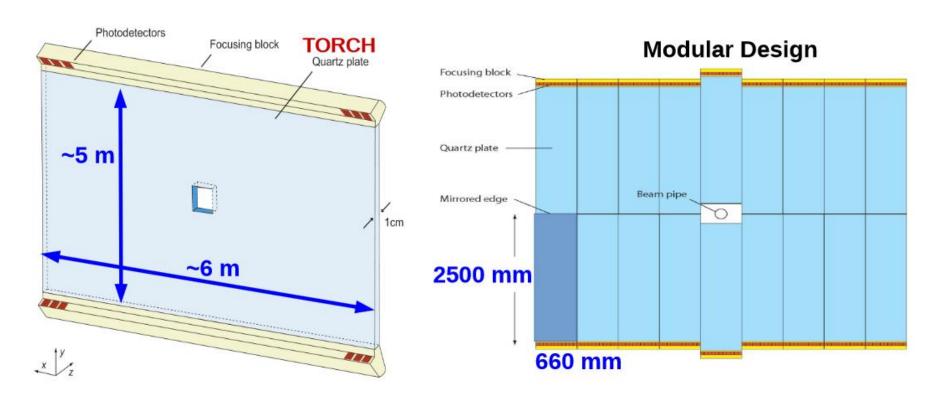






# TORCH (i)

A time of flight detector based on Cherenkov radiation

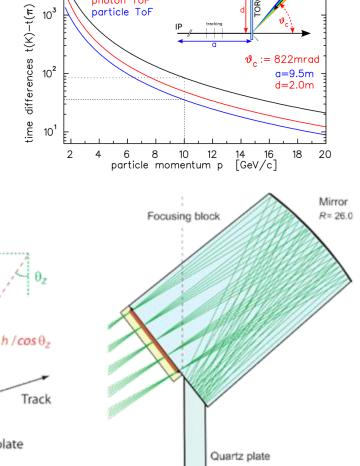




# TORCH (ii)

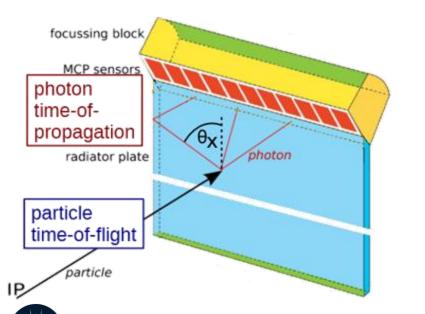
The focusing unit converts angles to positions allowing the measurement of propagation time

Multiple photons from the same track can improve the time accuracy



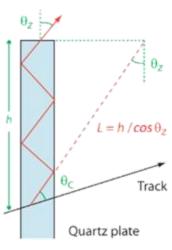
Δt ToF+ToP combined photon ToP

particle ToF



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a = 9.5 m

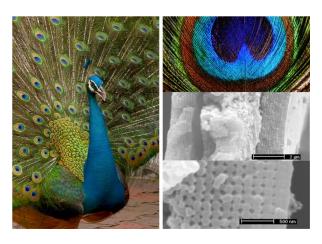
## **Photonic Crystals**

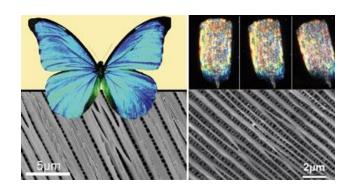




Optical micrograph of blue feather barbs www.pnas.org/cgi/doi/10.1073/pnas.12043831 09

#### Parrot (scarlet macaw) in a tropical forest

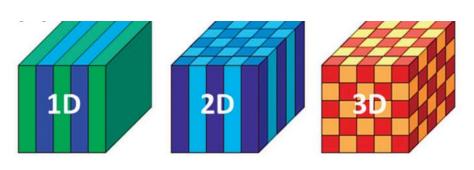




CERN Courier Aug23, 2005

Phys: Rev. 25-72, 1010902 (2005)

# **Photonic Crystals**



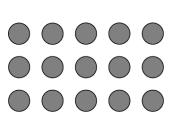
- Periodic arrangement of objects with two different refractive indices.
- Lattice constants comparable to the wavelength of light in the material
- electrons in pure semi-conductor : similar features to photons in photonic crystal

- Cherenkov radiation:
  - Physical origin is a mixture of conventional Cherenkov radiation and Transition radiation
  - Photon propagation in the crystal in the form of Bloch waves
  - All the properties can be derived from solving Maxwell's equations for periodic lattice

arXiV:0808:3519

- There is no Cherenkov Threshold
- Cherenkov cone can be forward or backward
- Starting from two media with refractive indices n<sub>1</sub> and n<sub>2</sub>, create a new effective refractive index n<sub>3</sub>.

r=cylinder radius a= lattice constant



## Summary

- The field of Particle Identification Detectors is an evolving field.
- They have contributed to some of the important discoveries in High Energy Physics in the last 50 years and they continue to be a crucial part of some of the recent experiments.
- The RICH detectors offer excellent Particle Identification capability for the hadrons since they can be designed to have very good single photon Cherenkov Angle resolution and large Photoelectron yield. Recent advances in photodetectors enhance the capability of these detectors.
- The particle ID using dE/dx, time-of-flight and Transition Radiation detectors continue to provide Particle Identification in different experiments.
- Particle identification is a crucial part of some of the Astroparticle physics experiments and long base line neutrino experiments



#### The end

# Thank you for your attention

Any (more) questions?



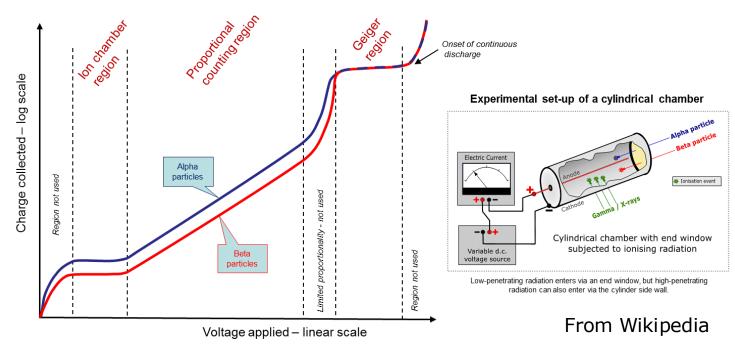
## **Detector technologies**

#### **Practical Gaseous Ionisation Detection Regions**

This diagram shows the relationship of the gaseous detection regions, using an experimental concept of applying a varying voltage to a cylindrical chamber which is subjected to ionising radiation. Alpha and beta particles are plotted to demonstrate the effect of different ionising energies, but the same principle extends to all forms of ionising radiation.

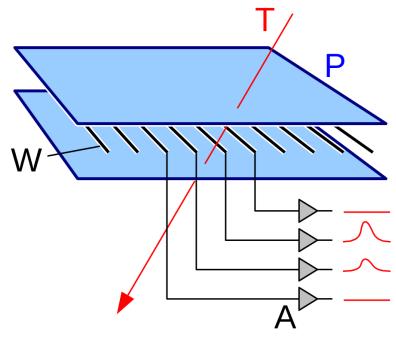
The ion chamber and proportional regions can operate at atmospheric pressure, and their output varies with radiation energy. However, in practice the Geiger region is operated at a reduced pressure (about 1/10<sup>th</sup> of an atmosphere) to allow operation at much lower voltages; otherwise impractically high voltages would be required. The Geiger region output does not differentiate between radiation energies.

#### Variation of ion pair charge with applied voltage

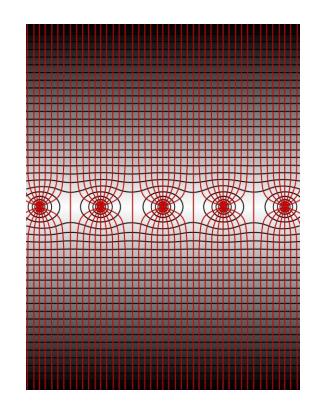




# Multiwire proportional chamber



By Michael Schmid - CC BY-SA 3.0, https://commons.wikimedia.org/w/index.php?curid=442715

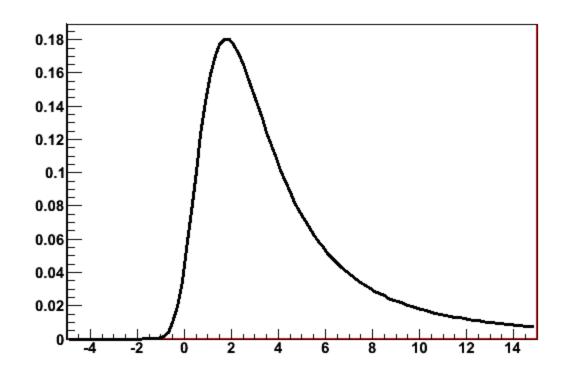


By Wiso - CC BY-SA 3.0, https://commons.wikimedia.org/w/index.php?curid=2934457



#### Landau distribution

#### Energy loss due to ionisation





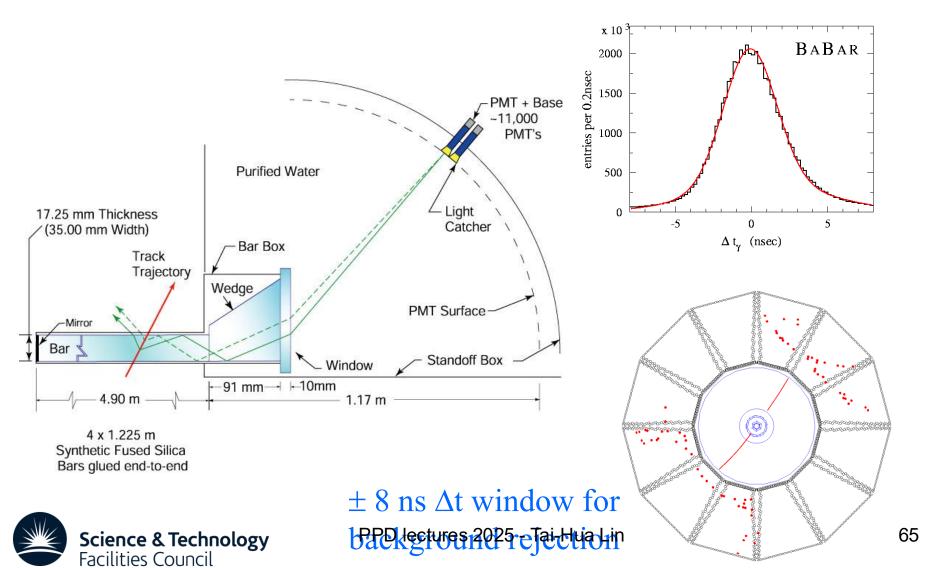
## Features of the PMTs and HPDs

- PMT: ➤ Typical Gain of MAPMT 300 K.
  - Excellent time resolution: 125 ps for example (Ex: used in underwater Cherenkov detectors).
  - Active area fraction: 40 %: Fraction of effective detection area. This can be Improved with a lens, but then one may loose some photons at the lens surface.
  - ➤ Recent developments: Flat panel pmts with 89 % active area fraction.

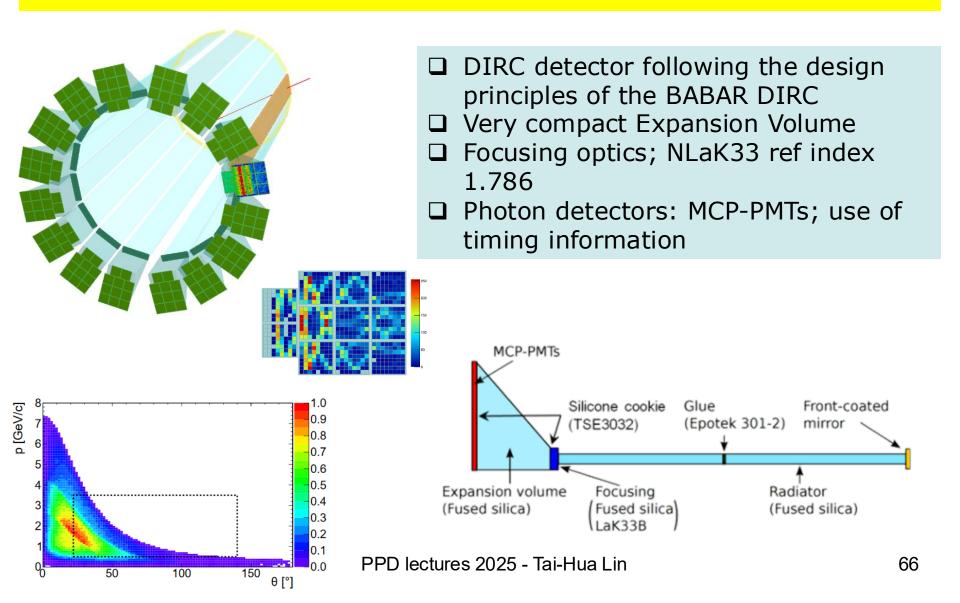
    New photocathodes with >45% QE at 400 nm
- HPD: > Typical gain 5K, but quite uniform across different channels.
  - > Excellent Single photon identification capability.
  - ➤ Active area fraction: 35 → 76 %



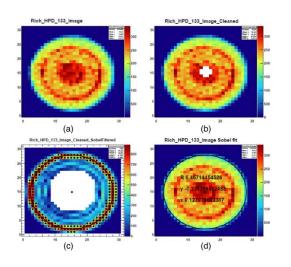
## **Babar DIRC**



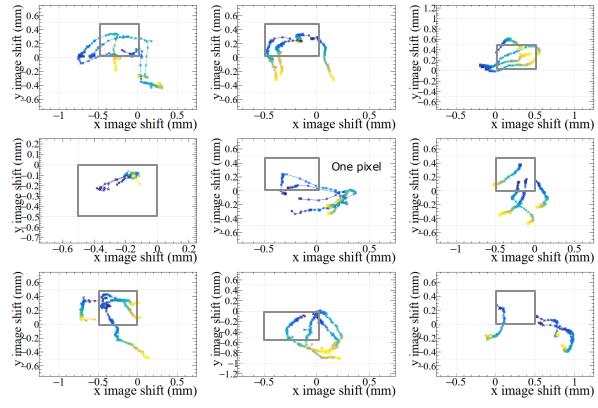
#### **Panda Barrel DIRC**



# **HPD** photocathode position

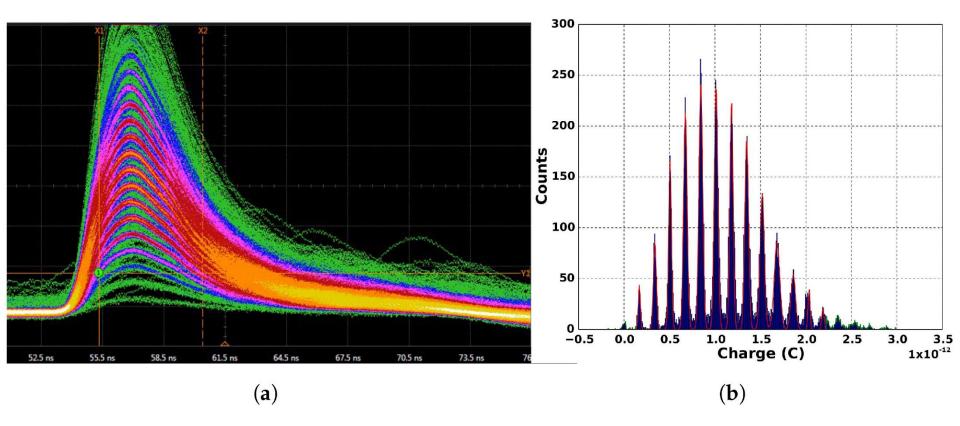


Automatic Online calibration for every run (max 1 hour)





## SiPM spectrum

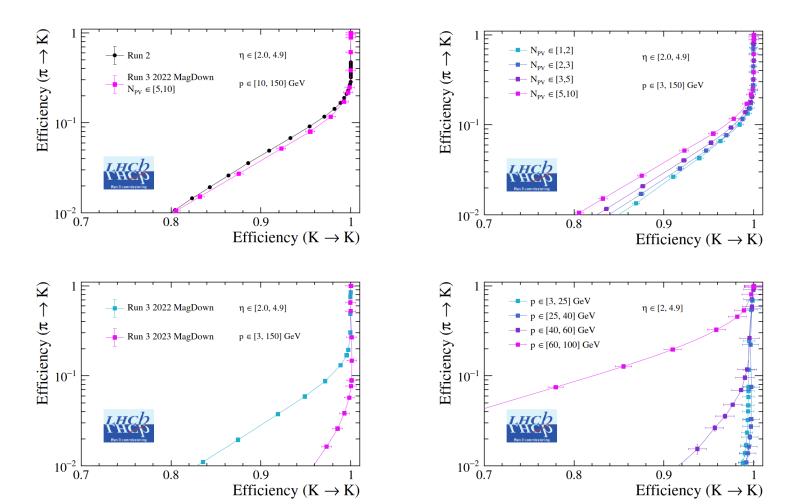


Hamamatsu SiPM S13360-3075CS

https://doi.org/10.3390/electronics10080961

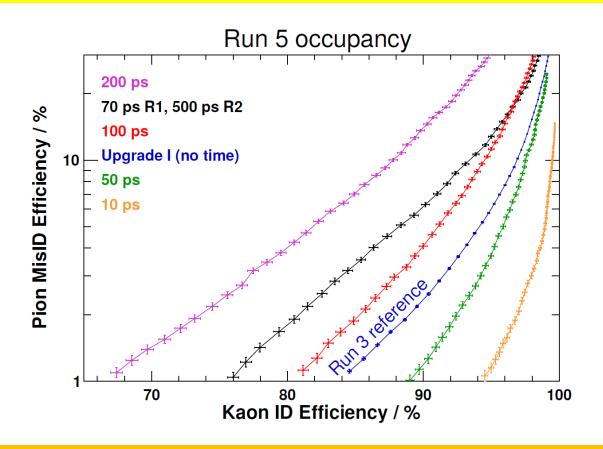


## Performance of new RICH





# Future (LHCb) Upgrades

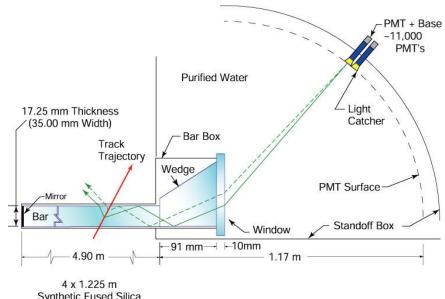


Improve pattern recognition using accurate timing of primary vertices



#### DIRC PRINCIPLE

- If  $n > \sqrt{2}$  some photons are always totally internally reflected for  $\beta \approx 1$  tracks.
- Radiator and light guide: Long, rectangular Synthetic Fused Silica ("Quartz") bars (Spectrosil: average  $\langle n(\lambda) \rangle \approx 1.473$ , radiation hard, homogenous, low chromatic dispersion)
- Photons exit via wedge into expansion region (filled with 6m<sup>3</sup> pure, de-ionized water).



Synthetic Fused Silica Bars glued end-to-end

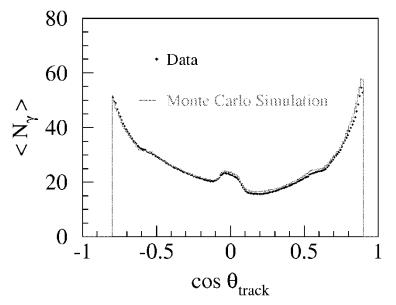
- Pinhole imaging on PMT array (bar dimension small compared to standoff distance). (10,752 traditional PMTs ETL 9125, immersed in water, surrounded by hexagonal "light-catcher", transit time spread ~1.5nsec, ~30mm diameter)
- DIRC is a 3-D device, measuring: x, y and time of Cherenkov photons, defining  $\theta_c$ ,  $\phi_c$ ,  $t_{\text{propagation}}$  of photon.





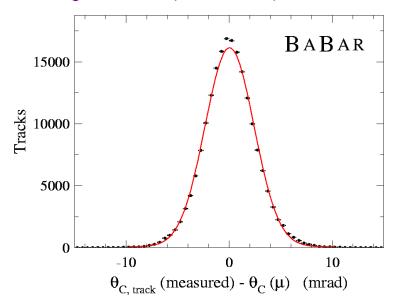
#### DIRC PERFORMANCE

Number of Cherenkov photons per track (di-muons) vs. polar angle:



Between 20 and 60 signal photons per track.

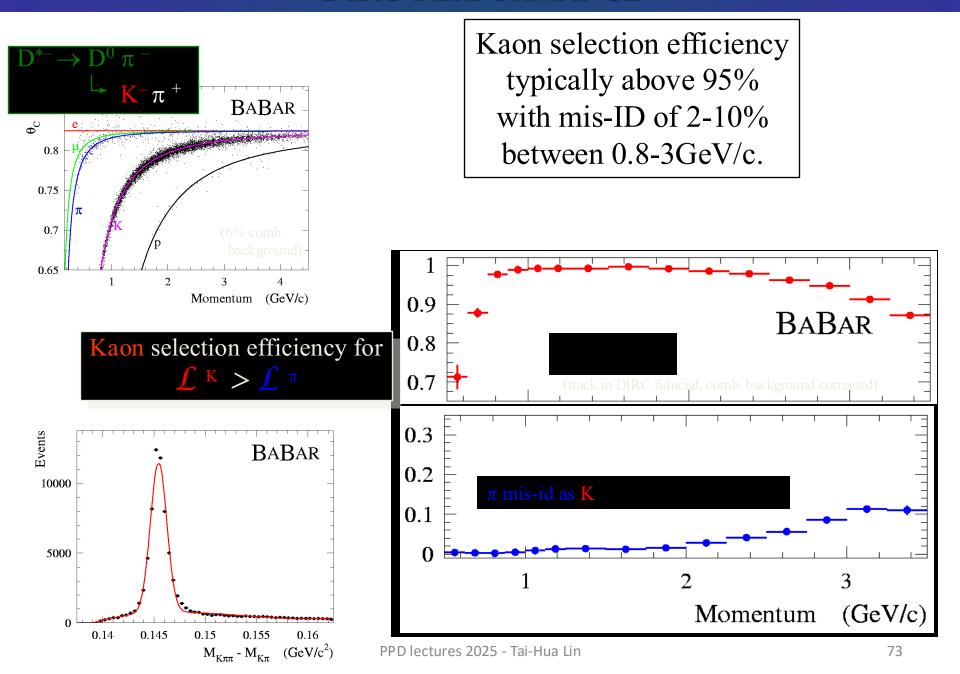
Resolution of Cherenkov angle fit per track (di-muons):



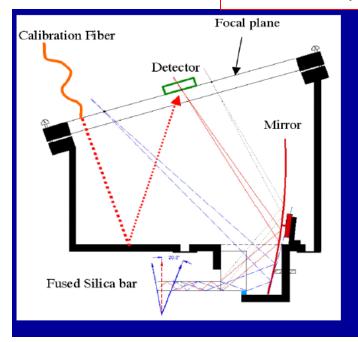
 $\sigma(\Delta\theta_c) = 2.4 \text{ mrad}$ 

Track Cherenkov angle resolution is within ~10% of design

## **DIRC PERFORMANCE**



### New Development: Focussing DIRC



$$v_{group}(red) > v_{group}(blue)$$

- Red photons arrive before blue photons
- Time of Propagation= PathLength/ v group
- Correct for Chromatic error from the mesurement of time of propagation.

#### → Future DIRC needs to be smaller and faster:

Focusing and smaller pixels can reduce the expansion volume by a factor of 7-10 Faster PMTs reduce sensitivity to background.

#### Additional benefit of the faster photon detectors:

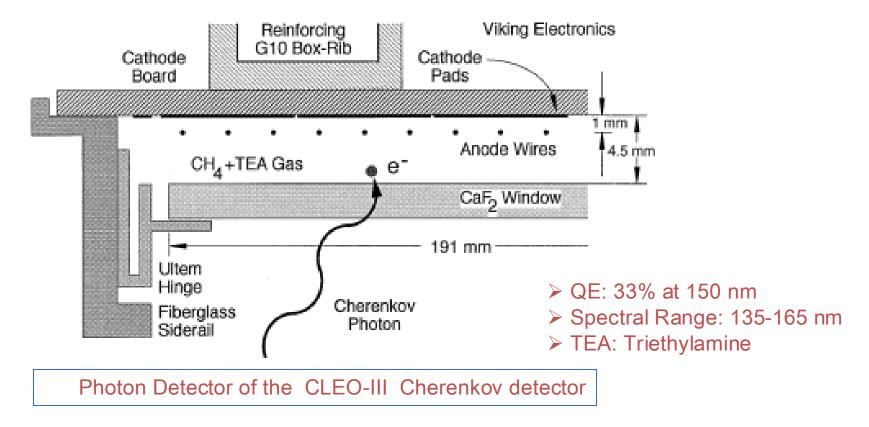
Timing resolution improvement: σ ~1.7ns (BABAR DIRC) → σ ≤150ps (~10x better) which allows measurement of <u>photon color</u> to correct the chromatic error of θ<sub>e</sub> (contributes σ ~ 5.4 mrads in BABAR DIRC)

#### Focusing mirror effect:

- Focusing eliminates effect of the bar thickness (contributes  $\sigma \sim 4$  mrads in BABAR DIRC)
- However, the spherical mirror introduces an aberration, so its benefit is smaller.

Ref: NIMA 595(2008)104-107

#### Gas Based Photon Detectors



- > photon passes through the CaF<sub>2</sub> and converts to photoelectron by ionizing a TEA molecule.
- ➤ The photoelectron drifts towards and avalanches near the anode wires, theirby inducing a charge signal on the cathode pads.

**CAPRICE Experiment** Balloon Experiment: RICH detector Pads 2 mm TMAE vapour Ethane gas 10 mm Strips Quartz window Cosmic TMAE: (tetrakis(dimethylamino) ethylene) MWPC C4F10-high purity gas spherical mirror

500 mm

PPD lectures 2025 - lal-Hua Lin

## Components of a Cherenkov Detector

➤ Main Components: • Radiator : To produce photons

• Mirror/lens etc. : To help with the transport of photons

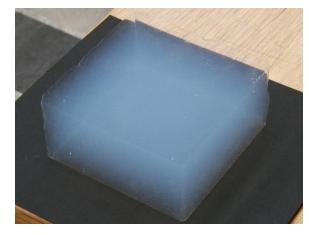
Photodetector : To detect the photons

➤ Radiator: Any medium with a Refractive Index.

	•		9 1	ı	
Lyamn				$\mathbf{r}$	70
Examp	$\mathbf{v}_{\mathbf{l}}$	ıau	ાહા	$\mathbf{U}$	0

Medium	n-1	Υth	Photons/m
He (STP)	3.5 10-5	120	3
CO <sub>2</sub> (STP)	4.1 10-4	35	40
Silica aerogel	0.025-0.075	4.6-2.7	2400-6600
water	0.33	1.52	21300
Glass	0.46-0.75	1.37-1.22	26100-33100

Aerogel: network of SiO<sub>2</sub> nano-crystals



$$\gamma = 1/\text{sqrt}(1-\beta^2)$$

> The atmosphere, ocean are the radiators in some Astro Particle Cherenkov Detectors

## LHCb-RICH Design

RICH1: Aerogel L=5cm p:2 $\rightarrow$ 10 GeV/c n=1.03 (nominal at 540 nm) C<sub>4</sub>F<sub>10</sub> L=85 cm p: < 70 GeV/c n=1.0014 (nominal at 400 nm)

**Upstream of LHCb Magnet** 

Acceptance: 25→250 mrad (vertical)
300 mrad (horizontal)

Gas vessel: 2 X 3 X 1 m<sup>3</sup>

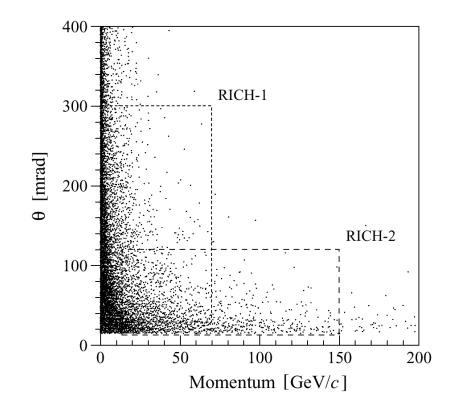
RICH2: CF<sub>4</sub> L=196 cm p: < 100 GeV/c n =1.0005 (nominal at 400 nm)

**Downstream of LHCb Magnet** 

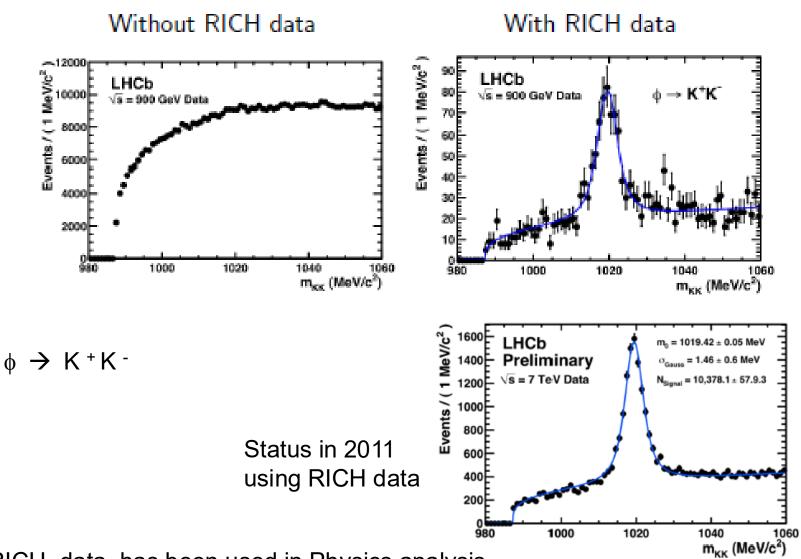
Acceptance: 15→100 mrad (vertical)

120 mrad (horizontal)

Gas vessel: 100 m<sup>3</sup>



# LHCb RICH Data in Physics Analysis



RICH data has been used in Physics analysis

## Example of LHCb-RICH PERFORMANCE

- Performance as seen in Simulated Data in 2010
- Yield: Mean Number of hits per isolated saturated track ( $\beta$ ~1).

Aerogel	C4F10	CF4
4.6	34.0	24

Single Photon Cherenkov Angle Resolutions in mrad.

Components and Overall (mrad)	Aerogel	C <sub>4</sub> F <sub>10</sub>	CF₄
Chromatic	2.65	0.84	0.48
Emission Point	0.34	0.61	0.27
Pixel Size	0.59	0.6	0.19
PSF	0.78	0.79	0.29
Overall RICH	2.84	1.45	0.65
Overall RICH+Tracks	2.86	1.50	0.76

- Chromatic: From the variation in refractive index.
- Emission Point: Essentially from the tilt of the mirrors.
- Pixel Size: From the granularity of the Silicon detector pixels in HPD
- PSF ( Point Spread Function):

From the spread of the Photoelectron direction as it travells inside the HPD,

(from the cross focussing in the electron optics)

#### Differential Cherenkov Detectors

- > Very small acceptance in  $\beta$  and direction of the charged particle. (Narrow range in velocity and direction intervals ).
- $\triangleright$  From the Cherenkov angle ( $\theta$ ) determine  $\beta$ .
- > Mostly used for identifying particles in the beam lines.
- ightharpoonup Resolution that can be achieved =  $\Delta \beta / \beta = (m_1^2 m_2^2) / 2 p^2 = \tan \theta \Delta \theta$

m<sub>1</sub>,m<sub>2</sub> (particle masses )<< p ( momentum)

- ➤ At high momentum, to get better resolution, use gas radiators which have smaller refractive index than solid radiators. Have long enough radiators to get sufficient signal photons in the detector.
- ➤ To compensate for Chromatic dispersion (n (E<sub>ph</sub>)), lens used in the path of the photons. (DISC: Differential Isochronous self- collimating Cherenkov Counter).
- $\triangleright$   $\Delta\beta/\beta$  from 0.011 to 4\* 10 -6 achieved.

## Differential Cherenkov Detectors

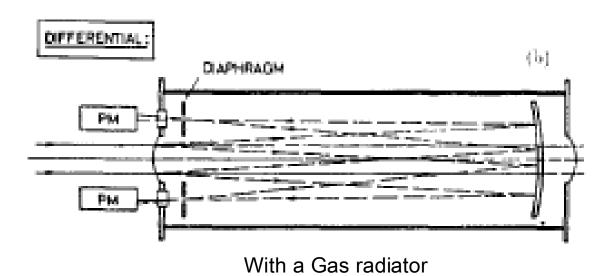


Table 2 Some differential Cherenkov counters

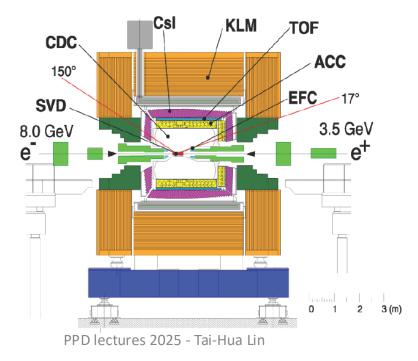
Type	Year	Length [m]	Angle [mrad]	Gas	Range for $(\pi - K)$ [GeV/c]	Remarks	Ref.
IHEP [1]	1968	5	23	He, N <sub>2</sub>	< 100	no optical	
[2]		10	12		< 200	correction	[3]
DISC	1964	2	44	CO <sub>2</sub>	< 100	corrected	[4]
FNAL DISC	1973	5.5	25	He	< 500	ld.	[2]
					$(< 100 \text{ for } \pi - \mu - e)$		
CEDAR W	1976	3.25	31	N <sub>2</sub>	< 150	Id	[5]
N		3.90	26	He	< 340		
HYPERON	1972	0.3	120	SF <sub>6</sub>	< 40	Id.	[2]
DISC				( < 100 for Σ−p)			
For comparison:							
LDISC	1976	0.05	640	FC88	< 5	corrected	[6]
				liquid		high aperture	

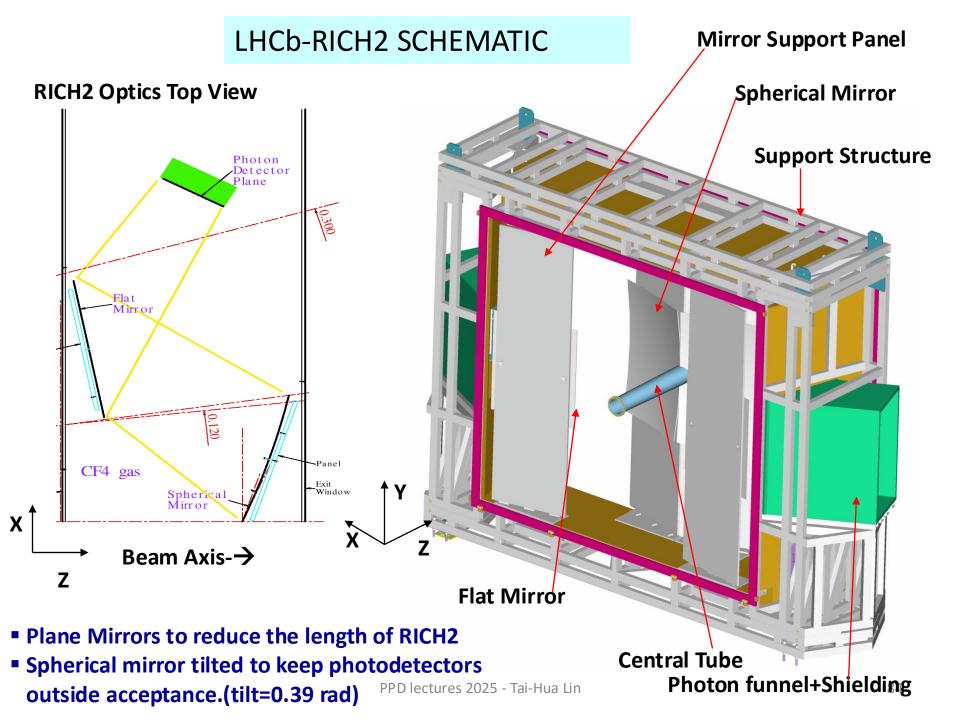
### Threshold Cherenkov Detectors

- > Can be used over a large area, for Example : For secondary particles in a fixed target or Collider experiment.
- $\triangleright$  E691 at Fermilab: To study decays of charm particles in the 1980's  $\Delta \beta/\beta = 2.3 * 10^{-5}$  using gas radiator.

➤ BELLE Experiment: To observe CP ciolation in B-meson decays at an electron-positron collider.

➤ BELLE: Continues to take Data.





## LHCb- RICH2 STRUCTURE



Entrance Window (PMI foam between two carbon fibre epoxy Skins)





RICH2

### **RICH detectors**

 $\rightarrow \Delta \beta / \beta = tan(\theta) * \Delta\theta_c = K \text{ where } \Delta\theta_c = < \Delta\theta > / sqrt(N_{ph}) + C$ 

where  $\Delta\theta$  is the mean resolution per single photon in a ring and C is the error contribution from the tracking, alignment etc.

- For example , for 1.4 m long CF<sub>4</sub> gas radiator at STP and a detector with N<sub>0</sub> = 75 cm<sup>-1</sup>  $K = 1.6 * 10^{-6}$  . (E=6.5 eV,  $\Delta E = 1 \text{ eV}$ )
- This is better than similar Threshold counters by a factor 125.
  This is also better than similar Differential counters by a factor 2.
- > Reason: RICH measures both hand to directly lent Differential and Threshold counters.
- $\triangleright$  Let u= sin<sup>2</sup> ( $\theta$ ) = 1- (1/n<sup>2</sup>) (m/p\*n) <sup>2</sup>

Number of standard deviations to discriminate between mass  $m_1$  an  $m_2$  =  $N_{\sigma} = (u_2-u_1) / (\sigma_u * sqrt(N))$  where  $\sigma_u : \Delta \theta$  converted into the parameter u. ( $\Delta \theta = error$  in single photon  $\theta$  measurement)

ightharpoonup At momentum p (= $\beta$  E), p= sqrt((m<sub>2</sub><sup>2</sup>-m<sub>1</sub><sup>2</sup>)/(2\* K \* N<sub> $\sigma$ </sub>)), for  $\beta \sim 1$ 

This equation can be used in the design of the RICH detectors.

➤ One the first large size RICH detector: in DELPHI at LEP.

# **History of Cherenkov radiation**

- The formula  $\cos(\theta) = 1/(n\beta)$  was already predicted by Heaviside in 1888
- ~1900: 'Blue glow' seen in fluids containing concentrated Radium (Marie & Pierre Curie)
- Pavel Alexeevich Cherenkov (1904-1990): Lebedev Physical Institute of the Russian Academy of Sciences.
- Discovery and Validation of Cherenkov Effect: 1934-37
- Full Explanation using Maxwell's equations: I.M. Frank and I.E. Tamm in 1937
- Nobel Prize in 1958: Cherenkov, Frank and Tamm.



# First experiments



Typical Apparatus used by Cherenkov to study the angular distribution of Cherenkov photons. (Incident  $\gamma$  ray produces electrons by Compton scattering in the liquid).

#### P. Cherenkov established that:

- Light Intensity is proportional to the electron path length in the medium.
- Light comes only from the 'fast' electrons above a velocity threshold, in his Apparatus.
- Light emission is prompt and the light is polarized.
- The wavelength spectrum of the light produced is continuous. No special spectral lines.
- The angular distribution of the radiation, its intensity, wavelength spectrum and its dependence on the refractive index agree with the theory proposed by his colleagues Frank and Tamm.

